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<b>Detailed Terms</b>	

**Diamond nitrogen-vacancy center as a probe of random fluctuations in a nuclear spin ensemble**Abdelghani Laraoui,<sup>1</sup> Jonathan S. Hodges,<sup>2</sup> Colm A. Ryan,<sup>3</sup> and Carlos A. Meriles<sup>1,\*</sup><sup>1</sup>*Department of Physics, City College of New York–CUNY, New York, New York 10031, USA*<sup>2</sup>*Department of Electrical Engineering, Columbia University, New York, New York 10023, USA*<sup>3</sup>*Department of Nuclear Science and Engineering, Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA*

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New schemes that exploit the unique properties of nitrogen-vacancy (NV) centers in diamond are presently being explored as a platform for high-resolution magnetic sensing. Here we focus on the ability of a NV center to monitor an adjacent mesoscopic nuclear spin bath. For this purpose, we conduct comparative experiments where the NV spin evolves under the influence of surrounding  $^{13}\text{C}$  nuclei or, alternatively, in the presence of asynchronous alternating current fields engineered to emulate bath fluctuations. Our study reveals substantial differences that underscore the limitations of the semi-classical picture when interpreting and predicting the outcome of experiments designed to probe small nuclear spin ensembles. In particular, our study elucidates the NV center response to bath fluctuations under common pulse sequences, and explores a detection protocol designed to probe time correlations of the nuclear spin bath dynamics. Further, we show that the presence of macroscopic nuclear spin order is key to the emergence of semi-classical spin magnetometry.

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**I. INTRODUCTION**

Nitrogen-vacancy (NV) centers in diamond are presently the focus of a broad cross-disciplinary research effort combining the fields of condensed matter, atomic physics, and precision metrology. Several unique properties, including the long spin coherence times and superb photostability at room temperature, make NV centers key to various proposals for quantum information processing<sup>1</sup> and quantum cryptography.<sup>2</sup> Central to the present study is the use of NVs for high-resolution magnetometry, an application facilitated by the ability to initialize, manipulate, and read out spins with high fidelity.<sup>3</sup> Initial experimental work has used single NVs to map the magnetic field created by a ferromagnetic particle with nanoscale resolution<sup>4</sup> and has demonstrated detection of synchronous, coil-induced alternating current (AC) fields<sup>5</sup> with sensitivity approaching  $30 \text{ nT/Hz}^{1/2}$ . Further, bulk and superficial NVs have been exploited to determine the local amplitude and direction of the magnetic field generated by current-carrying wire strips imprinted on the host diamond surface.<sup>6–9</sup>

Although the use of single NVs to control and probe few adjacent, strongly coupled nuclear spins is well documented,<sup>10</sup> extensions that target the monitoring of more numerous, weakly coupled nuclear ensembles are still pending. Motivating this line of research are applications in which the NV is envisioned as a high-resolution *nuclear spin* magnetometer capable of providing information on the local density, chemical composition, physical structure, and/or dynamical processes within “mesoscopic” ( $\sim 10 \text{ nm}$ ) regions of an “*a priori* unknown” sample. In an initial theoretical approach to the problem,<sup>11</sup> a mesoscopic, unpolarized nuclear spin ensemble at room temperature was modeled via a stochastic, time-dependent magnetic field interacting with a single NV center. The resulting description is “semi-classical,” in that the “back-action” of the nuclear spin ensemble, which generates NV-nuclei entanglement, is not considered.

Here we revisit this topic by observing a single NV center alternatively exposed to the surrounding “bath” of  $^{13}\text{C}$  spins

or an external, coil-generated magnetic field. Despite the ability of the semi-classical approach to explain the observed NV response to common pulse sequences (e.g., excitation-evolution, Hahn-echo, etc), our results expose fundamental limitations in its ability to predict the outcome of an arbitrary nuclear spin sensing protocol. We illustrate the problem through an experiment designed to probe the slow, long-term correlation function of the effective nuclear magnetic field acting on the NV center. Moreover, we tackle some common misconceptions when interpreting the NV dynamics in a fluctuating nuclear spin bath, derive analytical expressions that clarify on the role of adjacent (versus distant) nuclear spins, and investigate the conditions necessary to regain the limit of semi-classical magnetometry.

Our manuscript is organized in the following way: Sec. II briefly touches on some practical details, explains the pulse protocols used, and presents our experimental observations. Through quantum mechanical calculations and “disjoint cluster” simulations of the combined nuclear ensemble–NV system, Sec. III discusses the NV response in the presence of a slowly fluctuating nuclear “bath”. Finally, concluding remarks and technical details on some of the analytical formulae and numerical simulations are presented in Sec. IV and the Appendices, respectively.

**II. EXPERIMENTAL PROTOCOLS****A. Experimental methods**

The theory and practice of single-NV center detection and control is well described in the literature and will be only briefly reviewed. The negatively charged NV center addressed here comprises a substitutional nitrogen associated with a vacant, adjacent lattice site. The ground state is a spin triplet with a concomitant “crystal field” aligned along the [111] axis (or its crystallographic equivalents). The zero-field (or crystal field) splitting between magnetic sublevels  $|m_S = 0\rangle$  and  $|m_S = \pm 1\rangle$  is  $\Delta = 2.87 \text{ GHz}$ . Green light illumination produces, to a good approximation, a spin-conserving

transition to the first excited state—also a triplet—which, in turn, leads to broadband photoluminescence emission. Non-radiative intersystem crossing transitions to lower energy metastable states are strongly spin selective because the shelving rate from  $|m_S = 0\rangle$  is much smaller than those from  $|m_S = \pm 1\rangle$ , thus allowing for an optical readout of the spin state (i.e., the  $|m_S = 0\rangle$  state is brighter). The spin-selective crossing rates also lead to an almost complete pumping of the spin degree of freedom after  $\sim 1 \mu\text{s}$  of illumination.<sup>12</sup>

Experiments are carried out in a high-purity diamond crystal with NV concentration less than 10 ppb. We optically address individual centers using a purpose-built confocal microscope. Electron spin resonance (ESR) spectroscopy of the spin sublevels is carried out using a thin copper wire (20  $\mu\text{m}$ ) overlaid on the crystal surface. We apply a weak, direct current (DC) magnetic field  $B_0$  ( $\sim 4$  mT) collinear with the crystal field to break the degeneracy between the  $|m_S = \pm 1\rangle$  states. This allows us to selectively address one of the two possible transitions (e.g.,  $|m_S = 0\rangle \rightarrow |m_S = 1\rangle$ ), which virtually renders the NV a two-level system. When necessary, we use a four-turn coil in the vicinity of the host diamond to generate an auxiliary field  $b_A(t) = b_{A,0} + b_{A,1} \cos(\omega_A t + \varepsilon_A)$ ; a computer-controlled wave function generator is used to set amplitudes  $b_{A,0}$ ,  $b_{A,1}$ , frequency  $\omega_A$  and phase  $\varepsilon_A$ . The system is designed to introduce random amplitude and phase changes over a timescale of a few milliseconds, longer than the time necessary for a single NV observation ( $< 200 \mu\text{s}$ ) but much shorter than the averaging time ( $\sim 1$  minute) required to overcome the shot noise.

### B. Semi-classical response of the NV center to echo and Ramsey sequences

To investigate the effect of a nuclear spin ensemble on the NV signal, we start with the response of a typical color center to the Hahn-echo sequence  $(\pi/2)_y - \tau - (\pi)_y - \tau - (\pi/2)_y$  shown in Fig. 1(a). As reported in prior observations,<sup>13</sup> the signal exhibits periodic “collapses and revivals,” whose amplitude progressively decays with a time constant that approaches the NV “true” coherence time  $T_2$ . The time separation between consecutive revivals coincides with the  $^{13}\text{C}$  Larmor frequency at the applied magnetic field (4 mT in the present case), a key feature that motivates a simple, heuristic interpretation: As the nuclear spins surrounding the NV center precess around  $B_0$  with frequency  $\omega_N$ , their summed dipolar fields effectively create a random magnetic field  $b_N(t) = b_{N,0} + b_{N,1} \cos(\omega_N t + \varepsilon_N)$ , whose amplitude and phase slowly change with time. Then, the phase picked up by the NV center during the first half of the Hahn-echo sequence differs from that accumulated during the second half and depends on the relative phase of the nuclear field, thus resulting in a signal “collapse” after a time-ensemble average. If, however, the nuclear spins have enough time to complete one full rotation during  $\tau$ , the NV accumulates no net phase and an echo “revival” takes place.

