

There is good evidence that the deuteron ground state is primarily $1s$ ($n=1, l=0$). First, the lowest energy state in practically all the model potentials is an s state. Secondly, the magnetic moment of H^2 is approximately the sum of the proton and the neutron moments, indicating that $\underline{s}_n \parallel \underline{s}_p$ and no orbital motion of the proton relative to the neutron; this is also consistent with the total angular momentum of the ground state, $I=1$.

We therefore consider only the $l=0$ radial wave equation,

$$-\frac{\hbar^2}{m} \frac{d^2 u(r)}{dr^2} + V(r)u(r) = E u(r)$$

$\downarrow -E_B$

where $u(r) = rR(r)$ and m is the neutron (or proton) mass.

Notice this equation is mathematically equivalent to the 1-D wave equation except for the B.C., $u(0) = 0$.

Solutions:

$$u(r) = a e^{iKr} + b e^{-iKr} \quad r < r_0, \quad K = [m(V_0 - E_B)]^{1/2} / \hbar$$

$$u(r) = a' e^{K'r} + b' e^{-K'r} \quad r > r_0, \quad K' = \sqrt{mE_B} / \hbar$$

Applying the boundary conditions

- (i) $u(0) = 0$, so $R(0)$ is finite
- (ii) $u(r \rightarrow \infty) = 0$ since we are dealing with a bound state
- (iii) continuity of u and first-order derivative at $r=r_0$

one obtains

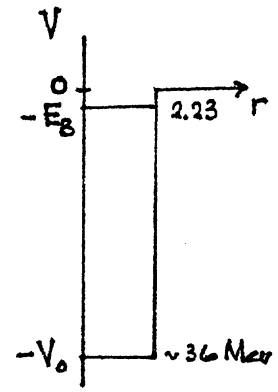
$$u(r) = c \sin Kr \quad r < r_0$$

$$a' = 0, \quad \therefore u(r) = b' e^{-K'r} \quad r > r_0$$

$$K \cot Kr_0 = -K' \quad \text{this is a relation between } V_0, E_B \text{ and } r_0$$

$$\text{or } \tan Kr_0 = -\frac{K}{K'} = -\left(\frac{V_0 - E_B}{E_B}\right)^{1/2}$$

Suppose $V_0 \gg E_B$, then RHS is large and $Kr_0 \sim \pi/2$ (we will see below that $V_0 \sim 36$ Mev as compared to $E_B \sim 2.23$ Mev).



Then
$$K \sim \frac{\sqrt{mV_0}}{\hbar} \sim \frac{\pi}{2r_0}$$

or
$$V_0 r_0^2 \sim \left(\frac{\pi}{2}\right)^2 \frac{\hbar^2}{m} \sim 1 \text{ Mev-barn}$$

We see from this result that a knowledge of E_B allows us to only determine the product $V_0 r_0^2$. From n-p scattering (discussed below) we will find $r_0 \sim 2F$, thus we obtain $V_0 \sim 36 \text{ Mev}$.

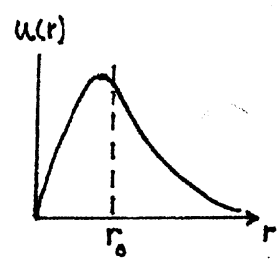
Since the interior wave function, $\sin Kr$, must match with the exterior wave function, $\exp(-Kr)$, the quantity Kr_0 must be slightly greater than $\pi/2$ (more accurate estimate gives 116° instead of 90°), so if we write

$$K = 2\pi/\lambda \sim \pi/2r_0$$

then the 'effective wavelength' λ is approximately $4r_0$ which suggests that much of the wave function is not in the interior region.

