Origin of oxidized mercury in the summertime free troposphere over the southeastern US

The MIT Faculty has made this article openly available. Please share how this access benefits you. Your story matters.

<table>
<thead>
<tr>
<th></th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>As Published</td>
<td><a href="http://dx.doi.org/10.5194/acp-16-1511-2016">http://dx.doi.org/10.5194/acp-16-1511-2016</a></td>
</tr>
<tr>
<td>Publisher</td>
<td>Copernicus GmbH</td>
</tr>
<tr>
<td>Version</td>
<td>Final published version</td>
</tr>
<tr>
<td>Accessed</td>
<td>Thu Feb 14 12:10:23 EST 2019</td>
</tr>
<tr>
<td>Citable Link</td>
<td><a href="http://hdl.handle.net/1721.1/102625">http://hdl.handle.net/1721.1/102625</a></td>
</tr>
<tr>
<td>Terms of Use</td>
<td>Creative Commons Attribution</td>
</tr>
<tr>
<td>Detailed Terms</td>
<td><a href="http://creativecommons.org/licenses/by/3.0/">http://creativecommons.org/licenses/by/3.0/</a></td>
</tr>
</tbody>
</table>
Origin of oxidized mercury in the summertime free troposphere over the southeastern US


1Department of Atmospheric Sciences, University of Washington, Seattle, WA, USA
2Environmental Program, Colorado College, Colorado Springs, CO, USA
3School of Science, Technology, Engineering and Mathematics, University of Washington-Bothell, Bothell, WA, USA
4Department of Earth, Atmospheric and Planetary Sciences, Massachusetts Institute of Technology, Cambridge, MA, USA
5Earth Observing Laboratory, National Center for Atmospheric Research, Boulder, CO, USA
6Department of Atmospheric and Oceanic Sciences, University of California, Los Angeles, CA, USA
7Atmospheric Chemistry Observations and Modeling Laboratory, National Center for Atmospheric Research, Boulder, CO, USA
8Rosenstiel School of Marine and Atmospheric Science, University of Miami, Miami, FL, USA
9Department of Chemistry, University of California, Irvine, CA, USA
10Department of Atmospheric and Oceanic Sciences, University of Colorado, Boulder, CO, USA
11Department of Physics, University of Helsinki, Helsinki, Finland

a now at: College of Engineering and Physical Sciences, University of New Hampshire, Durham, NH, USA

Correspondence to: V. Shah (vshah@uw.edu)

Received: 17 August 2015 – Published in Atmos. Chem. Phys. Discuss.: 5 October 2015
Revised: 12 January 2016 – Accepted: 15 January 2016 – Published: 10 February 2016

Abstract. We collected mercury observations as part of the Nitrogen, Oxidants, Mercury, and Aerosol Distributions, Sources, and Sinks (NOMADSS) aircraft campaign over the southeastern US between 1 June and 15 July 2013. We use the GEOS-Chem chemical transport model to interpret these observations and place new constraints on bromine radical initiated mercury oxidation chemistry in the free troposphere. We find that the model reproduces the observed mean concentration of total atmospheric mercury (THg) (observations: 1.49±0.16 ng m⁻³, model: 1.51±0.08 ng m⁻³), as well as the vertical profile of THg. The majority (65 %) of observations of oxidized mercury (Hg(II)) were below the instrument’s detection limit (detection limit per flight: 58–228 pg m⁻³), consistent with model-calculated Hg(II) concentrations of 0–196 pg m⁻³. However, for observations above the detection limit we find that modeled Hg(II) concentrations are a factor of 3 too low (observations: 212±112 pg m⁻³, model: 67±44 pg m⁻³). The highest Hg(II) concentrations, 300–680 pg m⁻³, were observed in dry (RH < 35 %) and clean air masses during two flights over Texas at 5–7 km altitude and off the North Carolina coast at 1–3 km. The GEOS-Chem model, back trajectories and observed chemical tracers for these air masses indicate subsidence and transport from the upper and middle troposphere of the subtropical anticyclones, where fast oxidation of elemental mercury (Hg(0)) to Hg(II) and lack of Hg(II) removal lead to efficient accumulation of Hg(II). We hypothesize that the most likely explanation for the model bias is a systematic underestimate of the Hg(0) + Br reaction rate. We find that sensitivity simulations with tripled bromine radical concentrations or a faster oxidation rate constant for Hg(0) + Br, result in 1.5–2 times higher modeled Hg(II) concentrations and improved agreement with the observations. The modeled tropospheric lifetime of Hg(0) against oxidation to Hg(II) decreases from 5 months in the base simulation to 2.8–1.2 months in our sensitivity simulations. In or-
der to maintain the modeled global burden of THg, we need to increase the in-cloud reduction of Hg(II), thus leading to faster chemical cycling between Hg(0) and Hg(II). Observations and model results for the NOMADSS campaign suggest that the subtropical anticyclones are significant global sources of Hg(II).

1 Introduction

Exposure to mercury affects the human nervous system, hinders cognitive development in children, and causes cardiovascular diseases in adults (Mergler et al., 2007; Karagas et al., 2012). In fish, mammals, and birds, mercury can adversely affect reproductive behavior (Scheuhammer et al., 2007). Although mercury is naturally present in our environment, human activities have increased its concentrations in the atmosphere and ocean by factors of 3 to 7 (Lamborg et al., 2002; Selin, 2009; Strode et al., 2010; Amos et al., 2013; Zhang et al., 2014) making mercury exposure a major public health concern.

In the troposphere, 90 % of mercury occurs in its elemental form (Hg(0)) in the gas phase, while the rest is in the oxidized (mercuric, +2) state (Hg(II)), either in the gas-phase or bound to particles (Gustin et al., 2013). The chemical forms of atmospheric Hg(II) have not been identified, but laboratory and theoretical studies suggest that they likely include HgCl₂, HgBr₂, HgBrBrO, HgBrNO₂, HgBrHO₂, HgO, HgSO₄, Hg(NO₃)₂, and Hg(OH)₂ (Gustin et al., 2013; Dibble et al., 2012; Huang et al., 2015). Both natural and anthropogenic processes emit mercury to the atmosphere, mostly as Hg(0). Atmospheric Hg(II) originates predominantly from the oxidation of Hg(0), with a minor contribution from direct anthropogenic emissions (Driscoll et al., 2013). Unlike Hg(0), Hg(II) is highly water-soluble and reactive, and is quickly scavenged from the atmosphere by rainwater or aerosol particles, such that 60 % of the global mercury deposition is estimated to occur by wet and dry deposition of Hg(II) (Selin et al., 2008; Holmes et al., 2010).

Atomic bromine (Br) has been observed as the main oxidant of atmospheric Hg(0) in the polar and the marine boundary layers (Lindberg et al., 2002; Ebinghaus et al., 2002; Laurie et al., 2003; Obrist et al., 2011), and laboratory studies (Ariya et al., 2002; Donohoue et al., 2006), theoretical calculations (Goedsite et al., 2004, 2012; Balabanov et al., 2005; Dibble et al., 2012; Shepler et al., 2007), and modeling studies (Holmes et al., 2006, 2010) suggest a predominant role of Br in the oxidation of Hg(0) in the global atmosphere. While other oxidants have been proposed, in particular OH and O₃ (Hall, 1995; Spicer et al., 2002; Pal and Ariya, 2004a, b; Sumner et al., 2005; Rutter et al., 2012), theoretical studies (Calvert and Lindberg, 2005; Goedsite et al., 2004; Hynes et al., 2009) suggest that the bimolecular reaction of Hg(0) with O₃ and OH is too slow in the atmosphere, and the fast rates observed in the laboratory could have been influenced by wall-mediated reactions or formation of aerosol particles (Tossell, 2006; Subir et al., 2011).

Measurements at a few high-elevation sites have shown episodic enhancements of Hg(II) concentrations (350–600 pg m⁻³) usually in low relative humidity and low CO air, suggesting efficient in situ production of Hg(II) in the free troposphere (Landis et al., 2005; Swartzendruber et al., 2006; Faïn et al., 2009; Sheu et al., 2010; Weiss-Penzias et al., 2015). Sillman et al. (2007) reported higher (60–248 pg m⁻³) Hg(II) concentrations at 3 km altitude than near the surface in aircraft flights off the Florida coast. Swartzendruber et al. (2009) found a large variability in Hg(II) concentrations (0–500 pg m⁻³) during five flights over the Pacific Northwest below 5 km altitude, with higher concentrations in free-tropospheric air with low aerosol concentrations. Lyman and Jaffe (2012) found enhanced Hg(II) concentrations of ~ 450 pg m⁻³ and depleted total mercury (THg, THg = Hg(0) + Hg(II)) concentrations (< 1 ng m⁻³) in a stratospheric intrusion, suggesting rapid oxidation of Hg(0) and loss of Hg(II) above the tropopause. During multiple year-round flights over Tennessee, USA, Brooks et al. (2014) found that Hg(II) concentrations at 2–4 km altitude were 10–30 times higher than those near the surface throughout the year. Typically, these past aircraft campaigns have focused on Hg(II) measurements over small spatial scales. The sparseness of free-tropospheric measurements of Hg(II) has hindered the validation of redox chemistry in global models of tropospheric mercury, which display large inter-model variability in wet and dry deposition (Bullock et al., 2008; Bullock et al., 2009; Travnikov et al., 2010).