As shown in Fig. 1(c), one can use the coil-generated field  $b_A(t)$  to induce an identical response. This response is more generic than an artificial situation: It could also, for example, come from physical motion of the NV center in a static gradient field. To eliminate the influence of  $^{13}\text{C}$  spins on the signal

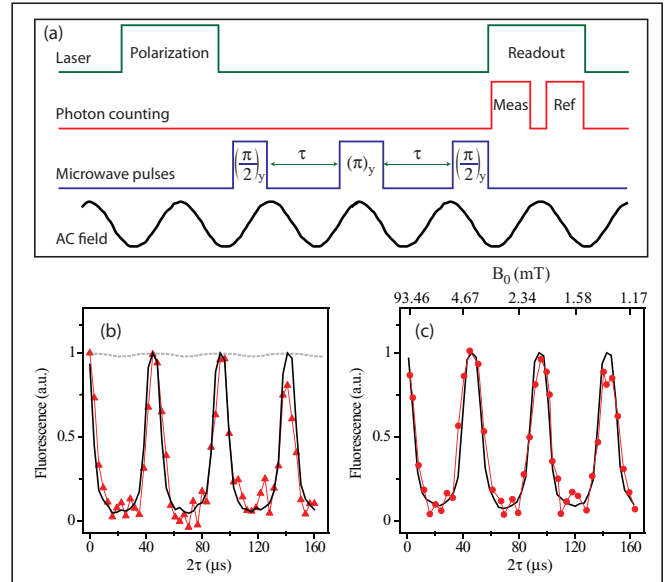


FIG. 1. (Color online) (a) Schematics of the detection protocol in a Hahn-echo sequence. An AC field of variable amplitude, phase, and frequency is used when necessary, as described in the text. (b)  $^{13}\text{C}$ -induced pattern of collapses and revivals for  $B_0 = 4$  mT. In this case, no AC field is applied. The solid line is a fit to the experimentally observed response (triangles) using Eq. (2) with  $\sqrt{\langle b_{N,-1}^2 \rangle} = 4.64 \mu\text{T}$  and  $\omega_N/2\pi = 42.5$  kHz. The faint, dashed line corresponds to Eq. (2) for  $\sqrt{\langle b_{N,-1}^2 \rangle} = 0.1 \mu\text{T}$  (see main text). (c) Hahn-echo sequence in the presence of an asynchronous AC field of rms amplitude  $\sqrt{\langle b_{A,-1}^2 \rangle} = 3.2 \mu\text{T}$  and frequency  $\omega_A/2\pi = 44.5$  kHz. The upper horizontal axis indicates the value of  $B_0$  at each time  $\tau$ . Good agreement is found between experiment (circles) and Eq. (3) (solid lines).

modulation, we progressively adjust the DC field  $B_0$  (see upper axis in Fig. 1(c)) so that the nuclear Larmor frequency  $\omega_N/2\pi = (\gamma_N B_0)/2\pi$  exactly matches the inverse of a desired evolution time  $\tau$  (i.e., nuclear spins undergo one full cycle at every given  $\tau$ ). The result is a train of  $b_A$ -induced collapses and revivals at the frequency  $\omega_A$ .

We now find an analytical expression that formally describes these ideas in the limit of infinitely short pulses. By detuning the  $|m_S = -1\rangle \rightarrow |m_S = 0\rangle$  transition from our microwave field, we can treat the NV spin as a two-level system. Within the  $m_S = \{0, 1\}$  sub-manifold, we thus define the normalized NV signal,

$$S = \langle \text{Tr} \{ \rho \sigma_z \} \rangle, \quad (1)$$

where  $\rho$  denotes the system density matrix and  $\sigma_z = |0\rangle\langle 0| - |1\rangle\langle 1|$  is the Pauli operator in the direction of the crystal field (assumed along the  $z$ -axis); brackets indicate time (or ensemble) average (see below). Within a time interval shorter than  $T_2$ , the NV response to a Hahn-echo sequence in the presence of a “classical” field  $b_A(t)$  is given by

$$S_C^{HE}(\tau) = \langle \cos \phi_2 \rangle, \quad (2)$$

where  $\phi_n$ ,  $n = 1, 2$ , denotes the phase picked up by the NV center during the  $n$ th interpulse interval and  $\phi_2 \equiv \phi_1 - \phi_2 =$

$(4\gamma_{\text{NV}}b_{A-1}/\omega_A) \sin^2(\omega_A\tau/2) \sin(\omega_A\tau + \varepsilon_A)$ ;  $\gamma_{\text{NV}}$  is the color center gyromagnetic ratio. In the limit  $\phi_n \ll 1$ , Eq. (2) yields

$$S_C^{HE}(\tau) \cong 1 - 8\langle(\gamma_{\text{NV}}b_{A-1}/\omega_A)^2 \sin^2(\omega_A\tau + \varepsilon_A)\rangle \sin^4(\omega_A\tau/2) \\ = 1 - 4\langle\gamma_{\text{NV}}^2 \langle b_{A-1}^2 \rangle / \omega_A^2 \rangle \sin^4(\omega_A\tau/2), \quad (3)$$

where the last expression holds for the case in which phase and amplitude fluctuations are independent and the phase distribution function is uniform over  $[0, 2\pi)$ . Equation (3) predicts revivals at times  $\tau_m = 2\pi m/\omega_A$ , with  $m$  an integer. As expected, the signal is immune to the DC field component  $b_{A,0}$ , but the depth of the modulation (or, in the more general case, the duration of the revival) can be controlled via  $b_{A-1}$ .

For future reference, we note that the present picture can be easily extended to a Ramsey sequence  $(\pi/2)_y - t - (\pi/2)_y$ : In this case, one imagines the nuclear field  $b_{N,0}$  (or, correspondingly,  $b_{A,0}$  in the coil-induced analog) changing randomly over time, with the concomitant signal interference and monotonic decay over a time  $T_2^*$  when averaged in a time ensemble. Analytically, and in the same approximation as in Eq. (3), we find

$$S_C^R(t) = \langle \cos \phi \rangle \cong 1 - (\gamma_{\text{NV}}^2 \langle b_{A-1}^2 \rangle / \omega_A^2) \sin^2(\omega_A t / 2) \\ - (\gamma_{\text{NV}}^2 \langle b_{A,0}^2 \rangle t^2 / 2), \quad (4)$$

where, as before, we assumed  $1 \gg \phi = (2\gamma_{\text{NV}}b_{A-1}/\omega_A) \sin(\omega_A t / 2) \cos(\omega_A t / 2 + \varepsilon_A) + \gamma_{\text{NV}}b_{A,0}t$  and  $t < T_2^*$ .

### C. Probing the noise correlation time

In line with recent experiments designed to probe the spin noise present in a small ensemble,<sup>14</sup> the above framework suggests that it should be possible to use the NV center to determine the average amplitude, central frequency, and time correlation of the acting random field. Although alternative routes are conceivable,<sup>8</sup> Fig. 2(a) shows a flexible approach comprising two echo-sequences of fixed, identical duration  $2\tau$  separated by a variable free evolution time  $\tilde{\tau}$ . This sequence shares some similarities and motivation with 2D exchange experiments,<sup>15</sup> such as NOESY, or hyperfine correlation protocols,<sup>16</sup> such as HYSORE or DEFENCE. In the presence of a random magnetic field  $b_A(t)$  and with proper phase cycling, our protocol yields a signal

$$S_C^{\text{Cor}}(\tau, \tilde{\tau}) = \langle \sin \phi_{12} \sin \phi_{45} \rangle, \quad (5)$$

where we have chosen the pulse phases to give a sinusoidal dependence to maximize sensitivity (see Appendix A). Similar to Eq. (2),  $\phi_n$ ,  $n = 1, \dots, 5$  denotes the accumulated phase during the  $n$ th free-evolution interval and  $\phi_{n,n+1} \equiv \phi_n - \phi_{n+1}$ . Equation (5) indicates that the present sequence can be used as a tool to characterize the time coherence of the field under investigation (so long as the field correlation time  $\tilde{\tau}_c$  is shorter than the NV longitudinal relaxation time  $T_1$ ). For the particular setting  $\omega_A \tilde{\tau} = 2m\pi$  with  $m$  integer, and in the case  $\tilde{\tau} < \tilde{\tau}_c \leq T_1$ , we obtain for random  $\varepsilon_A$  but *fixed*  $b_{A-1}$

$$S_C^{\text{Cor}}(\tau, \tilde{\tau}) = \frac{1}{2}[1 - J_0(2K)], \quad (6)$$

where  $J_0$  denotes the zero-order Bessel function and  $K \equiv (4\gamma_{\text{NV}}b_{A-1}/\omega_A) \sin^2(\omega_A\tau/2)$ . Figure 2(b) shows the predicted

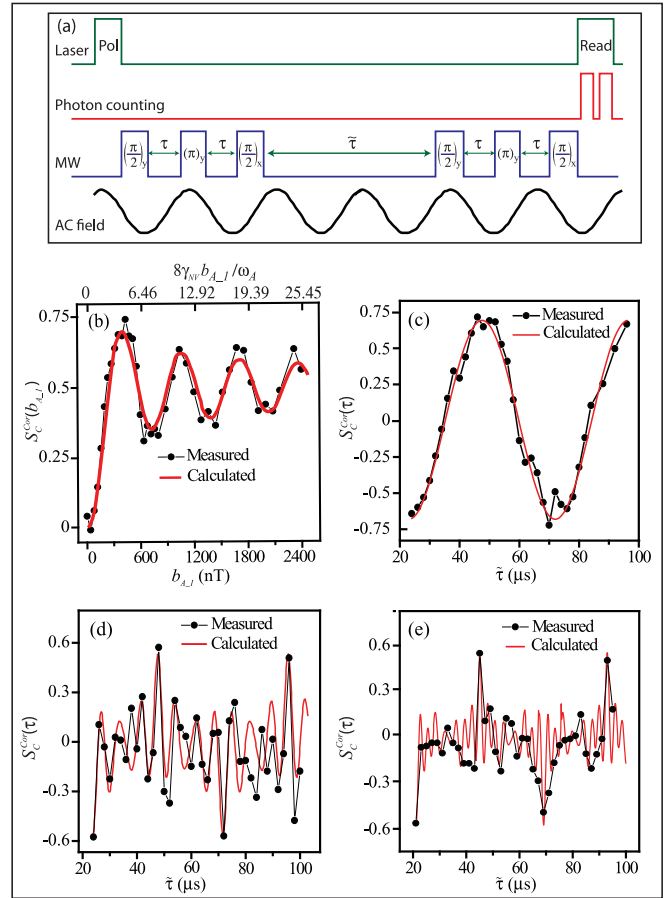


FIG. 2. (Color online) (a) Correlation protocol in the presence of an asynchronous AC field. (b) NV response as a function of AC field amplitude  $b_{A-1}$  for the case  $\omega_A(2\tau + \tilde{\tau}) = 2m\pi$ . Here, we choose  $\tilde{\tau} = 2\tau = 48 \mu\text{s}$ . (c–e) Signal as a function of  $\tilde{\tau}$  for  $b_{A-1} = 440 \text{ nT}$ ,  $2.2 \mu\text{T}$ , and  $4.4 \mu\text{T}$ , respectively. In (b) through (e), circles represent experimental points and solid curves are calculated using Eq. (5). In all cases, we use  $\omega_A/2\pi = 20.8 \text{ kHz}$  so that  $2\omega_A\tau = 2\pi$ . For the present timescale, we virtually eliminate the effect of  $^{13}\text{C}$  spins by setting the static field  $B_0 = 3.9 \text{ mT}$  so that  $\omega_N\tau = 2\pi$ .

