Nonlinear Radiating Instability of a Barotropic Eastern Boundary Current

The MIT Faculty has made this article openly available. Please share how this access benefits you. Your story matters.

<table>
<thead>
<tr>
<th></th>
<th></th>
</tr>
</thead>
<tbody>
<tr>
<td>As Published</td>
<td><a href="http://dx.doi.org/10.1175/JPO-D-12-0174.1">http://dx.doi.org/10.1175/JPO-D-12-0174.1</a></td>
</tr>
<tr>
<td>Publisher</td>
<td>American Meteorological Society</td>
</tr>
<tr>
<td>Version</td>
<td>Final published version</td>
</tr>
<tr>
<td>Accessed</td>
<td>Sun Dec 09 05:41:37 EST 2018</td>
</tr>
<tr>
<td>Citable Link</td>
<td><a href="http://hdl.handle.net/1721.1/85076">http://hdl.handle.net/1721.1/85076</a></td>
</tr>
<tr>
<td>Terms of Use</td>
<td>Article is made available in accordance with the publisher's policy and may be subject to US copyright law. Please refer to the publisher's site for terms of use.</td>
</tr>
<tr>
<td>Detailed Terms</td>
<td></td>
</tr>
</tbody>
</table>
Nonlinear Radiating Instability of a Barotropic Eastern Boundary Current

JINBO WANG
Massachusetts Institute of Technology/WHOI Joint Program, Cambridge, Massachusetts

MICHAEL A. SPALL
Woods Hole Oceanographic Institution, Woods Hole, Massachusetts

GLENN R. FLIERL AND PAOLA MALANOTTE-RIZZOLI
Massachusetts Institute of Technology, Cambridge, Massachusetts

(Manuscript received 11 September 2012, in final form 24 March 2013)

ABSTRACT

Linear and nonlinear radiating instabilities of an eastern boundary current are studied using a barotropic quasigeostrophic model in an idealized meridional channel. The eastern boundary current is meridionally uniform and produces unstable modes in which long waves are most able to radiate. These long radiating modes are easily suppressed by friction because of their small growth rates. However, the long radiating modes can overcome friction by nonlinear energy input transferred from the more unstable trapped mode and play an important role in the energy budget of the boundary current system. The nonlinearly powered long radiating modes take away part of the perturbation energy from the instability origin to the ocean interior. The radiated instabilities can generate zonal striations in the ocean interior that are comparable to features observed in the ocean. Subharmonic instability is identified to be responsible for the nonlinear resonance between the radiating and trapped modes, but more general nonlinear triad interactions are expected to apply in a highly nonlinear environment.

1. Introduction

Energetic mesoscale variability in the ocean interior has long been observed (Wyrtki et al. 1976; Stammer 1997) and motivated many studies concerning its origins. The most direct mechanism is baroclinic instability of the large-scale circulation. Gill et al. (1974) noted that the potential energy of large-scale wind-driven gyres in the ocean is several orders of magnitude larger than their kinetic energy. The vast stores of potential energy might be released through baroclinic instability to generate mesoscale eddies much stronger than the mean circulation (Gill et al. 1974; Robinson and McWilliams 1974). Although theory suggests that the eddy energy in weakly sheared zonal flows is limited by the shear of the mean flow (Pedlosky 1975), an introduction of a meridional component to the mean flow allows for eddy kinetic energy exceeding that of the mean (Robinson and McWilliams 1974; Spall 2000).

An alternative mechanism is related to the radiation of mechanical energy from swift oceanic boundary currents, such as the Gulf Stream. Many studies represent the Gulf Stream as a propagating northern boundary (Flierl and Kamenkovich 1975; Pedlosky 1977; Harrison and Robinson 1979; Malanotte-Rizzoli et al. 1987). These results identify important mechanisms governing the energy radiation from strong ocean currents. Talley (1983) derives the wave properties by solving for the stability of a steady zonal flow and shows that instability radiation will not occur unless there is a westward component in the zonal current or the far field is made baroclinic. The main argument is that the wave characteristics of eastward-traveling instabilities do not match the dispersion relation of the free Rossby waves in the far field.

Instability radiation occurs more easily for a nonzonal current. Pedlosky (1993) studied a baroclinic shear flow that is inclined with respect to a latitude circle. Although the study focuses on the generation of a boundary current
by the boundary trapping of two reflected unstable waves, it clearly demonstrates that a nonzonal flow is less stable and its instabilities can reach out to the far region. Kamenkovich and Pedlosky (1996, 1998a,b) explicitly studied the influence of nonzonality on jet instability and radiation. They found that even a slight nonzonality in the mean flow can generate radiating instabilities, which can significantly penetrate into the far field.

A meridional current is an extreme case of the nonzonality. The instabilities generated by a meridional western boundary current are able to radiate eastward even in the presence of realistic dissipation (Fantini and Tung 1987). Hristova et al. (2008) studied the radiating instabilities of meridional boundary currents and compared a western boundary current with an eastern boundary current. They showed that an eastern boundary current supports a greater number of radiating modes over a wider range of meridional wavenumbers than a western boundary current. However, these studies used a piecewise constant meridional velocity profile. This velocity profile reduces the stability equation to an ordinary differential equation with constant coefficients, which is easier to solve, but leads to unrealistic shortwave behaviors. Wang et al. (2012) extended these previous studies by considering a continuous velocity profile in a barotropic quasigeostrophic model and demonstrated that radiating instabilities can generate zonal striaions in the ocean interior. The nonlinear effects were shown to be important in energizing the instability radiation. As a follow-up study, we here discuss the mechanism that governs the nonlinear radiating instabilities of an eastern boundary current.

2. The model

We choose the simplest barotropic quasigeostrophic (QG) model, as used in Fantini and Tung (1987) and Wang et al. (2012), to focus on elementary nonlinear dynamics. The model is described by the barotropic vorticity equation:

\[ \psi = \psi_0 + \psi' \quad \text{and} \quad q = q_0 + q' = \nabla^2 \psi + \beta y + \nabla^2 \psi', \]

(2)

where the perturbation field is much weaker than the basic state. The basic state is assumed to be balanced by a steady external forcing

\[ \mathcal{F} = J(\overline{\psi}, \overline{q}) = -\overline{\psi}_x \nabla^2 \overline{\psi}_x + \overline{\psi}_y \nabla^2 \overline{\psi}_y + \beta \overline{\psi}_x, \]

(3)

where the subscripts denote partial derivatives. Here, \( \overline{\psi} \) is not the time-mean streamfunction, and \( \psi' \) can have a mean part; when we require means, we will denote them specifically by a time integral. For a basic state with only a zonal flow, no external forcing is needed as \( \overline{\psi}_x = 0 \). However, an external forcing is necessary to maintain a basic meridional flow. It takes the form \( \mathcal{F} = \beta \psi \), where \( (\overline{\psi}, \overline{q}) = (-\overline{\psi}_y, \overline{\psi}_x) \) is the steady-state barotropic flow, meaning a vorticity source or sink is needed to compensate the planetary vorticity change caused by the meridional movement of a water parcel.

