Tropical cyclone activity enhanced by Sahara greening and reduced dust emissions during the African Humid Period

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Tropical cyclones (TCs) can have devastating socioeconomic impacts. Understanding the nature and causes of their variability is of paramount importance for society. However, historical records of TCs are too short to fully characterize such changes and paleo-sediment archives of Holocene TC activity are temporally and graphically sparse. Thus, it is of interest to apply physical modeling to understanding TC variability under different climate conditions.

Here we investigate global TC activity during a warm climate state (mid-Holocene, 6,000 yBP) characterized by increased boreal summer insolation, a vegetated Sahara, and reduced dust emissions. We analyze a set of sensitivity experiments in which not only solar insolation changes are varied but also vegetation and dust concentrations. Our results show that the greening of the Sahara and reduced dust loadings lead to more favorable conditions for tropical cyclone development compared with the orbital forcing alone. In particular, the strengthening of the West African Monsoon induced by the Sahara greening triggers a change in atmospheric circulation that affects the entire tropics. Furthermore, whereas previous studies suggest lower TC activity despite stronger summer insolation and warmer sea surface temperature in the Northern Hemisphere, accounting for the Sahara greening and reduced dust concentrations leads instead to an increase of TC activity in both hemispheres, particularly over the Caribbean basin and East Coast of North America. Our study highlights the importance of regional changes in land cover and dust concentrations in affecting the potential intensity and genesis of past TCs and suggests that both factors may have appreciable influence on TC activity in a future warmer climate.

Significance

Our modeling study shows the crucial role of vegetation cover over the Sahara and reduced dust emission in altering tropical cyclone activity during the mid-Holocene (6,000 yBP). Our results also demonstrate how these regional changes in land cover and dust emission are able to affect areas far afield through changes of large-scale atmospheric circulation. Our study strongly suggests that an appropriate representation of land cover and dust emission is of paramount importance to be able to capture past—and potentially future—climate changes.

Author contributions: F.S.R.P. designed and performed research; K.A.E. and Q.Z. contributed analytic tools; F.S.R.P., M.C., and G.T.D. analyzed data; F.S.R.P., K.A.E., M.C., G.T.D., Q.Z., L.S., J.C.S., and J.P.D. contributed to the interpretation and the discussion of the results; and F.S.R.P. wrote the paper with contributions from all authors. The authors declare no conflict of interest.

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Table 1. Boundary conditions for each modeling experiment

<table>
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<th>Simulation</th>
<th>Orbital forcing</th>
<th>GHGs</th>
<th>Saharan vegetation</th>
<th>Saharan dust</th>
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<td>1850 AD</td>
<td>Desert</td>
<td>As PI</td>
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<td>6,000 yBP</td>
<td>As PI</td>
<td>As PI</td>
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<tr>
<td>MH_{GS+RD}</td>
<td>6,000 yBP</td>
<td>6,000 yBP</td>
<td>Shrub</td>
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<td>MH_{GS}</td>
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orbital forcing. However, proxy evidence has shown that the currently hyperarid Sahara was characterized by a mesic landscape during the MH (33–35) and that Sahara dust emissions were drastically reduced as a result (15). Previous modeling studies (30, 31) found that although the TC genesis potential during the MH decreased in the NH under stronger summer insolation, it increased in the SH under weaker summer insolation relative to the preindustrial (PI) climate. However, PMIP model simulations for the MH do not adequately reproduce the intensification and geographical expansion of the WAM (35) and ISM (36), as inferred from proxy archives, or the shift in the Intertropical Convergence Zone (ITCZ) position, which can have a pronounced influence on TC development (37). Recent modeling efforts have shown that accounting for extensive greening of the Sahara and the associated reduction of dust emissions allows for better simulation of the WAM strengthening (18, 38) and its teleconnections (39, 40). In particular, Muschitiello et al. (39) and Davies et al. (40) demonstrated the sensitivity of Arctic climate to Saharan vegetation during the termination of the African Humid Period. These studies highlight the important role played by vegetation cover and dust emission in shaping climate on both regional and global scales.

The aim of this work is to investigate the effects of changes in Saharan vegetation and the associated reduction in dust emission on TC activity in addition to the insolation forcing (standard MH PMIP simulation) compared with the PI climate. The results of this study also provide insight into the possible influence of dust and vegetation changes on TC activity as the Earth warms in response to increasing CO₂ concentrations.