The Nitrogen, Oxidants, Mercury, and Aerosol Distributions, Sources, and Sinks (NOMADSS) aircraft campaign was conducted over the southeastern US to determine the distribution of THg and Hg(II) at a regional scale. Here, we analyze these observations using the GEOS-Chem chemical transport model with the goals of examining the origin of Hg(II) in the free troposphere and testing the kinetics of the Br-initiated oxidation mechanism.

2 Observations and model used in this study

2.1 The NOMADSS campaign

The NOMADSS aircraft campaign was conducted over the southeastern US from 1 June to 15 July 2013 using the NSF/NCAR C-130 aircraft. We conducted 19 research flights of 4–7 h duration each out of Smyrna, Tennessee (36°1’N, 86°31’W) (Fig. 1). NOMADSS was part of the larger Southeast Atmosphere Study (SAS), a collaborative effort to characterize the atmospheric composition over the southeastern US (http://www.eol.ucar.edu/field_projects/sas). One of the objectives of the NOMADSS campaign, and the focus of this study, was to quantify the tropospheric distribution of mercury species. Mercury measurements were made using
the University of Washington’s Detector of Oxidized Mercury Species (UW-DOHGS), which is currently the only instrument capable of making simultaneous measurements of total and elemental mercury concentrations onboard an aircraft platform at a relatively high time resolution of 2.5 min (Swartzendruber et al., 2009; Lyman and Jaffe, 2012; Ambrose et al., 2013, 2015). In addition, the C-130 aircraft was equipped with other instruments summarized in Table 1.

This paper complements several other papers focused on the analysis of mercury measurements during NOMADSS. In particular, Gratz et al. (2015a) present an analysis of the high Hg(II) concentrations and BrO concentrations observed on one NOMADSS flight (research flight 6). Ambrose et al. (2015) use the NOMADSS observations to calculate enhancement ratios of Hg in the plumes of six coal-fired power plants, and compare them to the Hg emission ratios reported in the emission inventories. Song et al. (2014) combine the NOMADSS observations with results from the GEOS-Chem model to constrain Hg emissions from land and ocean sources. Gratz et al. (2015b) use the NOMADSS observations over Lake Michigan and a plume dispersion model to quantify the outflow of Hg from the Chicago/Gary industrial area.

2.2 The UW-DOHGS instrument

The UW-DOHGS is a dual channel instrument that simultaneously measures concentrations of THg and Hg(0). Ambrose et al. (2015) discuss the configuration of the instrument during NOMADSS, in-flight calibration as well as pre- and post-campaign laboratory tests. Briefly, the UW-DOHGS draws a fast flow of ambient air from a rear-facing aircraft inlet heated to 110 °C to facilitate transmission of Hg(II) compounds. Two Tekran® 2537B Hg vapor analyzers subsample the inlet flow at 1 L min⁻¹ (at 0 °C and 1 atm) through parallel channels. The Tekran® analyzers sample Hg(0) by Au-amalgamation pre-concentration on paired sample cartridges, with quantification by cold vapor atomic fluorescence spectrometry (CV-AFS). The sample integration time and measurement time resolution are both 2.5 min. In the THg channel, Hg(II) compounds are reduced to Hg(0) by passing the sampled air through a quartz pyrolyzer heated to 650 °C. In the Hg(0) channel, Hg(II) (in gas and particle-bound phases) is removed using either a quartz wool trap (first fourteen flights) or a cation exchange membrane (remaining five flights). Hg(II) concentrations are calculated from the difference between the THg and Hg(0) measurements. In comparison, the Tekran® 2537-1130-1135 speciation system uses KCl denuders to capture gas-phase oxidized mercury, which is subsequently thermally desorbed as elemental mercury for analysis (Landis et al., 2002). The measurement cycle of the Tekran® speciation system is 30 min or longer, compared to the 2.5 min cycle for the UW-DOHGS. The UW-DOHGS is not affected by O₃-interference, unlike the Tekran® system (Lyman et al., 2010; Ambrose et al., 2013; McClure et al., 2014). A limitation of UW-DOHGS is that the quartz wool traps can release Hg(II) in humid conditions (Ambrose et al., 2013, 2015), which decreased the number of Hg(II) observations in the boundary layer during the first 14 flights. This was not a problem on the later five flights when cation exchange membranes were used in place of quartz wool.

During NOMADSS, the 1σ uncertainty in THg and Hg(0) was 8–10%, and the detection limit (DL) was < 0.05 ng m⁻³. For Hg(II), the 1σ uncertainty varied between 38 pg m⁻³ (at THg of 1.2 ng m⁻³) and 55 pg m⁻³ (at THg of 2.8 ng m⁻³). The Hg(II) DL is calculated using the “same air” configuration, in which the Hg(II) filter is bypassed and both analyzers sample the same air downstream of the pyrolyzer in the THg channel (Ambrose et al., 2013, 2015). The 3σ Hg(II) DL for the campaign ranged between 58 and 228 pg m⁻³.

The UW-DOHGS instrument was operational during all 19 flights of the NOMADSS campaign, and concentrations of THg were measured continuously during each flight, except during calibration cycles. A total of 2381 (2.5 min average) observations of THg were made during the campaign. Hg(II) concentrations could be determined for only ~60% of the time (1503 observations), because of periodic in-flight calibration cycles and because of reduced retention efficiency of the Hg(II) traps during some flight segments in the moist boundary layer. For the entire NOMADSS campaign, 35% of the Hg(II) measurements (528 points out of 1503) were above the instrument’s DL. Here and in the rest of the paper, we use ADL (Above Detection Limit) observations to refer to Hg(II) measurements above the instrument’s DL and BDL (Below Detection Limit) for Hg(II) measurements below the DL. In the boundary layer (defined here as altitude < 2 km and water vapor > 8 g kg⁻¹), 87% of the 532 observations were BDL. In the free troposphere, 53% out of the 971 observations were BDL.
As more than half the Hg(II) observations during NOMADSS were BDL, we follow the recommendation of Helsel (2011) and use the robust regression on order statistics (ROS) method to assign values for BDL observations. The ROS method assumes a lognormal distribution for the observations, and estimates the distribution’s parameters using ADL measurements. The BDL values are then estimated using the fitted distribution. The ROS method is applied to calculate Hg(II) means and standard deviations reported in Tables 3 and 4 and Figs. 2–5.

2.3 The GEOS-Chem model

2.3.1 General description

The GEOS-Chem global 3-D atmospheric chemical transport model (www.geos-chem.org) is driven by meteorological fields from the NASA Global Modeling and Assimilation Office (GMAO) Goddard Earth Observing System (GEOS) Version 5 Forward Processing (FP). The GEOS-5 FP system consists of a general circulation model (GCM) coupled with a data assimilation (DAS) system (Rienecker et al., 2008), with a native horizontal resolution of 0.25° longitude and 72 vertical levels up to 0.01 hPa. The GEOS-5 FP meteorological fields are archived either at 1 or 3 h intervals, depending on the variables.

GEOS-Chem includes advection (Lin and Rood, 1996), convective transport (Wang et al., 2007), turbulent mixing (Lin and McElroy, 2010), wet deposition (Liu et al., 2001; Wang et al., 2011; Amos et al., 2012), and dry deposition (Wang et al., 1998; Zhang et al., 2011) of chemical species. The full-chemistry simulation includes an up-to-date chemical mechanism for gas-phase and heterogeneous reactions of HO\(_x\)-VOC-O\(_3\)-aerosols in the troposphere (Bey et al., 2001) with the recent addition of bromine chemistry (Parrella et al., 2012). We use here GEOS-Chem model version 9-02.

The GEOS-Chem Hg model simulates the emissions, chemistry, transport, and deposition of Hg(0) and Hg(II) in the atmosphere (Selin et al., 2007) with updates from Holmes et al. (2010), Amos et al. (2012), and Zhang et al. (2012), coupled with a 2-D ocean model (Strode et al., 2007; Sorensen et al., 2010) and a 2-D land model (Selin et al., 2008; Holmes et al., 2010). The model includes prescribed emissions from biomass burning and geogenic activity (Holmes et al., 2010). The global GEOS-Chem Hg simulation, with a resolution of 4° latitude × 5° longitude, was evaluated by Holmes et al. (2010), who found the modeled THg (1.71 ± 0.5 ng m\(^{-3}\)) to be in good agreement with observations (1.86 ± 1.0 ng m\(^{-3}\); correlation coefficient = 0.9) at 39 land-based sites across the globe. Simulated THg concentrations were about 10% higher than THg concentrations measured during three aircraft campaigns (INTEX-B, CARIBIC, and ARCTAS), but modeled vertical profiles were consistent with observations in the troposphere. Zhang et al. (2012) developed a nested-grid Hg simulation in GEOS-Chem, using a higher horizontal resolution (0.5° × 0.667°) over North America. They found that the average annual modeled wet deposition (7.2 ± 3.2 µg m\(^{-2}\) yr\(^{-1}\)) at 95 Mercury Deposition Network (MDN; http://nadp.sws.uiuc.edu/mdn/) sites was close to the observations (8.8 ± 3.6 µg m\(^{-2}\) yr\(^{-1}\)) and showed a correlation coefficient of 0.78. The modeled annual mean THg concentrations (1.42 ± 0.11 ng m\(^{-3}\)) were unbiased with respect to the observations (1.46 ± 0.11 ng m\(^{-3}\)) at 19 surface-based sites. While the modeled gaseous Hg(II) concentrations at these sites were higher than the observations by a factor of 1.5, the model captured the observed seasonal cycle with higher concentrations during spring and summer at most sites.

