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1 **Controls on the stability and composition of amphibole in the Earth's mantle**

2

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11 thermobarometry

12

13 **Abstract**

14 Presented here is a suite of new experiments aimed at quantifying the effects of pressure,
15 temperature, bulk composition and H₂O content on the stability and composition of amphibole in
16 the Earth's mantle. Experiments have been performed from 2-4 GPa and 950–1100 °C on fertile
17 and depleted mantle compositions. H₂O contents of most experiments are 0.65 wt.%. In the
18 fertile mantle composition, pargasitic amphibole is stable up to ~3.8 GPa at 1000 °C,
19 approximately 0.5 GPa higher than any previous study. The upper stability limit of amphibole in
20 depleted mantle is 0.7 GPa and 40 °C lower than in fertile mantle. The addition of 3 wt.% H₂O to
21 fertile mantle destabilizes amphibole by 0.5 GPa and 40 °C relative to the 0.65 wt.% H₂O
22 experiments. Compared to existing experiments on amphibole stability, these experiments

23 indicate that pargasitic amphibole may be stable in mantle lithosphere to almost 4 GPa (0.5 GPa
24 higher (15 km deeper) than previously thought). The extremely strong destabilizing effect of
25 H₂O suggests that deeper portions of the strongly fluid-fluxed mantle wedge may be amphibole-
26 free even at low temperatures near the slab-wedge interface. The molar alkali content of
27 amphibole is shown to be a linear function ($R^2 = 0.98$) of pressure and temperature and is
28 relatively insensitive to bulk compositional differences between fertile and depleted mantle. This
29 relationship is used to produce an empirical thermobarometer for pargasite-bearing spinel- and
30 garnet- lherzolites. Comparison to existing experimental data shows that this thermobarometer
31 has predictive ability over the pressure range of 1 to 4 GPa. Comparisons with pressure-
32 temperature estimates of garnet+amphibole peridotites further corroborate the applicability of
33 this thermobarometer for natural samples. Pressure estimates are presented for four examples of
34 metasomatized spinel peridotites otherwise lacking pressure information, and future avenues for
35 refinement of the thermobarometer are discussed.

36

37 **Introduction**

38 The Earth's uppermost mantle is pervasively metasomatized, as illustrated by dozens of
39 exhumed peridotite massifs and thousands of mantle xenoliths. Metasomatic agents come in
40 many flavors, from H₂O-rich fluids to silicate melts, but almost all metasomatized mantle
41 samples are enriched in H₂O and fluid-mobile elements. The most abundant of these elements
42 are the alkalis Na and K, so the most common new minerals formed during mantle
43 metasomatism are alkali-bearing hydrous silicates (amphiboles and micas). Because these
44 minerals both store H₂O when they form, and release it when they break down, they exert a

45 strong control on the geodynamic and geochemical behavior of the mantle lithosphere.
46 Specifically, the pressure-temperature (P-T) stability fields of amphiboles and micas play critical
47 roles in: (i) the extent to which H₂O is released into the overlying mantle wedge at subduction
48 zones or transported to the deep mantle; (ii) the water storage capacity of the uppermost mantle;
49 and (iii) the generation of alkaline, hydrous magmas in intraplate settings due to hydrous melting
50 triggered by the high-P-T breakdown of these phases during delamination or thermal
51 perturbation. Furthermore, the strong P-T-dependence of the composition of amphibole (e.g.
52 Niida and Green 1999) suggests that a calibration of these effects may yield a viable
53 thermobarometer for amphibole-bearing mantle samples.

54 In this study we present new experiments that rigorously assess the effects of pressure,
55 temperature, bulk composition and H₂O content on the stability and composition of amphibole in
56 the uppermost mantle. The need for such a study has become apparent over the past several
57 decades, during which many investigations of amphibole stability have been published,
58 producing a wide range of pressure- and temperature stability limits (Figure 1). Some of the
59 apparent problems have been identified but not quantified, and uncalibrated interlaboratory
60 differences make a collated analysis difficult. The two most significant problems, and those we
61 focus on in this paper, are the effects of bulk composition and H₂O content on amphibole
62 stability and composition. By isolating each variable over a range of pressures and temperatures,
63 we can identify the effects of all four variables (P, T, X, H₂O) and provide quantitative
64 information on some of the factors that influence the behavior of this notoriously complex
65 mineral group in mantle lithologies.

66

67 **Previous work**

68 Early work showed that the stability field of amphibole was not only pressure-invariant at high
69 temperatures but also nearly temperature-invariant at high pressures (Gilbert 1969; Kushiro
70 1970; Green 1973; Mysen and Boettcher 1975). Both the pressure and temperature stability
71 limits were shown to be dependent on several important factors:

72 *Bulk composition.* Mysen and Boettcher (1975) clearly showed that amphibole stability
73 was dependent on bulk composition. Subsequent work by Wallace and Green (1991) showed that
74 more alkali-rich, fertile mantle compositions stabilize amphibole to higher P and T than more
75 depleted compositions. Since then, it has been repeatedly shown that compositions rich in alkalis
76 tend to stabilize amphibole to the highest pressures and temperatures (e.g. Niida and Green 1999;
77 Fumagalli et al. 2009), but this effect has yet to be quantified.

78 *H₂O content.* It has long been recognized that the upper stability limit of hydrous phases
79 depends on the H₂O content of the system. Yoder and Kushiro (1969) suggested, based on a
80 study of phlogopite, that hydrous phases are most stable when the bulk H₂O content of the
81 system is equal to that of the hydrous phase in question (4 wt.% for phlogopite, 2 wt.% for
82 amphibole). However, the experimental literature suggests that even small amounts of excess
83 H₂O can leach amphibole-stabilizing alkalis into a fluid phase and thus induce amphibole
84 breakdown at lower-P-T conditions. An amphibole-bearing peridotite containing no other
85 hydrous phases will have a bulk H₂O content of only 0.2-0.3 wt.%, and Green et al. (2014) have
86 recently shown that bulk H₂O contents as low as 1 wt.% are sufficient to decrease amphibole
87 stability. This helps to explain the relatively small amphibole stability fields found by Kushiro
88 (1970; >5.8 wt.% H₂O), Green (1973; 6 wt.% H₂O in water-saturated experiments), Millhollen et

89 al. (1974; 5.7 wt.% H₂O), Mysen and Boettcher (1975; >8 wt.% H₂O), and Grove et al. (2006;
90 14.5 wt.% H₂O). This effect is the single largest control on amphibole stability in experiments
91 (Figure 1), yet it remains very poorly constrained.

92 *Oxygen fugacity.* Early studies on simple Fe³⁺-rich amphibole systems showed that
93 relatively huge fO_2 variations (5 log units) could change the stability field of Fe³⁺-rich
94 amphiboles by ~100 °C (Ernst 1962; Gilbert 1966). The implications for mantle samples,
95 however, are only minor because (1) amphiboles in equilibrium with peridotite are very poor in
96 iron (Mg# ~ 90), and (2) the total range in oxygen fugacity for all mantle samples within the
97 relevant pressure range is ~4 log units, from reduced Archean cratons to oxidized subduction
98 zones (Frost and McCammon 2008). We therefore do not consider the effects of fO_2 in this
99 study.

100 There has been significant recent interest in the geodynamic roles played by amphibole
101 and other hydrous minerals (e.g. Green et al. 2010; van Keken et al. 2011; Wada et al. 2012;
102 Walowski et al. 2015). Of particular interest are water storage and release during subduction, and
103 the depth to which amphibole is stable in the oceanic and continental lithosphere. These
104 problems require a rigorous and quantitative understanding of the controls on the stability of
105 amphibole in the upper mantle, which we aim to provide here. Having done this, we will
106 evaluate the potential for the compositional variability of mantle amphiboles to be harnessed as a
107 thermobarometer.

108

109 **Experimental Methods**

110

111 *Starting compositions*

112 To accurately quantify the behavior of amphibole in the Earth's uppermost mantle, it is
113 important to use realistic mantle compositions that span a typical range of mantle compositions.
114 To do this, we inspected the peridotite xenolith compositions in the GEOROC database
115 (<http://georoc.mpch-mainz.gwdg.de/georoc/>) and chose a fertile and a depleted composition
116 (Table 1). The fertile composition (FNa) was determined by taking the MgO content of the
117 primitive upper mantle estimate of Hart and Zindler (1986) – 37.8 wt.% – and adjusting the other
118 elements to produce a composition representative of peridotite xenoliths with that MgO content
119 (Figure 2). These adjustments were all small and were conducted to make the composition as
120 representative as possible of common xenolith compositions. Compositions were checked
121 against real xenolith compositions in terms of all elements incorporated into the starting mix to
122 ensure a realistic composition. The depleted composition (DNa) was also chosen to be
123 representative of mantle xenoliths, and corresponds approximately to the residue after 20% melt
124 extraction from a fertile peridotite as calculated by Walter (1999).

125 Both fertile and depleted compositions were modified by adding 1 wt.% of a metasomatic
126 component made of 65% H₂O, 30% Na₂O and 5% K₂O (values in wt.%). This metasomatic
127 component was based on compositional estimates for slab-derived fluids/melts at subduction
128 zones, which are H₂O- and alkali-rich (Stolper and Newman 1994; Keppler 1996; Scambelluri
129 and Philippot 2001; Grove et al. 2002). The slab-derived component may also contain other
130 components (e.g. Al, Si, Ca), particularly if it is a melt, but the major elements in peridotite will
131 be negligibly affected by 1 wt.% addition of the metasomatic component. A third bulk

132 composition (FNa3) was used to investigate the effect of H₂O on amphibole stability; FNa3
133 consists of composition FNa with additional H₂O added to give 3 wt.% H₂O in the bulk.

134 The resulting starting compositions (Table 1) are representative of metasomatized mantle
135 samples and sample both fertile and depleted compositions. Although some prior experimental
136 studies have investigated depleted peridotites (Mengel and Green 1989; Wallace and Green
137 1991; Fumagalli et al. 2009), the experiments were conducted on depleted compositions that had
138 30-60% olivine subtracted from their bulk. Olivine subtraction introduces additional uncertainty
139 due to assumptions made about the composition of olivine to remove, and results in bulk
140 compositions that have some depleted signatures (e.g. high Cr/Al and low Fe/Mg ratios) but
141 some enriched ones (e.g. high Na/Si and Al/Mg ratios). To avoid the uncertainty associated with
142 olivine subtraction, we have not performed this step on our depleted composition. This study is
143 therefore the first to present detailed information on amphibole stability and composition in a
144 depleted peridotite that has not been modified outside the range of natural depleted peridotite
145 compositions (Figure 2).

146

147 *Starting materials*

148 All starting materials were synthesized from high-purity oxides (SiO₂, TiO₂, Al₂O₃, Cr₂O₃,
149 Fe₂O₃, MnO, MgO, NiO), carbonates (Na₂CO₃, K₂CO₃), metallic Fe, and synthetic silicates
150 (CaSiO₃) and hydroxides (Mg(OH)₂). All materials except for the Fe- and H₂O sources were
151 dried in furnaces, then weighed out and ground together under ethanol for 3 hours using an agate
152 pestle and mortar. These were decarbonated at 850 °C for 1-2 days before FeO was added by
153 mixing Fe₂O₃ and metallic Fe together in appropriate molar proportions. The Fe-bearing mix was

154 then ground under ethanol for another hour, dried and conditioned for 3 days at 1050 °C and an
155 fO_2 of 0.5 log units above the quartz-fayalite-magnetite buffer. H₂O was added to mixes FNa and
156 DNa using synthetic brucite (Mg(OH)₂) to precisely control the H₂O content at 0.65 wt.%. A
157 final grinding under ethanol for one hour evenly distributed this H₂O source in the starting
158 material. The additional H₂O for mix FNa₃ was added by microsyringe to each capsule
159 immediately prior to welding. Weighing of the capsule before and after welding confirmed that
160 H₂O loss during the welding process was negligible (<0.00001 g).

161

162 *Experimental procedures*

163 All experiments were performed using Au capsules. These were fabricated from Au tubing with
164 an inner diameter of 0.13” and a wall thickness of 0.01”. Tubing segments $\frac{3}{8}$ ” long were cut
165 from tubing stock and triple-crimped on one end. This end was then welded using a PUK U3
166 TIG welder and flattened to produce a uniformly thick and flat capsule top. The capsule was then
167 filled with ~0.06 g of starting material, which was packed down into the capsule. The unfilled
168 portion of the capsule was cleaned to ensure a good weld. For FNa₃ runs, additional H₂O was
169 added at this point to yield 3 wt.% in the bulk, using a microsyringe and a scale to ensure precise
170 control over the amount of H₂O added. The bottom of the capsule was then triple-crimped and
171 welded shut. The sealed capsule was flattened in a die to produce a final capsule length of 0.15”,
172 then placed in an alumina ring and centered in a graphite furnace using MgO spacers.