As a reference experiment, we take the PI climate as simulated by an Earth System Model [EC-Earth version 3.1 (41)] with an atmospheric horizontal resolution of 1.125° x 1.125° and 62 vertical levels. We then analyze four sets of sensitivity experiments performed by Pausata et al. (38) (Methods). These four experiments are (i) the standard MH_{PMIP} in which the MH insolation and greenhouse gas boundary conditions have been used following the PMIP protocol; (ii) the MH_{GS+RD} where both the Sahara greening (GS) and the dust reduction (RD) are considered; (iii) the MH_{GS} in which Saharan land cover is set to shrub and dust is kept at PI values as in MH_{PMIP}; and (iv) the MH_{RD} where the dust concentrations are reduced by as much as 80% (Table 1; also figures 1 and S1 in Pausata et al. (38)), whereas the Sahara land cover is set to PI as in MH_{PMIP} (Table 1).

We use a dynamical downscaling technique (5) for simulating large numbers of TCs in each experiment (Methods). We also

![Fig. 1](https://example.com/fig1.png)

**Fig. 1.** (A) Climatological track density for the preindustrial simulation. (B) Tropical cyclone annual frequency for the Atlantic (ATL), Eastern (EP) and Western (WP) Pacific, Northern Indian Ocean (NIO), and Northern (NH) and Southern (SH) Hemispheres. The red horizontal bars show the observed climatological number of storms for the period 1980–2015. (C and D) Changes in track density for the MH_{PMIP} (C) (orbital changes only) and the MH_{GS+RD} (D) (orbital changes + green Sahara + reduced dust) experiments relative to the preindustrial. Only values that are significantly different at the 5% level using a local (grid-point) t test are shaded. The contour intervals (dashed, negative anomalies; solid, positive anomalies) follow the color bar scale. The box in A indicates the area where the track density and the number of storms per year have been scaled by a factor of 10. In this particular region, scaling is applied to the track density of all experiments to visually highlight its pattern and its changes over the North Atlantic due to the large model underestimation.
investigate the most important environmental factors (thermodynamic and kinematic) that affect TC genesis and track density (Methods) to better understand the causes of the changes in tropical storminess. For each experiment, we analyze a 30-y portion of the integration following an initial spin-up phase (Methods).

The global TC climatology in the PI simulation (Fig. 1A) is in excellent agreement with observations in both the Western Pacific and Northern Indian Oceans (Fig. S1). However, TCs are underestimated in the Eastern Pacific and particularly in the North Atlantic. Whereas the TCs are underestimated in the SH, they are overestimated in the NH. This could be in part related to the NH cold bias in our model (41). It is common for global models to underrepresent the activity of North Atlantic TCs, but the model used here has a larger bias than other climate models (5). However, the observed seasonal and interannual variabilities of Atlantic TCs are well captured by our model and by climate models in general (5). Furthermore, our analysis focuses on relative changes in TCs and environmental parameters favorable to TC genesis. An underlying assumption in our analysis is that the cold bias in the North Atlantic will not have a large effect on the predicted changes in tropical cyclogenesis. Further investigation of the causes of this bias is needed.

Our model experiments show that the orbital forcing changes alone lead to an increase in TC landfalls over South Asia and a widespread decrease of TCs offshore in the Western Pacific (MHPMIP, Fig. 1C). A similar pattern of an increase in TC density that occurs onshore toward the Caribbean Sea and a decrease offshore is simulated in the tropical North Atlantic (Fig. 1C). Over the entire NH the number of TCs is unchanged or slightly reduced in the MHPMIP simulation compared with the PI (Fig. 1B). In the SH the number of TCs increases, in overall agreement with previous studies (30, 31) that analyzed environmental factors favorable to TC development using PMIP simulations. Accounting for both the GS and RD (MGSRD, Fig. 1C) the impact of dust is limited to the Atlantic region (compare Figs. 1C and 2B), whereas the effects of the GS are worldwide (compare Figs. 1C and 2A). The global effect seen in the MHGS simulation (relative to MHPMIP) is most likely related to the large changes in the WAM (38) and the Walker circulation (Fig. S2) induced by the GS. Alternately, the changes in the WAM and Walker circulation in the MHRD simulation are similar to those induced by the orbital forcing alone (38). Hence, the effect of RD alone under desert Sahara conditions (MHRD) affects only the Caribbean and eastern coastal regions of the United States. This is where dust reduction has a direct radiative effect, resulting in a pronounced warming of sea surface temperatures (SSTs) (Fig. S3).

To better understand the causes behind the simulated changes in TC activity, we examine a genesis potential index (GPI) (Methods) and the environmental parameters that describe how favorable the climate state is for tropical cyclogenesis (Fig. S4). The GPI is a synthesis of kinematic (wind shear and absolute vorticity) and thermodynamic factors (potential intensity and moist entropy deficit) that affect tropical cyclogenesis. A description of the role of each environmental factor on the genesis of TCs is provided in Methods.