2.3.2 Updates to Hg emissions in GEOS-Chem

We have updated the global anthropogenic Hg emissions to the United Nations Environment Programme (UNEP)/Arctic Monitoring and Assessment Program (AMAP) 2010 (http://www.amap.no/mercury-emissions/datasets), and over North America we use the U.S. EPA National Emissions Inventory (NEI) 2011 (http://www.epa.gov/ttnchie1/net/2011inventory.html) and Environment Canada’s National Pollutant Release Inventory (NPR) 2011 emission inventories. The specification of anthropogenic emissions is assumed to be 90% Hg(0) : 10% Hg(II) from all anthropogenic sources based on Zhang et al. (2012) and Kos et al. (2013). For 2013, the GEOS-Chem global emission of mercury is 1850 Mg a\(^{-1}\) from anthropogenic sources, 5025 Mg a\(^{-1}\) from natural sources, and 5025 Mg a\(^{-1}\) from natural sources.
sources and re-emission, and 225 Mg a\(^{-1}\) from biomass burning. For 1 June–15 July 2013, the anthropogenic, natural, and biomass burning emissions over North America are 130, 870, and 22 kg d\(^{-1}\), respectively.

### 2.3.3 Updates to Hg chemistry in GEOS-Chem

The oxidation of Hg(0) is modeled as a two-step reaction initiated by Br radicals as originally implemented in GEOS-Chem by Holmes et al. (2010), based on the work ofGoodsite et al. (2004); Donohoue et al. (2006), and Balabanov et al. (2005):

\[
\begin{align*}
\text{Hg}(0) + \text{Br} &\rightarrow \text{HgBr} & (\text{R1}) \\
\text{HgBr} + \text{M} &\rightarrow \text{Hg}(0) + \text{Br} + \text{M} & (\text{R2}) \\
\text{HgBr} + \text{Br} &\rightarrow \text{Hg}(0) + \text{Br}_2 & (\text{R3}) \\
\text{HgBr} + \text{X} &\rightarrow \text{Hg}(\text{II}) & (X = \text{OH, Br, HO}_2, \text{NO}_2, \text{BrO}) & (\text{R4})
\end{align*}
\]

We use the recently corrected rate constant for the HgBr dissociation Reaction (R2) (Goodsite et al., 2012). Holmes et al. (2010) had assumed Br and OH as the second-step oxidants (Reaction R4). In addition, based on the recommendations of Dibble et al. (2012), we have updated this mechanism to include HO\(_2\), NO\(_2\), and BrO as the second-step oxidants. The rates constants used here are

\[
\begin{align*}
\frac{k_1}{M} &= 1.46 \times 10^{-32} \times \left(\frac{T}{298}\right)^{-1.86} [\text{M}] \text{ cm}^{-3} \text{ molecule}^{-1} \text{ s}^{-1} \\
&\quad \text{(Donohoue et al., 2006)} \\
\frac{k_2}{s} &= 2.67 \times 10^{44} \times \left(\frac{7292}{T}\right) \left(\frac{T}{298}\right)^{1.76} \times k_1 \text{ s}^{-1} \\
&\quad \text{(Goodsite et al., 2012)} \\
\frac{k_3}{s} &= 3.9 \times 10^{-11} \text{ cm}^{-3} \text{ molecule}^{-1} \text{ s}^{-1} \\
&\quad \text{(Balabanov et al., 2005)} \\
\frac{k_4}{s} &= 2.5 \times 10^{-10} \times \left(\frac{T}{298}\right)^{-0.57} \text{ cm}^{-3} \text{ molecule}^{-1} \text{ s}^{-1} \\
&\quad \text{(Goodsite et al., 2012; Dibble et al., 2012)}
\end{align*}
\]

The reduction of Hg(II) to Hg(0) is assumed to occur in cloud droplets in the presence of sunlight. The reduction is assumed to be proportional to the photolysis frequency of NO\(_2\), and the coefficient is estimated by constraining the model results with the observed mean burden of THg in the troposphere (Holmes et al., 2010). The uptake of Hg(II) by sea-salt particles is simulated as a kinetic mass transfer process (Holmes et al., 2010). For non sea-salt aerosol particles, gas-particle partitioning is simulated as an equilibrium process based on an empirical relationship (Amos et al., 2012). The modeled wet and dry deposition of gas-phase Hg(II) is analogous to that of HNO\(_3\), and of particle-bound Hg(II) to that of sulfate particles.

### 2.3.4 Simulations conducted for this study

In this study, the GEOS-Chem Hg model is run in a one-way nested-grid configuration, with the native horizontal resolution (0.25\(^\circ\) \times 0.3125\(^\circ\)) over North America, and a coarser resolution (2\(^\circ\) \times 2.5\(^\circ\)) globally. Both the coarse- and fine-grid models share the same vertical resolution with 47 layers (13 layers in the bottom 2 km, and 16 layers between 2 and 10 km). The global Hg model is spun-up with a 3-year simulation, and is then used to generate the initial and boundary conditions for the nested-domain, and to calculate the annual global mercury budget. Monthly concentrations of Br, BrO, OH, HO\(_2\), NO\(_2\), O\(_3\), and aerosols are obtained from a 1-year 4\(^\circ\) \times 5\(^\circ\)  global HO\(_4\)–NO\(_4\)–VOC–O\(_3\)–BrO\(_4\)–aerosols tropospheric chemistry GEOS-Chem simulation. We run the nested 0.25\(^\circ\) \times 0.3125\(^\circ\) Hg model for North America from 1 June to 15 July 2013. For comparison to the NOMADSS aircraft observations, the model is sampled along the flight track at the 2.5 min time step of the UW-DOHGS instrument.

In addition to the BASE simulation with the Hg(0) + Br chemistry described above (Sect. 2.3.3), we perform five sensitivity simulations (Table 2), and focus on two of them. In the first sensitivity simulation (FastK) we increase the GEOS-Chem Br and BrO concentrations by a factor of 3 in the 45\(^\circ\) S to 45\(^\circ\) N latitude band between 750 hPa and the tropopause. Parrella et al. (2012) found that the GEOS-Chem BrO is within the BrO measurement uncertainties of the GOME-2 satellite in the polar regions and at mid-latitudes, but in the tropics the GEOS-Chem BrO is too low by a factor of 2–4 throughout the year (Fig. 5 of Parrella et al., 2012). Furthermore, recent aircraft-based observations of BrO in the tropical and subtropical free troposphere of the Southern Hemisphere reported concentrations which were 2–4 times higher than predicted by the GEOS-Chem model (Wang et al., 2015).

In the second sensitivity simulation (FastK), we replace the Donohoue et al. (2006) rate constant for Reaction (R1) by the faster rate constant measured by Ariya et al. (2002). Following Dibble et al. (2012), we proportionally increase the rate constant for the backward Reaction (R2) as follows:

\[
\begin{align*}
\frac{k_{1,\text{FastK}}}{M} &= 3.6 \times 10^{-12} \times \left(\frac{[\text{M}]_{T,\text{p}}}{[\text{M}]_{273.15 \text{ K,1 atm}}}\right) \text{ cm}^{-3} \text{ molecule}^{-1} \text{ s}^{-1} \\
&\quad \text{(Ariya et al., 2002)} \\
\frac{k_{2,\text{FastK}}}{s} &= k_2 \times \frac{k_{1,\text{FastK}}}{k_1} \text{ s}^{-1} \\
&\quad \text{(Goodsite et al., 2012; Dibble et al., 2012)}
\end{align*}
\]

Relative to the BASE simulation, \(k_{1,\text{FastK}}\) and \(k_{2,\text{FastK}}\) are 10 times larger near the surface and 5 times larger in the upper troposphere (10–12 km). For both these sensitivity simulations, the in-cloud reduction rate is increased such that the tropospheric burden and lifetime of THg is similar to the
The vertical and horizontal distributions of THg and Hg(II) concentrations observed during NOMADSS are presented in Figs. 2 and 3. We exclude from these figures and the rest of our analysis fresh pollution plumes, where NOx, SO2, or C2H4 exceed 2 ppbv. This eliminates 6 % of the THg and 4 % of the Hg(II) measurements. The mean and standard deviation of the observed THg concentrations was 1.49 ± 0.16 ng m⁻³ (Fig. 3a). THg concentrations decrease slightly with altitude, from 1.54 ng m⁻³ near the surface to 1.38 ng m⁻³ at 6–7 km altitude (Fig. 2a), which is in agreement with previous aircraft-based measurements of THg over North America (Talbot et al., 2007; Swartzendruber et al., 2009; Mao et al., 2010). The variability in THg concentration is small, with standard deviations at different levels ranging from 6 to 15 % of the mean concentrations. The weak vertical gradient and the low standard deviation of the THg concentrations are consistent with the long tropospheric lifetime of THg.