173 Experiments were conducted at the MIT experimental petrology laboratory with a 0.5”
174 solid medium piston-cylinder apparatus (Boyd & England, 1960) using the hot-piston-in
175 technique (Johannes et al., 1971). The pressure medium was BaCO₃, for which no pressure

176 correction was deemed necessary based on pressure calibration of this apparatus using the
177 reaction: anorthite + gehlenite + corundum = Ca-Tschermak's pyroxene, and the spinel to garnet
178 transition in the CMAS peridotite analog system (1500 °C, 2.5 GPa) (Longhi 2005). Pressures are
179 considered accurate to ±50 MPa. Temperatures were measured and controlled with type-D
180 (W₉₇Re₃-W₇₅Re₂₅) thermocouples linked to a Eurotherm 831 controller. No EMF correction was
181 applied for pressure effects on the thermocouple. Reported temperatures are considered accurate
182 to ±10 °C. The thermal structure of the run assembly in this apparatus was previously determined
183 to have a 20 °C temperature gradient between the thermocouple junction and the hotspot of the
184 furnace (for which a correction was applied to the run temperature), and a temperature gradient
185 across the capsule of ~10 °C (Medard et al., 2008). Each experiment was initiated by manually
186 raising the pressure to 1 GPa at room temperature. The temperature was then increased to 865 °C
187 at a rate of 100 °C/min. As the BaCO₃ cell compressed, the pressure was manually raised to keep
188 it at 1 GPa. The run was held at 865 °C for 6 minutes, during which time the pressure was raised
189 to the final run pressure. The temperature was then ramped to the final run temperature at a rate
190 of 50 °C/min, with further manual pumping of the pressure system to maintain the pressure at the
191 final run pressure. The pressure typically stabilized after ~30 minutes. Run durations at these
192 conditions were 4-7 days (Table 2). The experiments were terminated by turning off the power to
193 the furnace. The capsule was extracted and a small hole was drilled to check for excess H₂O,
194 then the capsule was cut into 4 thin slices using a diamond wire saw. These slices were vacuum-
195 impregnated with epoxy, mounted in an epoxy disc, polished with alumina, and carbon-coated to
196 prepare them for electron microprobe analysis.

197

198 *Analytical methods*

199 All samples were analyzed using the JEOL JXA-8200 Superprobe at MIT. Backscattered
200 electron imaging, energy-dispersive spectrometry and chemical mapping were used to inspect the
201 mineralogy and appearance of the run. Phase compositions were obtained by wavelength-
202 dispersive spectrometry, using a focused electron beam (spot size $\sim 1 \mu\text{m}$) with a 15 kV
203 accelerating voltage and a 10 nA beam current. Count times for Na were varied to assess the
204 significance of undercounting due to Na migration. A count time of 10 s was ultimately used for
205 Na, which provided sufficiently high counts for consistent, suitably precise analyses without any
206 measurable Na loss. The rest of the elements were all counted for 40 s (no K migration problems
207 were observed). Data were reduced using the CITZAF correction package (Armstrong, 1995),
208 then manually inspected for data outliers produced by the beam partially hitting minerals other
209 than the one intended for analysis. This was a particularly acute problem for the highly poikilitic
210 garnets, which is reflected by the small number of points analyzed for garnet in some of the
211 experiments (Supplementary Table 4).

212

213 **Results**

214 Experimental conditions and phase assemblages are given in Table 2. Chemical analyses of
215 experimentally produced amphiboles are given in Table 3. Chemical analyses of all other phases
216 are given in the supplementary information.

217

218 *Equilibrium, quench alteration, and oxygen fugacity in experiments*

219 Experimental durations exceeding a few days were enough to achieve equilibrium. This was
220 confirmed by several observations: (i) internal chemical homogeneity of solid phases; (ii)
221 constancy of solid phase compositions and equilibrium crystal textures across the entire charge;
222 and (iii) systematic and predictable variability of solid phase compositions between experiments
223 at different pressures and temperatures. In the 1100 °C experiments, quench crystals were
224 common but easily identified by their acicular habit (Figure 3d) and melt-like chemistry (e.g.
225 quench amphiboles with Mg# = 77, cf. equilibrium amphiboles with Mg# ~ 90). Only
226 equilibrium phase compositions are reported here.

227 Four of the experiments had spinel, orthopyroxene and olivine present, making oxygen
228 fugacity estimates possible using the oxybarometer of Ballhaus et al. (1991). Two of these
229 experiments (D288 and D293) gave an oxygen fugacity of QFM + 0.4, while the other two
230 (D292 and D304) indicated more reducing conditions of QFM –1 to –2. This range is broadly
231 similar to that estimated by Grove et al. (2006) for very similar experiments (QFM –1 to +2). We
232 therefore infer that the oxygen fugacity of this suite of experiments was within this range. This
233 indicates a bulk $Fe^{3+}/\Sigma Fe$ of ≤ 0.15 (Kress and Carmichael 1991). The presence of a small
234 fraction of Fe^{3+} is important for understanding some of the substitution reactions controlling the
235 phase chemistry in these experiments – see below.

236

237 *Phase relations*

238 Figure 3 shows the typical appearance of experimental charges at different conditions. Figure 4
239 shows the pressure-temperature stability fields for the minerals observed in the three
240 experimental series. Melt was present in all experiments, as determined by the presence of glass

241 wisps connecting minerals in all runs, and quench crystals in 1100 °C runs. Melt pockets were
242 plucked out by the polishing process but were identified in backscattered electron images as open
243 spaces with few crystals; these areas were routinely observed in the hottest portion of the charge,
244 especially in 1050 and 1100 °C runs.

245 A free fluid phase was observed when capsules were drilled open after each experiment,
246 but this does not necessarily mean that a free fluid phase was present at the temperature and
247 pressure of the experiment. In compositions FNa and DNa (cf. FNa3), and at higher
248 temperatures, it is likely that sufficient melt was present to incorporate all of the H₂O. For a bulk
249 H₂O content of 0.65 wt.% and a high-pressure melt H₂O solubility of 20-30% (Hodges 1972;
250 Pirard and Hermann 2015), only 2-3% melt is required to incorporate all of the H₂O even if no
251 amphibole is present. The presence of amphibole (2 wt.% H₂O) reduces the melt fraction
252 required to incorporate all H₂O without the need for a free fluid phase. It is therefore likely that
253 in the FNa and DNa experiments the fluid observed during the drilling of capsules after each run
254 was dissolved in the melt at run conditions but exsolved upon quenching, producing
255 interconnected fluid and shards or wisps of glassed melt. In FNa3 experiments, there was almost
256 certainly a free fluid phase at run conditions due to the much higher bulk H₂O content.

257 Olivine is observed in all experiments. Crystals are hypidiomorphic to xenomorphic,
258 equant, and range in size from ~5 μm in the 950 °C experiment up to 30 μm in 1100 °C
259 experiments. In runs of all pressures, some crystals contain tiny (<< 1 μm) Al-Cr-rich inclusions,
260 which are remnants of spinels trapped by olivine growth during the pressure-temperature
261 ramping at the start of each experiment.

262 Orthopyroxene is observed in all experiments. Crystals are hypidiomorphic to
263 idiomorphic, stubby to elongate, and range from 10 μm long in the 950 $^{\circ}\text{C}$ experiment up to 70
264 μm long in 1100 $^{\circ}\text{C}$ experiments. They commonly contain small ($\sim 1\text{-}3\ \mu\text{m}$) inclusions of
265 clinopyroxene.

266 Clinopyroxene is observed in all experiments. However, in the highest-temperature
267 (1100 $^{\circ}\text{C}$) experiments, clinopyroxene is only observed in the coolest portions of the charge,
268 indicating that these experiments were very close to the upper temperature limit for
269 clinopyroxene. Crystals are hypidiomorphic to idiomorphic, equant in shape, and range from ~ 5
270 μm in the 950 $^{\circ}\text{C}$ experiment up to $\sim 15\ \mu\text{m}$ in the 1100 $^{\circ}\text{C}$ experiments.

271 Garnet is observed in all of the highest-pressure experiments. For composition FNa,
272 garnet is observed at all pressures greater than 2 GPa. However, in composition DNa, garnet is
273 only observed at pressures greater than 2.5 GPa. This is due to the much higher Cr/Al ratio of
274 composition DNa (Table 1), which stabilizes spinel over garnet to higher pressures (O'Neill
275 1981; Carroll Webb and Wood 1986; Klemme 2004; Ziberna et al. 2013). Crystals are large (up
276 to 100s of μm), hypidiomorphic to perfectly idiomorphic, and extremely poikilitic, containing
277 inclusions of all other mineral phases. The strong favoring of growth over nucleation and the
278 very low availability of inclusion-free areas for analysis are likely responsible for the greater
279 compositional scatter observed in garnet analyses compared to other solid phases
280 (Supplementary Table 4).

281 Spinel is observed in very few experiments. This is because at low pressures amphibole
282 provides a sufficiently large sink for Al that another aluminous phase is not required (at high
283 pressure, the lower abundance and lower Al content of amphibole is accompanied by garnet

284 stability). Amphibole and spinel only coexist in one experiment (Figure 4). Spinel is therefore
285 observed only in the highest-temperature DNa and FNa3 experiments, in which amphibole is
286 significantly destabilized. This is consistent with the observations of Fumagalli et al. (2009), who
287 performed experiments up to only 1000 °C and so never observed any spinel. In experiments that
288 do have spinel, crystals are rare (<1% of the charge), small (1-3 μm), hypidiomorphic and
289 equant.

290 Amphibole is observed in most experiments. As shown by Figure 4, its upper pressure
291 stability limit approaches temperature invariance, while its upper temperature stability limit
292 approaches pressure invariance. Its upper pressure limit is largely controlled by the progressive
293 up-pressure stabilization of garnet and the jadeitic clinopyroxene component, which sequester
294 many of the elements necessary for amphibole stability (notably Al, Cr, and Na). Because
295 amphibole is a hydrous mineral, its upper temperature limit is controlled by the dehydration
296 reactions that release H₂O into the melt phase. Amphibole crystals are stubby to elongate and
297 hypidiomorphic, ranging from ~5 μm long in the 950 °C experiment, up to 20 μm long in some
298 1050 °C experiments.

299

300 *Phase compositions*

301 All phases show systematic compositional changes as a function of pressure and temperature, as
302 well as compositional differences between the two bulk compositions. In general, phases in the
303 DNa experiments have higher Mg# and Cr₂O₃, and lower TiO₂, than in the FNa experiments.
304 Major elements that are much less abundant in the DNa bulk composition (Ca, Al) are only
305 mildly depleted in the minerals; the bulk depletion of these elements is instead recorded by the

306 lower abundance of Ca- and Al-bearing minerals (clinopyroxene and garnet). Most of the same
307 pressure-temperature compositional trends are observed in both the FNa and DNa experiments.
308 The few that are not resolved in the DNa experiments (MnO and Mg# in opx, cpx, and
309 amphibole) are largely due to the small number of experiments on the DNa composition. The
310 chemical systematics described below are based on the FNa experiments, for which the coverage
311 of pressure-temperature space is much more thorough. Some of the most significant pressure
312 effects are shown in Figure 5 for the 1000 °C FNa runs.

313 Olivine and orthopyroxene compositions become progressively Mn- and Fe-poor, and
314 Mg-rich, with increasing pressure and temperature. Orthopyroxene additionally becomes
315 progressively Ca-, Al- and Cr-poor, and Si-, Na- and (Mg+Fe)-rich with increasing pressure.
316 This is due to two main substitution reactions: (1) the high-pressure destabilization of the
317 Tschermak-type component and other associated trivalent components: $2 (\text{Al,Cr,Fe})^{3+} \Leftrightarrow$
318 $(\text{Mg,Fe})^{2+} + \text{Si}^{4+}$; and (2) The slightly decreased solubility of Ca in the orthopyroxene structure
319 at high pressure: $\text{Ca}^{2+} \Leftrightarrow (\text{Mg,Fe})^{2+}$. The small Na increase in orthopyroxene with increasing
320 pressure is due to the stabilization of a minor Na-bearing component at high pressure. The effects
321 of temperature on orthopyroxene composition are largely the inverse of the pressure effects:
322 higher temperatures stabilize the Tschermak-type and other trivalent components, and increase
323 the solubility of Ca in the orthopyroxene structure (the latter representing the up-temperature
324 closing of the orthopyroxene–clinopyroxene two-phase region). As a result, higher temperatures
325 make orthopyroxene more Ca-, Al-, and Cr-rich, and more (Mg+Fe)-poor. The effect of
326 temperature on Si is not well resolved because the expected decrease in Si concentration is of the
327 same order as the analytical uncertainty.