After substituting Eqs. (2) and (3), Eq. (1) becomes

\[ \partial_t q' + J(\overline{\psi}, q') + J(\psi', \overline{q}) + J(\psi', q') = \mathcal{F}', \]

(4)

where \( \mathcal{F}' \) represents frictional effects on perturbations. This equation is solved numerically to investigate the influence of nonlinearity on the instability of a meridional current. To study linear stability, the quadratic term in \( \psi' \) is neglected, resulting in the linear stability equation

\[ \partial_t q' + J(\overline{\psi}, q') + J(\psi', \overline{q}) = \mathcal{F}'. \]

(5)

In the following sections, we first study the linear and then the nonlinear stability problem of an eastern boundary current by numerically solving Eqs. (5) and (4), respectively.

3. Linear radiating instability

a. Stability equation

Let us consider a basic state with a parallel meridional eastern boundary current \( \overline{v}(x) \). We set \( x = -L \) at the western boundary and \( x = 0 \) at the eastern boundary. The linear stability Eq. (5) can be nondimensionalized using the cross-stream length scale \( L_0 \) and the velocity scale \( V \) of the boundary current. After dropping primes without causing confusion, the linearized Eq. (5) becomes

\[ (\partial_t + \overline{v} \partial_x) \nabla^2 \psi + \beta \psi_x - \overline{v} \psi_y - \mathcal{F}(\psi) = 0, \]

(6)
where $\beta^* = \beta L^2_b/V$ is the nondimensional beta, and $f(\psi)$ is the frictional damping on perturbations, which can be specified as Laplacian diffusion of vorticity $V \cdot A_{H} \nabla q$, where $A_{H}$ is the horizontal viscosity. Nonzero $A_{H}$ will be used in the nonlinear simulations. Previous studies, for example, Fantini and Tung (1987) and Hristova et al. (2008), further simplify it to an ordinary differential equation for example, Fantini and Tung (1987) and Hristova et al. (2008).

The simplified stability equation for an inviscid, barotropic, quasigeostrophic, meridional current. Previous studies, including Wang et al. (2012): linear and nonlinear theories. The profile is similar to the boundary current to aid the comparison between stabilities with infinitely large wavenumbers.

This is the stability equation for an inviscid, barotropic, quasigeostrophic, meridional current. Previous studies, for example, Fantini and Tung (1987) and Hristova et al. (2008), further simplify it to an ordinary differential equation by choosing a step function to represent the boundary current $\overline{v}$. The simplified stability equation becomes easier to solve, but there is no shortwave cutoff. The discontinuous velocity profile generates a delta function in the background vorticity and supports instabilities with infinitely large wavenumbers.

In this study, we use a continuous profile to represent the boundary current to aid the comparison between linear and nonlinear theories. The profile is similar to Wang et al. (2012): $\overline{v}(x) = -V \text{sech}^2 \left( \frac{x - x_0}{L_b} \right)$, where $x_0$ is the location of the center of the boundary current.

Two methods are used to solve the linear problem with a continuous velocity $\overline{v}(x)$. The linear eigenvalue problem represented by Eq. (8) with proper boundary conditions discussed below is solved by a shooting method. The second method considers an initial value problem for both inviscid and viscous cases. The linear model equation in Eq. (6) is initialized with a sine wave, whose growth rates are calculated by fitting the time series of the square root of the domain-integrated energy to an exponential curve. The boundary conditions for the initial value problem are no-normal-flow at the eastern and western solid boundaries and periodic conditions in the meridional direction. The boundary conditions needed for the shooting method are discussed as follows.

### b. Boundary conditions and radiating modes

The no-normal-flow boundary condition is applied at the solid eastern boundary,

$$\phi(0) = 0.$$  \hspace{1cm} (10)

In the regions away from the boundary current, say $x = -L_e$, $\overline{v}$ approaches zero, so that perturbation eigenfunctions satisfy the free Rossby wave dispersion relation and their zonal structures is proportional to $e^{ikx}$, where the zonal wavenumber $k$ satisfies

$$k^2 + \frac{\beta^*}{L_e} k + \overline{f}^2 = 0.$$  \hspace{1cm} (11)

The boundary condition at $x = -L_e$ is the radiation condition

$$\phi_L(-L_e) = ik\phi(-L_e),$$  \hspace{1cm} (12)

where $k$ is one of the two solutions of Eq. (11). It is chosen in such a way that it either represents a decay structure or an outward (westward in this case) propagation of wave energy. The two solutions of $k$, say $k^{(1)}$ and $k^{(2)}$, have opposite-signed imaginary parts since their product, $k^{(1)}k^{(2)} = \overline{f}^2$ is real. The eigenfunction $\phi(x)$ of an eastern boundary current decays westward, requiring the imaginary part of $k$, that is, $\nu_k$, to be negative.

Negative $\nu_k$ also corresponds to westward Rossby wave group velocity in the limit of $\nu_k \rightarrow 0$ (Hristova et al. 2008).

Extra analysis is required to distinguish radiating modes from trapped modes, as they both decay into the far field. The decaying structure of an unstable radiating mode is formed because the packets of perturbations take finite time to propagate away into the interior with relatively unchanged amplitude, meanwhile grow exponentially near the energy source. The amplitude of an unstable radiating mode is always larger over the jet region than in the interior. A necessary condition used to identify a radiating mode is the so-called phase speed condition, that is, the dispersion relation of a radiating mode has to match the dispersion relation of the free waves in the interior, Rossby waves in this case (Talley 1983).

### c. Results

Figure 1 shows the linear growth rate $\nu_k$ and frequency $\omega$, as a function of wavenumber $l$, where $\nu_k$ and $\omega$ are the imaginary and real parts of $c$ and $\omega = lc$, respectively. Note that there are two unstable modes for a sech$^2$ velocity profile on an f plane, which are a sinusoidal mode and a varicose mode, corresponding to symmetric and
antisymmetric eigenfunction structures, respectively (Lipps 1962). In our case, the beta effect and side boundary effect alter the original eigenfunctions resulting in modified sinuous and varicose modes, which are indicated by the solid and dashed lines in Fig. 1.

Figure 1a shows that the most unstable mode is a sinuous mode (solid line) at $l = 1.05$. The modified sinuous mode has a shortwave cutoff at $l = 1.88$. No unstable sinuous modes are found for $l < 0.706$. Unstable modes over the longwave range are modified varicose modes (dashed line). Instabilities exist for $0.125 < l < 1.04$. The calculation by the initial value method for the inviscid case is shown by the stars. The good agreement between the lines and the symbols validates both methods. Frictional and nonlinear terms are added to the linear numerical model used in the initial value problem to study nonlinear effect in the next section.