For Hg(II), the observed mean concentration was 212 ± 112 pg m⁻³, for ADL measurements (35 % of measurements) (Table 3). When we include BDL values using the ROS method (see Sect. 2.2), the mean Hg(II) concentration is 110 ± 103 pg m⁻³ (Table 3). In order to assess the overall mean distribution of Hg(II) during NOMADSS, Figs. 2b and 3c display the Hg(II) statistics that include BDL values estimated with the ROS method. Figure 2b shows that observed Hg(II) concentrations increased from 40 pg m⁻³ at 0–1 km altitude to 200 pg m⁻³ at 6–7 km (Fig. 2b). Large enhancements in Hg(II) concentrations, of up to 500 pg m⁻³, were observed at 5–7 km during two research flights (RF) over Texas (RF-06 and RF-09, box 1 in Figs. 2b and 3d). Concentrations of up to 680 pg m⁻³ were observed at 1–3 km on one flight over the Atlantic Ocean (RF-16). We will discuss these flights in more detail in Sect. 6. Table S1 in the Supplement presents a summary of Hg(II) observations on all flights. Previous Hg(II) measurements at high-elevation ground sites and from aircraft showed that high free tropospheric Hg(II) concentrations were generally observed in air masses that were dry and clean (Landis et al., 2005; Swartzendruber et al., 2006, 2009; Fain et al., 2009; Sheu et al., 2010). To examine whether this relation holds for the NOMADSS observations, we classify all THg and Hg(II) observations
Table 3. Chemical characteristics of NOMADSS observations classified in four air mass categories.

<table>
<thead>
<tr>
<th></th>
<th>All observations</th>
<th>low RH/low CO</th>
<th>low RH/high CO</th>
<th>high RH/low CO</th>
<th>high RH/high CO</th>
</tr>
</thead>
<tbody>
<tr>
<td>No of THg observations</td>
<td>2381</td>
<td>233</td>
<td>551</td>
<td>212</td>
<td>1385</td>
</tr>
<tr>
<td>THg observations (ng m⁻³)</td>
<td>1.49 ± 0.16</td>
<td>1.35 ± 0.15</td>
<td>1.48 ± 0.11</td>
<td>1.44 ± 0.20</td>
<td>1.53 ± 0.15</td>
</tr>
<tr>
<td>No of Hg(II) observations</td>
<td>1503</td>
<td>184</td>
<td>414</td>
<td>159</td>
<td>746</td>
</tr>
<tr>
<td>(ADL)</td>
<td>(528)</td>
<td>(132)</td>
<td>(244)</td>
<td>(47)</td>
<td>(105)</td>
</tr>
<tr>
<td>Hg(II) all observations (pg m⁻³)</td>
<td>110 ± 103</td>
<td>239 ± 141</td>
<td>146 ± 81</td>
<td>108 ± 123</td>
<td>48 ± 57</td>
</tr>
<tr>
<td>(ADL)</td>
<td>(212 ± 112)</td>
<td>(289 ± 136)</td>
<td>(189 ± 76)</td>
<td>(249 ± 140)</td>
<td>(155 ± 73)</td>
</tr>
<tr>
<td>Altitude (km)</td>
<td>2.6 ± 1.9</td>
<td>4.6 ± 1.8</td>
<td>3.9 ± 1.2</td>
<td>2.9 ± 2.2</td>
<td>1.7 ± 1.5</td>
</tr>
<tr>
<td>RH (%)</td>
<td>49 ± 27</td>
<td>16 ± 10</td>
<td>15 ± 9</td>
<td>68 ± 17</td>
<td>66 ± 13</td>
</tr>
<tr>
<td>CO (ppbv)</td>
<td>107 ± 33</td>
<td>65 ± 4</td>
<td>97 ± 13</td>
<td>65 ± 3</td>
<td>124 ± 30</td>
</tr>
<tr>
<td>O₃ (ppbv)</td>
<td>55 ± 14</td>
<td>52 ± 16</td>
<td>63 ± 17</td>
<td>43 ± 16</td>
<td>54 ± 10</td>
</tr>
<tr>
<td>NOₓ (ppbv)</td>
<td>158 ± 156</td>
<td>55 ± 32</td>
<td>67 ± 34</td>
<td>44 ± 39</td>
<td>232 ± 170</td>
</tr>
<tr>
<td>CH₂O (ppbv)</td>
<td>1.8 ± 1.4</td>
<td>0.5 ± 0.3</td>
<td>0.7 ± 0.3</td>
<td>0.9 ± 0.4</td>
<td>2.7 ± 1.3</td>
</tr>
</tbody>
</table>

The four air mass categories are based on thresholds of RH = 35 % and CO = 75 ppbv. The mean and standard deviation for each category are indicated. a Number of 2.5 min THg samples. b Total number of 2.5 min Hg(II) samples, including samples below the detection limit (BDL). The number in parenthesis indicates the number of 2.5 min Hg(II) samples above the detection limit (ADL). c Mean Hg(II) concentration and standard deviation for all observations, including BDL as estimated using the ROS method. d Mean Hg(II) concentration and standard deviation for ADL observations.

Figure 2. Vertical profiles of (a) THg and (b) Hg(II) during NOMADSS. Individual 2.5 min observations are shown with grey circles (for THg and Hg(II) measured using the quartz wool filter) and blue squares (for Hg(II) measured using the cation exchange membrane filter). Observations above the detection limit (ADL) are indicated as filled circles/squares, while observations below the detection limit (BDL) are shown as open circles/squares and are estimated using the regression on order statistics (ROS) method. The means and standard deviations calculated for 1 km vertical bins are shown for the observations (black diamonds and error bars) and the BASE GEOS-Chem simulation (green line and shading). The numbers on the right hand side of each panel indicate the number of 2.5 min observations in each 1 km bin. For Hg(II), the second number in parenthesis indicates the number of ADL observations. The areas marked as “1” and “2” highlight measurements of high Hg(II) concentrations and are referenced in the text and Fig. 3.

Based on the measured relative humidity (RH) and CO concentrations (Table 3). We use thresholds of RH = 35 % and CO = 75 ppbv to classify the observations into four categories: “high RH/high CO”, “high RH/low CO”, “low RH/high CO”, and “low RH/low CO”. When RH or CO measurements were not available (18 % of the observations), we use the GEOS-Chem simulated RH values and CO concentrations. The ROS procedure is used to assign values for BDL observations separately for each category.

Using this method, Hg(II) observations partition into the four categories as follows (Table 3): “low RH/low CO” (12 % of observations), “low RH/high CO” (28 %), “high RH/low CO” (10 %), “high RH/high CO” (50 %). The highest mean concentrations of Hg(II) are found in the “low RH/low CO” category (239 pg m⁻³ for all observations, 289 pg m⁻³ for ADL observations). Observations in this category had the lowest average THg concentrations (1.35 ng m⁻³), and were relatively clean with low mixing ratios for CH₂O (0.5 ppbv), NOₓ (55 pptv), O₃ (52 ppbv), and a mean RH of 16 %. These air samples were observed mainly during the high-altitude (5–7 km) flights over Texas (RF-06 and RF-09), and at 1–3 km over the Atlantic on RF-16. The chemical characteristics and HYbrid Single-Particle Lagrangian Integrated Trajectory (HYSLIGHT) (Draxler and Hess, 1998) back trajectories for these “low RH/low CO” air masses indicate subsidence from the clean subtropical upper troposphere (Sect. 6).

The “low RH/high CO” category displays moderate enhancements in Hg(II) concentrations, with a mean of 146 pg m⁻³ (189 pg m⁻³ for ADL observations). Back trajectories (not shown here) indicate transport from high latitudes (> 60°N) accompanied by subsidence. The mean concentrations of NOₓ (67 pptv) and CH₂O (0.7 ppbv) were low and similar to values for the “low RH/low CO” category, indicating relatively clean air masses but originating from the
mid- and upper-troposphere at high latitudes instead of tropical latitudes as in the “low RH/low CO” category.