328 Clinopyroxene shows more complex behavior than orthopyroxene, with three important
329 substitution reactions producing several effects. (1) The dominant substitution reaction is the
330 stabilization of a jadeite (or other Na-bearing) component with increasing pressure: $\text{Ca}^{2+} +$
331 $(\text{Mg,Fe})^{2+} \Leftrightarrow \text{Na}^+ + (\text{Al,Cr,Fe})^{3+}$. With increasing pressure, this produces a decrease in Ca and
332 Mg, and an increase in Na and Cr (Figure 5). (2) The high-pressure destabilization of the
333 Tschermak-type component (as in orthopyroxene) offsets the up-pressure increase in Al from the
334 Na-bearing component, and complicates the behavior of Fe, which, due to its multiple valence
335 states, is involved on both sides of both reactions. These two competing substitutions explain
336 why Al only increases slightly with pressure while Na increases a lot, and why the Mg# of
337 clinopyroxene is not strongly correlated with pressure (Figure 5). As in orthopyroxene, the
338 Tschermak-type component is stabilized at higher temperatures, producing higher Al and lower
339 (Mg+Fe) concentrations. (3) The third substitution reaction is the substitution $\text{Ca}^{2+} \Leftrightarrow (\text{Mg,Fe})^{2+}$,
340 which produces an up-temperature decrease in clinopyroxene Ca content and offsets the
341 temperature effects of the substitution of Ca for Mg and Fe related to the enhanced Ca-
342 Tschermak solubility in clinopyroxene.

343 Amphibole shows highly systematic compositional variability as a function of pressure
344 (Figure 5) and temperature. Amphiboles are broadly pargasitic ($\text{NaCa}_2(\text{Mg}_4\text{Al})\text{Si}_6\text{Al}_2\text{O}_{22}(\text{OH})_2$),
345 with some K in the A site and Fe in the M1, M2, and M3 sites. With increasing pressure,
346 amphiboles show a significant substitution towards richterite ($\text{Ca}^{2+} + (\text{Al,Fe,Cr})^{3+} + 2 \text{Al}^{3+} \Leftrightarrow$
347 $\text{Na}^+ + \text{Mg}^{2+} + 2 \text{Si}^{4+}$), with minor substitutions towards magnesiokatophorite ($\text{Ca}^{2+} + \text{Al}^{3+} \Leftrightarrow \text{Na}^+$
348 $+ \text{Si}^{4+}$) and possibly edenite ($(\text{Al,Fe,Cr})^{3+} + \text{Al}^{3+} \Leftrightarrow \text{Mg}^{2+} + \text{Si}^{4+}$). Mg# increases with
349 temperature and pressure, and high temperatures lead to lower Si and Ca, and higher Na. The up-

350 temperature increase in Na appears to be due to a coupled substitution with both Al and Fe³⁺. K
351 contents increase substantially with both pressure and temperature, akin to the behavior of Na.
352 Notably, the Na₂O and K₂O contents of amphiboles in DNa, FNa and FNa3 experiments are
353 almost indistinguishable, despite varying alkali and H₂O contents in the bulk (see discussion).

354 Garnet compositions become increasingly pyrope-rich at higher temperatures and
355 grossular-poor at higher pressures. Slightly higher Fe and lower Al contents at high pressures
356 suggest a minor Al³⁺ ↔ Fe³⁺ substitution. There is no resolvable change in Mg# with pressure,
357 as both MgO and FeO are mildly increased at high pressures. Garnets in the DNa experiments
358 have ~0.7 mol% lower Al₂O₃ and correspondingly higher Cr₂O₃ than the equivalent experiments
359 on the FNa composition.

360 Spinel is only present in one FNa3 and three DNa experiments, and shows no systematic
361 compositional change with pressure, temperature or bulk composition. The Mg# ranges from 62
362 to 68; the Cr# (100Cr/(Cr+Al)) ranges from 49 to 59.

363

364 **Discussion**

365

366 *Compositional controls on the stability of amphibole in the mantle*

367 A number of compositional variables control the stability and composition of amphibole in the
368 Earth's mantle. We will explore the most important of these and, where possible, quantify their
369 influence on amphibole stability.

370 Experiments by Green and coworkers showed that higher bulk abundances of elements
371 compatible in amphibole but less so in other mantle minerals (alkalis and perhaps Ti) will

372 increase amphibole's pressure-temperature stability (Wallace and Green 1991). Fertile mantle
373 compositions should therefore stabilize amphibole to higher pressures and temperatures than
374 depleted ones. This idea is reinforced by our experiments: the upper stability limit of amphibole
375 in the depleted composition is 0.7 GPa and 40 °C lower than in the fertile composition (Figure
376 6).

377 However, the idea of fertile and depleted peridotite is complex and depends on the nature
378 of the metasomatizing agent responsible for refertilization. Depletion and refertilization of
379 peridotites by the removal or addition of basaltic melts have somewhat predictable effects, but
380 the addition of slab-derived fluids and slab melts (of sediments and/or basalts) at subduction
381 zones produces a much wider range of compositions. The resulting effects on amphibole stability
382 will depend on the nature of the added material and can vary in many ways.

383 The recent work of Pirard and Hermann (2015) provides a compositionally extreme but
384 plausible example of this variability. They have shown that an ultradepleted peridotite/dunite that
385 is refertilized by a Ca-poor, alkali-rich sediment melt will stabilize amphibole to higher pressures
386 (> 4.5 GPa) than typical fertile peridotite (< 4 GPa, this study). This is probably due to the
387 absence of clinopyroxene in low-Ca bulk compositions: in Ca-richer peridotites, the up-pressure
388 stabilization of the jadeitic clinopyroxene component ($\text{NaAlSi}_2\text{O}_6$) limits the stability of Na-
389 bearing amphibole. In the absence of clinopyroxene, there is no major Na-bearing phase to
390 replace amphibole at higher pressures, so a very Na+K-rich, Ca-poor amphibole
391 (magnesiokatophorite) remains stable at high pressure, as long as there are sufficiently high
392 alkali contents in the bulk composition.

393 This is just one facet of a much larger problem. Work by Konzett and Ulmer (1999) and
394 Fumagalli et al. (2009) on the same bulk composition suggests that in systems sufficiently K-rich
395 to stabilize phlogopite there are actually two stability fields for amphibole (Figure 7): a more
396 sodic amphibole at lower pressure, and a more potassic amphibole at much higher pressure
397 resulting from the high-pressure breakdown of phlogopite. In systems too K-poor to stabilize
398 phlogopite ($K_2O < \sim 0.1$ wt.%), the K_2O released upon amphibole breakdown goes into the melt
399 at high temperatures, and clinopyroxene (0.1-0.4 wt.% K_2O at 6-8 GPa; Konzett and Ulmer
400 1999) and trace grain-boundary phases at lower temperatures.

401 A third issue is the stability of kaersutitic (Ti-rich) and Ti-pargasitic amphiboles, formed
402 by low-pressure, high-temperature metasomatism of peridotites by hydrous, alkaline, Ti-rich
403 basaltic melts. Furthermore, although Ti-rich (several percent TiO_2) amphiboles are generally
404 restricted to the spinel stability field, a Ti-component (0.5-0.7 wt.%) in amphibole remains up to
405 very high pressures, (e.g. amphiboles in PKP and MARID xenoliths from kimberlites; Konzett
406 and Ulmer 1999). The effect of Ti on amphibole stability has yet to be resolved, but the overall
407 effect seems to be a stabilizing one.

408 The effects of most of these variables on amphibole stability have yet to be quantified.
409 However, we can make several statements about the effects of bulk composition on the stability
410 of amphibole (see Figure 8 for a schematic summary), based on prior work and new insights
411 provided by this study:

412 (1) High alkali contents stabilize amphibole to higher pressures and temperatures
413 (Wallace and Green 1991; this study)

414 (2) High bulk Na/Ca ratios ($\text{Na}_2\text{O}/\text{CaO} > \sim 0.5$?) stabilize a highly sodic amphibole to
415 higher pressures at the expense of clinopyroxene (Pirard and Hermann 2015; this study)

416 (3) While the highly sodic amphiboles produced by high bulk Na/Ca ratios (i.e. low Ca)
417 are very pressure-stable, these high-Na/Ca amphiboles have lower temperature stability limits
418 than pargasitic amphiboles. In simple systems the practically Ca-free sodic amphibole
419 glaucophane is only stable up to 800 °C (Tropper et al. 2000); magnesiokatophorite ($\text{Na}_2\text{O}/\text{CaO}$
420 ~ 2) is stable in ultra-Ca-depleted peridotites to ~ 1000 °C (Pirard and Hermann 2015); and
421 pargasite ($\text{Na}_2\text{O}/\text{CaO} < 0.5$) is stable in fertile and depleted peridotites to at least 1100 °C (this
422 study). Higher bulk Ca contents stabilize amphibole to higher temperatures, although there does
423 not seem to be any benefit to increasing bulk Ca content beyond that necessary to stabilize
424 pargasite.

425 (4) Ti (and possibly Cr) stabilizes amphibole, but has unclear effects on its upper stability
426 limit (Wallace and Green 1991; this study). The range of natural variability in the abundance of
427 other elements (Al, Si, Fe and Mg) appears not to have significant effects on amphibole stability.

428 (5) Free H_2O destabilizes amphibole by leaching amphibole-stabilizing elements into a
429 fluid phase. See section “Effects of H_2O on the stability and composition of amphibole”, below,
430 for further details.

431 The effects of these variables have yet to be fully quantified, but the results of our
432 experimental study allow the development of a framework for interpreting natural samples,
433 and suggest avenues for future experimental investigations into amphibole stability.

434

435 *Controls on the composition of amphibole in the mantle: developing a thermobarometer for*
436 *metasomatized peridotites*

437 The experiments presented here show that amphibole composition follows an extremely
438 systematic and predictable dependence on pressure and temperature. The variations in the molar
439 abundance of Na₂O + K₂O in amphibole between pressures of 2 and 3.7 GPa and temperatures of
440 950 and 1100 °C can be expressed as a linear function of pressure and temperature (Figure 9):

441

$$442 \quad (\text{Na}_2\text{O} + \text{K}_2\text{O})_{\text{amph}} = -2.8616 + 1.0333 P + 0.0037 T \quad R^2 = 0.98 \quad (1)$$

443

444 where (Na₂O + K₂O)_{amph} is the mol% abundance of Na₂O + K₂O in amphibole on an anhydrous
445 basis, P is in GPa and T is in °C. The estimated error variance is 0.007. The mol% calculations
446 for amphibole do not include H₂O or halogens, allowing the conversion to be made directly from
447 electron microprobe data. The temperature sensitivity is relatively small, but the pressure effect
448 is large, so if temperature is known, pressure can be precisely predicted from amphibole
449 composition. This effectively tracks the up-pressure stabilization of the (K,Na)-richterite
450 component in pargasitic-edenitic amphiboles, and so holds potential as a geobarometer for
451 ultramafic rocks containing such amphiboles.

452 Figure 10 shows the predictive capability of this model for literature data on
453 experimentally produced amphiboles in peridotite bulk compositions. For relatively typical
454 fertile and depleted peridotites, this model provides very good fit even outside the calibration
455 range (< 2 GPa, < 950 °C); if the same regression is performed against all data, the fit is almost
456 identical, even outside the calibration range. Furthermore, the predictive capability of this model

457 is stronger than the scatter of the data suggests. This is because a similar amount of scatter is
458 observed even for individual studies that vary only pressure and temperature (e.g. Wallace and
459 Green 1991; Niida and Green 1999). These studies each scatter symmetrically around the model
460 prediction - so the model is accurately predicting the compositional variability - but the scatter is
461 not systematic with respect to pressure or temperature. This is unexpected, because if only
462 pressure and temperature are being varied then it is a thermodynamic requirement that amphibole
463 composition can only vary along two vectors in composition space (one each for pressure and
464 temperature), which is observed in our experiments and accounted for by our model. It is
465 therefore likely that much of the scatter observed in the literature data is due to experimental
466 irreproducibility or analytical uncertainty in those experiments, and that the model uncertainties
467 given in Figure 10 are maximum – and probably unrealistically high – errors. Whereas the mean
468 absolute error (MAE) for all of the experiments in Figure 10 is 0.24 GPa, the MAE for the
469 experiments from this study is only 0.06 GPa. The true uncertainty for high-quality microprobe
470 analyses of amphiboles is between these values.

471 The rest of the scatter can be ascribed to systematic differences between our experiments
472 and those plotted in Figure 10. These differences can help to constrain the range of conditions
473 over which our model may be applied to experimental and natural amphiboles. Three main
474 differences are observed:

475 1. H₂O. High abundances of free H₂O significantly decrease the stability field of
476 amphibole by leaching alkalis into a free fluid phase. It has been suggested that this leaching
477 decreases the alkali content of amphiboles at a given pressure and temperature (Green et al.
478 2014). However, this is not observed in our experiments (both the FNa and FNa3 experiments fit

479 our model equally well), and is not resolved by literature data for bulk H₂O contents lower than
480 ~6 wt.% (Figure 10c). The main effect of alkali leaching appears to be on the stability of
481 amphibole rather than its composition. However, for extremely high H₂O contents (the 14.5
482 wt.% H₂O experiments of Grove et al. (2006) and Green et al. (2014)), amphiboles are
483 systematically depleted in alkalis relative to our model prediction (Figure 10c), indicating a small
484 leaching effect on amphibole composition at very high bulk H₂O.