Figure 1b shows the frequency $\omega_r$. The lines indicate $\omega_r = l c_r$, while the symbols represent $\omega_r = -\beta k_x/(k_y^2 + \ell^2)$. According to the phase speed condition, the two frequencies at each $l$ should be equal in order for instabilities to radiate. For the modified varicose mode (dashed line), the circles fall on the dashed line over the wavenumber range $l < 0.46$, meaning the longwave modes with $l < 0.46$ are expected to radiate. Similarly, long sinuous modes (solid line) with $l < 0.74$ are also expected to radiate, but only over a very narrow wavenumber range $0.706 < l < 0.74$. The two critical wavenumbers separating the radiating and trapped modes are $l = 0.46$ for the modified varicose mode and $l = 0.74$ for the modified sinuous mode (the two dotted vertical lines in Fig. 1b).

One example of the eigenfunctions for radiating mode and trapped mode are shown in Fig. 2. The eigenfunctions show a wavy structure for the radiating mode, but a fast-decay structure for the trapped mode. The linear analysis shows that radiating instabilities occur over the longwave end for each mode (sinuous or varicose). It is qualitatively consistent with Kamenkovich and Pedlosky (1996) and Hristova et al. (2008). Although radiating modes are able to transfer energy away from the unstable region to affect the interior, they have smaller growth rates than the most unstable mode in this inviscid linear theory. Friction can suppress the unstable inviscid radiating modes, leaving the significance of the radiating mode in question.

4. Nonlinear radiating instability

The exponential growth of an initially infinitesimal perturbation slows down when the perturbation becomes finite and starts to feed back into the mean. Once an initial small perturbation develops to finite amplitude, linear theory fails and nonlinear interaction becomes important for the perturbation development. The elementary mechanism for nonlinear interaction is the triad resonance, which describes that a triad of waves are resonant if their phases satisfy $\theta_1 \pm \theta_2 \pm \theta_3 = 0$, which is equivalent to both wavenumbers and frequencies satisfying $l_1 \pm l_2 \pm l_3 = 0$ and $\omega_1 \pm \omega_2 \pm \omega_3 = 0$ (Phillips 1960). Here we are interested in identifying the effect of nonlinear interaction on the radiating instabilities from the point of view of elementary triad resonance. To
single out the nonlinear process among a wave triad, we weakly force the model to generate only one linearly unstable mode (denoted as the primary mode henceforth), suppress other modes by friction, and look for a wave triad that is resonant. A case with a nonresonant wave triad is presented in section 4d as a comparison.

The model is described by Eq. (4), where the damping is provided by Laplacian diffusion of vorticity $F = \nabla \cdot A_H \nabla q$ (after dropping primes). Here, $A_H$ is strongly increased at the western boundary to remove energy and enstrophy as used in Fox-Kemper (2003):

$$A_H = A_H^w - (A_H^w - A_H^e) \exp \left( \frac{x}{\alpha L_x} \right),$$

where $\alpha$ controls the decay scale. This function changes from approximately $A_H^w$ at the western boundary to $A_H^e$ at the eastern boundary. Different values of $\alpha$ are tested to confirm that our conclusions are not sensitive to $\alpha$ as long as $A_H$ is approximately $A_H^w$ over the eastern boundary current region. The frictional effect in the interior is not energetically important as the perturbation field has a large horizontal scale (see the appendix for the scaling argument). Bottom friction was also tested to confirm that our conclusion about the nonlinear interaction between trapped and radiating modes does not depend on the specific form of friction (not shown). Cases with $\alpha = 0.15$ are presented here. We use $A_H^w = 10^4 \text{m}^2\text{s}^{-1}$ and $A_H^e = 100 \text{m}^2\text{s}^{-1}$ in this study. The perturbation energy budget over the eastern boundary current region varies for different values of $A_H$.

The model domain is an $L_x \times L_y$ meridional channel discretized using $N_x \times N_y$ grid points, which are specified in each following section. The boundary conditions are periodic in the meridional direction, with no-normal-flow and slip conditions along the solid walls.

The velocity profile is described by Eq. (9) with $V = 0.11\ \text{m}\text{s}^{-1}$, $L_b = 50\ \text{km}$, and $x_0 = -100\ \text{km}$. The speed $[O(10\ \text{cm}\text{s}^{-1})]$ and the width of the boundary current are consistent with observations (Hickey 1979; Davis 1985; Brink and Cowles 1991). Other velocity profiles were also tested without altering our final conclusions.

In the following, we first describe the methodology for identifying discrete unstable modes, then show the nonlinear evolution of the discrete modes in two scenarios. One scenario has a resonant wave triad, and the other has no resonant wave triad. Results for different scenarios are also discussed in terms of energetics.

a. Linear growth rates

To single out a discrete unstable mode, we first examine the linear growth rates calculated based on Eq. (5) for the specified boundary current using the initial value method with friction considered. This calculation uses $\beta = 1.8 \times 10^{-11} \ (\text{m} \text{s}^{-1})$, $L_x = L_y = 5000\ \text{km}$, and $N_x = N_y = 256$. The resulting spatial resolution $19.53\ \text{km}$ is sufficient for the most unstable mode, which has a meridional wavelength about $350\ \text{km}$.

The model is initialized with an infinitesimal sine wave in $y$ with a discrete wavenumber $\ell_n = 2\pi/L_y$, where $n$ is an integer varying from 8 to 20. The corresponding wavelengths $\lambda_n = 2\pi/\ell_n$ range from 625 to $250\ \text{km}$ given $L_y = 5000\ \text{km}$.

Figure 3 shows the growth rate as a function of meridional wavelength. There are longwave and shortwave cutoffs at wavelength about 550 and $250\ \text{km}$, respectively. The resolved most unstable mode has wavelength about $350\ \text{km}$. According to the calculation in Fig. 1b, the critical wavenumber dividing the radiating and trapped modes is $l = 0.46$ for the modified sinuous mode and $l = 0.74$ for the modified varicose mode, corresponding to two critical wavelengths $\lambda = 683$ and $\lambda = 425\ \text{km}$, respectively. As a result, modes with a wavelength smaller than $425\ \text{km}$ are trapped ones and those with a wavelength larger than $683\ \text{km}$ are able to radiate. This is true for both modified varicose and sinuous modes.

b. Case I, nonlinear resonance

In this experiment, we use the same set of parameters as those used in section 4a, but with $L_y = 700\ \text{km}$ and $N_y = 32$ (the spatial resolution is reduced to $21.88\ \text{km}$ from $19.53\ \text{km}$). We can pick a limited number of unstable modes by reducing spectral resolution $2\pi/L_y$ because instabilities are confined in a narrow range of wavelength from 250 to $550\ \text{km}$. Given $L_y = 700\ \text{km}$, the longest mode resolved in the model has a wavelength $\lambda_1 = 700\ \text{km}$. The second and third modes have wavelength $\lambda_2 = 350$ and $\lambda_3 = 233\ \text{km}$, respectively. The
growth rate curve in Figure 3 shows that only the second mode \( n = 2 \) has a positive growth rate. In the following, we set the model to support only one unstable mode, and test the two proposed scenarios. Hereafter we refer the mode with wavenumber \( \ell_n = 2\pi \ell / L_y \) as \( M_n \).