The GPI changes in the MHPMIP simulation indicate more favorable conditions for tropical cyclogenesis on the western and eastern coasts of Central America, in the South China Sea, and over most of the tropical regions in the SH (Fig. 3A). Less favorable conditions are simulated over most of the tropical Pacific and the eastern Atlantic Ocean. The changes in GPI relative to the PI simulation are in agreement with the changes in simulated TCs using the downscaling technique, suggesting that the GPI is a reliable measure of TC activity under both modern and MH climate conditions. The GPI anomalies in the MHPMIP simulation are also very similar to the ensemble mean changes in GPI estimated from 10 PMIP models (compare figures 3 and 9a in ref. 30).

Accounting for the combined effects of GS and RD (MGSRD) yields an anomaly pattern that is very similar to the one simulated in the MHPMIP, but substantially amplified (compare Fig. 3E with 5C). The GPI analysis confirms the global impacts of GS (Fig. 3E), whereas the effects of the RD are more limited to the tropical Atlantic (Fig. 3G).

An in-depth analysis of each component characterizing the GPI reveals that, in general, the most important environmental factors driving the changes in TC activity in the MH simulations are the wind shear and the potential intensity (Fig. 3 and Fig. S5). The striking increase in TC density over the Caribbean Sea in the MGS+RD simulation is most likely due to the combination of decreased wind shear and enhanced midlevel entropy

![Fig. 2. (A and B) Changes in track density for the MHGS (A) (orbital changes + green Sahara) and the MHPMIP (B) (orbital changes + reduced dust) experiments relative to the MHPMIP. Only values that are significantly different at the 5% level using a local (grid-point) t test are shaded. The contour intervals (dashed, negative anomalies; solid, positive anomalies) follow the color bar scale. Results are scaled by a factor of 10 over the North Atlantic (Fig. 1A).](image-url)
content (i.e., humidity level closer to saturation) (Fig. 1D and Figs. S5–S7), in addition to the large potential intensity increase. The absolute vorticity also increases over the entire tropical Atlantic poleward of 10°N (Fig. S8). The combined effect of these factors, largely induced by the strengthening of the WAM, boosts the TC activity over the eastern tropical Atlantic in the MH GS+RD experiment relative to the PI simulation.

In the SH, the changes in potential intensity, and to lesser extent the entropy changes, play the most important roles (Fig. 3 and Fig. S8). We infer from this that the variations in available thermodynamic energy imparted from changes in both the SST and the atmospheric environment are crucial for increasing SH TC activity in the MH simulation, despite the large SST cooling relative to the PI. This result further demonstrates that SST by itself is an inadequate indicator of storm strength, as shown in previous work (42).

In summary, our results show that during the MH—a period characterized by increased NH summer insolation and extensive greening of the Sahara desert—TC activity may have been amplified globally, and in particular over the western Atlantic, compared with TC activity in the PI climate state. These results appear to be at variance with findings from previous studies of MH TCs obtained using PMIP simulations. The latter show slightly more favorable TC conditions in the SH but less favorable conditions in the NH. The apparent dissimilarity arises from the fact that only orbital changes were considered in previous studies of the MH climate (30, 31), whereas here we show that including the GS and the consequent reduction in dust emission are critical influences on TC activity.

Previous studies (30, 31) that investigated changes in TC genesis potential for past climate did not have access to the 6-h model output required for downscaling, so that it was not feasible to analyze simulated storms directly. Those previous studies were based exclusively on analysis of the indexes and environmental parameters that are favorable to TC development. The underlying assumption in such work is that the GPI, which reliably captures the net effect of the factors that influence TC genesis under modern climate conditions, also performs well under different climate states (30, 31). Our results show that this assumption may be reasonable: The GPI anomalies in the MH experiments closely match the simulated changes in TC density obtained using a downscaling model. This is in agreement with results from Camargo et al. (43), who showed that genesis indexes that take into account the saturation deficit and potential intensity (as in the case of the GPI) are best suited to capture changes in global TC frequency under different climate states.