Most of the remaining Hg(II) observations (50 % of observations) fall in the “high RH/high CO” category. These observations were collected mainly in the continental boundary layer, with moderately high concentrations of CO (68 ppbv), CH₂O (2.7 ppbv), and NO₃ (232 pptv). In this category, the mean concentration of THg is 1.53 ng m⁻³, while that of Hg(II) is 48 pg m⁻³ (155 pg m⁻³ for ADL observations, which account for only 14 % of observations in this category). The “high RH/low CO” category has lower CO (68 ppbv), CH₂O (0.9 ppbv), NO₃ (44 pptv), and O₃ (43 ppbv). These samples were observed mostly near the marine boundary layer over the Atlantic Ocean (RF-14 and RF-16), but some were observed at high altitudes possibly in air detrained from clouds. In this category, the mean concentration of THg is 1.44 ng m⁻³, and that of Hg(II) is 108 pg m⁻³ (249 pg m⁻³ for ADL observations, 30 % of the observations).

The observations of Hg(II) during NOMADSS are similar to previous observations of Hg(II) in terms of their magnitude, vertical profiles, and airmass characteristics. During five summertime flights in the Pacific Northwest, Swartzendruber et al. (2009) also found RGM (gaseous component of Hg(II)) to be generally below the instrument’s DL of 80–160 pg m⁻³. They found high RGM air masses (200–500 pg m⁻³) on two flights at 2–6 km altitude. These air masses had low aerosol extinction coefficient, indicating either larger production of RGM at higher altitudes or loss of RGM in the presence of aerosol particles (Swartzendruber et al., 2009). On 28 flights from August 2012 to June 2013, Brooks et al. (2014) found that mean RGM and PBM (particle-bound component of Hg(II)) concentrations at 0–6 km over Tennessee, USA were 34 and 30 pg m⁻³, respectively. Highest concentrations on each flight were always observed at 2–4.5 km altitude. RGM concentrations exhibited a minimum in the winter months and a maximum in the summer months. On one flight in June 2013, the vertical profile of RGM showed a steep increase from the surface (∼5 pg m⁻³) to 4 km altitude (120 pg m⁻³) followed by decrease between 4 and 6 km (70 pg m⁻³) (Brooks et al., 2014). At the high elevation Mt. Bachelor Observatory (2.7 km a.s.l.) in Oregon, the mean RGM concentration during May–August 2005 was 43 pg m⁻³ (Swartzendruber et al., 2006). Low RGM concentrations (<50 pg m⁻³) were seen in boundary layer air during the day, and higher concentrations were seen in free-tropospheric air during the night. The highest 10 % of nighttime RGM concentrations were between 160 and 600 pg m⁻³, and were measured in air with low RH and low CO, in similarity to the observations of high Hg(II) in our “low RH/low CO” category during NOMADSS.

4 Comparison to the BASE GEOS-Chem Hg simulation

THg concentrations simulated in the BASE GEOS-Chem model (1.51 ± 0.08 ng m⁻³) agree well with observations (1.49 ± 0.16 ng m⁻³) (Fig. 3a, b). The model captures the lower THg concentrations observed over central Texas, South Carolina, and part of the Atlantic Ocean, but overestimates THg concentrations over the Ohio River Valley by ∼20 %, possibly because of recent reductions in power plant emissions of mercury not captured in the 2011 NEI inventory used in our study. The model reproduces the vertical profile of observed THg, decreasing from 1.57 ng m⁻³ (observed: 1.54 ± 0.15 ng m⁻³) near the surface to 1.43 ng m⁻³ (observed: 1.38 ± 0.19 ng m⁻³) at 6–7 km altitude. However, we note that the variability of modeled THg in each 1 km bin (shaded green area in Fig. 2a) is a factor of 2 to 5 lower than the observed variability. Overall, the GEOS-Chem THg concentrations are within ±10 % of the measured concentrations for 70 % of the observations, which indicates that the model provides reasonable simulations of the emissions and mean lifetime of THg in the lower and middle troposphere during NOMADSS.
Table 4. Modeled THg and Hg(II) concentrations in the three GEOS-Chem Hg simulations for the four air mass categories.

<table>
<thead>
<tr>
<th>All observations</th>
<th>“low RH/low CO”</th>
<th>“low RH/high CO”</th>
<th>“high RH/low CO”</th>
<th>“high RH/high CO”</th>
</tr>
</thead>
<tbody>
<tr>
<td>Observed Hg(II) (pg m⁻³)</td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>All</td>
<td>110 ± 103</td>
<td>239 ± 141</td>
<td>146 ± 81</td>
<td>108 ± 123</td>
</tr>
<tr>
<td>(ADL)</td>
<td>212 ± 112</td>
<td>289 ± 136</td>
<td>189 ± 76</td>
<td>249 ± 140</td>
</tr>
<tr>
<td>BASE model Hg(II) (pg m⁻³)</td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>All</td>
<td>49 ± 41</td>
<td>99 ± 48</td>
<td>76 ± 36</td>
<td>20 ± 25</td>
</tr>
<tr>
<td>(ADL)</td>
<td>67 ± 44</td>
<td>96 ± 51</td>
<td>75 ± 34</td>
<td>14 ± 18</td>
</tr>
<tr>
<td>(BDL)</td>
<td>39 ± 44</td>
<td>105 ± 37</td>
<td>78 ± 39</td>
<td>22 ± 27</td>
</tr>
<tr>
<td>3× Br model Hg(II) (pg m⁻³)</td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>All</td>
<td>62 ± 72</td>
<td>162 ± 104</td>
<td>91 ± 67</td>
<td>28 ± 41</td>
</tr>
<tr>
<td>(ADL)</td>
<td>98 ± 94</td>
<td>176 ± 116</td>
<td>94 ± 76</td>
<td>19 ± 30</td>
</tr>
<tr>
<td>(BDL)</td>
<td>43 ± 94</td>
<td>124 ± 47</td>
<td>86 ± 51</td>
<td>31 ± 45</td>
</tr>
<tr>
<td>FastK model Hg(II) (pg m⁻³)</td>
<td></td>
<td></td>
<td></td>
<td></td>
</tr>
<tr>
<td>All</td>
<td>80 ± 98</td>
<td>208 ± 144</td>
<td>128 ± 95</td>
<td>18 ± 25</td>
</tr>
<tr>
<td>(ADL)</td>
<td>125 ± 120</td>
<td>216 ± 160</td>
<td>128 ± 89</td>
<td>16 ± 17</td>
</tr>
<tr>
<td>(BDL)</td>
<td>55 ± 120</td>
<td>189 ± 88</td>
<td>129 ± 102</td>
<td>19 ± 28</td>
</tr>
</tbody>
</table>

| Observed THg (ng m⁻³) |                 |                 |                 |                 |
| All                | 1.49 ± 0.16     | 1.35 ± 0.15     | 1.48 ± 0.11     | 1.44 ± 0.20     | 1.53 ± 0.15      |
| BASE model THg (ng m⁻³) |           |                 |                 |                 |                 |
| All                | 1.51 ± 0.08     | 1.43 ± 0.06     | 1.50 ± 0.05     | 1.40 ± 0.03     | 1.55 ± 0.07      |
| 3× Br model THg (ng m⁻³) |       |                 |                 |                 |                 |
| All                | 1.51 ± 0.11     | 1.40 ± 0.09     | 1.50 ± 0.08     | 1.36 ± 0.04     | 1.55 ± 0.10      |
| FastK model THg (ng m⁻³) |     |                 |                 |                 |                 |
| All                | 1.52 ± 0.09     | 1.44 ± 0.07     | 1.52 ± 0.06     | 1.40 ± 0.04     | 1.55 ± 0.08      |

* Model values corresponding to ADL observations.  b Model values corresponding to BDL observations.

Because the majority of Hg(II) measurements are below the 58–228 pg m⁻³ DL of the UW-DOHGS instrument, we consider ADL and BDL observations separately. Table 4 shows that when observed Hg(II) are BDL (65 % of measurements), the mean Hg(II) concentrations predicted with the BASE simulation (39 ± 34 pg m⁻³) are indeed below the instrument’s DL. For ADL observations, the BASE model underestimates observations by a factor of 3 (model: 67 ± 44 pg m⁻³; observations: 212 ± 112 pg m⁻³). The model predicts an increase in Hg(II) concentrations with altitude which is much smaller than observed (Fig. 2b). While underestimating the observed mean concentrations, the model captures the factor of 3–6 increase in Hg(II) in the “low RH” air mass categories relative to the “high RH” categories (Tables 3 and 4).

Figure 5 shows scatter plots between observed and modeled Hg(II) concentrations for the “low RH/low CO” and “low RH/high CO” categories. We find that the GEOS-Chem BASE simulation has a large negative (−47 to −58 %) normalized mean bias (NMB = \( \frac{\sum_i (M_i - O_i)}{\sum_i O_i} \)), where \( O_i \) and \( M_i \) are observed and simulated values, and for \( O_i < DL \), \( \sum_i O_i \) is calculated using the ROS procedure. Furthermore, about 60 % of modeled Hg(II) values are within a factor of two of the observations (FAC2 = fraction of points where \( 0.5 \leq M_i / O_i \leq 2 \), for \( O_i \leq DL \), we assume \( 0.5 \leq M_i / O_i \leq 2 \) if \( M_i < 2 \times DL \)). If we consider only ADL observations, FAC2 decreases to 28–39 %.