Figure 4 shows the streamfunction of \( M_1 (\lambda_1 = 700 \text{ km}) \) and \( M_2 (\lambda_2 = 350 \text{ km}) \). Here, \( M_1 \) is a radiating mode with a streamfunction that decays very slowly westward. The shorter wave mode \( M_2 \), however, is trapped around the boundary current region.

The frequencies for the first four modes are listed in Table 1. It is clear that \( \omega_2 + 2\omega_1 = 0 \), meaning the frequency relationship between \( M_1 \) and \( M_2 \) satisfies the requirement for resonance. We also notice that the frequencies of \( M_2, 3, 4 \) are negative, corresponding to negative phase speed. The longer mode \( M_1 \), however, has a positive frequency (equivalently a positive phase speed), indicating it is a retrograde mode.

Here, \( M_1 \) and \( M_2 \) satisfy the nonlinear resonance criteria as \( \ell_2 = 2\ell_1 \) and \( \omega_2 + 2\omega_1 = 0 \). We now initialize the nonlinear model with random noise, and integrate it in time. Different linear modes can interact and the resonance between \( M_1 \) and \( M_2 \) is expected.

Figure 5 shows an example of the time series of the perturbation streamfunction at a fixed station in the boundary region and two snapshots of perturbation streamfunction at two developing stages. Three different developing stages are clearly distinguishable in the time series. The amplitude of the streamfunction over stage I \((t = 0 \ldots 4.5)\) is negligible compared to the finite amplitude at later time. After a period of exponential growth, the second stage (stage II) is reached and sustained between about \( t = 4.5 \) and \( t = 38 \). There is another apparent transition period around \( t = 38 \), after which perturbations reach the third stage (stage III) with a stronger oscillation.

One should notice that distinguishable stages result from the weakness of the forcing and the slow development of the instability. The forcing scale and the model spectral resolution are reduced to such a level that the model supports only one unstable mode. The slow development of the instability allows for the clear illustration of the physical processes. The system will reach equilibrium significantly faster if the forcing scale or the spectral resolution is increased for the system to support larger growth rates or more than one unstable mode. In a test experiment (not shown) a 9% increase in the amplitude of the basic state shortens the time scale for the emergence of stage II to 0.6 from the original 4.5.

![FIG. 4. Streamfunctions at an arbitrary time corresponding to \( \psi_1 \) with \( \lambda_1 = 700 \text{ km} \) and \( \psi_2 \) with \( \lambda_2 = 350 \text{ km} \). They are normalized by their maximum value as they are linear modes. Only the subdomain from \( x = -2500 \text{ km} \) to 0 is plotted to show a clearer streamfunction structure.](image)

**Table 1. Frequencies and growth rates of the first four modes.**

<table>
<thead>
<tr>
<th>Mode index</th>
<th>( \lambda ) (km)</th>
<th>( \omega_r ) (cpy)</th>
<th>( \omega_i ) (cpy)</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>700</td>
<td>1.44</td>
<td>-0.1554</td>
</tr>
<tr>
<td>2</td>
<td>350</td>
<td>-2.88</td>
<td>0.2630</td>
</tr>
<tr>
<td>3</td>
<td>233</td>
<td>-6.48</td>
<td>-0.1555</td>
</tr>
<tr>
<td>4</td>
<td>175</td>
<td>-10.08</td>
<td>-1.1945</td>
</tr>
</tbody>
</table>

![FIG. 5. (a) The time series of the streamfunction at a fixed station in the boundary region. The time axis is nondimensionalized by \( 2\pi / \omega_i \) in which \( \omega_i \) of \( M_2 \) is chosen. The streamfunction snapshots at (b) \( t = 20 \) and (c) \( t = 50 \) are shown. Numerals I, II and III indicate different stages. Unit \( \text{m}^2 \text{s}^{-1} \) is used in all panels.](image)
Two snapshots of the perturbation streamfunction explain the different stages shown in the time series. During stage II, only \( M_2 \) is energetically significant. The existence of \( M_1 \) with energy extending into the interior is not evident in stage II but clearly shown in stage III (Fig. 5c).

To identify the mechanism that governs the nonlinear perturbation development, we study the time evolution of enstrophy for each mode. The Fourier coefficients for each mode are saved at each time step. The enstrophy for each discretized mode is \( |\tilde{q}(n,t)|^2 \), where \( \tilde{q} \) is the complex Fourier coefficient of vorticity, \( n \) represents the mode index, and \( t \) represents time. The time series of the enstrophy for each mode is shown in Fig. 6.

The initial linear development for each mode can be clearly observed in the time series of enstrophy because the parameters used in the integration produce small linear growth rates. Before \( t = 0.7 \), \( M_2 \) (solid line) is the only growing mode, which is expected from the linear study. It grows fastest and reaches nonlinear equilibrium at about \( t = 4.5 \).

The growth of \( M_4 \) (dashed line) after the initial period of decay is caused by the self interaction of \( M_2 \) as \( M_4 \) is the superharmonic of \( M_2 \) in terms of wavenumber. This is essentially a special case of the triad interaction described in Phillips (1960). However, \( M_4 \) is only a forced mode by \( M_2 \) without resonance because their frequencies do not satisfy the resonance relation. The initial decay of \( M_4 \) indicates the initial adjustment of the model to the random noise initialization.

The time series of \( M_1 \) and \( M_3 \) are similar, but different from those of \( M_2 \) and \( M_4 \). Here, \( M_1 \) and \( M_3 \) decay for \( t < 4.5 \) with approximately the same negative growth rate, which is consistent with the linear result shown in Fig. 3. They, however, start to grow at \( t = 4.5 \), when \( M_2 \) reaches equilibrium. Linearly decaying modes become nonlinearly unstable, and the linearly unstable mode is eventually equilibrated by nonlinear interaction and friction.