Our simulations of the effects on tropical cyclones of enhanced vegetation cover and reduced dust emissions in the Sahara are difficult to test with existing paleotempest records. At present, such records do not yet provide a geographically and temporally complete view of Holocene storm activity for model evaluation. However, a reconstruction of TC activity from Vieques, Puerto Rico indicates increased intense TC activity during times when other proxy evidence suggests an intensified WAM over the last 5,500 y (20). Furthermore, one of the intervals with the highest TC activity in the Vieques reconstruction occurred between 2,500 y and 1,000 y BP, when the flux of Saharan dust was low compared with that in the subsequent millennium (44). Other

Fig. 3. Storm-season [June to October (JASO NH); January to April (JFMA SH)] changes in genesis potential index (GPI; A, C, E, and G) and potential intensity ($V_{PI}$; B, D, F, and H) for each sensitivity experiment relative to the PI reference simulation. The contour intervals (dashed, negative anomalies; solid, positive anomalies) follow the color bar scale. Only values that are significantly different at the 5% level using a local (grid-point) k-s test are shaded. The zero-anomaly line is omitted for clarity.
evidence supportive of the results presented here includes a recent study (45) highlighting the link between reduced dust loading during the early and mid-Holocene and the northward expansion of the ITZC, which in turn can intensify TC activity in the western North Atlantic (37). The message from these paleoclimatic studies is that the Sahara dust layer appears to suppress TC activity, consistent with the inverse relationship between Saharan dust emission and Atlantic TC activity found in instrumental records (9, 46). Our finding that a notable increase in TC activity in the North Atlantic is strongly associated with a reduction in airborne dust concentrations is therefore consistent with the physics of tropical climates and the known environmental history of the Sahara.

At present, our model results cannot be directly substantiated by paleotempest records given their sparsity. Nevertheless, our study provides a testable prediction of the relationship between climatic conditions in the Sahara and the global behavior of TCs that can be empirically evaluated as more proxy reconstructions of TC activity become available. Additional modeling studies studying TC activity and accounting for vegetation and dust feedbacks are needed to supplement our findings. It would also be important to investigate TC behavior under future anthropogenic-induced climate change, in which the vegetation and dust feedbacks have been generally overlooked. The societal importance of this proposed relationship in the context of anthropogenic climate change should serve as a stimulus for obtaining additional proxy records. The characterization of TC behavior under future climate change is important to investigate, as TCs will affect many parts of the world, including the United States, and their potential destructive force must be better understood.

Methods

Experimental Design. In the mid-Holocene control (MHcontrol) experiment, the boundary conditions were set at preindustrial values—except for the orbital forcing and greenhouse gases—following the PMIP protocol (32). The changes in the Earth's orbital parameters in the MH increase the amplitude of the seasonal cycle in NH insolation by ~5% compared with present-day values. The MH sensitivity studies analyzed in this work in which we change in turn the vegetation over the Sahara domain (evergreen shrub: 11°–33°N and 15°W–35°E) and reduce the dust concentrations by up to 80% are described in detail in Pausata et al. (38). The 80% RD was applied over a broad area around the Sahara desert from the nearby Atlantic Ocean to the Middle East and throughout the troposphere (up to 150 hPa). A smoothing filter was used to avoid excessively abrupt transitions in dust concentrations (figure S10). Above 150 hPa, the dust reduction was more evenly applied due to the fact that aerosol particles are uniformly distributed at those elevations. The change in dust concentration, although based on available proxy records off the coast of northwestern Africa (14, 15), and vegetation cover are not meant to provide a faithful representation of the MH conditions over the Sahara and nearby regions, because no accurate vegetation reconstruction is available at the moment. They are instead designed to more easily disentangle the effects of land surface cover and dust loading on atmospheric circulation and specifically on TC activity.

The dust distribution used in this study and in Pausata et al. (38) was taken from the Community Atmosphere Model (CAM) (48), which is used in the Coupled Model Intercomparison Project (CMIP) phase 5. However, the dust distribution shows very high concentrations over eastern Africa that are not present in observations, in which dust loading peaks over Chad and the western Sahara desert. To test our model results to the spatial pattern of the dust distribution we create new prescribed dust concentrations and distribution based on the long-term monthly mean (1980–2015) output from the MERRAero dataset, which assimilates satellite data (Fig. S9) (https://gmao.gsfc.nasa.gov/reanalysis/merra/MERRAero/). The changes in TC activity using the dust (and reduced) files inferred from MERRAero do not show appreciable differences from the case in which the CAM-derived dust is used (compare Fig. 1 and Fig. S10).

The greening of the Sahara corresponds to a reduction in the surface albedo from 0.3 to 0.15 and an increase in the leaf area index from 0.2 to 2.6 (compare Fig. 1 in ref. 38). The dust reduction leads to a decrease in the dust aerosol optical depth of almost 60% and in the global total aerosol optical depth of 0.02 (fig. 1 in Pausata et al. [38]). Initial conditions for the MH experiments were taken from the preindustrial spin-up run, and the simulations were then run until the climate reaches quasi-equilibrium after about 100–200 y, depending on the experiment. We focus on the equilibrium responses, and we analyze 30 y of each sensitivity experiment.