and observed responses as a function of  $b_{A-1}$  when  $2\omega_A\tau = 2\pi$ , whereas Figs. 2(c)–2(e) extend the comparison to the more general case  $\omega_A\tilde{\tau} \neq 2m\pi$  for various values of  $b_{A-1}$ . When  $b_{A-1}$  fluctuates and  $\omega_A(2\tau + \tilde{\tau}) \neq 2m\pi$ , an analytical expression can be attained if  $\phi_n \ll 1$ . One finds

$$S_C^{\text{Cor}}(\tau, \tilde{\tau}) \cong \langle \phi_{12}\phi_{45} \rangle \sim (16\gamma_{\text{NV}}^2 \langle b_{A-1}^2 \rangle / \omega_A^2) \\ \times \sin^4(\omega_A\tau/2) \cos[\omega_A(2\tau + \tilde{\tau})], \quad (7)$$

where, once again, we assumed  $\tilde{\tau} < \tilde{\tau}_c, T_1$ .

With a natural abundance of only  $\sim 1\%$ ,  $^{13}\text{C}$  spins experience relatively weak homonuclear dipolar couplings on the order of  $\sim 100 \text{ Hz}$ , implying that if these were the only noncommuting dynamics, once formed, random coherences should persist for several milliseconds. Similar to the experiment of Fig. 1, one should thus be able to anticipate the effect of the  $^{13}\text{C}$  bath in the above sequence with our engineered AC field  $b_A(t)$ . We do this in Fig. 3(a) for a

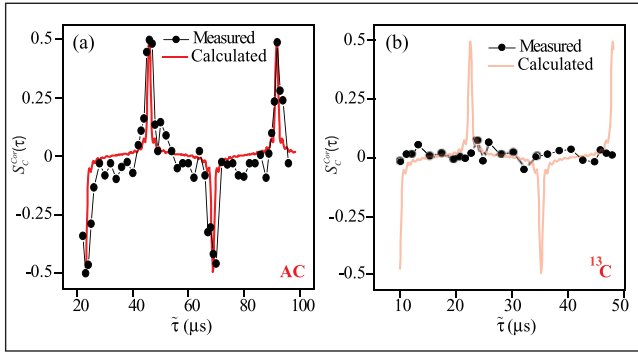


FIG. 3. (Color online) (a) Correlation signal as a function of  $\tilde{\tau}$  for an asynchronous AC field of rms amplitude  $\sqrt{\langle b_{A-1}^2 \rangle} = 4.4 \mu\text{T}$  and frequency  $\omega_A/2\pi = 21.7 \text{ kHz}$ . We set  $\tau = 23 \mu\text{s}$  for  $B_0 = 4 \text{ mT}$ . (b) Analog protocol adapted to  $^{13}\text{C}$  spins. Maintaining  $B_0 = 4 \text{ mT}$ , we set  $\tau = 11.5 \mu\text{s}$  so that  $2\omega_N\tau = 2\pi$ . The faint, solid line is the expected pattern assuming  $\sqrt{\langle b_{N-1}^2 \rangle} = 4.4 \mu\text{T}$  [see Fig. 1(a)] and frequency  $\omega_N/2\pi = 43.5 \text{ kHz}$ .

root mean square (rms) field amplitude  $\sqrt{\langle b_{A-1}^2 \rangle}$  of  $\sim 4 \mu\text{T}$ , as determined from a comparison between the coil- and  $^{13}\text{C}$ -induced echo signals of Fig. 1. In good agreement with our calculations, we find a strong correlation signal whose period coincides with  $\Delta\tilde{\tau} = 2\pi/\omega_A$  as predicted by Eq. (7). We note, however, that, in this case,  $4\gamma_{\text{NV}}\sqrt{\langle b_{A-1}^2 \rangle}/\omega_A \gg 1$  largely exceeds the NV linear range [ $b_{A-1} < 300 \text{ nT}$  in Fig. 2(b)]. This causes interference between sinusoidal [Fig. 2(c)] and rapidly oscillating responses [Fig. 2(d)–2(e)] and leads to a pattern of sharp positive (and negative) peaks correctly reproduced by the numerical computation of Eq. (5) (solid line).

In stark contrast with these observations, however, Fig. 3(b) shows that no distinguishable signal is present when  $^{13}\text{C}$  spins are the source of the fluctuating field, a result we confirmed with observations in several distinct NV centers. This response (or, rather, the lack thereof) is intriguing, particularly given the short timescale of the experiment [where  $\tilde{\tau} \sim 2\tau$  and  $(4\tau + \tilde{\tau}) \ll T_2$ ]. Although one can argue that the effective nuclear field (and, therefore, the amplitude of the resulting modulation pattern) could possibly be smaller than that assumed, we emphasize that  $\sqrt{\langle b_{A-1}^2 \rangle} \sim 4 \mu\text{T}$  largely “saturates” the NV response [see Fig. 2(b)]. With the present signal-to-noise ratio, one can thus rule out fluctuating fields down to  $\sim 100 \text{ nT}$ , which, in turn, is much too small a value to reproduce the collapses and revivals of Fig. 1 [see dashed, faint line in Fig 1(b)]. These observations point to an incomplete description of the relevant NV nuclear spin dynamics under a semi-classical model.

### III. DISCUSSION

#### A. Quantum response of the NV center

To more thoroughly describe the problem, we develop a fully quantum mechanical formulation of the dynamics. We start by defining a Hamiltonian that explicitly takes into account the presence of  $M$  surrounding nuclear spins. Taking

a secular approximation with the zero-field splitting defining the quantization axis, we write

$$H = \Delta S_z^2 - \gamma_{\text{NV}} S_z B_0 + \sum_{q=x,y,z} \sum_{j=1}^M S_z A_{zq}^{(j)} I_q^{(j)} + \omega_{N,0} \sum_{j=1}^M I_z^{(j)} + H'_N. \quad (8)$$

The first two terms denote the NV crystal field and Zeeman contributions, whereas the third term expresses the coupling between the NV center and neighboring nuclear spins  $I^{(j)}$ . Nuclear Zeeman and spin-spin interactions are included through the fourth and fifth terms, respectively. Consistent with proposed sensing applications, in which the NV is only “weakly” coupled to the system under consideration, we assume  $A_{zq} \leq \omega_{N,0}$ . This simplification may indeed apply to the  $^{13}\text{C}$  environment of a small (but non-negligible) fraction of single NVs in our sample (e.g., for  $B_0 \sim 4 \text{ mT}$ , this condition is already met if the nearest  $^{13}\text{C}$  spin is at a distance of  $0.8 \text{ nm}$  or greater; higher fields reduce this distance further). Also, we limit our discussion to times much shorter than  $T_2$ , which allows us to ignore homonuclear dipolar interactions (i.e., we assume  $H'_N \sim 0$ ).

Starting from Eq. (1) and after initializing the system to  $|m_S = 0\rangle\langle m_S = 0| \otimes \rho_N$ , one can show that for an arbitrary pulse sequence of the form  $(\pi/2)_y - t_1 - \text{pulse1} - t_2 - \text{pulse2} - \dots - t_m - (\pi/2)_{-y}$ , the signal is given by

$$S_Q(t) = \left\langle \text{Re} \left\{ \prod_{j=1}^M \text{Tr}_r \left\{ \rho_N^{(j)} [(U_T^\dagger)_{10}(U_T)_{10} + (U_T^\dagger)_{00} \times (U_T)_{11} + (U_T^\dagger)_{00}(U_T)_{10} + (U_T^\dagger)_{10}(U_T)_{11}] \right\} \right\} \right\rangle, \quad (9)$$

where  $\text{Tr}_r$  indicates reduced trace over nuclear spin states,  $U_T$  is the time evolution operator (excluding projection and excitation pulses), and  $(U_T)_{\alpha\beta} \equiv \langle m_S = \alpha | U_T | m_S = \beta \rangle$  with  $\alpha, \beta = \{0, 1\}$  (see Appendix B for details). Equation (9) assumes that all microwave pulses act selectively within the  $m_S = \{0, 1\}$  manifold and that  $\rho_N$  can be expressed as the tensor product  $\rho_N^{(1)} \otimes \rho_N^{(2)} \otimes \rho_N^{(3)} \dots \otimes \rho_N^{(M)}$  of individual nuclear spins (a condition we will revisit later in the manuscript).