We can estimate the relaxation constant (or decay length) in the exterior region, $r > r_0$

$$\frac{1}{K'} = \frac{\hbar}{\sqrt{mE_B}} \sim 4.3 \text{ F}$$



This means that the two nucleons in H^2 spend a large fraction of their time at $r > r_0$, the classically forbidden region of negative kinetic energy. We can calculate the root-mean-square radius of the deuteron wave function,

$$R_{rms}^2 \equiv \frac{\int_0^\infty dr r^2 R^2(r)}{\int_0^\infty dr R^2(r)} = \frac{\int_0^\infty r^2 dr r^2 R^2(r)}{\int_0^\infty r^2 dr R^2(r)}$$

If we put $R(r) \sim e^{-K'r}$ for all r which should result in an overestimate, one gets

$$R_{rms}^2 = \frac{\hbar}{\sqrt{2mE_B}} = 3 \text{ F}$$

which can be compared with

$$(1.4 \times A^{1/3})^2 \sim 3.11 \quad \text{or} \quad (1.2 \times A^{1/3})^2 \sim 2.29 \quad A=2$$

B. Neutron-proton Scattering

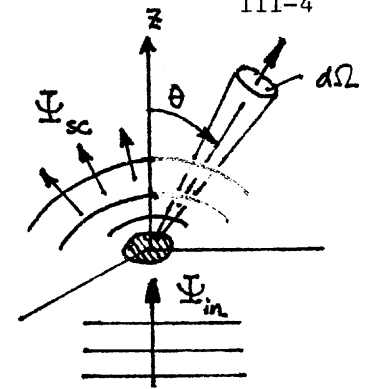
One can also obtain information about the neutron-proton interaction by studying the scattering of neutrons by a hydrogen sample. To describe the scattering process we imagine the incident neutrons are represented by a plane wave of the form

$$\Psi_{in} = b e^{i(kz - \omega t)}$$

where $k = \sqrt{2\mu T}/\hbar$, μ is the reduced mass of the incoming particle.

T is the kinetic energy in center-of-mass coordinates (CMCS).

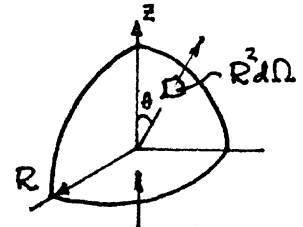
Incident flux $J_{in} = v \Psi_{in}^* \Psi_{in} = v |b|^2$, v is relative speed.



We look for the scattered wave in the form of a spherical outgoing wave,

$$\Psi_{sc} = f(\theta) b \frac{e^{i(kr - \omega t)}}{r}$$

where $f(\theta)$ is called the scattering amplitude.



$$dN = (\Psi_{sc}^* \Psi_{sc} v) R^2 d\Omega$$

No. particles scattered into area $R^2 d\Omega$ per unit time.

$$\frac{d\sigma}{d\Omega} = \frac{dN/d\Omega}{J_{in}} = |f(\theta)|^2$$

angular differential scattering cross section

$$\sigma = \int d\Omega |f(\theta)|^2$$

total scattering cross section

The quantity we want to calculate is $f(\theta)$. Stating the problem in another way, we want to solve the wave equation

$$-\frac{\hbar^2}{2\mu} \nabla^2 \Psi + V(r)\Psi = T\Psi$$

subject to the B.C.

$$\Psi \xrightarrow{r \rightarrow \infty} b \left[e^{ikz} + f(\theta) \frac{e^{ikr}}{r} \right]$$

We proceed by considering a partial wave analysis:

Since $V(r)$ is spherically symmetric it is most convenient to discuss the solution in spherical coordinates. We already know that the general solution is of the form

$$R(r) P_l^{(m)}(\cos\theta) e^{im\varphi}$$

which, because the potential does not depend on φ , reduces to

$R_l(r) P_l(\cos\theta)$. Thus we expand

$$\Psi(r, \theta) = \sum_{l=0}^{\infty} \psi_l(r) P_l(\cos\theta)$$

one can show that $e^{ikz} = e^{ikr \cos \theta} = \sum_l F_l(r) P_l(\cos \theta)$,
 then $F_0(r) = \sin kr / kr$

[For arbitrary l , one finds $F_l(r) \sim i^l (2l+1) \frac{\sin(kr - l\pi/2)}{kr}$.
 Furthermore, we will write $f(\theta) = \sum_l f_l P_l(\cos \theta)$.