485 2. K₂O. This model is calibrated against experiments in which amphibole is the only
486 major alkali-bearing mineral and the only significant site for potassium. In more K-rich systems,
487 phlogopite is also stable, which sets up exchange reactions for alkalis (mostly K) between
488 amphibole and phlogopite, with their own dependences on pressure and temperature. These are
489 not accounted for by our model, so it is unsurprising that the phlogopite-bearing experiments of
490 Mengel and Green (1989) and Fumagalli et al. (2009) are each systematically offset from our
491 model predictions (Figure 10d: Mengel and Green = 0.43 wt.% K₂O; Fumagalli et al. = 0.54
492 wt.% K₂O). For phlogopite-free systems (all experiments below 0.1 wt.% bulk K₂O; Figure 10d),
493 the bulk K₂O content has no effect, but the model is not calibrated for phlogopite-bearing
494 peridotites. We therefore propose that this thermobarometer should be restricted to phlogopite-
495 free assemblages or those with only accessory phlogopite.

496 3. CaO. The extremely Ca-poor composition of Pirard and Hermann (2015; bulk CaO =
497 0.29 wt.%) stabilizes a much more alkali-rich amphibole (Figure 10a). In most experiments on
498 amphibole in peridotite, the M4 site is dominated by Ca because the bulk CaO/Na₂O ratio is high
499 (> 4 in all compositions listed in Table 1 except that of Pirard and Hermann). However, in the
500 experiments of Pirard and Hermann, this ratio is only 0.2, so the stable amphibole is Ca-poor,

501 (Na,K)-rich magnesiokatophorite. This indicates that our model is currently only applicable to
502 pargasite-edenite-type amphiboles, which require a sufficiently high bulk CaO, or
503 CaO/(Na₂O+K₂O) ratio, to be stable. Our experiments on the depleted mantle composition DNa
504 (CaO = 1.69 wt.%; CaO/(Na₂O+K₂O) = 3.7) satisfy this requirement, as evidenced by their good
505 fit to the model. Our model is therefore applicable to most amphibole-bearing depleted
506 peridotites, with the possible exception of ultradepleted compositions.

507 Overall, this model is applicable to amphibole-bearing peridotites from 1 – 3.7 GPa and
508 800 – 1150 °C. It is calibrated for phlogopite-free systems with olivine, orthopyroxene and
509 clinopyroxene, with or without an additional aluminous phase. Minor accessory phlogopite is
510 unlikely to have an adverse effect, but if phlogopite is a major phase then its effect on the
511 distribution of K may affect pressure estimates. The thermobarometer should also be applicable
512 to clinopyroxene-free systems – our highest-temperature experiments are almost clinopyroxene-
513 free – but sufficient Ca must be present to stabilize pargasite-edenite-type amphiboles over more
514 Ca-poor, alkali-rich compositions.

515

516 *Effects of H₂O on the stability and composition of amphibole*

517 Many experiments conducted on peridotites have had much larger quantities of H₂O than is
518 realistic for metasomatic processes likely to occur in much of the mantle. The water storage
519 capacity of the solid mantle is high at low temperatures (>1 wt.%, and up to 2 wt.% in fully
520 chloritized peridotite up to 840 °C; Grove et al. 2006; Till et al. 2012; Schmidt and Poli 2014),
521 and at higher temperatures the presence of H₂O triggers the production of melts, which at mantle
522 pressures have H₂O solubilities > 20 % (Hodges 1972; Pirard and Hermann 2015). The presence

523 of a significant free fluid phase during the production of amphibole in peridotites might be
524 important in the immediate vicinity of the slab-mantle interface but is unlikely for mantle
525 metasomatism further into the mantle wedge and in other parts of the Earth's mantle. This is an
526 important experimental consideration because the presence of free fluid destabilizes amphibole
527 by leaching alkalis from solid phases (Green et al. 2010, 2014).

528 One of the obstacles in quantifying the effects of free fluid on amphibole stability is that
529 many other variables have effects that remain unquantified. Different bulk compositions,
530 experimental procedures, analytical methods, and often uncontrolled or unmonitored oxygen
531 fugacities make it difficult to make significant quantitative statements beyond very broad
532 observations (e.g. Figure 1). However, the FNa and FNa3 studies presented here are identical
533 except for the abundance of H₂O, so they can be used to quantitatively determine the effects of
534 bulk H₂O content on amphibole stability.

535 Figure 11 shows the amphibole stability field from compositions FNa (0.65 wt.% H₂O),
536 FNa3 (3 wt.% H₂O), and the experiments of Grove et al. (2006; 14.5 wt.% H₂O). The Grove et
537 al. experiments and ours were both performed using identical experimental and analytical
538 procedures and so remove interlaboratory calibration issues. The Grove et al. experiments were
539 also performed on a fertile peridotite composition, but with lower alkali contents similar to those
540 in composition DNa (this study). It is possible that some of the decrease in amphibole stability in
541 these experiments relative to the FNa and FNa3 experiments is due to the lower alkali contents in
542 the Grove et al. bulk composition. The estimate provided here for the destabilizing effect of H₂O
543 at the highest H₂O contents (Figure 11b) should therefore be considered a maximum.

544 The effect of H₂O on the composition of amphibole is less clear. Green et al. (2010,
545 2014) noted that the presence of a fluid phase significantly decreases the Na₂O content of
546 clinopyroxene. A similar effect was proposed for amphibole, but the effect of up to 3 wt.% H₂O
547 is essentially unresolvable within analytical uncertainty (Figure 4 in Green et al. 2014). Indeed,
548 the amphiboles in our FNa₃ experiments (3 wt.% H₂O) show no resolvable alkali depletion
549 relative to the FNa experiments (0.65 wt.% H₂O). However very high H₂O contents (>6 wt.%
550 H₂O) do produce resolvably alkali-depleted amphiboles (Figure 10c).

551 Our experiments show that amphibole compositions from experiments with a few percent
552 H₂O are similar to those with <1 wt.% H₂O, but the amphibole stability fields are not. Future
553 experiments aimed at exploring the stability and composition of not only amphibole but also
554 other alkali-bearing phases should ideally use bulk H₂O contents of less than 1 wt.%. An
555 exception would be for areas of extremely high fluid flux, such as the immediate supra-slab zone
556 and regions of deep fluid penetration in deeply faulted oceanic lithosphere. These strongly
557 hydrated regions are repeatedly fluxed by H₂O and so have locally low residual alkali contents
558 that will not stabilize amphibole to such high pressures and temperatures during subsequent
559 subduction.

560

561 *Applications to natural samples: uses and limitations*

562 To test our new thermobarometer on natural samples, we have applied it to the peridotites of the
563 Ulten Zone in the Upper Austroalpine of NE Italy, to which many existing thermobarometers
564 have been applied (Nimis and Morten 2000). Because most high-quality geobarometers are
565 based on garnet equilibria (e.g. Nimis and Grütter 2010), it is desirable to have

566 amphibole+garnet-bearing peridotites for this comparison. Equilibrium garnet and pargasitic to
567 edenitic amphibole rarely occur together in peridotite samples. However, the peridotites of the
568 Ulten Zone contain both phases, and are particularly amenable to the rigorous use of our
569 thermobarometer because the amphibole is commonly present as a disseminated equilibrium
570 phase (i.e. not just in veins). These rocks were spinel peridotites from the mantle wedge that
571 became entrained by the subducting plate, metasomatized, and brought into the garnet stability
572 field at relatively low temperatures (Nimis and Morten 2000). Pressure-temperature estimates
573 typically range from 750-850 °C and 2.0-2.7 GPa (Rampone and Morten 2001). For the three
574 amphibole-bearing peridotites analyzed by Nimis and Morten (2000) in their own
575 thermobarometric analysis, our thermobarometer gives pressure estimates of 2.0-2.2 GPa at 850
576 °C, or 2.3-2.6 GPa at 750 °C. These estimates are in good agreement with those from existing
577 thermobarometers and corroborate the utility and applicability of our new amphibole-based
578 thermobarometer.

579 One of the potential uses for our new thermobarometer is the determination of
580 equilibrium pressure conditions for metasomatized spinel peridotites, for which high-quality
581 barometers remain elusive. The observation that the alkali-in-amphibole thermobarometer
582 presented here reproduces experimental pressures down to 1 GPa (Figure 10b) suggests that this
583 may be coupled with existing thermometers to provide reasonable pressure estimates for spinel
584 peridotites.

585 We have calculated pressures for four studies of metasomatized spinel peridotites from
586 various tectonic settings. We chose only examples of peridotites in which the amphibole is
587 pargasitic and present as a disseminated phase. Amphiboles in or closely associated with veins

588 may provide erroneous pressure predictions because they are produced by an inherently
589 disequilibrium process: the diffusive reaction between a hydrous melt and a host rock of very
590 different composition. Conversely, peridotites with disseminated amphibole of consistent
591 composition typically represent complete reaction between a hydrous fluid and a host peridotite,
592 for which experiments provide a close proxy, and on which our thermobarometer may be used.

593 (1) St Paul's Rocks, equatorial Atlantic. This enormous oceanic core complex, which has
594 exposed mylonitized spinel peridotites above sea level, was studied in detail by Melson et al
595 (1972). Sample 18-900 from their study is a relatively fertile peridotite with mildly enriched
596 alkali contents and a spinel harzburgite lithology, and contains abundant pargasite and only
597 accessory phlogopite (bulk $K_2O = 0.11$ wt.%). Two-pyroxene thermometry from associated
598 peridotites (Roden et al. 1984) yields pressure-corrected temperature estimates of 960-1030 °C.
599 These temperatures combined with the amphibole composition from sample 18-900 yield a
600 pressure estimate of 1.5-1.7 GPa, consistent with estimates based on postulated geotherms from
601 Melson et al. (1972).

602 (2) Zabargad Island, Red Sea. Another peridotite-rich island exposed in an (incipient)
603 ocean, Zabargad hosts a wide range of mantle lithologies. Its fertile lherzolite units are
604 pervasively metasomatized and contain ubiquitous disseminated pargasite of quite consistent
605 composition (Agrinier et al. 1993). This metasomatic event is thought to predate the alteration
606 events associated with the exhumation of Zabargad. Based on the abundance of heavy Li
607 isotopes (Brooker et al. 2004), and Os-isotope melt extraction ages (Snow and Schmidt 1999),
608 the metasomatic event that produced these pargasites is thought to be related to subduction
609 processes during the Pan-African orogeny. Using an estimated equilibration temperature of 1000

610 °C from pyroxene thermometry (Piccardo et al. 1988), we calculate an equilibration pressure of
611 2.1 +/- 0.1 GPa based on the composition of amphiboles from 6 peridotite samples.

612 (3) Victoria Land, Antarctica. Spinel lherzolite xenoliths from Baker Rocks contain
613 disseminated and vein pargasites that have recently been studied by Bonadiman et al. (2014).
614 These lherzolites are not very Na-enriched or deeply sourced ($\text{Na}_{\text{cpx}} = 0.030\text{-}0.066$ apfu),
615 allowing us to use the simple Brey and Köhler (1990) two-pyroxene thermometer, which
616 provides accurate temperature estimates for these general conditions (Nimis and Grütter 2010).
617 Focusing on the disseminated pargasite, we iterated temperatures obtained from the Brey and
618 Köhler thermometer – which requires a pressure term for its prediction – with pressure estimates
619 using our alkali-in-amphibole thermobarometer to obtain corrected temperature estimates of
620 990–1050 °C and pressure estimates of 2.2–2.6 GPa for six spinel lherzolites. Although this
621 pressure may appear too high for the spinel stability field, these lherzolites have very high bulk
622 $\text{Cr}_2\text{O}_3/\text{Al}_2\text{O}_3$ ratios, which increase the pressure of the spinel-garnet transition (O'Neill 1981;
623 Carroll Webb and Wood 1986; Klemme 2004; Zibera et al. 2013). The $\text{Cr}_2\text{O}_3/\text{Al}_2\text{O}_3$ ratio of
624 composition FNa (this study, spinel-garnet transition between 2 and 2.5 GPa) is 0.10. The ratio
625 for composition DNa (this study, spinel-garnet transition greater than 2.5 GPa) is 0.23. The ratios
626 for the Bonadiman (2014) samples with the highest pressures predicted from our
627 thermobarometer are > 0.4 , indicating a very high pressure ($>>2.5$ GPa) for the spinel-garnet
628 transition. Our pressure predictions are therefore consistent with these samples having last
629 equilibrated in the spinel stability field.

630 (4) Avacha volcano, Kamchatka arc. Mantle xenoliths reported from this volcano sample
631 the sub-arc mantle and are predominantly metasomatized harzburgites (Arai et al. 2003). A

632 precise temperature estimate for these harzburgites is not given by the authors, but is reported as
633 900 – 1000 °C as estimated from two-pyroxene thermometry. For a temperature of 900 °C, the
634 composition of amphibole in the metasomatized harzburgite yields a pressure of 1.1 GPa, but for
635 a temperature of 1000 °C, this estimate goes up to 1.5 GPa. This highlights the value of
636 combining the alkali-in-amphibole thermobarometer with a precise geothermometer to achieve
637 the most reliable pressure estimates.