The growths of \( M_1 \) and \( M_3 \) are obviously caused by their nonlinear interactions with \( M_2 \). The nonlinear interaction between \( M_2 \) and \( M_1 \) is what we expected as their wavenumbers and frequencies satisfy the nonlinear resonance relation (refer to Table 1). Here, \( M_1 \) starts to be resonant with the primary wave \( M_2 \), after \( M_2 \) obtains finite amplitude and equilibrates when the energy input from the background mean jet balances the energy loss to dissipation and radiation (quantified in the next section). The growth of \( M_3 \) is caused by the interaction between \( M_1 \) and \( M_2 \) without resonance. During the nonlinear growth, \( M_1 \) and \( M_3 \) have amplitudes that are too small to feed back into the primary wave \( M_2 \). The nonlinear growth rates of \( M_1 \) and \( M_3 \) stay constant, until they reach nonnegligible amplitudes around \( t = 38 \).

The fact that \( M_1 \) does not grow from the start but from \( t = 4.5 \) indicates an important parameter restriction regarding the resonant interaction in the weakly nonlinear regime. For there to be a resonant interaction, the growth rate of \( M_2 \) and the decay rate of \( M_1 \) must be small. Mathematically, the growth rates of the unstable modes have to be small to define fast and slow time scales as done in Pedlosky (1970); intuitively, the growth of each mode should be slower than its oscillation for a wave triad to have enough time to interact and exchange energy before their amplitudes grow or decay significantly. This condition is violated by the fast growth of \( M_2 \) during \( t < 4.5 \) (Table 1) resulting in a nonresonance condition between \( M_1 \) and \( M_2 \). Instead, the resonance starts after \( M_2 \) stops growing at \( t = 4.5 \). \( M_1 \) starts to grow because the energy gain of \( M_1 \) through nonlinear energy transfer overcomes the energy loss by the direct action of friction.

In the final equilibrium after \( t = 38 \), the whole system reaches a new balance in which the main players are \( M_2 \) and \( M_1 \); \( M_3 \) and \( M_4 \) are several orders of magnitude smaller than the two main modes because they are passively forced modes. As the model is weakly forced, the final state is still weakly nonlinear and the interaction between \( M_1 \) and \( M_3 \) is still clear. Otherwise, resonance loses its meaning in a very nonlinear regime as resonant interaction is no longer a privileged interaction over others.

In summary, there is one primary wave, \( M_2 \), growing linearly during the initial adjustment period. A secondary instability occurs after the primary linear instability reaches a nonlinear equilibrium and is driven by the nonlinear resonance between the primary wave \( M_2 \) and the secondary wave \( M_1 \). In the particular circumstance in this study, the secondary wave is the subharmonic of the primary wave. The instability between two harmonics is often referred to as subharmonic instability, which is well studied in the context of internal
gravity waves (McComas and Bretherton 1977, e.g.), but not for barotropic shear instability of meridional currents. Nonlinear self interaction of $M_2$ and the triad interaction that involves $M_1$ and $M_2$ also produce growth for other modes with larger wavenumbers, for example, $M_3$ and $M_4$. Their amplitudes, however, are smaller than $M_1$ and $M_2$.

The cause of the three stages observed in the time series of perturbation streamfunction (Fig. 5a) is clear in the time evolution of the enstrophy for individual modes. During stage I, the primary wave $M_2$ grows exponentially to gain finite amplitude. During stage II, $M_2$ stops growing and starts to generate the secondary instability resulting in the growth of $M_1$. During stage III, system reaches equilibrium with two energetic modes, $M_1$ and $M_2$.

**c. Energetics**

The processes mentioned in the previous section can be additionally illuminated by diagnosing the energy budgets of the whole system and each normal mode.

Figure 7 shows the time series of the perturbation kinetic energy integrated over the whole domain (thick line) and 1000-km-wide eastern boundary domain (thin line) both normalized by $EKE_{total}$. The discrepancy between these two lines implies the influence of the radiating mode.

\[
\psi = N/2 \sum_{n=-N/2}^{N/2} \psi_n = N/2 \sum_{n=-N/2}^{N/2} \phi_n(x,t) \exp(i n y) \quad \text{and} \\
q = N/2 \sum_{n=-N/2}^{N/2} q_n = N/2 \sum_{n=-N/2}^{N/2} \zeta_n(x,t) \exp(i n y), \tag{14}
\]

where $\phi_n = \phi^* n$ to make the variable $\psi$ and $q$ real (the asterisk represents complex conjugate), the $\ell$ represents the wavenumber of the longest model-resolved wave defined as $\ell = 2\pi/L_y$, and $\zeta_n = (\partial_{xx} - n^2 \ell^2)\phi_n$. The linear Jacobians in Eq. (4) become

\[
J(\psi, \nabla^2 \psi + \beta y) + J(\overline{\psi}, \nabla^2 \psi) = \sum_{n=-N/2}^{N/2} \sum_{j=-N/2}^{N/2} \left[ i \ell (\frac{\partial \phi_m}{\partial x} \zeta_j - \frac{\partial \phi_j}{\partial x} \zeta_m) \right] e^{i(j+m)y}. \tag{15}
\]

The discretized Jacobian for perturbations becomes

\[
J(\psi, q) = \sum_{m=-N/2}^{m=N/2} \sum_{j=-N/2}^{j=N/2} \left( i \ell \left( \frac{\partial \phi_m}{\partial x} \zeta_j \right) \right) \exp(i(j+m)y). \tag{16}
\]

The terms with $j + m = n$ will act as new forcings to the $n$th mode $\psi_n$. As a result, Eq. (16) can be rewritten as

\[
J(\psi, q) = \sum_{n=-N/2}^{N/2} \sum_{m=-N/2}^{N/2} J_n \exp(i n y) \quad \text{and} \\
J_n = \sum_{m=-N/2}^{m=N/2} i \ell \left[ (n-m) \frac{\partial \phi_m}{\partial x} \zeta_{n-m} - m \phi_m \frac{\partial \zeta_{n-m}}{\partial x} \right], \tag{17}
\]

where $J_n \exp(i n y)$ represents the new forcing on the $n$th mode by nonlinear interaction. For example, for a truncated model with only mode $n = 0, \pm 1, \pm 2$, the nonlinear term for each mode becomes

\[
J_0 = i \ell \frac{\partial}{\partial x} (\phi_1^0 \zeta_1 - \phi_1^0 \zeta_1^0) + i 2 \ell \frac{\partial}{\partial x} (\phi_2^a \zeta_2 - \phi_2^a \zeta_2^0), \tag{18}
\]

\[
J_1 = i \ell \left( \frac{\partial \phi_0}{\partial x} \zeta_1 - \phi_1 \frac{\partial \zeta_0}{\partial x} + i \ell \left( \frac{2 \partial ^2 \phi_0}{\partial x^2} \zeta_2 + \phi_1 \frac{\partial ^2 \zeta_2}{\partial x^2} - \frac{\partial \phi_0}{\partial x} \zeta_1^0 - 2 \phi_2 \frac{\partial \zeta_2}{\partial x} \right) \right), \tag{19}
\]

\[
J_2 = i 2 \ell \left( \frac{\partial \phi_0}{\partial x} \zeta_2 - \phi_2 \frac{\partial \zeta_0}{\partial x} \right) + i \ell \left( \frac{\partial \phi_1}{\partial x} \zeta_1 - \phi_1 \frac{\partial \zeta_1}{\partial x} \right). \tag{20}
\]

The terms in $J_0$ represent the feedback of the perturbations on the basic state. The terms in the first set of
Table 2. The kinetic energy budgets integrated over the eastern boundary current region (−1000 km < x < 0) at different stages. All values are normalized by the total energy input by Reynolds stress at their own stage. The longwave mode is weak at the first stage, but it becomes significant at the second stage. It consequently alters the route of energy transfer.