For the particular case of the Hahn-echo sequence, we have  $U_T = U(\tau)R_y(\pi)U(\tau)$ , where  $R_y(\pi)$  denotes the  $\pi$ -rotation operator along the  $y$ -axis and  $U(\tau)$  is the free evolution operator during the interpulse interval  $\tau$ . Equation (9) then yields

$$S_Q^{HE}(\tau) = \left\langle \text{Re} \left\{ \prod_{j=1}^M \text{Tr}_r \left\{ \rho_N^{(j)} (U^\dagger)_{00}(U^\dagger)_{11}(U)_{00}(U)_{11} \right\} \right\} \right\rangle. \quad (10)$$

To model random fluctuations in the spin-1/2 nuclear bath, we write  $\rho_N^{(j)}$  in the form  $\rho_N^{(j)} = 1/2 + 2P_N^{(j)}I_z^{(j)} + 2T_N^{(j)}I_x^{(j)}\cos\varepsilon_j + 2T_N^{(j)}I_y^{(j)}\sin\varepsilon_j$ , where  $P_N^{(j)}$ ,  $T_N^{(j)}$  and  $\varepsilon_j$  are stochastic, independent parameters that characterize the polarization and transverse coherence of individual nuclei at the beginning of the echo sequence. After some algebra, we

find (see Appendix C)

$$S_Q^{HE}(\tau) \cong \left\langle 1 - 2 \sin^2(\omega_{N_0} \tau / 2) \left\{ \sum_{j=1}^M \tilde{\theta}_j^2 \sin^2(\omega_{N_1}^{(j)} \tau / 2) + 4 \sum_{\substack{j,k=1 \\ j \neq k}}^M T_N^{(j)} T_N^{(k)} \tilde{\theta}_j \tilde{\theta}_k \sin(\omega_{N_1}^{(j)} \tau / 2) \sin(\omega_{N_1}^{(k)} \tau / 2) \right. \right. \\ \left. \left. \times \sin[(\omega_{N_0} + \omega_{N_1}^{(j)}) \tau / 2 - \tilde{\varphi}_j - \varepsilon_j] \sin[(\omega_{N_0} + \omega_{N_1}^{(k)}) \tau / 2 - \tilde{\varphi}_k - \varepsilon_k] \right\} \right\rangle, \quad (11)$$

where we define  $\vec{\omega}_{N_1}^{(j)} \equiv (A_{zx}^{(j)}, A_{zy}^{(j)}, A_{zz}^{(j)} + \omega_{N_0})$ ,  $\tan \tilde{\theta}_j \equiv [(\omega_{N_1,x}^{(j)})^2 + (\omega_{N_1,y}^{(j)})^2]^{1/2} / \omega_{N_1,z}^{(j)}$  and  $\tan \tilde{\varphi}_j \equiv \omega_{N_1,y}^{(j)} / \omega_{N_1,x}^{(j)}$ . Furthermore, we assume the nuclear Zeeman interaction is much stronger than the transverse hyperfine fields, such that  $\tilde{\theta}_j \ll 1$  for all nuclear spins  $j$ . After ensemble average, the second sum cancels and Eq. (11) reduces to

$$S_Q^{HE}(\tau) \cong 1 - 2 \sin^2(\omega_{N_0} \tau / 2) \sum_{j=1}^M \tilde{\theta}_j^2 \sin^2(\omega_{N_1}^{(j)} \tau / 2) \\ \approx 1 - 2 \sin^4(\omega_{N_0} \tau / 2) \sum_{j=1}^M \tilde{\theta}_j^2. \quad (12)$$

To facilitate comparison with the semi-classical result, we consider the nuclear spins as classical dipoles and express  $b_{N_1}$  in Eq. (3) as the superposition of contributions from classical magnetic moments of transverse amplitude  $T_N^{(j)}$ . We expect this classical transverse component to precess at the Larmor frequency of the static field and couple to the defect through the hyperfine terms. We start by expressing the summed field at the defect site in the form

$$\gamma_{NV} b_{N_1}(t) \\ = \sum_{j=1}^M T_N^{(j)} [A_{zx}^{(j)} \cos(\omega_{N_0} t + \varepsilon_j) + A_{zy}^{(j)} \sin(\omega_{N_0} t + \varepsilon_j)] \\ \cong \omega_{N_0} \sum_{j=1}^M T_N^{(j)} \tilde{\theta}_j \cos(\omega_{N_0} t + \varepsilon_j - \tilde{\varphi}_j), \quad (13)$$

which allows us to calculate the accumulated phase [see Eq. (2)]. Using the same approximation as in Eq. (3) and after taking the ensemble average, we find

$$S_C^{HE}(\tau) \cong \left\langle 1 - 8 \sin^4(\omega_{N_0} \tau / 2) \right. \\ \left. \times \left[ \sum_{j=1}^M T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0} \tau + \varepsilon_j - \tilde{\varphi}_j) \right]^2 \right\rangle \\ = 1 - 4 \sin^4(\omega_{N_0} \tau / 2) \sum_{j=1}^M (T_N^{(j)})^2 \tilde{\theta}_j^2. \quad (14)$$

Interestingly, Eq. (14) predicts a pattern of collapses and revivals virtually identical to the quantum mechanical formula [Eq. (12)]. Yet, the underlying physics is fundamentally different: While the echo modulations present in Eq. (14) depend on nuclear spins having a non-zero transverse polarization, the quantum expression [Eq. (12)] is completely insensitive to the

exact state of the bath at the time of the experiment. Formally, the origin of the modulation arises from Eq. (10), which compares evolution under  $U_{00}$  followed by  $U_{11}$  with evolution in the inverse order. With an anisotropic hyperfine interaction, these two operators do not, in general, commute, thus leading to a decay of the echo signal. However, if  $\omega_{N_0} \tau = 2m\pi$ , then  $U_{00} = \exp(2m\pi i I_z^{(j)}) = 1$  for all nuclei, and a revival takes place regardless of the explicit form of  $\rho_N$ . We therefore conclude that collapses and revivals are not just a consequence of classical precession of the statistical polarization of the bath. Echo modulations persist even if the state of surrounding nuclear spins is known. An important corollary is that echo collapses and revivals must be present even when neighboring nuclei are perfectly polarized before the Hahn-echo sequence is applied.

## B. The effect of the $^{13}\text{C}$ distribution

While the above discussion indicates that collapses and revivals occur for any separable initial nuclear spin configuration, a question relevant in nuclear magnetometry experiments concerns the role played by adjacent (versus distant) spins in an ensemble. Although it is true that dipolar interactions with *individual* spins decay as the inverse cube of the distance  $r$  to the center, the number of nuclei between  $r$  and  $r+dr$  typically increases quadratically. The combined effect, well known in nuclear magnetic resonance, gives rise to slowly decaying, “long-range” dipolar fields, which have proven useful to couple nuclear spins over macroscopic distances.<sup>17–20</sup> To more precisely identify the location of nuclei contributing the most to the echo modulations in Eq. (12), we write

$$\sum_{j=1}^M \tilde{\theta}_j^2 \rightarrow \int_{\text{Nuclear Ensemble}} d^3 r \kappa(\vec{r}) \frac{(A_{zx}^2 + A_{zy}^2)}{\omega_{N_0}^2} \approx \frac{2\pi \kappa C^2}{3r_{\min}^3}. \quad (15)$$

The last expression in Eq. (15) assumes a uniform nuclear spin density  $\kappa(\vec{r}) = \kappa$  and uses  $A_{zx} / \omega_{N_0} = C \sin(2\theta) \cos \varphi / r^3$  and  $A_{zy} / \omega_{N_0} = C \sin(2\theta) \sin \varphi / r^3$  with the proportionality constant  $C \equiv -(3\mu_0 \gamma_{NV} \gamma_N \hbar / 8\pi \omega_{N_0})$ ;  $r_{\min}$  denotes the radius of the shell containing the nearest non-zero-spin nuclei. As usual,  $\mu_0$  is the vacuum magnetic permeability,  $\gamma_N$  denotes the nuclear gyromagnetic ratio,  $\hbar$  is Planck’s constant divided by  $2\pi$ , and  $\theta$  and  $\varphi$  are the polar and azimuthal angles formed by the nuclear-NV vector in the laboratory reference frame (with the  $z$ -axis collinear with the crystal field). After replacing in Eq. (12), we find

$$S_Q^{HE}(\tau) \sim 1 - \frac{4\pi \kappa C^2}{3r_{\min}^3} \sin^4(\omega_{N_0} \tau / 2). \quad (16)$$

Equation (16) exposes a rapid fall-off of the depth of the modulation with distance: For example, for natural abundance  $^{13}\text{C}$  spins in the diamond lattice and at the same 4-mT field, we get  $4\pi\kappa C^2/(3r_{\min}^3) \sim 1$  for  $r_{\min} \sim 2$  nm; these modulations, however, virtually decrease by an order of magnitude if  $r_{\min}$  doubles, implying that relevant nuclear spins lie within a small  $\sim 5$ -nm-radius sphere around the NV center. On a related comment, we mention that in an experiment in which a  $\sim 20$ -nm-diameter diamond nanostructure probes an organic sample,  $4\pi\kappa C^2/(3r_{\min}^3)$  is of order  $10^{-2}$ , indicating that the

interaction with surrounding protons<sup>11</sup> will not lead to a detectable pattern of collapses and revivals. We also note that the constant  $C$  can be scaled by adjusting  $B_0$ ; higher fields can make  $C$  small enough to suppress these effects for the majority of NV centers in natural abundance  $^{13}\text{C}$  diamond.