638 Studies on all four of the localities discussed above have lamented the lack of an
639 appropriate barometer for spinel peridotites. The pressures we have estimated here are consistent
640 with all available information (spinel stability field, bulk composition, estimated geotherms), and
641 as we improve our understanding of the factors controlling amphibole stability, these estimates
642 will become increasingly precise. For the assemblage olivine + orthopyroxene + pargasitic
643 amphibole ± clinopyroxene ± garnet/spinel, this thermobarometer already provides useful
644 pressure estimates from 1 to 4 GPa. Of course, the mantle is a complex place, and the unusual
645 compositional features observed in some mantle samples need to be incorporated into this model
646 to extend its use to unusual compositions. Of particular interest moving forward are the effects of
647 TiO₂, oxygen fugacity, halogens, bulk Na/K ratio, and amphibole-phlogopite equilibria.

648

649 *Implications for water budgets in subduction zones and continental lithosphere*

650 A comprehensive understanding of hydrous phase stability is necessary in the study of
651 subduction zone and lithospheric dynamics, providing constraints on fluid transport and melt
652 generation, the water storage capacity and water content of the upper mantle, and intraplate
653 melting. In addition to providing a new tool for constraining the pressure-temperature conditions

654 of metasomatized peridotites, our experimental results also extend the amphibole stability field
655 in fertile peridotite to ~0.5 GPa higher pressure than previous studies (Figure 12), and also
656 indicate a large stability field for amphibole in truly depleted peridotites.

657 The mantle wedge immediately above the subducting slab is continuously fluxed by slab-
658 derived fluids, and is therefore likely to be strongly alkali-leached. It is highly unlikely that
659 amphibole is stable in this region to pressures above ~2 GPa (Grove et al. 2006). Parts of the
660 mantle deep within the subducting slab itself may retain amphibole up to the stability limit in
661 depleted peridotite (3-3.2 GPa).

662 Finally, it is worth noting that these new constraints provide updated inputs into models
663 of water storage and release in thermally perturbed lithosphere. The stability of amphibole to
664 higher pressure increases the overall water storage capacity of the lithosphere for a range of
665 cratonic geotherms (Figs. 7 and 12). This may lead to deeper dehydration melting processes and
666 thereby deeper generation of intraplate alkaline magmas.

667

668 **Conclusions**

669 1. The upper stability limit of pargasitic amphibole in fertile mantle peridotite has been
670 redefined as 3.7–4.0 GPa, ~0.5 GPa higher (15 km deeper) than any previous studies. At
671 ~1000 °C, this corresponds to conditions found in the sub-cratonic lithospheric mantle,
672 indicating that pargasitic amphibole is stable to greater depths in the cratonic lithosphere.
673 This indicates an expanded reservoir for water storage in the mantle, and extends the
674 pressure range over which thermal perturbations may trigger melting due to water release
675 associated with high-temperature amphibole breakdown.

- 676 2. Amphibole is stable to substantially lower pressure (3.0-3.2 GPa) in depleted peridotites.
677 This supports previous studies that suggested that bulk composition affects amphibole
678 stability, and will help models of hydrous phase stability to incorporate bulk composition.
- 679 3. The abundance of water in the peridotite system has a huge effect on the stability of
680 amphibole, decreasing the upper pressure stability limit by ~0.2 GPa per wt.% H₂O. This
681 has significant implications for the stability of amphibole at the interface between
682 subducting slabs and the overlying mantle wedge, where high concentrations of water
683 ascending through progressively depleted mantle may destabilize amphibole at very low
684 pressures (<2 GPa). Strongly fluid-fluxed portions of the deep mantle wedge are
685 therefore likely to be amphibole-free.
- 686 4. A new thermobarometer has been developed for amphibole-bearing peridotite. It is based
687 on the alkali content of amphibole, and may be used at 1-4 GPa over a range of fertile
688 and depleted peridotite compositions and bulk H₂O contents.

689

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873

874
875 **Figure Captions**

876
877 **Fig. 1** Effect of H₂O on the upper pressure-temperature stability limit of pargasitic-edenitic
878 amphibole in peridotite, from published experiments. Experiments with <5 wt.% H₂O stabilize
879 amphibole to much higher pressures and temperatures than much wetter experiments. Stability
880 fields from Kushiro (1970), Green (1973), Millhollen et al. (1974), Mysen and Boettcher (1975),
881 Mengel and Green (1989), Wallace and Green (1991), Niida and Green (1999), Grove et al.
882 (2006), Fumagalli et al. (2009), Tumiati et al. (2013)

883
884 **Fig. 2** Compositions of starting materials used in this study (triangles) and prior studies (squares)
885 compared to alkali-enriched (Na₂O > 0.4 wt.%, K₂O > 0.03 wt.%) peridotite samples from the
886 GEOROC database (open circles). The grey field represents the depletion/re-enrichment trend
887 populated by many peridotitic mantle xenoliths that are not alkali-enriched. See Table 1 for
888 guide to symbols for experimental compositions

889

890 **Fig. 3** Back-scattered electron images of run products. A and B: Run D274 (FNa, 1050 °C, 3
891 GPa), showing the typical appearance of the capsule section (A) and each phase (B). C: Run
892 D296 (FNa, 1000 °C, 3.7 GPa), showing the presence of amphibole at 3.7 GPa. D: Run D293
893 (DNa, 1100 °C, 2 GPa), showing the appearance of spinel. Also shown in this panel is the typical
894 appearance of quench amphibole, which grew from the melt phase during quenching of the
895 experiment. Quench amphibole was identified in several 1100 °C experiments, based on its
896 elongate habit, presence in voids, and melt-like chemistry ($Mg\# = 77$)

897
898 **Fig. 4** Phase diagrams for the three experimental series conducted in this study. All runs also
899 contain olivine + orthopyroxene + clinopyroxene. See tables 1 and 2 for compositions of FNa,
900 FNa3 and DNa. Curve indicates the upper pressure-temperature limit of amphibole stability. Grt-
901 in represents the low-pressure stability limit of garnet. Size of symbols represents the estimated
902 maximum experimental uncertainties (± 10 °C, ± 0.05 GPa)

903
904 **Fig. 5** Compositions of experimental phases (in mol%) as a function of pressure at 1000 °C for
905 the FNa bulk composition. $Mg\# = 100 * \text{molar Mg}/(\text{Mg}+\text{Fe})$

906
907 **Fig. 6** Amphibole stability limit in fertile (FNa) vs. depleted (DNa) peridotite - this study

908
909 **Fig. 7** Stability fields of amphiboles in a K-rich peridotite (0.54 wt.% K_2O , composition “B +
910 Phl” in Table 1), from Konzett and Ulmer (1999; $P = 6 - 9.5$ GPa, $T = 1100 - 1400$ °C; squares)
911 and Fumagalli et al. (2009; $P = 1.5 - 5$ GPa, $T = 700 - 1000$ °C; circles). Filled and open symbols

912 indicate the presence or absence of amphibole, respectively. High-temperature/low-pressure
913 extension of Na-amphibole stability is estimated from existing studies cited in Table 1 (estimated
914 uncertainty = ± 50 °C, 0.2 GPa). Red curve: hotter portions of the Nazca slab surface pressure-
915 temperature profile from Syracuse et al. (2010; model D80). Sub-Cascadian and sub-Mexican
916 slab surfaces also pass through the K-richterite stability field. Blue curve: 40 mW/m³ geotherm,
917 representing the coldest geotherms through continental lithosphere (Lee et al., 2010)

918

919 **Fig. 8** Schematic of controls on the upper pressure-temperature stability of amphibole in mantle
920 peridotite. Not shown here are the effects of halogens, which are not addressed in this study and
921 remain poorly constrained – fluorine in particular appears to stabilize amphiboles to higher
922 temperatures and possibly pressures (Foley 1991), but much more work is required to fully
923 constrain the effects of anions on amphibole stability. The effect of fO_2 is small for relevant
924 oxygen fugacities and amphibole compositions

925

926 **Fig. 9** Correlation between pressure, temperature and molar Na₂O + K₂O in amphiboles from this
927 study. Plotted are all 19 amphibole-bearing experiments from FNa, FNa3 and DNa runs; no
928 systematic deviation from this correlation is observed in any of the three series. Grid shows the
929 planar fit defined by Equation (1). Grid increments are 0.1 GPa and 10 °C. Only 13 experiments
930 are visible here; the other 6 plot just below the grid

931

932 **Fig. 10** Predictive capabilities of the alkali-in-amphibole thermobarometer. (a) Predicted vs.
933 measured molar Na₂O + K₂O contents of experimental amphiboles synthesized in peridotite

934 compositions. Filled yellow circles: data from this study; filled black circles: literature data. (b)
935 Predicted vs. measured pressures for the same amphiboles as (a). Grey ellipse highlights data
936 from the extremely Ca-depleted bulk composition used by Pirard & Hermann (2015) - see text
937 for details. (c) and (d) Difference between measured and modeled $\text{Na}_2\text{O} + \text{K}_2\text{O}$ contents of
938 amphiboles as a function of bulk H_2O (c) or K_2O (d). The experiments with 14 wt.% bulk H_2O
939 all have 0.03 wt.% bulk K_2O ; the experiments with 0.43 wt.% bulk K_2O have 0.2-0.4 wt.% bulk
940 H_2O ; the experiments with 0.54 wt.% bulk K_2O have 1.5-5.0 wt.% bulk H_2O . Data from Mengel
941 and Green (1989), Wallace and Green (1991), Niida and Green (1999), Grove et al. (2006),
942 Fumagalli et al. (2009), Green et al. (2014), and new analyses from the Grove et al. (2006)
943 experiments. Data from Pirard and Hermann (2015) are excluded from error estimates – see text
944

945 **Fig. 11** Effect of bulk H_2O content on the upper pressure-temperature stability of amphibole in
946 fertile peridotite. (a) Upper stability curves for experimental series FNa (0.65 wt.% H_2O) and
947 FNa3 (3 wt.% H_2O), plus that of Grove et al. (2006; 14.5 wt.% H_2O), produced in the same lab
948 for a fertile peridotite with slightly lower alkali contents and much higher H_2O . (b)
949 Representation of data in (a) to show effect of H_2O on pressure- (solid black line) and
950 temperature- (dashed blue line) stability. Error bars represent absolute uncertainty for pressure
951 and maximum uncertainty for temperature. Curves should not be extrapolated to lower H_2O
952 contents because no destabilizing effect has been resolved for H_2O contents significantly less
953 than 1 wt.%
954

955 **Fig. 12** Stability field of amphibole in fertile and depleted peridotite from this study (solid black
 956 lines) compared to previous experimental studies (dashed grey lines; references as in Figure 1).
 957 Grey region = range of cratonic geotherms determined from xenoliths, broadly corresponding to
 958 heat flows of 40-50 mW m⁻² (Lee et al. 2011). Black region = range of slab surface temperatures
 959 for Mexico and the Cascades from the D80 model of Syracuse et al. (2010)

960

961 **Table Captions**

962

963 **Table 1** Starting compositions (wt.%) used in mantle experiments that provide information on
 964 amphibole stability

965

966 **Table 2** Experimental conditions and phase assemblages

967

968 **Table 3** Amphibole compositions (wt.%) from experiments

Table 1 Starting compositions (wt.%) used in mantle experiments that provide information on amphibole stability

	This study		Other studies									
	FNa ¹	DNa ²	MPY-40 ³	Tin-40 ⁴	HPY-40 ⁵	"C" ⁵	M&B A ⁶	M&B B	M&B C	M&B D	H&Z ⁷	HZ2 ⁸
SiO ₂	44.65	43.72	47.15	47.5	47.9	47.2	45.7	43.7	45.1	44.82	46.2	46.84
TiO ₂	0.15	0.05	0.28	0.13	1.18	0.25	0.05	0.2	0.13	0.52	0.18	0.19
Al ₂ O ₃	3.87	1.79	7.28	5.35	5.91	5.6	1.6	4	3.92	8.21	4.06	4.28
Cr ₂ O ₃	0.37	0.42	0.75	0.75	0.72	1.08					0.4	0.42
FeO	8.14	7.97	7.27	7.52	8.81	6.69	5.9	8.89	8.19	9.77	7.56	4.14
MnO	0.14	0.13	0.12	0.18	0.13	0.07	0.09	0.12	0.14	0.19	0.1	0.1
MgO	37.70	42.82	30.57	32.8	28.8	33.59	42.8	37.4	38.81	26.53	37.82	39.8
CaO	3.38	1.69	5.63	4.97	5.14	4.43	0.7	3.5	2.66	8.12	3.22	3.39
Na ₂ O	0.62	0.40	0.66	0.3	0.95	0.35	0.09	0.38	0.27	0.89	0.33	0.35

K ₂ O	0.08	0.06	0	0.03	0.22	0.43	0.04	0.01	0.02	0.03	0.03	0.03
NiO	0.25	0.30	0.29	0.43	0.13	0.3	0.26	0.24	0.25	0.2	0.28	0.29
H ₂ O added	0.65 & 3	0.65	0.6-0.7	0.2 & 1	0.3 & 10	0.15-0.4	11-25	8-38	14-38	13-36	0-14.5 ⁷	0-14.5
Natural / Synthetic	Syn	Syn	Syn	Syn	Syn	Syn	Nat	Nat	Nat	Nat	Syn	Syn
Symbol in Figure 2	Open triangle	Filled triangle	Yellow square	Dark blue square	Black square	Grey square	Green triangle	Green triangle	Green triangle		Open square	

¹Representative fertile mantle composition from the GEOROC database of mantle xenoliths, plus 1 wt.% of a metasomatic fluid composed of 65 wt.% H₂O, 30 wt.% N

²Representative depleted mantle determined from the GEOROC database of mantle xenoliths, plus 1 wt.% of a metasomatic fluid composed of 65 wt.% H₂O, 30 wt.% N

³MORB pyrolite minus 40 % olivine, from Niida and Green (1999)

⁴Tinaquillo lherzolite minus 40 % olivine, from Wallace and Green (1991)

⁵Composition C from Mengel and Green (1989): average depleted spinel peridotite from the North Hessian Depression, NW Germany, plus 1.5 % phlogopite, minus 60 %

⁶Compositions M&B A, B, C and D are natural lherzolites from Mysen and Boettcher (1975).