<table>
<thead>
<tr>
<th>Budget terms</th>
<th>Stage II Total</th>
<th>Stage II Stage III</th>
<th>Stage II Stage III</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>n = 1</td>
<td>n = 2</td>
<td>n = 1</td>
</tr>
<tr>
<td>Rey</td>
<td>1</td>
<td>0.0002</td>
<td>0.997</td>
</tr>
<tr>
<td>Fric</td>
<td>-1.0083</td>
<td>-0.0001</td>
<td>-0.98</td>
</tr>
<tr>
<td>Flux</td>
<td>-0.0016</td>
<td>-0.0001</td>
<td>-0.004</td>
</tr>
<tr>
<td>Non</td>
<td>—</td>
<td>10^{-6}</td>
<td>—</td>
</tr>
</tbody>
</table>

parentheses in $J_1$ and $J_2$ represent the interaction of $M_1$ and $M_2$ with the basic state, respectively. The terms in the second set of parentheses in $J_1$ represent the interaction between $M_1$ and $M_2$ acting as a new forcing to $M_1$. The terms in the second set of parentheses in $J_2$ indicate that the self-interaction of $M_1$ imposes a new forcing to $M_2$.

After substituting Eqs. (14), (15), and (17), Eq. (4) becomes

$$0 = \int_L \left[ -R \phi_n \frac{\partial^2 \phi_n}{\partial x^2} + \beta \frac{\partial \phi_n}{\partial t} + \frac{\beta}{2} \| \phi_n \|^2 - A_H \left( \phi_n \frac{\partial^3 \phi_n}{\partial x^3} - \frac{\partial \phi_n}{\partial x} \frac{\partial^2 \phi_n}{\partial x^2} - 2n^2 \frac{\partial^2 \phi_n}{\partial x^2} \frac{\partial \phi_n}{\partial x} \right) \right] dx dt$$

$$+ \int_L \left[ \beta \phi_n \frac{\partial^2 \phi_n}{\partial x^2} + \frac{\beta}{2} \| \phi_n \|^2 - A_H \left( \phi_n \frac{\partial^3 \phi_n}{\partial x^3} - \frac{\partial \phi_n}{\partial x} \frac{\partial^2 \phi_n}{\partial x^2} - 2n^2 \frac{\partial^2 \phi_n}{\partial x^2} \frac{\partial \phi_n}{\partial x} \right) \right] dx dt$$

where $A_H$ is assumed to be a constant. This assumption simplifies but does not change the perturbation energy equation shown next because $A_H$ mainly act on perturbations with large amplitude and small scales over the boundary current region where $A_H$ is approximately constant.

The energy equation for the $n$th mode is $(\phi_n G_n + \phi_n G_{-n})/2$. After some algebra (see appendix for details), the energy equation for the $n$th mode is

$$\frac{\partial \xi_n}{\partial t} + J_n + \beta \frac{\partial \phi_n}{\partial x} = intl \left( \phi_n \frac{\partial^3 \phi_n}{\partial x^3} - \frac{\partial \phi_n}{\partial x} \frac{\partial^2 \phi_n}{\partial x^2} - 2n^2 \frac{\partial^2 \phi_n}{\partial x^2} \frac{\partial \phi_n}{\partial x} \right) + A_H \left( \frac{\partial^2 \xi_n}{\partial x^2} - n^2 \xi_n \right),$$

At stage II, the only dominate mode $M_2$ is responsible for 99.7% of the total energy input and for about the same amount of energy that is dissipated. Here, $M_1$ has negligible contribution either in energy input or loss because of its small amplitude. The nonlinear energy transfer between these two modes is infinitesimal.

At stage III, both $M_1$ and $M_2$ play important roles in the energy balance. About 30% of the total energy input is done by $M_1$ and 70% is done by $M_2$. Overall, friction accounts for 77.5% of the energy removal for the boundary region. Most (73.5%) of this 77.5% is due to $M_2$. The remaining energy removal (out of the boundary region) is done by radiation, mainly because of the divergent term related to the beta effect $(\partial/\partial x)(\beta/2)||\psi||^2$ (not shown). The energy radiation is done by $M_1$.

The energy budget for each mode is far from being closed without considering the nonlinear energy transfer between the two modes. More energy is needed for $M_1$ to be steady. However, the nonlinear energy transfer has
no net contribution to the total energy budget because the energy gain in \( M_1 \) corresponds to an equal energy loss in \( M_2 \). The total energy budget is closed within 3.5% of the total energy gain.

The quantity \( M_1 \) has a positive contribution to the energy input during stage III. In the linear inviscid calculation, the growth rate of \( M_1 \) is positive: \( \omega_i = 0.045 \) corresponding to \( l = 2\pi(L_0/L_y) = 0.45 \) (Fig. 1), which means that \( M_1 \) can draw energy from the basic state. In the linear viscous calculation, the growth rate of \( M_1 \) becomes negative (Table 1) because the energy loss by friction surpasses the energy gain through Reynolds stress, which is still positive. In the nonlinear simulation, however, the energy gain of \( M_1 \) through both nonlinear energy transfer and Reynolds stress surpasses the energy drain by friction to yield a positive growth rate. Once \( M_1 \) reaches a finite amplitude, the energy gain of \( M_1 \) by the Reynolds stress becomes nonnegligible.

In addition, the characteristics of \( M_1 \) can be altered by its nonlinear interaction with \( M_2 \). We test this by an extended experiment, in which the model is truncated to have only the basic state and \( M_1 \), starts from the state at \( t = 59.4 \) with the same parameters and stops at \( t = 63 \).