### C. The effect of $^{13}\text{C}$ spin order

We return to Eq. (11), rewritten below for presentation purposes in the form

$$S_Q^{HE}(\tau) = 1 - 2 \sin^4(\omega_{N_0}\tau/2) \sum_{j=1}^M \tilde{\theta}_j^2 - 8 \sin^4(\omega_{N_0}\tau/2) \left\langle \left[ \sum_{j=1}^M T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0}\tau - \tilde{\varphi}_j - \varepsilon_j) \right]^2 - \sum_{j=1}^M [T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0}\tau - \tilde{\varphi}_j - \varepsilon_j)]^2 \right\rangle, \quad (17)$$

As before, we used the approximation  $\omega_{N_0}^{(j)} \sim \omega_{N_0}$ . Comparison with Eqs. (13) and (14) shows that the term in brackets can be interpreted as resulting from a ‘‘classical’’ nuclear spin field. To examine the role of nuclear spin order, consider the situation in which the nuclear spins exhibit spatial correlations relative to each other while remaining overall asynchronous with respect to the experimental repeat timing. At a given time, suppose, for example, that the parameters  $\varepsilon_j$  are chosen so as to generate the maximum possible dipolar field at the NV site. Here we select a near-optimum configuration where  $\langle T_N^{(j)} \rangle = T_N$  for all  $j$ , and  $\varepsilon_j$  satisfies  $\varepsilon_j(\theta, \phi) = -[\phi + \pi f(\theta) + \varepsilon_0]$  with  $f(\theta) = 0$  if  $\theta \leq \pi/2$  and  $f(\theta) = 1$  otherwise (i.e., the phase grows with the azimuthal angle and spins at opposite sides of the equator point in opposite directions);  $\varepsilon_0$  is a constant, arbitrary initial phase that fluctuates over the many repeats of one observation (i.e., the nuclear bath remains asynchronous). Assuming, as before, a uniform spin density  $\kappa(\vec{r}) = \kappa$  and with the correspondence  $\sum_{j=1}^M \rightarrow \int_{\text{Nuclear ensemble}}$ , one finds after ensemble average

$$[S_Q^{HE}(\tau)]_{\text{Order}} \sim 1 - \sin^4\left(\frac{\omega_{N_0}\tau}{2}\right) \times \left\{ \frac{4\pi\kappa C^2}{3r_{\min}^3} + \left[ 3\pi T_N \kappa C \ln\left(\frac{r_{\max}}{r_{\min}}\right) \right]^2 - \frac{4\pi\kappa C^2 T_N^2}{r_{\min}^3} \right\}, \quad (18)$$

where the subscript emphasizes the special case of assumed nuclear spin order. Expanding Eq. (18) to get the sum of independent terms, we identify two types of contributions: The second and fourth terms [originating from the first and third sums in Eq. (17), respectively] are ‘‘short range,’’ in the sense that they become negligible when  $r_{\min}$  exceeds a few nanometers, as discussed above [Eq. (16)]. The third term, however, is ‘‘long-range,’’ because it grows logarithmically with the size of the nuclear ensemble characterized by  $r_{\max}$  (assumed to be much greater than  $r_{\min}$ ).

Because the nuclear spin order can be simply specified as a well-chosen set of  $\varepsilon_j(\theta, \phi)$ , we use the semi-classical formulae derived above to help us interpret the meaning of this contribution: Starting from the upper half of Eq. (14) and assuming the same relative phases  $\varepsilon_j(\theta, \phi)$ , we find the expression

$$[S_C^{HE}(\tau)]_{\text{Order}} \sim 1 - \sin^4\left(\frac{\omega_{N_0}\tau}{2}\right) \left[ 3\pi T_N \kappa C \ln\left(\frac{r_{\max}}{r_{\min}}\right) \right]^2, \quad (19)$$

implying that  $[S_Q^{HE}(\tau)]_{\text{Order}} \rightarrow [S_C^{HE}(\tau)]_{\text{Order}}$  when  $r_{\min}$  is sufficiently large.

Although Eqs. (18) and (19) are strictly valid only for our chosen *ordered* nuclear spin configuration, spin order leads in general to qualitatively different, nonlocal contributions that may ultimately dominate the signal response. Changing from ‘‘short-’’ to ‘‘long-range’’ can be loosely characterized by a critical value  $r_{\text{crit}}$  dependent on the specific geometry of the problem and defined by the conditions

$$\sum_{j=1}^M [T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0}\tau - \tilde{\varphi}_j - \varepsilon_j)]^2 < \left[ \sum_{j=1}^M T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0}\tau - \tilde{\varphi}_j - \varepsilon_j) \right]^2,$$

and

$$\sum_{j=1}^M [T_N^{(j)} \tilde{\theta}_j \sin(\omega_{N_0}\tau - \tilde{\varphi}_j - \varepsilon_j)]^2 \ll 1. \quad (20)$$

We conclude, therefore, that nuclear spin order marks the emergence of ‘‘long-range’’ fields and leads to ‘‘classical’’ spin magnetometry [Eq. (13)] when all contributing nuclear spins lie beyond  $r_{\text{crit}}$ .

For completeness, we mention that similar mechanisms are at work in the case of a Ramsey sequence (see Appendix D). In analogy to Eq. (11), one can derive a quantum mechanical expression for the Ramsey signal  $S_Q^R(t)$  that explicitly takes into account the longitudinal and transverse couplings with the surrounding nuclear spin bath [Eq. (D3)]. In the absence of a refocusing pulse,  $S_Q^R(t)$  decays monotonically over a time  $T_2^*$  [see Eq. (D4)] in a way that, nonetheless, does not depend on the rms amplitude of the nuclear field fluctuations [compare with Eq. (4)]. Rather than a superposition of signals with slightly different frequencies—our intuitive, classical interpretation of the signal decay—the NV response falls off as a consequence of quantum interference arising from differences in the nuclear spin evolution introduced by the NV center itself [operators  $U_{00}$  and  $U_{11}$  in Eq. (D1)]. A practical outcome is that one cannot “burn a hole”<sup>21</sup> in the magnetic resonance spectrum of a single NV center: As shown in Fig. 4, a 100- $\mu$ s-long, very low power pulse collapses one entire peak of the  $^{14}\text{N}$  triplet. Similar to the Hahn-echo case, the Ramsey

decay is caused by short-range interactions and has a quantum (i.e., nonclassical) origin. Finally, it is not difficult to prove that a classical description of the interplay between the NV center and surrounding nuclear spins re-emerges in the presence of nuclear spin order (see Appendix D).

#### D. Quantum approach to the correlation protocol

We now tackle the pulse protocol of Fig. 2(a) used for measuring time correlations of the magnetic field. On the basis of the expressions derived for the Hahn-echo protocol, one expects non-negligible signatures from the “spin noise” terms in the second sum of Eq. (11) when the experiment is designed to detect long-term correlations of the nuclear bath. Unfortunately, and unlike the Ramsey or Hahn-echo experiments, the pulse sequence is complex enough to make impractical the derivation of analytical formulae. We can, nonetheless, estimate the amplitude of such fluctuations. For this purpose, we consider the second sum in Eq. (11),

$$\begin{aligned}
 V^{HE} &\equiv 8 \sin^2(\omega_{N,0}\tau/2) \sum_{\substack{j,k=1 \\ j \neq k}}^M T_N^{(j)} T_N^{(k)} \tilde{\theta}_j \tilde{\theta}_k \sin\left(\omega_{N,1}^{(j)}\tau/2\right) \sin\left(\omega_{N,1}^{(k)}\tau/2\right) \\
 &\quad \times \sin\left(\left(\omega_{N,0} + \omega_{N,1}^{(j)}\right)\tau/2 - \tilde{\varphi}_j - \varepsilon_j\right) \sin\left(\left(\omega_{N,0} + \omega_{N,1}^{(k)}\right)\tau/2 - \tilde{\varphi}_k - \varepsilon_k\right) \\
 &\approx 8 \sin^4(\omega_{N,0}\tau/2) \sum_{\substack{j,k=1 \\ j \neq k}}^M T_N^{(j)} T_N^{(k)} \tilde{\theta}_j \tilde{\theta}_k \sin(\omega_{N,0}\tau - \tilde{\varphi}_j - \varepsilon_j) \sin(\omega_{N,0}\tau - \tilde{\varphi}_k - \varepsilon_k),
 \end{aligned} \tag{21}$$

and calculate the rms amplitude  $\sigma_V^{HE} = \sqrt{\langle (V^{HE})^2 \rangle}$ . Retaining terms to second order in  $\tilde{\theta}$ , a lengthy but straightforward calculation yields

$$\sigma_V^{HE}(\tau) \approx 8 \sin^4\left(\frac{\omega_{N,0}\tau}{2}\right) T_N^2 \sqrt{\frac{3}{4} \sum_{\substack{j,k=1 \\ j \neq k}}^M \tilde{\theta}_j^2 \tilde{\theta}_k^2}, \tag{22}$$

where we assumed for simplicity  $\langle (T_N^{(j)})^2 \rangle = T_N^2$  for all spins  $j$ . By comparison with Eq. (12), we conclude that when a collapse takes place ( $\omega_{N,0}\tau \sim \pi$ ), bath fluctuations introduce changes in the NV signal as large as the depth of the collapse itself.