⁷Estimated primitive mantle from Hart and Zindler (1986). Used in Grove et al. (2006; 14.5 wt.% H₂O), Green et al. (2010, 2014; 0-1.45 wt.% H₂O), and Till et al. (2011)

⁸Hart & Zindler (1986) primitive mantle composition, but with significant Fe depletion. From Green et al. (2010; 2014)

⁹Mildly depleted lherzolite -30 % olivine +5 wt.% phlogopite +0.4 wt.% Na₂O. From Konzett and Ulmer (1999); used in KNCFMASH by Fumagalli et al. (2009) and T

¹⁰Pargasite-rich spinel peridotite mylonite from St. Paul's Rocks containing 5.7 wt.% H₂O bound in hydrous minerals. From Millhollen et al. (1974).

¹¹Mixture of 75% San Carlos olivine + 25% sediment melt. From Pirard and Hermann (2015).

969

970

Table 2 Experimental conditions and phase assemblages

Run	Starting Material	Duration (hr)	T (°C)	P (GPa)	Bulk H ₂ O (wt%)	Anhydrous phases	Hydrous phases
D278	FNa	167	950	3	0.65	Ol, opx, cpx, grt	amph
D263	FNa	165	1000	2	0.65	Ol, opx, cpx	amph
D264	FNa	163	1000	2.5	0.65	Ol, opx, cpx, grt	amph
D267	FNa	168	1000	3	0.65	Ol, opx, cpx, grt	amph
D281	FNa	165	1000	3.5	0.65	Ol, opx, cpx, grt	amph
D296	FNa	163	1000	3.7	0.65	Ol, opx, cpx, grt	amph
D310	FNa	147	1000	4	0.65	Ol, opx, cpx, grt	-
D265	FNa	167	1050	2	0.65	Ol, opx, cpx	amph
D272	FNa	168	1050	2.5	0.65	Ol, opx, cpx, grt	amph
D274	FNa	168	1050	3	0.65	Ol, opx, cpx, grt	amph
D285	FNa	142	1050	3.5	0.65	Ol, opx, cpx, grt	amph

D289	FNa	169	1050	3.7	0.65	Ol, opx, cpx, grt	-
D282	FNa	142	1100	2	0.65	Ol, opx, cpx	amph
D280	FNa	162	1100	2.5	0.65	Ol, opx, cpx, grt	amph
D286	FNa	95	1100	3	0.65	Ol, opx, cpx, grt	-
D283	FNa	120	1100	3.5	0.65	Ol, opx, cpx, grt	-
D298	FNa3 ¹	170	1000	3	3	Ol, opx, cpx, grt	amph
D305	FNa3	141	1000	3.2	3	Ol, opx, cpx, grt	amph
D308	FNa3	167	1000	3.5	3	Ol, opx, cpx, grt	-
D299	FNa3	139	1050	2	3	Ol, opx, cpx	amph
D302	FNa3	148	1050	2.5	3	Ol, opx, cpx, grt	amph
D309	FNa3	144	1050	3	3	Ol, opx, cpx, grt	-
D311	FNa3	146	1075	2.5	3	Ol, opx, cpx, grt	-
D304	FNa3	142	1100	2	3	Ol, opx, cpx, sp	-
D295	DNa	120	1000	2	0.65	Ol, opx, cpx	amph
D297	DNa	170	1000	3.2	0.65	Ol, opx, cpx, grt	-
D291	DNa	163	1000	3.5	0.65	Ol, opx, cpx, grt	-
D288	DNa	167	1050	2.5	0.65	Ol, opx, cpx, sp	amph
D290	DNa	166	1050	3	0.65	Ol, opx, cpx, grt	amph
D293	DNa	140	1100	2	0.65	Ol, opx, cpx, sp	-
D292	DNa	168	1100	2.5	0.65	Ol, opx, cpx, sp	-

ol = olivine; *opx* = orthopyroxene; *cpx* = clinopyroxene; *grt* = garnet; *sp* = spinel; *amph* = amphibole

All runs also contain a melt phase

¹FNa3 consists of composition FNa with additional H₂O added to give 3 wt.% H₂O in the bulk

971

972

Table 3 Amphibole compositions (wt.%) from experiments

Run No.	Bulk	T (°C)	P (GPa)	n ¹	SiO ₂	SiO ₂ / σ ²	TiO ₂	TiO ₂ / σ	Al ₂ O ₃	Al ₂ O ₃ / σ	Cr ₂ O ₃	Cr ₂ O ₃ / σ	FeO	FeO / σ	MnO	MnO / σ	MgO	MgO / σ	CaO	CaO / σ
D278	FNa	950	3	9	46.5	0.9	0.54	0.05	12.0	0.7	1.41	0.12	3.67	0.09	0.08	0.02	19.8	0.6	10.5	0.4
D263	FNa	1000	2	12	45.4	0.5	0.54	0.04	14.1	0.2	1.18	0.12	3.69	0.10	0.08	0.03	18.6	0.2	11.4	0.2
D264	FNa	1000	2.5	12	45.6	0.3	0.57	0.04	13.9	0.4	1.28	0.13	3.71	0.08	0.08	0.01	18.6	0.2	11.3	0.2
D267	FNa	1000	3	13	46.9	0.4	0.61	0.05	12.0	0.3	1.19	0.21	3.44	0.09	0.07	0.01	19.5	0.3	10.4	0.2
D281	FNa	1000	3.5	17	49.2	0.4	0.68	0.05	10.5	0.2	1.11	0.10	3.24	0.10	0.06	0.02	19.8	0.4	9.10	0.2

D296	FNa	1000	3.7	12	48.3	0.9	0.66	0.08	10.0	0.7	1.07	0.17	3.41	0.10	0.07	0.02	21.2	0.5	8.8	0
D265	FNa	1050	2	11	45.9	0.3	0.54	0.03	14.0	0.2	1.20	0.13	3.72	0.13	0.08	0.01	18.6	0.2	11.4	0
D272	FNa	1050	2.5	15	45.9	0.5	0.62	0.03	13.7	0.4	1.22	0.18	3.72	0.07	0.09	0.02	18.7	0.4	11.0	0
D274	FNa	1050	3	14	46.9	0.7	0.74	0.06	11.8	0.3	1.31	0.12	3.61	0.08	0.08	0.02	19.0	0.4	10.1	0
D285	FNa	1050	3.5	8	47.3	0.5	0.72	0.06	10.4	0.1	0.95	0.07	3.54	0.07	0.06	0.02	20.7	0.3	8.7	0
D282	FNa	1100	2	17	44.0	0.6	0.64	0.04	14.3	0.3	1.67	0.15	3.48	0.06	0.09	0.01	19.6	0.5	10.7	0
D280	FNa	1100	2.5	14	44.0	0.6	0.59	0.03	14.4	0.3	1.29	0.12	3.64	0.08	0.08	0.01	19.7	0.5	10.3	0
D298	FNa3	1000	3	12	45.6	0.6	0.75	0.10	11.8	0.3	1.25	0.19	3.55	0.05	0.06	0.01	20.2	0.4	10.2	0
D305	FNa3	1000	3.2	11	46.2	0.6	0.74	0.11	11.6	0.5	1.25	0.14	3.61	0.08	0.07	0.02	20.3	0.5	9.8	0
D299	FNa3	1050	2	18	44.2	0.7	0.59	0.03	14.3	0.2	1.45	0.16	3.71	0.07	0.07	0.01	19.5	0.4	11.1	0
D302	FNa3	1050	2.5	16	44.2	0.6	0.61	0.06	14.1	0.5	1.35	0.20	3.64	0.06	0.06	0.01	19.4	0.3	10.7	0
D295	DNa	1000	2	11	46.2	0.5	0.40	0.02	11.6	0.3	2.14	0.26	3.21	0.13	0.08	0.01	21.1	0.8	10.7	0
D288	DNa	1050	2.5	13	45.0	0.4	0.40	0.03	12.7	0.1	1.85	0.19	3.83	0.12	0.08	0.02	19.7	0.2	10.4	0
D290	DNa	1050	3	19	45.2	0.6	0.49	0.05	12.6	0.4	1.86	0.25	4.15	0.21	0.07	0.01	19.6	0.4	10.0	0

¹Number of analyses

²One standard deviation of the analyses for this oxide

³Mg# = 100 * molar Mg/(Mg+Fe)

973

974

975

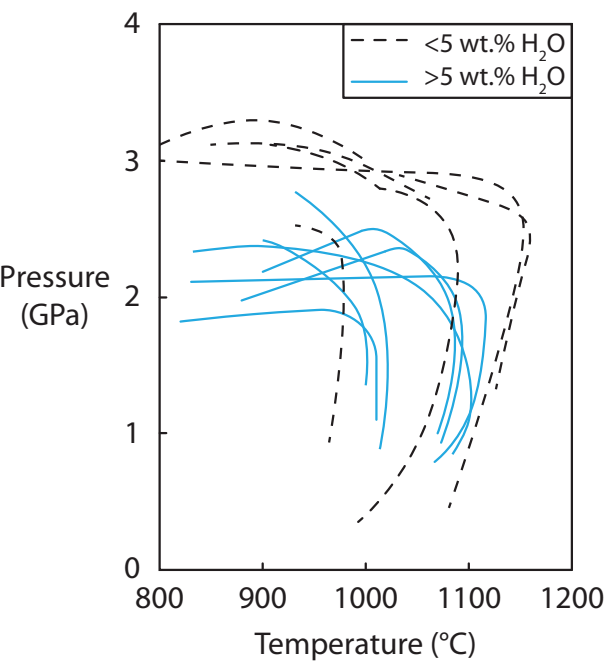


Figure 1

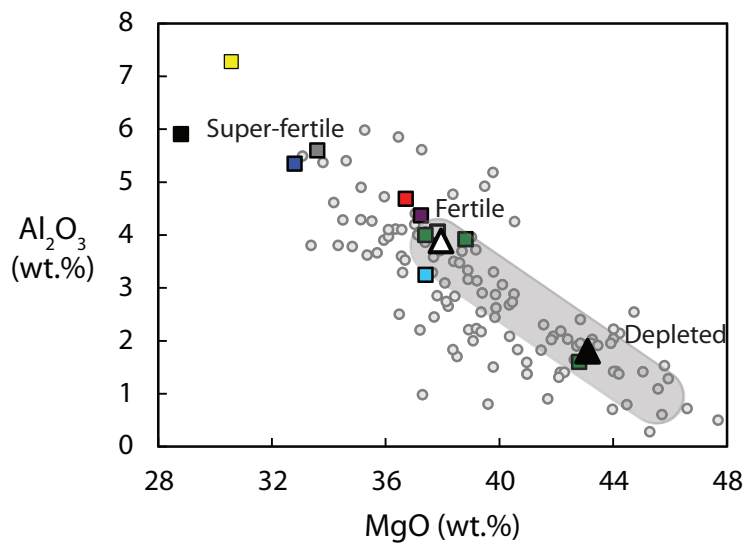
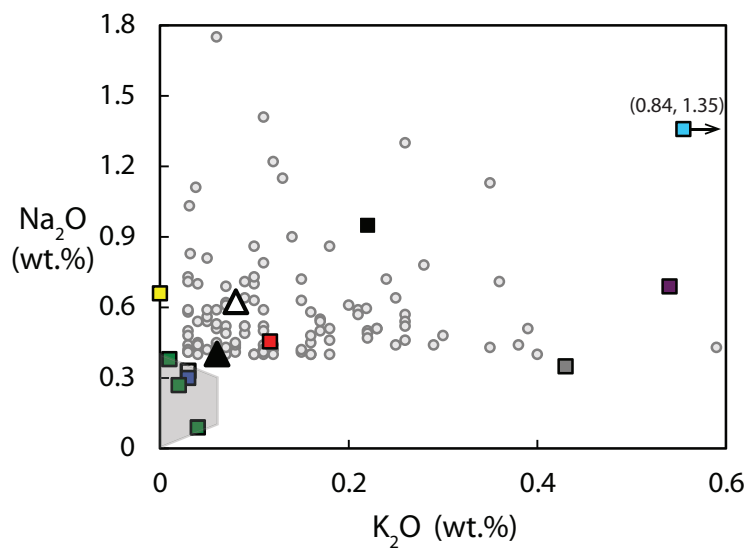