Figure S2 (in the Supplement) displays scatterplots for the “high RH” categories. For ADL observations, the NMB is between −75 and −94 % for these categories (Fig. S2 and Table 4). We note that for these two “high RH” categories a conclusive evaluation of the model performance is difficult, however, because of the large fraction of BDL observations.

Considering the systematic model underestimate of observed Hg(II) concentrations, particularly in dry air masses where reduction and wet deposition are suppressed, we hypothesize that the bias in the Hg(II) concentrations is because the model simulated oxidation of Hg(0) to Hg(II) is too slow. We test this hypothesis in the next section by examining two sensitivity simulations.

5 Sensitivity studies with enhanced Hg(0) oxidation

As described in Sect. 2.3.4, we conduct two sensitivity simulations: in the 3× Br simulation, we evaluate the model’s response to an increase in Br concentrations, and in the FastK simulation, we evaluate the model’s response to a faster oxidation rate constant. The results of these simulations are summarized in Table 4 and detailed comparisons to observations are presented in Figs. 4 and 5. Both the 3× Br and FastK
Figure 4. Mean vertical profiles of (a) THg and (b) Hg(II) concentrations averaged in 1 km vertical bins during NOMADSS. Observations are shown with black diamonds (error bars indicate the standard deviation). BDL Hg(II) observations are estimated using the ROS method for each vertical bin. The means and standard deviations for the three model simulations are shown as lines and shading: BASE (green), 3× Br (blue), and FastK (orange).

Simulations reproduce the observed mean concentration (Table 4) and vertical profile (Fig. 4a) of THg as we compensate the increase in Hg(0) oxidation with an increase in Hg(II) reduction rate to maintain the THg burden.

The 3× Br and FastK simulations predict a stronger vertical gradient in Hg(II) concentrations, in closer agreement with observations (Fig. 4). Above 5 km, Hg(II) concentrations in the 3× Br (165 ± 104 pg m⁻³) and FastK (184 ± 156 pg m⁻³) simulations show significantly better agreement with observations (189 ± 103 pg m⁻³) relative to the BASE simulation (97 ± 46 pg m⁻³) (Fig. 4b). For the “low RH/low CO” category, average modeled Hg(II) concentrations increase from 99 ± 48 pg m⁻³ (BASE) to 162 ± 104 pg m⁻³ (3× Br) and 208 ± 144 pg m⁻³ (FastK), compared to the observed 239 ± 141 pg m⁻³. The modeled NMB for all Hg(II) observations in the “low RH/low CO” category decreases from −58 % (BASE simulation) to −32 % (3× Br) and −12 % (FastK), while the FAC2 index increases from 62 % (BASE) to 69 % (3× Br) and 80 % (FastK) (Fig. 5a–c). However, the sensitivity simulations cannot reproduce the high Hg(II) concentrations observed on RF-16 at 1–3 km (blue circles in Fig. 5). We present a detailed discussion of this flight in Sect. 6.

In the “low RH/high CO” category, the 3× Br Hg(II) concentrations (91 ± 67 pg m⁻³) are ∼20 % higher than the BASE model (76 ± 36 pg m⁻³), but lower than the observed concentrations (146 ± 81 pg m⁻³) (Table 4). In the FastK simulation, Hg(II) concentrations increase to 128 ± 95 pg m⁻³, improving the NMB (FastK: −11 %, 3× Br: −37 %, BASE: −47 %) (Fig. 5d–f). For this category, the 3× Br Hg(II) concentrations are not much higher than the BASE simulation because in the 3× Br we use higher Br concentration only between 45° N and 45° S, whereas most of the “low RH/high CO” air masses originated from high latitudes. The FAC2 index is higher in the FastK simulation (75 %) compared to the BASE (58 %) and 3× Br (61 %) simulations.

Overall, the factor of 2 decrease in model bias for Hg(II) with the sensitivity simulations particularly for the “low RH/low CO” air samples suggests that the oxidation of Hg(0) in the upper troposphere is 3 (3× Br) to 5 (FastK) times faster than what was considered previously in the GEOS-Chem model.

We ran two additional simulations that included oxidation by OH/O₃ and by BrO, respectively, in addition to the Br-initiated oxidation of the BASE version, to test whether the underestimate in the modeled oxidation results from exclusion of these oxidants. We add the following oxidation reactions to the model:
Hg(0) + O₃ → Hg(II) \quad \text{(R5)}
Hg(0) + OH → Hg(II) \quad \text{(R6)}
Hg(0) + BrO → Hg(II), \quad \text{(R7)}

where \( k_5 = 3.0 \times 10^{-20} \text{cm}^2\text{molecule}^{-1} \text{s}^{-1} \) (Hall, 1995), \( k_6 = 8.7 \times 10^{-14} \text{cm}^2\text{molecule}^{-1} \text{s}^{-1} \) (Sommar et al., 2001), and \( k_7 = 3.0 \times 10^{-14} \text{cm}^2\text{molecule}^{-1} \text{s}^{-1} \) (Spicer et al., 2002).

The inclusion of Hg(0) + O₃ and Hg(0) + OH pathways decreases the global tropospheric chemical lifetime of Hg(0) for June–July 2013 from 3.2 months (BASE simulation) to 0.6 months, which is also lower than the lifetimes in the 3 × Br (1.9 months) and the FastK (0.7 months) simulations. However, the Hg(0) + O₃ and Hg(0) + OH reaction pathways have a relatively small effect on the modeled Hg(II) concentrations in the upper troposphere. The 5–7 km Hg(II) concentrations in this simulation (108 ± 79 pg m⁻³) are similar to the BASE simulation (97 ± 45 pg m⁻³), but the Hg(II) concentrations below 5 km (71 ± 51 pg m⁻³) are higher than the BASE simulation (39 ± 33 pg m⁻³). The increase in oxidation with the Hg(0) + O₃ and Hg(0) + OH pathways is mostly in the lower troposphere, whereas in the upper troposphere, faster oxidation is compensated by faster reduction. While the oxidation rates of Hg(0) + O₃ and Hg(0) + OH reactions have large uncertainties (see Sect. 1), our results indicate that the underestimate in Hg(II) concentrations in the BASE model persists despite the inclusion of the Hg(0) + O₃ and Hg(0) + OH oxidation pathways.

When the Hg(0) + BrO oxidation pathway is added, the June–July global tropospheric chemical lifetime of Hg(0) is 0.9 months. The simulated 5–7 km Hg(II) concentrations (115 ± 75 pg m⁻³) are higher than those in the BASE simulation, but not as high as the 3 × Br and FastK simulations, and much lower than the observations. The NMB for the “low RH/low CO” air masses is −47 % for this simulation, compared to −32 % for the 3 × Br simulation and −12 % for the FastK simulation. Thus, although including the Hg(0) + BrO reaction with the rate from Spicer et al. (2002) brings the model closer to the observations, it does not completely explain the model underestimate of Hg(II) concentrations.

6 Case studies of individual flights

We analyze in more detail three NOMADSS flights during which high concentrations of Hg(II) were observed. For RF-06 and RF-09 over Texas, we trace the origin of the high Hg(II) to transport from the subtropical Pacific anticyclone, while for RF-16 off the South Carolina Coast, enhanced Hg(II) concentrations appear to have been produced in the subtropical Atlantic anticyclone. The time series of the observations and model results along the flight tracks are shown in Figs. 6–8.

One of the goals of RF-06 was to sample in dry air masses with potentially enhanced Hg(II) concentrations. The meteorological forecasts indicated the presence of such an air mass at 6 km altitude over west Texas. After sampling in the boundary layer, the aircraft ascended to 6.8 km altitude, measuring 182–347 pg m⁻³ of Hg(II) between 18:02 and 19:47 UTC (Fig. 6a and d). The Hg(II) DL on this flight was 114 pg m⁻³, and quartz wool was used as the Hg(II) filter on the instrument’s Hg(II) channel. The mean concentration of THg in this air mass was 1.35 ng m⁻³, about 20 % lower than the THg concentrations measured during the rest of the flight (Fig. 6b). The observed low RH (11 %), CO (65 ppbv), O₃ (39 ppbv), NOₓ (53 pptv), CH₂O (204 pptv), and C₃H₈ (36 pptv) indicate a clean and dry air mass (Fig. 6d). The 7-day HYSPLIT back trajectories (Fig. 6e) indicate that the air mass was transported from the subtropical Pacific anticyclone 3 days before it was sampled over Texas (see also Gratz et al., 2015a). The back trajectories show subsidence of the air mass from the upper troposphere (10–12 km) before transport out of the Pacific anticyclone.

Figure 6c shows that the 3 × Br simulation captures the enhancement in Hg(II) concentrations observed over Texas (observations: 251 pg m⁻³, 3 × Br: 306 pg m⁻³), while the BASE model (146 pg m⁻³) is too low and the FastK model (367 pg m⁻³) too high. Outside the region with the high Hg(II) concentration, all three model simulations calculate Hg(II) concentrations that are BDL, and are in agreement with the observations. During this flight, observations indicate high BrO concentrations of 1.9 ± 0.35 pptv (Fig. 6d) (Gratz et al., 2015a). The GEOS-Chem modeled BrO concentrations in this air mass (0.3–0.4 pptv) are a factor of 5 lower than the observations (Fig. 6d), and thus our assumption of higher Br concentrations in the 3 × Br simulation is more consistent with these measurements (Sect. 2.3.4).