Figure 8 shows the energy of \( M_1 \) (Fig. 8a) and the streamfunction of \( M_1 \) at \( t = 59.4 \) (Fig. 8b) and \( t = 63 \) (Fig. 8c). Since there is no energy support from \( M_2, M_1 \) decays (Fig. 8a), which is qualitatively consistent with the prediction by the linear theory (Fig. 3). The structures of \( M_1 \) before (Fig. 8b) and after (Fig. 8c) the truncation are different in terms of both zonal tilt and the detailed structures over the boundary current region. Because of the nonlinear interaction with \( M_2, M_1 \) can alter its structure so that the energy transfer from the mean switches from negative, which is predicted by the linear viscous theory, to positive. This phenomenon is clear in the energy budget in Table 2, where the rey (the energy transfer from the mean by the Reynolds stress) of \( M_1 \) is positive and accounts for 30% of the total energy input.

This experiment shows that, in a strict sense, Fourier modes in the weakly nonlinear simulations are slightly different from their linear equivalence. The unstable mode can alter the long, stable mode. The longwave mode at finite amplitude is not sustainable without the support of the shortwave mode through nonlinear wave-wave interaction.

Another nonlinear effect is that the finite amplitude perturbations feed back into the mean [the term \( J_0 \) in Eq. (18)] to change the structure of the basic boundary current. The modified boundary current will certainly change the stability properties and yield different eigenmodes. The new eigenmodes do not satisfy the requirement for triad-resonance, so that the two harmonics decouple, leading to a weaker eddy field. The boundary current will then relax to its unaffected state, generating again the triad resonance to start a new cycle. This mechanism can be detected from the minipanel in Fig. 6 where the energy in different modes oscillates during stage III. The oscillations of \( M_1 \) and \( M_2 \) are almost out of phase, which can be a sign of changing state of resonance between the two modes. Given less resonance between the two modes, \( M_1 \) decays and \( M_2 \) grows for the less nonlinear energy transfer from \( M_2 \) to \( M_1 \). Similarly, \( M_1 \) grows and \( M_2 \) decays during more resonance state. This oscillation can be dynamically similar to the finite-amplitude oscillations shown in Pedlosky (1970), however, the detailed study of the oscillation mechanism is beyond the scope of this paper.

We summarize the route of the energy transfer at stage III by the diagram shown in Fig. 9 as follows. The mean boundary jet is supported by an external forcing. It becomes unstable and transfers energy into the eddy field. Here, 70% of this energy goes to the most unstable, trapped mode \( M_2 \), and 30% goes to the long radiating mode \( M_1 \). Of the energy that goes into \( M_2 \), 57% is dissipated locally, and 13% is injected into \( M_1 \) through nonlinear energy transfer. The quantity \( M_1 \) drains 39%
of total energy, of which dissipation accounts for 14% and radiation for 25%. Overall, 57% of the total energy loss is due to the trapped mode and 43% is due to the radiating mode. Although the exact energy partition shown here should not be regarded as an accurate guideline for the eastern boundary currents in the real ocean, this simulation clearly demonstrates that the linearly decaying radiating mode and the most unstable trapped mode become almost equally important in the energy budget for an eastern boundary region.

d. Case 2, nonresonant triad

We keep the same model setup used in the previous section, but we increase $\beta$ from $1.8 \times 10^{-11}$ to $2 \times 10^{-11}$ (m s)$^{-1}$. The change of $\beta$ alters the frequency of the radiating mode $M_1$, so that the frequency requirement for nonlinear resonance, $\omega_2 = 2\omega_1$, is not satisfied.

Figure 10 shows the time series of the perturbation streamfunction at a random location in the boundary current (Fig. 10a), a snapshot of the perturbation field in the equilibrium state (Fig. 10b), and the time evolutions of enstrophy for the first four modes (Fig. 10c). The evolution of enstrophy shows that the unstable mode $M_2$ quickly stands out from the initial random noise and dominates the system after that. Here, $M_1$ and $M_3$ decay until their values reach machine precision, and $M_4$ is sustained by the nonlinear self-interaction of $M_2$ but with a negligible amplitude.

The frequencies for the first four modes are listed in Table 3, showing a substantial mismatch between $\omega_2$ and $2\omega_1$. As a result, the resonance criterion is not satisfied, and no significant nonlinear interaction happens between these two modes.

The energy budget for this case is similar to the stage II in the previous case, so is not repeated here.

5. Conclusions and discussion

The energy radiation from swift oceanic currents can potentially contribute to the energy in the ocean interior. The orientation of the currents plays an important role in the radiating capability of the instabilities generated by unstable currents. Motivated by previous studies, we investigated the nonlinear radiating instability of a barotropic eastern boundary current.

It is found that the linearly decaying and radiating modes can resonate with an unstable trapped mode to become nonlinearly unstable. These long radiating waves play an important role in the energy budget of an eastern boundary current, as well as in exporting energy into the ocean interior. In our experiment, 25% of total energy gained by instabilities is radiated into the ocean interior.

In our specific model setup, the nonlinear process is related to subharmonic instability, which is caused by the nonlinear resonance between two harmonics. We identify this process by reducing model spectral resolution to single out one unstable mode, then by varying the nondimensional beta to look for wave triad that is

<table>
<thead>
<tr>
<th>Mode index</th>
<th>$\lambda$ (km)</th>
<th>$\omega_1$ (cpy)</th>
</tr>
</thead>
<tbody>
<tr>
<td>1</td>
<td>700</td>
<td>2.16</td>
</tr>
<tr>
<td>2</td>
<td>350</td>
<td>-2.88</td>
</tr>
<tr>
<td>3</td>
<td>233</td>
<td>-6.48</td>
</tr>
<tr>
<td>4</td>
<td>175</td>
<td>-10.08</td>
</tr>
</tbody>
</table>
resonant. A nonresonant case is also presented as a comparison.

It is noteworthy that the longwave mode is a linearly unstable in an inviscid environment but is suppressed by friction in a viscous environment. It vanishes without nonlinear energy transfer in a viscous setup, as the total energy sink through the radiation and frictional dissipation is greater than its energy gain from the basic state. Its structure is also modified when it nonlinearly interacts with the unstable shortwave mode.

One application of this study is to explain the observed quasi-zonal striations in the ocean interior discovered from satellite altimeter data (Maximenko et al. 2005), some of which extend westward from the eastern boundary. Wang et al. (2012) demonstrated that the nonlinear radiating instabilities of an eastern boundary current can produce zonal striations in the ocean interior with a magnitude that is comparable to the observed values. The interior velocity produced by the radiating instability is about 5 cm s$^{-1}$, which is comparable to 6.9 cm s$^{-1}$ found in Maximenko et al. (2005) for mid-latitudes (20°–40°). The main point is that even a very idealized eastern boundary current can produce nearly zonal interior flows roughly consistent with observations.

Several phenomena not considered here might also be important for eastern boundary current stability and energy radiation. Baroclinicity can become important and increases the nonlinearity of the system by adding an additional energy source. It is unclear what the characteristics of the radiating baroclinic instabilities are, and whether they obey similar mechanisms so one could extend previous studies of the baroclinic instability of a meridional boundary current by considering a continuous velocity profile with a vertical shear, and also by nonlinear simulations. Furthermore, it is unknown what the characteristics of the radiating instabilities are, and whether they obey similar mechanisms so one could extend previous studies of the baroclinic instability of an eastern boundary current in a baroclinic environment (Spall 2010) and by localizing instabilities.