TC Dynamical Downscaling and Genesis Indexes. Here we provide a brief overview of the TC dynamical downscaling technique adopted in this study; an extensive description can be found in Emanuel et al. (5). This technique incorporates both thermodynamic and kinematic statistics derived from the coupled-model simulations to drive a detailed, coupled ocean–atmosphere tropical cyclone model, producing a large number (~10−5-105) of synthetic TCs. These are then used to characterize the TC climatology of each sensitivity experiment. The synthetic TCs are generated through the following three steps. Genesis. The genesis of the synthetic tropical cyclones is based on “random seeding”: The storms are initiated at points that are randomly distributed in space and time, under the following conditions of warm-core vortices with peak wind speeds of only 6 m/s and almost no midlevel humidity anomaly in their cores. These random “seeds” are distributed everywhere and at all times, except at the equator (~±3° latitude). This technique does not account for storms initiating distant to the extent that can favor TC genesis, such as easterly waves. The algorithm includes a parameterization of the detrimental effect of environmental wind shear on TCs.

Tracks. We use a “beta-and-advection” model to predict storm tracks using only the large-scale background wind fields. The beta-and-advection model uses synthetic wind time series at 850 hPa and 250 hPa, represented as Fourier series of random phase with the monthly means, variances, and covariances obtained using daily data from EC-Earth and a geostrophic turbulence power-law distribution of kinetic energy.

Intensity. We use the “Coupled Hurricane Intensity Prediction System” (“CHIPS”), which is a coupled air-sea TC model to predict the wind field of each storm. This model has been largely used to produce skillful, real-time hurricane intensity forecasts. The model takes as inputs EC-Earth’s thermodynamic state and wind shear, as described in detail in Emanuel (49). CHIPS also account for the variable midtropospheric temperature and relative humidity, reading in EC-Earth’s monthly mean entropy at 600 hPa.

We also separately calculate the potential intensity (Vp), which is determined by the thermal structure of the atmosphere and is a proxy for the upper limit of thermodynamically achievable intensity for the mature phase of a TC. The Vp depends on the temperatures at the surface (SST) and the convective outflow level (Tc).

\[ V_p = \left[ \frac{C_p}{C_f} \right] SST - T_c \left( k_s - k \right) \]

where \( C_p/C_f \approx 1 \) is the ratio of the exchange coefficients for enthalpy and drag and \( k_s - k \) is the difference between the sea-surface saturation-specific enthalpy \( (k_s) \) and that of the overlying marine boundary layer \( (k) \). Finally, we adopted the GPI to collectively take into account the most important environmental factors (thermodynamic and kinematic) that affect TC genesis. We adopted the GPI defined in Emanuel (50), with the exception that it incorporates the findings of Tippett et al. (51), who showed that vorticity does not appear to be rate limiting outside of very low latitudes. In doing so, we are also consistent with previous paleoclimatic studies on MH TCs (30, 31), in which they use the so-called “clipped vorticity,”

\[ GPI = \frac{b \min \left[ \eta, 4 \times 10^{-5} \right] \eta}{x_m \left( 25 + V_{shear} \right)} \]

where \( \eta \) is the absolute vorticity calculated at 850 hPa, \( V_{shear} \) is the 200- to 850-hPa wind shear value, \( x_m \) is the moist entropy deficit, \( V_{shear} \) is the maximum potential intensity a TC can theoretically achieve, and \( b \) is an empirically derived normalization factor to calibrate the GPI to preindustrial values.

Role of Each GPI Component on TC Developments. The wind shear and the absolute vorticity play the two key kinematic factors included in the GPI. The wind shear is the vertical shear of the horizontal winds between 200 hPa and 850 hPa. High wind shear causes asymmetries in the developing cyclone that reduce the efficiency of drawing warm and moist air from the ocean and also allow inflow of relatively cooler and drier air from the top, which weakens
cyclogenesis. The vorticity acts as a spin-up mechanism that favors cyclogenesis formation. The potential intensity and the moist entropy deficit are the thermodynamic factors included in the GPl: The potential intensity indicates the maximum wind speed permissible by the thermodynamic environment. The moist entropy difference between the midtroposphere and the boundary layer provides information on the duration needed for an initial perturbation to moisten the midtroposphere before intensification occurs. Higher entropy difference signifies a longer duration. For clarity we have plotted in Model Description and Experiment Design the reciprocal of the entropy parameter, which means that higher values correspond to more favorable conditions for TC development.


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