The transport pattern of dry air from the subtropical Pacific upper troposphere to the southeastern US persisted for several days and was sampled again on 24 June 2013 during flight RF-09 (Fig. 7a). High Hg(II) concentrations of 200–480 pg m⁻³ were measured between 16:00 and 19:30 UTC as the C-130 aircraft flew five constant-altitude legs between 2.5 and 6.7 km over eastern Texas. The Hg(II) DL on this flight was 94 pg m⁻³, and quartz wool was used as the Hg(II) filter. Low RH (19 %) and low O₃ concentrations (50 ppbv) accompanied the high Hg(II) (Fig. 7c and d). BrO measurements are not available for this flight due to interference of clouds with the instrument. The HYSPLIT back trajectories show transport from the subtropical Pacific anticyclone (Fig. 7e), similar to RF-06. As the C-130 aircraft descended to 1.2 km, high concentrations of Hg(II) (up to 360 pg m⁻³) were observed in the continental boundary layer between 19:30 and 20:00 UTC. This was associated with high THg concentrations of 2.0 ng m⁻³, and was likely due to emissions of both Hg(II) and Hg(0) from nearby sources.
Figure 6. Case study for RF-06 on 19 June 2013. (a) The C-130 flight track with circles color-coded based on observed Hg(II) concentrations (locations along the flight track with no Hg(II) observations are filled in black, and below the detection limit (BDL) observations are filled in white). The background map displays the GEOS-Chem Hg(II) concentrations at 450 hPa for the 3×Br simulation between 17:00–20:00 UTC. The time series of the observations of THg and Hg(II) are shown in panels (b, c). The THg and Hg(II) measurements are represented by filled diamonds with the uncertainty represented by the gray shading. The dashed line represents the DL, and BDL observations are plotted at DL/2 with open circles. Modeled concentrations of THg and Hg(II) are displayed for the BASE (green line), 3×Br (blue), and FastK (orange) simulations. Flight times where high Hg(II) concentrations were observed are highlighted. The time series of flight altitude, RH, CO, O₃, and BrO (observations and model) are shown in panel (d). Panel (e) shows the 7-day HYSPLIT back trajectories for the highlighted sections of the flight. The contours show the NCEP/NCAR Reanalysis 500 hPa geopotential heights on 18Z for the day of the flight. The hatched areas show regions with greater than an average of 10 mm day⁻¹ of surface precipitation for 7 days before the flight.

The mean THg concentrations for RF-09 calculated by the three simulations are between 1.38 and 1.42 ng m⁻³, and are close to the mean of the observations (Fig. 7b). Between 16:00 and 19:00 UTC, the Hg(II) concentrations calculated with the FastK model (302 pg m⁻³) are closer to the observations (321 pg m⁻³) than with the BASE (127 pg m⁻³) and 3×Br (253 pg m⁻³) models (Fig. 7c). However, between 19:00 and 19:30 UTC, the modeled Hg(II) concentrations for the three simulations are considerably lower (40–60 pg m⁻³) than the observed concentrations (300 pg m⁻³). The simulated RH (54 %) during this time was also higher than the observed RH (25 %). A comparison of the observed and modeled vertical profiles of isoprene (not shown) indicates an overestimate in the modeled local boundary layer depth, which could explain the difference between the model and the observations for this section of the flight.

For both RF-06 and RF-09, the GEOS-Chem simulations predict that the high observed Hg(II) concentrations were produced in the upper troposphere of the Pacific anticyclone. The upper troposphere is characterized by fast oxidation of Hg(0) to Hg(II) due to cold temperatures and higher Br concentrations. Furthermore, the cloud-free conditions within the anticyclones prevent removal of Hg(II) by deposition and aqueous reduction, leading to accumulation of Hg(II) (Sect. 8).

6.2 RF-16 (8 July 2013)

During RF-16, the aircraft flew to the South Carolina coast, with the goal of measuring the vertical distribution of mercury over the ocean (Fig. 8a). Of the eight constant altitude legs flown over the Atlantic, four were in the free troposphere, between 1.0 and 4.5 km, and four were in the marine boundary layer (MBL). THg concentrations showed a slight decrease from the MBL (1.58 ng m⁻³) to the top of the vertical profile (1.37 ng m⁻³) (Fig. 8b). Hg(II) was mostly BDL in the MBL, but was high in the free troposphere reaching up to 680 pg m⁻³ with a mean concentration of 450 pg m⁻³ (Fig. 8c). The Hg(II) DL on this flight was 91 pg m⁻³, and the Hg(II) filter was a cation exchange membrane. The free-tropospheric air had a low concentration of CO (65 ppbv) and low RH (33 %) (Fig. 8d). The CHBr₃ concentration was about 1 pptv in the free tropospheric air, compared to about 2 pptv in the MBL. In consistency with the low RH and CHBr₃, the HYSPLIT back trajectories show slow subsidence of the air mass from 4–6 km altitude in the subtropical Bermuda anticyclone (Fig. 8e).
The simulated THg concentrations in the MBL (1.34–1.38 ng m\(^{-3}\)) are lower than the observations (1.58 ng m\(^{-3}\)), suggesting an underestimate in the ocean emission flux (Song et al., 2015) or an overestimate in the deposition flux (Fig. 8b). None of the model simulations capture the enhancements in Hg(II) observed between 17:00 and 18:15 UTC, and again between 19:30 and 20:15 UTC (Fig. 8c). During this flight, observed BrO concentrations remained below the instrument’s DL of 0.9 pptv, and the modeled BrO concentrations were 0.1–0.3 pptv. We performed an additional simulation increasing the modeled BrO concentrations in the free troposphere of the Bermuda anticyclone to 0.9 pptv (with a proportional increase in Br radical concentrations) for the FastK simulation (FastK + 0.9BrO). The resulting Hg(II) concentrations along the flight track increased to 150–500 pg m\(^{-3}\), in better agreement with the observations (Fig. 8c). The oxidation of Hg(0) in this air mass was considerably faster than the 3 × Br or FastK simulations, which suggests that the uncertainties in both the Br radical concentrations and in the oxidation rate constant can simultaneously affect the overall bias in the modeled Hg(II) in certain areas.

7 Links to previous studies

The above comparison of the simulated Hg(II) concentrations with the NOMADSS observations shows that the Hg(0) oxidation based on the standard rate constants (Goodsite et al., 2004, 2012; Donohoue et al., 2006) and the GEOS-Chem calculated Br concentrations is too slow. Increasing the free tropospheric Br concentrations by a factor of 3 or considering the higher rate constants of Ariya et al. (2002) leads to significant improvement in the model results in the free troposphere compared to the NOMADSS observations. By contrast, Weiss-Penzias et al. (2015) found that the GEOS-Chem simulated Hg(II) concentrations in the free troposphere were 2.5 times higher on average than the observed concentrations at five high-elevation sites in western USA and Taiwan. Differences in the instruments and models used in their study and ours make it difficult to directly compare our findings. Weiss-Penzias et al. (2015) used the Tekran\(^®\) 2537-1130-1135 system, which can underestimate Hg(II) in the presence of O\(_3\) (Lyman et al., 2010; McClure et al., 2014). Their GEOS-Chem model is based on the reaction kinetics described in Holmes et al. (2010) and does not include updates described in Sect. 2.3.3. Importantly, the rate constant for the dissociation of HgBr (Reaction R2) has since been corrected, and is now a factor of 10–20 higher than the previous value. The faster rate of dissociation of HgBr decreases the modeled Hg(II) concentration at 600 hPa by a factor of about 1.5.

The relatively high DL of the UW-DOHGS instrument makes the NOMADSS observations unsuitable for an evaluation of faster oxidation in the boundary layer. However, three previous studies in the tropical and mid-latitude MBL have reported similar findings that the standard Hg(0) + Br oxidation kinetics are too slow to reproduce the observed Hg(II) concentrations, as discussed below.

Sprovieri et al. (2010) observed the diurnal cycle in RGM concentrations over the Adriatic Sea with daily enhancements of 20–40 pg m\(^{-3}\) at midday. Using a box model, the authors found that they could reasonably reproduce the observations with the Ariya et al. (2002) rate constant for the Hg(0) + Br reaction (Reaction R1), but only if the HgBr thermal dissociation (Reaction R2) was neglected. In their study, Br was considered to be the sole second-step oxidant (Reaction R4). If HO\(_2\), NO\(_2\), and BrO were to be included as second-step oxidants (Dibble et al., 2012), as we do in our FastK simulation, it would provide an additional pathway for the oxidation of HgBr to Hg(II), and partly overcome the slowing effect of the HgBr thermal dissociation.