Figure 2

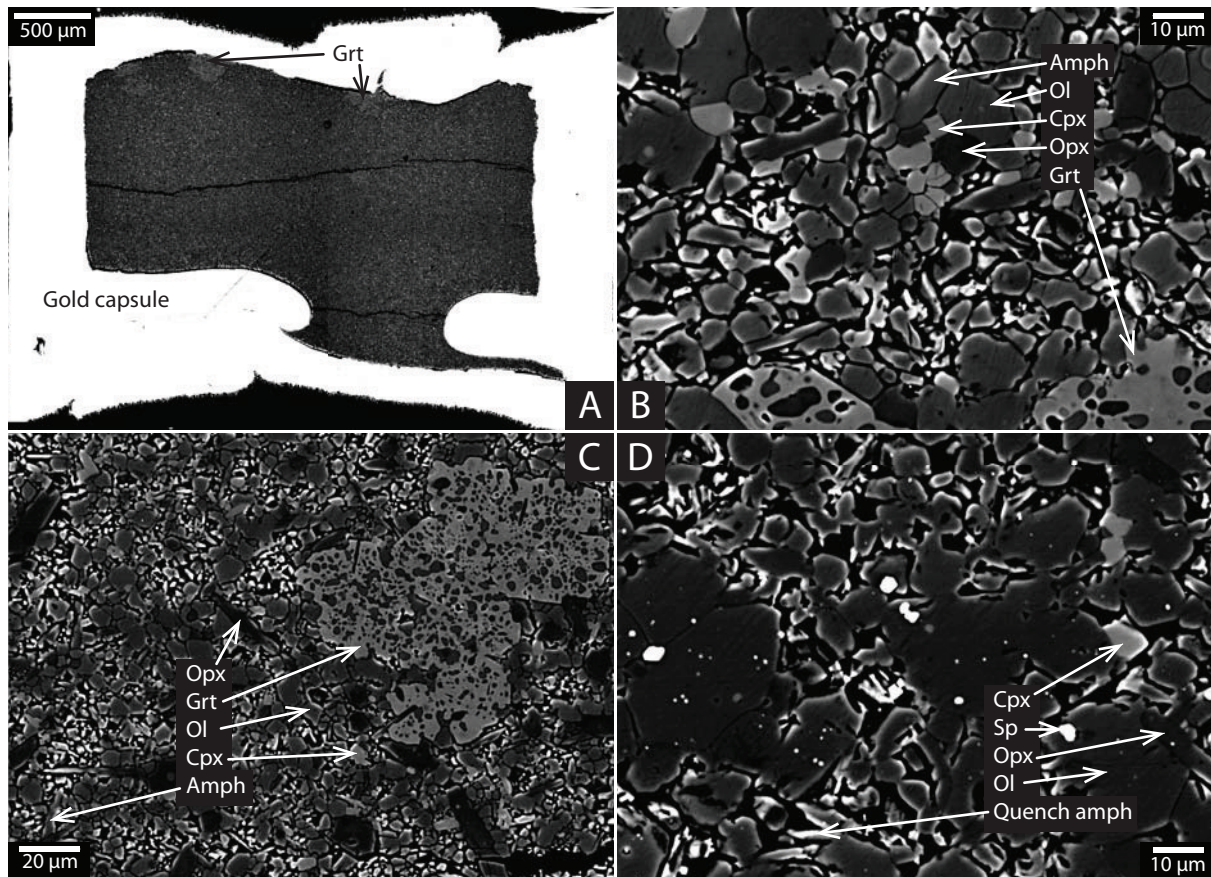


Figure 3

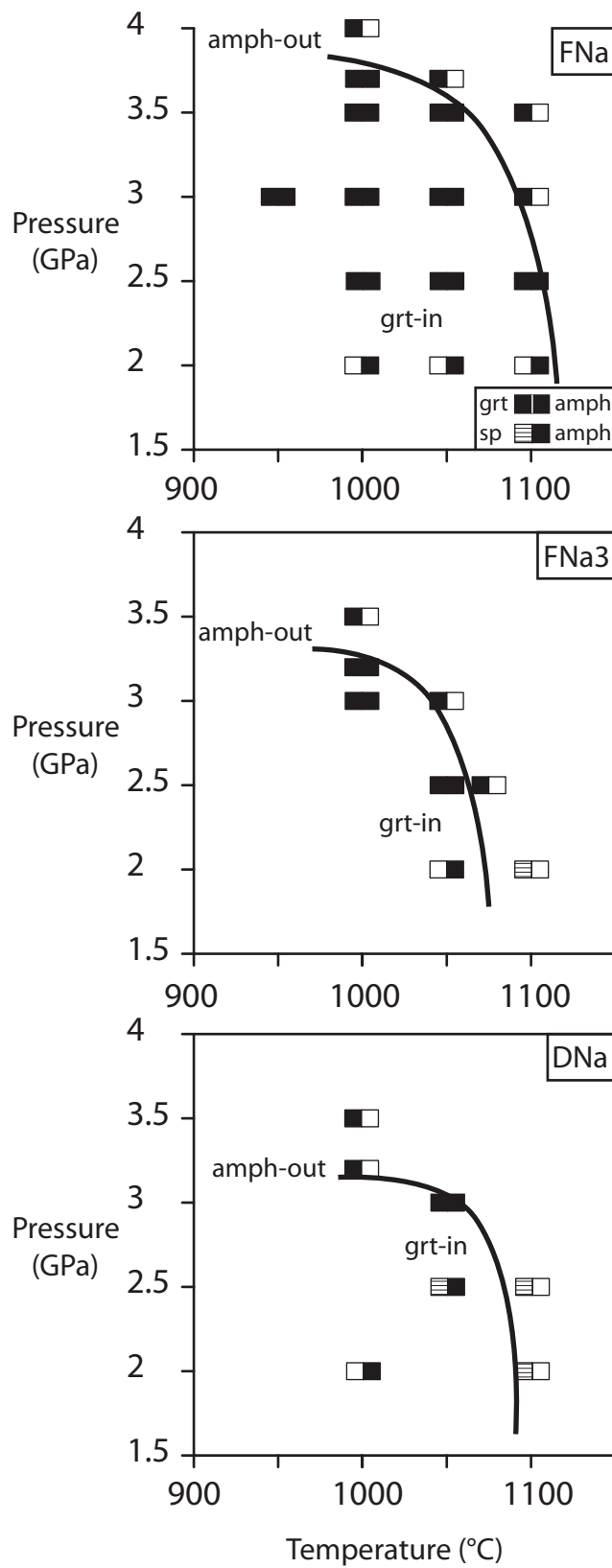


Figure 4

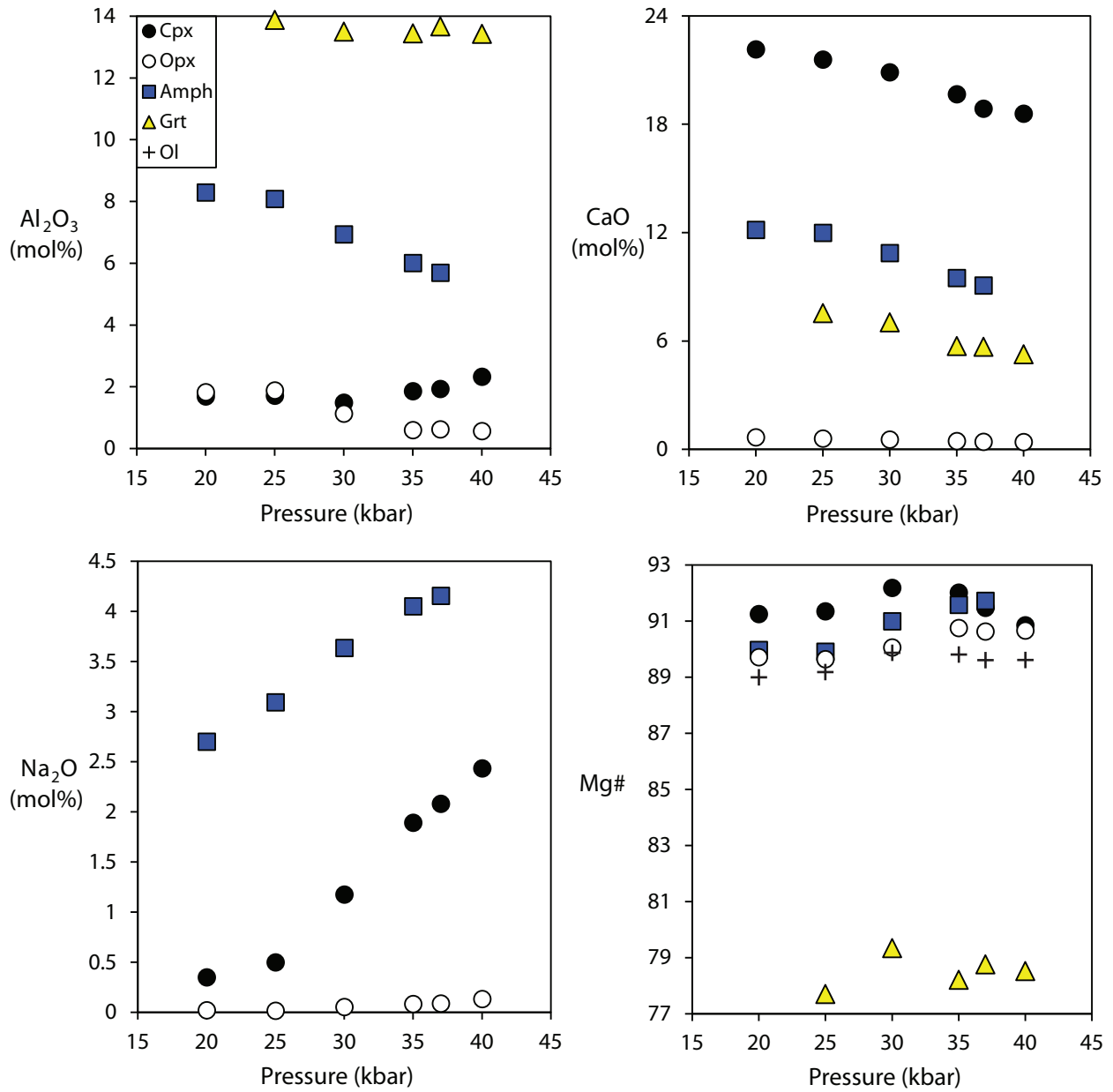


Figure 5

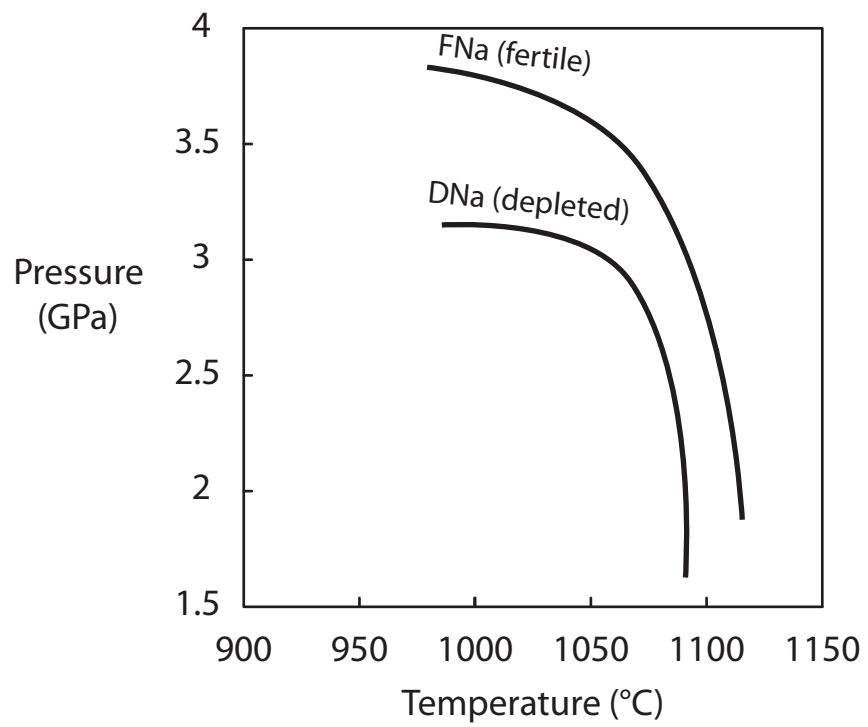


Figure 6

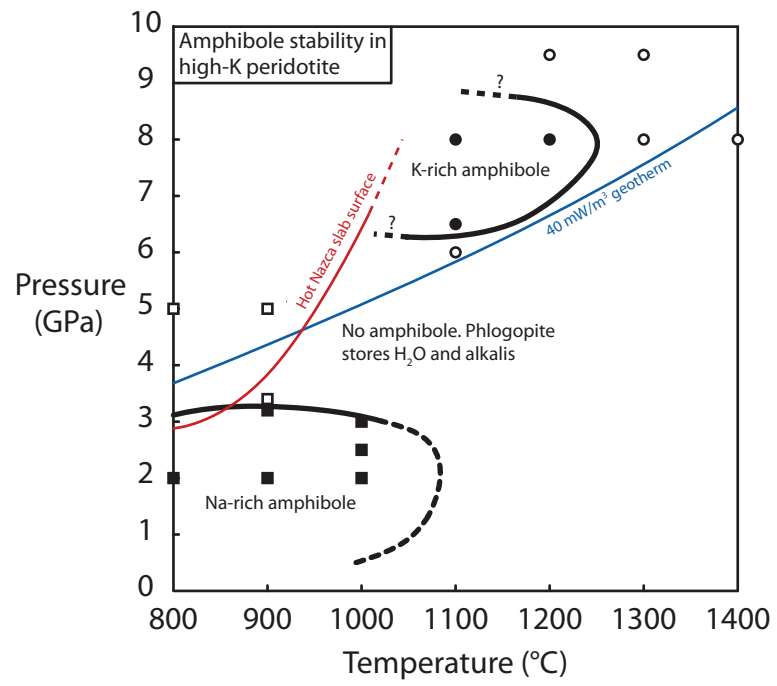


Figure 7