Acknowledgments. We thank Joe Pedlosky and Hristina Hristova for helpful discussions. The comments from two anonymous reviewers have improved the manuscript greatly. J. Wang would like to thank the rest of his thesis committee Ken Brink and Markus Jochum. J. Wang was supported by the MIT/WHOI education office and the MIT Y-S Fellowship when this study was done, and by NASA Grant NNX12AD47G when this manuscript is prepared. M. Spall is supported by Grant OCE-0926656. G. Flierl is supported by Grant OCE-0752346.

APPENDIX

The Energy Budget for the $n$th Mode

The vorticity equation of the $n$th mode $G_n$, that is, Eq. (22), is

$$
\frac{\partial}{\partial t} \left( \frac{\partial^2 \phi_n}{\partial x^2} - n^2 \ell^2 \phi_n \right) + J_n + \beta \frac{\partial \phi_n}{\partial x} = \text{inl} \left[ \phi_n \frac{\partial^2 \psi}{\partial x^2} - \nabla \left( \frac{\partial^2 \phi_n}{\partial x^2} - n^2 \ell^2 \phi_n \right) \right]$$

$$+ A_H \left( \frac{\partial^4 \phi_n}{\partial x^4} - 2n^2 \ell^2 \frac{\partial^2 \phi_n}{\partial x^2} + n^4 \ell^4 \phi_n \right).$$

(A1)

The energy equation can be derived by multiplying $\phi_n^* \theta_*$ on both sides of $G_n$. Note that the energy equation in real space is $0.5(\phi_n^* G_n + \phi_n G_n^*)$, which is equivalent to $0.5(\phi_{-i} G_n + \phi_n G_{-i})$ where $\phi_{-i}$ is $\phi_n^*$ and $G_{-i} = G_n^*$. It is also the real part of $\phi_n^* G_n$ denoted as $\Re(\phi_n^* G_n)$.

Some terms of $\phi_n^* G_n$ are reorganized as (after dropping the subscript $n$ in $\phi_n$, and using the subscripts to denote partial derivatives):
\[
\phi^* (\phi_{xx} - n^2 \ell^2 \phi_x) = \phi^* (\phi_{xxx} - n^2 \ell^2 \phi_{xx}) = -(\phi^* \phi_{xx} + n^2 \ell^2 \phi^* \phi_{x}) + (\phi^* \phi_{xx}),
\]
\[
\phi^* \int \left( \phi_{xx} - \nu \phi_{xx} + n^2 \ell^2 \nu \phi \right) = \int \phi^* \phi_{xx} - \int (\nu \phi^* \phi_{xx} + \nu \phi \phi_{x} + \nu \nu (\phi^* \phi_x + n^2 \ell^2 \phi^* \phi)), \quad \text{and}
\]
\[
\phi^* A_H (\phi_{xxx} - 2n^2 \ell^2 \phi_{xx} + n^4 \ell^4 \phi) = A_H (\phi^* \phi_{xxx} - \phi_{xx}^* \phi_{xx} - 2n^2 \ell^2 \phi^* \phi_{x}) + A_H (\phi_{xx}^* \phi_{xx} + 2n^2 \ell^2 \phi^* \phi_{x} + n^4 \ell^4 \phi^* \phi).
\]

The energy equation for the nth mode in real space \(0.5(\phi_{-n} G_{n} + \phi_{n} G_{-n})\) can be written as:

\[
E_t = R \left[ \phi^* \phi_{xx} + \frac{\beta}{2} \| \phi \|^2 + n \nu \phi^* \phi_x - A_H (\phi^* \phi_{xxx}) 
\right]
\]
\[
- \phi_{xx}^* \phi_{xx} - 2n^2 \ell^2 \phi^* \phi_x 
\]
\[
- A_H (\| \phi_x \|^2 + 2n^2 \ell^2 \| \phi_x \|^2 + n^4 \ell^4 \| \phi \|^2) + \Re (\phi^* J_n),
\]

where \(E = 1/2(\| \phi_x \|^2 + n^2 \ell^2 \| \phi \|^2), \| \phi^2 \| = \phi^* \phi, \) and \(\Im(A)\) represents the imaginary part of a complex number \(A\).

Taking the time and domain integration over the eastern boundary current region

\[
\langle \cdot \rangle = \int_t \int_{L_x}^{0} 
\int_{-L_y}^{0} = L_y \int_{-L_y}^{0} 
\]

for a meridionally periodic domain, the energy equation then becomes

\[
0 = \int_t \int_{x=-L_x}^{L_x} \left[ \phi^* \phi_{xx} + \frac{\beta}{2} \| \phi \|^2 - A_H (\phi^* \phi_{xxx}) - \phi_{xx}^* \phi_{xx} - 2n^2 \ell^2 \phi^* \phi_x \right] + \Re (\phi^* J_n),
\]

where the boundary conditions: \(\phi = 0\) and \(\phi_x = 0\) at the side boundary \(x = 0\), and \(\nu = 0\) in the interior \((L_x = 20L_b\) in this case) are used. We use ‘flux’ to denote the contribution of fluxes by pressure, radiation, and friction, “rey” the effect of Reynolds stress, “fric” the energy dissipation by friction, and “non” the nonlinear energy transfer. After substituting the subscript \(n\) for \(\phi\), the energy budget equation becomes Eq. (23).

The energy equation can be further simplified by dropping the frictional fluxes, which are negligible in the interior comparing with the beta effect because of the high-order horizontal derivatives. The parameter measuring the importance of friction is

\[
E = O \left( \frac{A_H L_x^3}{\beta} \right),
\]

which is \(O(10^{-4})\) given \(\beta = 1.8 \times 10^{-11} \text{ (m s)}^{-1}, A_H = 100 \text{ (m^3 s^{-1})}, \) and the typical meridional scale of the perturbation field \(L = 700 \text{ km in the interior}. \) Here, \(E\) becomes even smaller if the interior perturbation zonal scale \(L = 2000 \text{ km is used}. \) As a result, the frictional fluxes in the interior are negligible, the energy equation then becomes

\[
0 = \int_t \int_{x=-L_x}^{L_x} \left[ \phi^* \phi_{xx} + \frac{\beta}{2} \| \phi \|^2 - \phi_{xx}^* \phi_{xx} - 2n^2 \ell^2 \phi^* \phi_x \right] + \Re (\phi^* J_n),
\]

The energy budget discussed in the text is calculated based on Eq. (A2), but the one based on Eq. (A3) is approximately the same.

REFERENCES