To better understand the role of these fluctuations in the correlation protocol of Fig. 2(a), we conducted a numerical simulation that takes into account interactions with hundreds of  $^{13}\text{C}$  spins randomly distributed over a virtual diamond lattice [Fig. 5(a)]. To bring the computing time down to realistic values, we followed a “disjoint cluster” approach, wherein dipolar couplings between distinct nuclear spins in the lattice are taken into account only for clusters of nuclear spins, up to a predefined cluster size threshold.<sup>22</sup> Similar to Fig. 3(a),  $S_Q^{\text{Cor}}(\tau = \pi/\omega_{N,0}, \tilde{\tau})$  exhibits periodic maxima (or minima) when  $\tilde{\tau} = 2m\pi/\omega_{N,0}$  (or  $\tilde{\tau} = (2m+1)\pi/\omega_{N,0}$ ). To expose the influence of homonuclear dipolar couplings on the overall pattern more clearly, we conducted an analogous computation

that ignored the effect of  $H'_N$  altogether. As expected, we find that  $^{13}\text{C}$ - $^{13}\text{C}$  interactions are only responsible for a slow decay over a time  $T_{2,N}$  and can, as before, be ignored when describing the main signal features.

Despite the absence of analytical formulae, one can interpret these results within the framework of our prior discussion. We start by noting that  $S_Q^{\text{Cor}}$  depends on the  $^{13}\text{C}$  spatial distribution and markedly changes from one NV to the next [Figs. 5(a), 5(b)]. Overall, however,  $S_Q^{\text{Cor}}$  never exceeds a few percent of the allowed range (from  $-1$  to  $1$ ), a value below our observation limit [of order  $\sim 10\%$ , see Fig. 3(b)] and thus undetectable in our present experimental conditions. In agreement with Eq. (22), this small signal amplitude is not indicative of weak nuclear fields. Rather than a smooth, sinusoidal oscillation,  $S_Q^{\text{Cor}}$  features sharp crests and valleys, implying that the NV response is far from the linear regime (see Fig. 2 and main narrative in Sec. II). We surmise, therefore, that the null response in Fig. 3(b) is truly the result of a weak correlation between the nuclear spin bath configurations at the beginning of the first and second encoding segments of the detection protocol ( $t = 0$  and  $t = 2\tau + \tilde{\tau}$ , respectively). The latter, in turn, seems to be a consequence of the NV center entanglement with neighboring nuclear spins [an effect we gauged above via the second sum in Eq. (11)]; that is, the very encoding process alters the bath configuration and thus diminishes the correlation between two consecutive observations. Similar to Eq. (18) [and Eq. (D5)



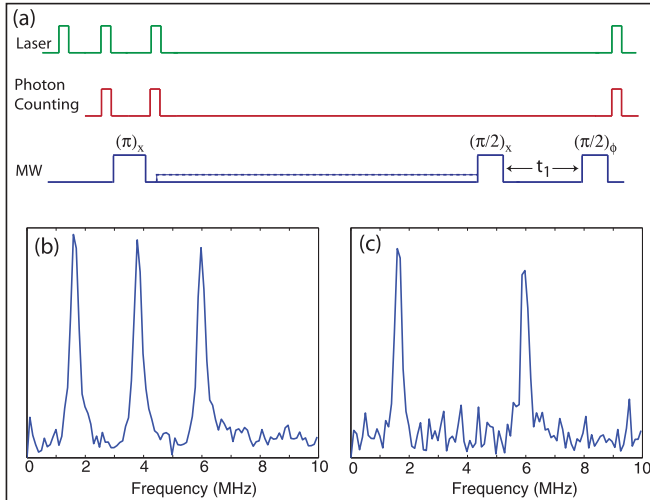


FIG. 4. (Color online) Microwave hole burning experiment shows an inability to burn a hole in the NV line. (a) Schematics of the pulse sequence: The NV center is polarized and a bright- and dark-state photon counting rate reference is taken. Then, an optional, long pulse is applied; the pulse power ( $\sim 80$  kHz) corresponds to less than one-fifth the line width of an individual resonance in the triplet. Finally, a Ramsey-style readout is performed. The phase of the second  $90^\circ$  pulse in the Ramsey sequence is incremented in TPPI fashion to mimic an offset frequency.<sup>21</sup> Appropriate Fourier transform processing of the Ramsey data reveals the high-resolution NV spectrum. (b) In the case of no saturation pulse, we see the expected  $^{14}\text{N}$  triplet. (c) Application of the saturation pulse collapses the entire center peak, rather than burning a hole.

in Appendix D], we speculate that the classical response [Fig. 3(a)] resurfaces when the state of individual nuclear spins correlates, at least partly, with that of their neighbors and when all spins are located beyond  $r_{\text{crit}}$ . The details of this transition will likely be the subject of further work.

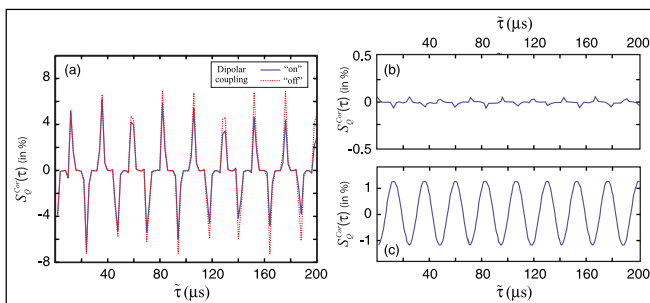


FIG. 5. (Color online) (a) “Disjoint cluster” simulation of the NV response to the correlation protocol of Fig. 2(a). Whether or not homonuclear dipolar couplings are included in the simulation (solid and dashed lines, respectively), the calculated pattern resembles that in Fig. 3(a), although of much smaller amplitude. For these calculations, we considered a total of 834  $^{13}\text{C}$  nuclei randomly distributed over a  $21 \times 21 \times 21$ -unit-cell diamond lattice. (b) Simulations over different NV centers display strong signal variability. Calculations over diverse  $^{13}\text{C}$  environments indicate, however, that  $S_Q^{\text{Cor}}$  never exceeds a few percent of the allowed maximum. (c) As expected,  $S_Q^{\text{Cor}}$  transitions to a sinusoidal shape when  $^{13}\text{C}$  spins are restricted to more than 10 unit cells away from the NV center.

#### IV. CONCLUSION

The use of NV centers as a platform toward mesoscopic nuclear spin magnetometry relies on detection protocols designed to interrogate large numbers of nuclear spins weakly coupled to the center. Throughout the present study, we investigated the effect of these couplings on the response of NV centers to common pulse sequences, namely the Ramsey and Hahn-echo protocols. We found strong phenomenological similarities with the case where, rather than interacting with neighboring nuclear spins, the NV center evolves in the presence of a classical magnetic field designed to emulate nuclear spin bath fluctuations.

In the limit where nuclear spins are independent from each other, a closer examination reveals that these similarities are only superficial, and that fundamentally different mechanisms are at work. For the particular case of the Hahn-echo sequence, our analysis suggests that the pattern of collapses and revivals must persist, for example, if the nuclear bath is initialized to the fully polarized state. Contrary to widespread perception, we also find that field inhomogeneities produced over time by a fluctuating nuclear spin environment are not the only cause for the signal decay in a Ramsey protocol. Instead, our analysis indicates that both collapses and revivals and Ramsey decoherence are inherently quantum mechanical in nature: Rather than a decay due to random phase accumulation, the NV center becomes nonobservable because of entanglement with the nuclear spin bath. This process results from anisotropic interactions effectively constrained to a small, nanometer-diameter volume and will thus be negligible if probed nuclear spins are sufficiently removed from the center. In this limit, we find that the combined NV-center/nuclear-spin system can be described well by semi-classical equations when the bath exhibits long-range order.

Even though both semi-classical and quantum descriptions ultimately predict identical Hahn-echo or Ramsey patterns, the seemingly subtle underlying differences have a dramatic effect in more general scenarios. We illustrated the problem through a pulse sequence designed to provide information on the long-term correlations of a fluctuating magnetic field acting on the NV center. Although our approach succeeds in exposing the central frequency, rms amplitude and correlation time of a classical, random magnetic field, it fails to extract the same information when we consider nuclear spins close to the center as the source of the field. We interpret our observations as an unwelcome but unavoidable consequence of the interrogation process itself, whose influence on the bath evolution alters the time correlation it would otherwise exhibit.

Disjoint cluster simulations of the integrated probe-bath system suggest that not all information is lost and that the nuclear spin correlation pattern should be observable if higher signal-to-noise ratio conditions are reached. In this regard, one possibility could be, for example, the use of diamond crystals configured in the form of a solid-immersion lens, where much higher NV fluorescence count rates have been reported.<sup>23,24</sup> Such observations would shed light on the nuclear bath dynamics without the limitations imposed by NV decoherence. In particular, one could, for example, tailor the protocol to gauge directly  $^{13}\text{C}$  spin relaxation caused by varying NV illumination during the free-evolution time  $\tilde{\tau}$ .<sup>25</sup> Finally, and

although the initial state of the bath considered here only takes into account single spin coherences [Eq. (10)], one can envision modified sequences in which stochastic multiple-quantum coherences of the nuclear system are systematically converted into single-quantum coherences before observation. This class of schemes could be useful, for instance, to provide information on the mesoscale distribution of  $^{13}\text{C}$  spins (whose relatively weak couplings do not lead to resolved splittings in the NV spectrum).