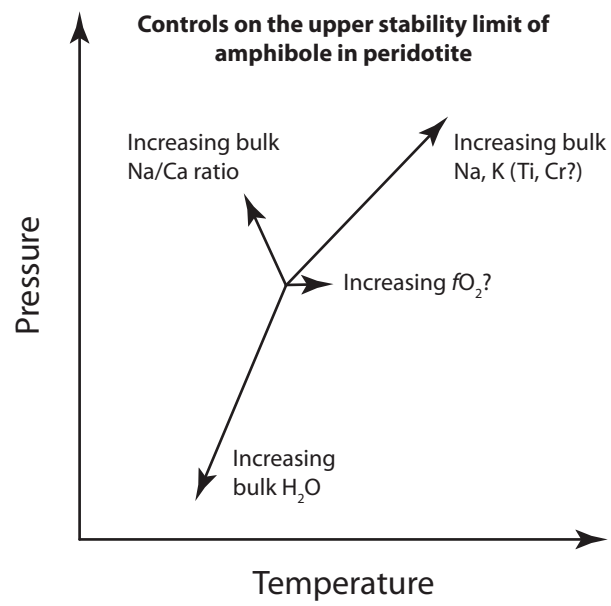


Figure 8

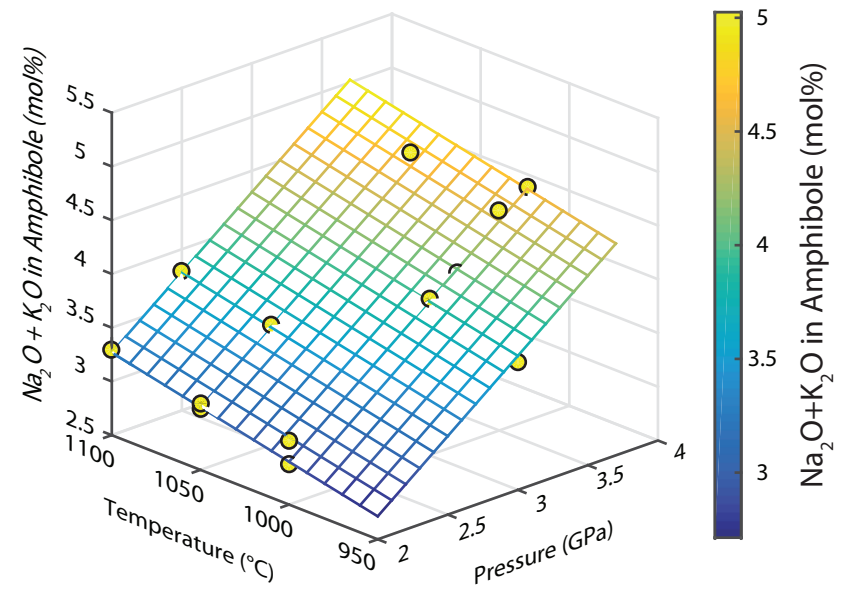


Figure 9

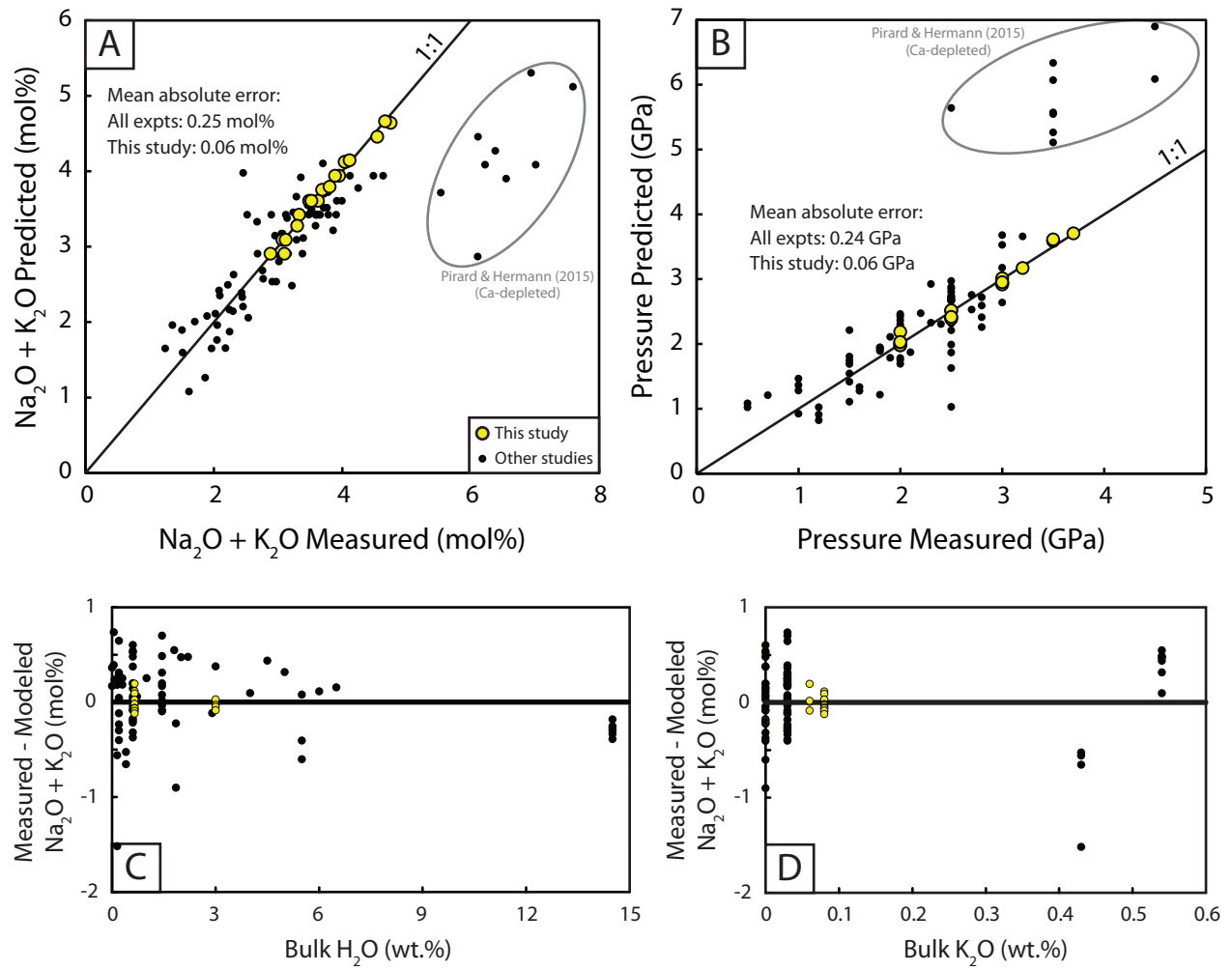


Figure 10

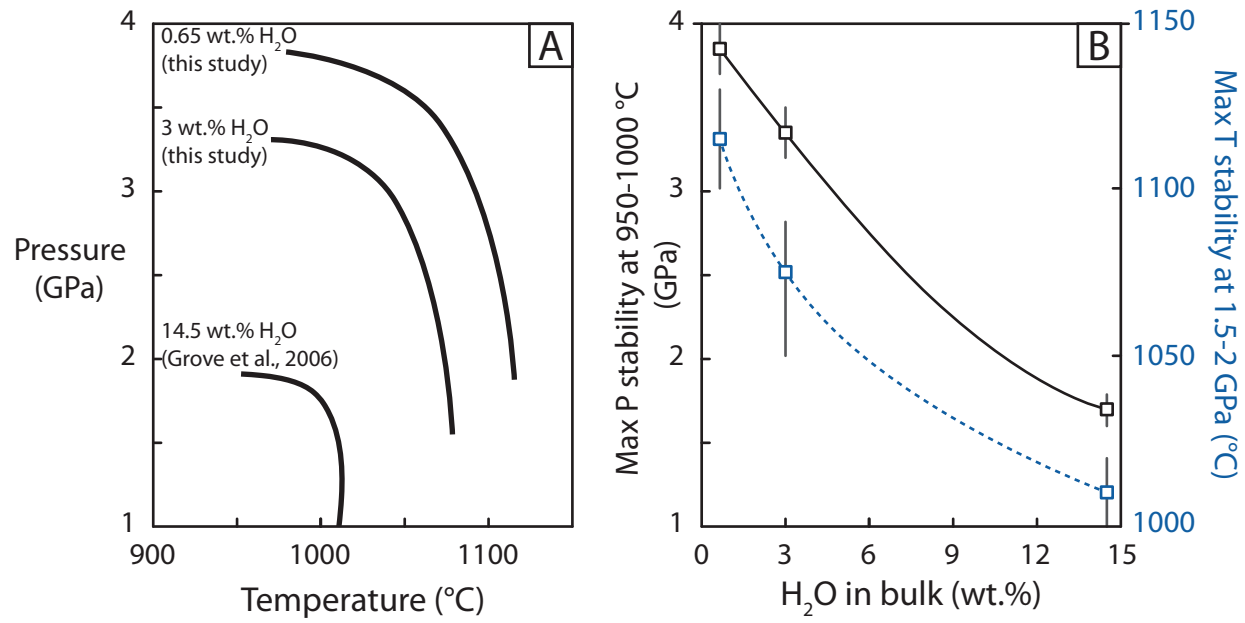


Figure 11

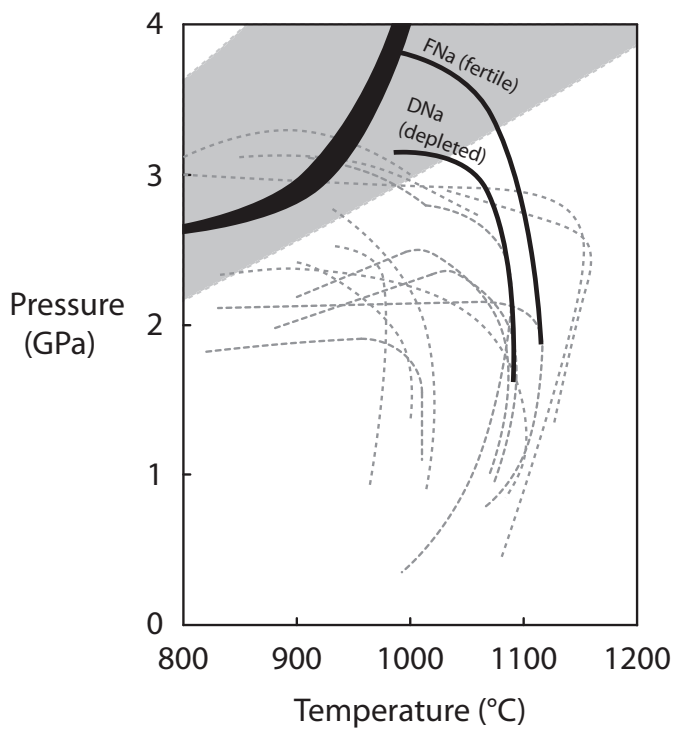


Figure 12