In analyzing RGM observations over the Galápagos Islands in the equatorial Pacific, Wang et al. (2014) showed that the inclusion of HO\(_2\) and NO\(_2\) as second-step oxidants in the Hg(0) + Br reaction scheme based on Goodsite et al. (2012) and Dibble et al. (2012) was necessary to simulate the observed magnitude of the midday peaks in RGM concentrations. The box modeling study assumed peak daytime BrO concentrations of 0.2 pptv, which is similar to the annual average GEOS-Chem BrO concentration of 0.14 pptv. However, because of the uncertainties in the Tekran\(^®\) 2537-1130-1135 system measurements, the actual concentrations of Hg(II) could possibly be higher than the observed concentrations (Wang et al., 2014; Lyman et al., 2010; Ambrose et al., 2013).

Tas et al. (2012) present an analysis of mercury depletion events associated with extremely high BrO concentrations (20–80 pptv) over the Dead Sea in Israel. They found that
the standard Br-initiated mechanism was insufficient to reproduce the rate of depletion of Hg(0) and showed that oxidation of Hg(0) by BrO, at a rate close to that reported by Spicer et al. (2002), was necessary to explain the loss of Hg(0). If the FastK kinetics are considered, the Br-initiated pathway by itself can explain a large fraction of observed depletion rate of Hg(0). Although the BrO oxidation pathway cannot be entirely ignored (as discussed in Sect. 5), its importance, in this case, would be much smaller than what was found by Tas et al. (2012). Overall, these MBL studies are consistent with our analysis of the NOMADSS free tropospheric observations in implying much faster tropospheric oxidation of Hg(0) than currently assumed.

8 Implications of faster oxidation in the GEOS-Chem model

The global annual tropospheric mercury budgets for the three simulations are presented in Fig. 9. Despite faster oxidation in the 3 × Br and FastK simulations, we maintain the same global burden of THg by increasing the Hg(II) reduction rate (Sect. 2.3.4). Thus, the lifetime of THg against reduction is similar in all three simulations (~ 8.5 months). While we acknowledge that the modeled burden and lifetime of THg are affected by the uncertainty in the emission and deposition fluxes of mercury (Lin et al., 2006; Selin, 2009), we choose to maintain these in the three simulations because it allows us to focus on the model’s sensitivity to redox kinetics.

The tropospheric oxidation of Hg(0) to Hg(II) increases from 10 900 Mg a⁻¹ in the BASE simulation, to 18 300 Mg a⁻¹ for 3 × Br (factor of 1.7 increase) and 42 100 Mg a⁻¹ (factor of 3.9 increase) for the FastK simulation. The lifetime of Hg(0) against oxidation to Hg(II) decreases from 5 months in the BASE simulation to 2.8–1.2 months in the 3 × Br and FastK simulations. The tropospheric burden of Hg(II) increases by 33% in the 3 × Br simulation and by 66% in the FastK simulations.

Compared to the BASE simulation, the lifetime of Hg(II) against reduction decreases from 35 days (BASE simulation) to 19 days (3 × Br) and 8 days (FastK) simulations, respectively. This results in faster cycling between Hg(0) and Hg(II). Globally, 48% of the Hg(II) formed in the troposphere in the BASE simulation is reduced back to Hg(0) (Fig. 9a), whereas in the 3 × Br and FastK simulations, that fraction increases to 68–88% (Fig. 9b and c). The faster reduction in the simulations with faster oxidation implies that reduction plays a predominant role in controlling the burden and distribution of Hg(II) in the atmosphere. In view of our poor understanding of Hg(II) reduction in the atmosphere (Subir et al., 2011), we suggest that further laboratory and field measurements be conducted to constrain this process.

In this study, we chose to increase the reduction rate of Hg(II) to maintain the global burden of THg when oxidation is enhanced in 3 × Br and FastK simulations. An alternative approach would have been to keep the reduction rate constant and increase the global emissions of THg by factors of 2 to 4. Such a large increase exceeds the range of reported emission estimates (Selin, 2009; Song et al., 2015; Pirrone et al., 2010), and thus we did not consider it. Moreover, this would lead to a 2 to 4-fold decrease in the lifetime of THg against deposition which is not supported by the nearly constant vertical profiles of THg observed in the troposphere during NOMADSS and other aircraft campaigns (Holmes et al., 2010).

Figure 10 shows the simulated distribution of Hg(II) at 450 hPa (6.5 km) for 1 June–15 July 2013. In all three simulations, Hg(II) is enhanced in the subtropical anticyclones, with larger enhancements in the 3 × Br and the FastK simulations (Fig. 10a–c). A vertical profile in the Pacific subtropical anticyclone (Box 1, Fig. 10d) shows that the increase occurs throughout the free troposphere (Fig. 10d). Over the eastern US (Fig. 10e), the free tropospheric Hg(II) vertical profiles are similar for the three simulations below 8 km, where on average the increase in oxidation is matched by the increase in aqueous phase reduction in clouds. Above 8 km, the absence of liquid clouds prevents reduction from compensating for faster oxidation, and Hg(II) concentrations in the 3 × Br and FastK simulations are higher. Thus, the increase in the global Hg(II) mass burden in the 3 × Br and FastK simulations is concentrated largely in the subtropical anticyclones, and, to a smaller extent, in the extra-tropical upper troposphere.
Two processes maintain the high modeled Hg(II) concentrations in the subtropical anticyclones. First, the anticyclones are characterized by large-scale sinking motion which transports higher Hg(II) concentrations from the upper troposphere where fast oxidation of Hg(0) results from higher Br concentrations and lower temperatures slowing the thermal dissociation of the HgBr intermediate (Reaction R2). The locations of the 450 hPa enhancements in Hg(II) concentrations displayed in Fig. 10 are thus largely associated with the descending branches of the Hadley circulation. Note that the model predicts larger Hg(II) concentrations in the Southern Hemisphere subtropics where the winter Hadley circulation is stronger. Second, the sinking air in the anticyclones suppresses cloud formation and precipitation, thereby preventing loss of Hg(II) by reduction and wet deposition. This leads to efficient accumulation of Hg(II) in the subtropical anticyclones, even at lower altitudes. The model predicts low 450 hPa Hg(II) concentrations in regions with high cloud cover and precipitation (such as the Western Pacific), and in regions with low insolation, such as the Southern (winter) Hemisphere polar region.

The model simulates frequent episodes of high Hg(II) concentrations at 6.5 km over Texas during the summer of 2013 (Fig. 11a). From 1 June to 15 July, simulated midday (noon–3 p.m.) Hg(II) concentrations were higher than 250 pg m$^{-3}$ for 15 days in the FastK simulation (3×Br: 12 days), two of which were the days when RF-06 and RF-09 were conducted. RF-10 also flew at 7 km (425 hPa) over northern Texas and Oklahoma (Fig. 1) on 27 June 2013, but observed Hg(II) concentrations remained BDL (DL: 134 pg m$^{-3}$) during the high altitude leg. The 3×Br and FastK modeled Hg(II) concentrations over this region during RF-10 were 136 and 208 pg m$^{-3}$, respectively, significantly lower than those simulated for RF-06 and RF-09. The highest Hg(II) at 6.5 km over Texas and Oklahoma occurs when the wind over the southern US is southwesterly (Fig. 11b) transporting Hg(II) from the semi-permanent Pacific anticyclone, whereas the lowest concentrations occur when the transport is from the north (Fig. 11c). Our results suggest that the transport of Hg(II) produced in the Pacific anticyclone could be an important source of Hg(II) over southeastern US, potentially influencing wet deposition in the region if these high Hg(II) air masses are exposed to deep convection.

9 Conclusions
In this work, we have analyzed aircraft-based measurements of atmospheric mercury species (THg and Hg(II)) made during the NOMADSS campaign that took place over the southeastern US during summer 2013. The THg observations show a weak vertical gradient with higher concentrations in
that the NOMADSS observations are most consistent with 3× Br, and the FastK simulations, respectively, implying a greater importance of reduction to chemistry of mercury in the atmosphere. In the subtropical anticyclones, the 3× Br and FastK simulations predict a 3 to 5-fold enhancement in Hg(II) concentrations at 450 hPa relative to the global average Hg(II) concentration. These subtropical anticyclones are dry, cloud-free regions which provide ideal conditions for accumulation of Hg(II). The high Hg(II) in the Pacific anticyclone is periodically transported over southern US during summer and could be an important source of mercury wet deposition in the region. Future measurements in the subtropical anticyclones can help us gain deeper insights into the pathways of Hg(0) oxidation in the atmosphere.

The Supplement related to this article is available online at doi:10.5194/acp-16-1511-2016-supplement.

Acknowledgements. This material is based upon work supported by the National Science Foundation under grant no. 1217010 to D. A. Jaffe, L. Jaegle and N. E. Selin, grant no. 1215712 to J. Stutz, and grant no. 1216743 to C. A. Cantrell and R. L. Mauldin III. The authors thank participants from National Center for Atmospheric Research’s Earth Observing Laboratory and Research Aviation Facility for their support in the planning and execution of the NOMADSS campaign.

Edited by: A. Dastoor

References


V. Shah et al.: Origin of oxidized mercury in the summertime free troposphere