### ACKNOWLEDGMENTS

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### APPENDIX A: DERIVATION OF EQ. (5)

Starting from Eq. (1) and using  $\rho_{\text{NV}}(0) = |0\rangle\langle 0| = (1/2)(1 + \sigma_z)$ , we find at the end of the correlation protocol  $(\pi/2)_y - \tau - (\pi)_y - (\pi/2)_x - \tilde{\tau} - (\pi/2)_y - \tau - (\pi)_y - (\pi/2)_x$

$$\rho_{\text{NV}}(2\tau + \tilde{\tau}) = \frac{1}{2}[1 + \sigma_y \cos \phi_{12} \cos \phi_3 - \sigma_x(\sin \phi_{12} \cos \phi_{45} + \sin \phi_3 \cos \phi_{12} \sin \phi_{45}) + \sigma_z(\sin \phi_{12} \sin \phi_{45} - \sin \phi_3 \cos \phi_{12} \cos \phi_{45})], \quad (\text{A1})$$

where we used the notation in the main text and  $\phi_3 = (2\gamma_{\text{NV}}b_{A,1}/\omega_A) \sin(\omega_A \tilde{\tau}/2) \cos(\omega_A \tau + \omega_A \tilde{\tau}/2 + \varepsilon_A^{(3)})$ , with  $\varepsilon_A^{(3)}$  denoting the magnetic field phase at the beginning of the  $\tilde{\tau}$  interval;  $\sigma_x, \sigma_y, \sigma_z$  denote, as usual, the set of Pauli matrices.

On the other hand, the modified protocol  $(\pi/2)_y - \tau - (\pi)_x - (\pi/2)_{-x} - \tilde{\tau} - (\pi/2)_y - \tau - (\pi)_y - (\pi/2)_x$  yields

$$\rho_{\text{NV}}(2\tau + \tilde{\tau}) = \frac{1}{2}[1 - \sigma_y \cos \phi_{12} \cos \phi_3 + \sigma_x(-\sin \phi_{12} \cos \phi_{45} + \sin \phi_3 \cos \phi_{12} \sin \phi_{45}) + \sigma_z(\sin \phi_{12} \sin \phi_{45} + \sin \phi_3 \cos \phi_{12} \cos \phi_{45})]. \quad (\text{A2})$$

Therefore, if the proper phase cycling is introduced from one repeat to the next, we obtain

$$S_C^{\text{Cor}}(\tau, \tilde{\tau}) = \langle \sin \phi_{12} \sin \phi_{45} \rangle \quad (\text{A3})$$

in agreement with Eq. (5). For completeness, we mention that for the conditions of the present experiment we find (both experimentally and numerically) that  $\langle \sin \phi_3 \cos \phi_{12} \cos \phi_{45} \rangle \sim 0$ . In practice, this makes the phase cycling unnecessary. Also, we note that other correlation functions (including  $\langle \cos \phi_{12} \cos \phi_{45} \rangle$ ,  $\langle \sin \phi_{12} \cos \phi_{45} \rangle$ , etc.) can be obtained with a proper selection of the relative phases within the same general pulse protocol.

### APPENDIX B: DERIVATION OF EQ. (9)

We start with equation (1) of the main text for the signal at the end of the experiment:  $S_Q = \text{Tr}\{\sigma_z R_{90}^\dagger U_T R_{90}(|0\rangle\langle 0| \otimes \rho_N) R_{90}^\dagger U_T^\dagger R_{90}\}$ . We assume an initial state  $|0\rangle\langle 0| \otimes \rho_N$ ;  $U_T$  is the evolution operator for the pulse sequence of the total electron-nuclear space excluding the  $90^\circ$  excitation and measurement pulses  $R_{90}, R_{90}^\dagger$ , respectively. As all other pulses in the sequences presented herein, these pulses are restricted to the two-dimensional subspace  $\{m_S = 0, 1\}$  of the electron spin. The excitation and measurement pulses rotate the final observable from  $\sigma_z$  to  $\sigma_x$  and the initial electron state from  $|0\rangle\langle 0|$  to  $\frac{1}{2}[|0\rangle\langle 0| + |0\rangle\langle 1| + |1\rangle\langle 0| + |1\rangle\langle 1|]$ . Putting these into Eq. (1) and using the fact that in the outer-product form  $\sigma_x = |0\rangle\langle 1| + |1\rangle\langle 0|$ , we have

$$\begin{aligned} S &= \frac{1}{2} \text{Tr}[(|0\rangle\langle 1| + |1\rangle\langle 0|)U_T(|0\rangle\langle 0| + |0\rangle\langle 1| + |1\rangle\langle 0| + |1\rangle\langle 1|)\rho_N U_T^\dagger] \\ &= \frac{1}{2} \text{Tr}_r\{(U_T)_{00}\rho_N(U_T^\dagger)_{01} + (U_T)_{00}\rho_N(U_T^\dagger)_{11} + (U_T)_{01}\rho_N(U_T^\dagger)_{01} + (U_T)_{01}\rho_N(U_T^\dagger)_{11} \\ &\quad + (U_T)_{10}\rho_N(U_T^\dagger)_{00} + (U_T)_{10}\rho_N(U_T^\dagger)_{10} + (U_T)_{11}\rho_N(U_T^\dagger)_{00} + (U_T)_{11}\rho_N(U_T^\dagger)_{10}\}. \end{aligned} \quad (\text{B1})$$

In Eq. (B1), we identify four complex conjugate pairs, allowing us to take only twice the real part of the second line of the sum. For example:

$$(\text{Tr}_r\{(U_T)_{00}\rho_N(U_T^\dagger)_{01}\})^* = \text{Tr}_r\{[(U_T)_{00}\rho_N(U_T^\dagger)_{01}]^\dagger\} = \text{Tr}_r\{(U_T)_{10}\rho_N(U_T^\dagger)_{00}\}. \quad (\text{B2})$$

Thus,

$$S = \text{Re}\{\text{Tr}_r[(U_T)_{10}\rho_N(U_T^\dagger)_{00} + (U_T)_{10}\rho_N(U_T^\dagger)_{10} + (U_T)_{11}\rho_N(U_T^\dagger)_{00} + (U_T)_{11}\rho_N(U_T^\dagger)_{10}]\}, \quad (\text{B3})$$

which takes us to Eq. (9) of the text under the assumption of a tensor-product nuclear initial state.

**APPENDIX C: DERIVATION OF EQ. (11).**

For presentation purposes, we rewrite Eq. (10) in the form  $S_Q^{HE}(\tau) = \langle \text{Re} \{ \prod_{j=1}^M S_j^{HE} \} \rangle$ , where we defined

$$S_j^{HE} \equiv \text{Tr}_r \{ \rho_N^{(j)} (U^\dagger)_{00} (U^\dagger)_{11} (U)_{00} (U)_{11} \} = \text{Tr}_r \{ \rho_N^{(j)} e^{i\omega_{N,0} I_z^{(j)} \tau} e^{i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} e^{-i\omega_{N,0} I_z^{(j)} \tau} e^{-i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} \}. \quad (\text{C1})$$

After some algebra and using  $|\pm\rangle$  to denote the two eigenfunctions of  $I_z$  (nuclei are assumed to have spin  $I = 1/2$ ), one can show that

$$\begin{aligned} \langle + | e^{i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} | + \rangle &= (\langle - | e^{i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} | - \rangle)^* = \cos^2 \left( \frac{\tilde{\theta}_j}{2} \right) e^{i\tilde{\omega}_{N-1}^{(j)} \tau/2} + \sin^2 \left( \frac{\tilde{\theta}_j}{2} \right) e^{-i\tilde{\omega}_{N-1}^{(j)} \tau/2} \\ \langle + | e^{i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} | - \rangle &= -(\langle - | e^{i\tilde{\omega}_{N-1}^{(j)} \cdot \vec{I}^{(j)} \tau} | + \rangle)^* = i \sin \tilde{\theta}_j \sin \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) e^{-i\tilde{\varphi}_j}, \end{aligned} \quad (\text{C2})$$

and from here we get

$$\begin{aligned} S_j^{HE} &= 1 - 2 \sin^2 \tilde{\theta}_j \sin^2 \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin^2 \left( \frac{\omega_{N,0} \tau}{2} \right) + i P_N^{(j)} \sin^2 \tilde{\theta}_j \sin^2 \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin(\omega_{N,0} \tau) \\ &\quad - 4i T_N^{(j)} \sin \tilde{\theta}_j \sin \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin \left( \frac{\omega_{N,0} \tau}{2} \right) \left\{ \cos^2 \left( \frac{\tilde{\theta}_j}{2} \right) \sin \left[ (\omega_{N,0} + \omega_{N-1}^{(j)}) \frac{\tau}{2} - \tilde{\varphi}_j - \varepsilon_j \right] \right. \\ &\quad \left. + \sin^2 \left( \frac{\tilde{\theta}_j}{2} \right) \sin \left[ (\omega_{N,0} - \omega_{N-1}^{(j)}) \frac{\tau}{2} - \tilde{\varphi}_j - \varepsilon_j \right] \right\} \end{aligned} \quad (\text{C3})$$

For  $\tilde{\theta}_j \ll 1$ , Eq. (C3) takes the form

$$\begin{aligned} S_j^{HE} &= 1 - 2\tilde{\theta}_j^2 \sin^2 \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin^2 \left( \frac{\omega_{N,0} \tau}{2} \right) + i P_N^{(j)} \tilde{\theta}_j^2 \sin^2 \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin(\omega_{N,0} \tau) - \\ &\quad - 4i T_N^{(j)} \tilde{\theta}_j \sin \left( \frac{\omega_{N-1}^{(j)} \tau}{2} \right) \sin \left( \frac{\omega_{N,0} \tau}{2} \right) \sin \left[ (\omega_{N,0} + \omega_{N-1}^{(j)}) \frac{\tau}{2} - \tilde{\varphi}_j - \varepsilon_j \right] \end{aligned} \quad (\text{C4})$$

Noting that  $S_j^{HE}$  has the form  $S_j^{HE} = (1 - a_j) + i b_j$  with  $a_j, b_j < 1$ , we find

$$\begin{aligned} \text{Re} \left\{ \prod_{j=1}^M S_j^{HE} \right\} &\approx (1 - a_1) \cdots (1 - a_M) - b_1 b_2 (1 - a_3) \cdots (1 - a_M) - (1 - a_1) b_2 b_3 (1 - a_4) \cdots (1 - a_M) \\ &\quad - \cdots - (1 - a_1) \cdots (1 - a_l) b_l (1 - a_{l+1}) \cdots (1 - a_r) b_r (1 - a_{r+1}) \cdots (1 - a_M) - \cdots + \text{higher order terms}, \end{aligned} \quad (\text{C5})$$

which leads to

$$\begin{aligned} S_Q^{HE}(\tau) &\cong \left\langle 1 - 2 \sin^2 (\omega_{N,0} \tau/2) \left\{ \sum_{j=1}^M \tilde{\theta}_j^2 \sin^2 (\omega_{N-1}^{(j)} \tau/2) + 4 \sum_{\substack{j,k=1 \\ j \neq k}}^M T_N^{(j)} T_N^{(k)} \tilde{\theta}_j \tilde{\theta}_k \sin (\omega_{N-1}^{(j)} \tau/2) \sin (\omega_{N-1}^{(k)} \tau/2) \right. \right. \\ &\quad \left. \left. \times \sin [(\omega_{N,0} + \omega_{N-1}^{(j)}) \tau/2 - \tilde{\varphi}_j - \varepsilon_j] \sin [(\omega_{N,0} + \omega_{N-1}^{(k)}) \tau/2 - \tilde{\varphi}_k - \varepsilon_k] \right\} \right\rangle. \end{aligned} \quad (\text{C6})$$

**APPENDIX D: RAMSEY SEQUENCE**

Starting from Eq. (9), we find for a Ramsey sequence

$$S_Q^R(t) = \left\langle \text{Re} \left\{ \prod_j^M \text{Tr}_r \{ (U^\dagger)_{00} (U)_{11} \rho_N^{(j)} \} \right\} \right\rangle, \quad (\text{D1})$$

where  $U$  is the free evolution operator over time  $t$ . Using  $\rho_N^{(j)} = 1/2 + 2P_N^{(j)}I_z^{(j)} + 2T_N^{(j)}I_x^{(j)}\cos\varepsilon_j + 2T_N^{(j)}I_y^{(j)}\sin\varepsilon_j$ , we obtain

$$\begin{aligned} \text{Tr}_r\{(U^\dagger)_{00}(U)_{11}\rho_N^{(j)}\} &= \cos^2\left(\frac{\tilde{\theta}}{2}\right)\cos\left(\frac{(\omega_{N,0}-\omega_{N,1}^{(j)})t}{2}\right) + \sin^2\left(\frac{\tilde{\theta}}{2}\right)\cos\left(\frac{(\omega_{N,0}+\omega_{N,1}^{(j)})t}{2}\right) \\ &\quad + 2iP_N^{(j)}\left\{\cos^2\left(\frac{\tilde{\theta}}{2}\right)\sin\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(j)})t}{2}\right] + \sin^2\left(\frac{\tilde{\theta}}{2}\right)\sin\left[\frac{(\omega_{N,0}+\omega_{N,1}^{(j)})t}{2}\right]\right\} \\ &\quad - 2iT_N^{(j)}\sin\tilde{\theta}_j\sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right)\cos\left(\frac{\omega_{N,0}t}{2}-\tilde{\varphi}_j-\varepsilon_j\right). \end{aligned} \quad (\text{D2})$$

After replacing in Eq. (D1) and in the approximation discussed in Appendix C, we find

$$S_Q^R(t) = \left\langle A - \frac{1}{2}\sin\left(\frac{\omega_{N,0}t}{2}\right)\sum_{j=1}^M B_j \tilde{\theta}_j^2 \sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right) - 2\sum_{\substack{j,k \\ j \neq k}}^M C_{jk} (P_N^{(j)}P_N^{(k)}\Gamma_{jk} + T_N^{(j)}T_N^{(k)}\mathbf{K}_{jk} - 2T_N^{(j)}P_N^{(k)}\Lambda_{jk}) \right\rangle, \quad (\text{D3})$$

where

$$A = \prod_{j=1}^M \cos\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(j)})t}{2}\right]; \quad B_j = \prod_{\substack{k=1 \\ k \neq j}}^M \cos\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(k)})t}{2}\right]; \quad C_{jk} = \prod_{\substack{l=1 \\ l \neq j,k}}^M \cos\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(l)})t}{2}\right];$$

and

$$\begin{aligned} \Gamma_{jk} &= \left(1 - \frac{\tilde{\theta}_j^2}{2}\right)\sin\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(j)})t}{2}\right]\sin\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(k)})t}{2}\right] + \frac{\tilde{\theta}_j^2}{2}\sin\left[\frac{(\omega_{N,0}+\omega_{N,1}^{(j)})t}{2}\right]\sin\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(k)})t}{2}\right] \\ \mathbf{K}_{jk} &= \tilde{\theta}_j\tilde{\theta}_k\sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right)\sin\left(\frac{\omega_{N,1}^{(k)}t}{2}\right)\cos\left(\frac{\omega_{N,0}t}{2}-\tilde{\varphi}_j-\varepsilon_j\right)\cos\left(\frac{\omega_{N,0}t}{2}-\tilde{\varphi}_k-\varepsilon_k\right) \\ \Lambda_{jk} &= \tilde{\theta}_j\sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right)\cos\left(\frac{\omega_{N,0}t}{2}-\tilde{\varphi}_j-\varepsilon_j\right)\sin\left[\frac{(\omega_{N,0}-\omega_{N,1}^{(k)})t}{2}\right]. \end{aligned}$$

Assuming independent, uncorrelated nuclear spins and after ensemble average,  $\Gamma_{jk}$ ,  $\mathbf{K}_{jk}$ , and  $\Lambda_{jk}$  cancel and we can rewrite Eq. (D3) in the simpler form

$$S_Q^R(t) = A - \frac{1}{2}\sin\left(\frac{\omega_{N,0}t}{2}\right)\sum_{j=1}^M B_j \tilde{\theta}_j^2 \sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right) \sim \exp\left[-\left(\frac{t}{T_2^*}\right)^2\right]\left[1 - \frac{1}{2}\sin\left(\frac{\omega_{N,0}t}{2}\right)\sum_{j=1}^M \tilde{\theta}_j^2 \sin\left(\frac{\omega_{N,1}^{(j)}t}{2}\right)\right], \quad (\text{D4})$$

where  $T_2^*$  is a parameter characterizing the time decay resulting from  $A$  and  $B_j$ .

When spin order is present and assuming  $r_{\min} > r_{\text{crit}}$  (see main text),  $\sum_{\substack{j,k=1 \\ j \neq k}}^M \rightarrow (\sum_{j=1}^M)^2$  and Eq. (D3) can be cast in the form

$$[S_Q^R(t)]_{\text{Order}} \approx \exp\left[-\left(\frac{t}{T_2^*}\right)^2\right]\left\{1 - \frac{1}{2}\left\langle\left(\sum_{j=1}^M P_N^{(j)}A_{zz}^{(j)}t\right)^2\right\rangle - \sin^2\left(\frac{\omega_{N,0}t}{2}\right)\left\langle\left[\sum_{j=1}^M T_N^{(j)}\tilde{\theta}_j\cos\left(\frac{\omega_{N,0}t}{2}-\varepsilon_j-\tilde{\varphi}_j\right)\right]^2\right\rangle\right\}. \quad (\text{D5})$$

Starting from  $S_C^R(t) = \langle\cos\phi(t)\rangle$  with  $\phi(t) = \int_0^t dt' \gamma_{\text{NV}}b_N(t')$  (Eq. (4) in the main text), we regain Eq. (D5) if we make use of the correspondence with the field produced by classical magnetic moments

$$\begin{aligned} \gamma_{\text{NV}}b_{N,0} &\rightarrow \sum_{j=1}^M P_N^{(j)}A_{zz}^{(j)} \\ \gamma_{\text{NV}}b_{N,1}\cos(\omega_{N,0}t + \varepsilon_N) &\rightarrow \sum_{j=1}^M T_N^{(j)}[A_{zx}^{(j)}\cos(\omega_{N,0}t + \varepsilon_j) + A_{zy}^{(j)}\sin(\omega_{N,0}t + \varepsilon_j)] \\ &\cong \omega_{N,0}\sum_{j=1}^M T_N^{(j)}\tilde{\theta}_j\cos(\omega_{N,0}t + \varepsilon_j - \tilde{\varphi}_j), \end{aligned} \quad (\text{D6})$$

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