Then

$$\psi_0 \underset{r \rightarrow \infty}{\sim} b \left[\frac{\sin kr}{kr} + f_0 \frac{e^{ikr}}{r} \right]$$

$$= \frac{b}{2ikr} \left[(1 + 2ikf_0) e^{ikr} - e^{-ikr} \right] \equiv R_0(r) = \frac{u_0(r)}{r} \quad (*)$$

\downarrow outgoing \downarrow incoming

For $r > r_0$, the wave equation is

$$-\frac{\hbar^2}{2\mu} \frac{d^2 u_0}{dr^2} = T u_0, \quad k = \sqrt{2\mu T / \hbar^2}$$

$$u_0(r) = A \sin(kr + \delta_0) \quad \text{s-wave phase shift}$$

$$= \frac{A e^{-i\delta_0}}{2i} (e^{ikr} e^{2i\delta_0} - e^{-ikr}) \quad (**)$$

Comparing (*) and (**) one obtains

$$1 + 2ikf_0 = e^{2i\delta_0}$$

$$f_0 = \frac{e^{2i\delta_0} - 1}{2ik} = e^{i\delta_0} \frac{\sin \delta_0}{k} \quad (+)$$

Thus,

$$\frac{d\sigma}{d\Omega} = \frac{\sin^2 \delta_0}{k^2} = \chi^2 \sin^2 \delta_0, \quad \sigma = 4\pi \chi^2 \sin^2 \delta_0$$

We will define the scattering length as $a \equiv \lim_{k \rightarrow 0} (-f_0)$.

Since f_0 can not $\rightarrow \infty$ as $k \rightarrow 0$ (otherwise $\sigma \rightarrow \infty$), (+) shows that δ_0 must $\rightarrow 0$ as $k \rightarrow 0$, so

$$a = \lim_{k \rightarrow 0} \left[-e^{i\delta_0} \frac{\sin \delta_0}{k} \right] = -\frac{\delta_0}{k}$$

We can therefore obtain

$$\lim_{k \rightarrow 0} \frac{d\sigma}{d\Omega} = a^2$$

and $\lim_{k \rightarrow 0} \sigma = 4\pi a^2$

$\delta > 0$ (attractive pot.)

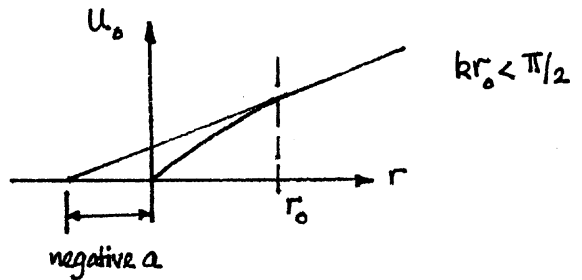
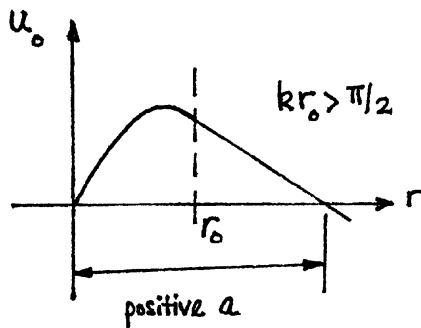
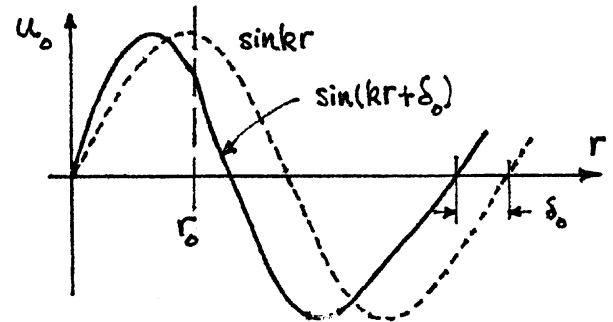
$\delta < 0$ (repulsive)

Significance of a as an extrapolation distance:

Notice that

$$u_0 = A \sin k(r-a), \quad r > r_0$$

$$\underset{kr_0 \rightarrow 0}{\sim} Ak(r-a)$$



[this case corresponds to
a potential which can
give rise to a bound
state]

[this potential can give
rise to a virtual state]

We now apply these results to n-p scattering. In this case we have the same equation as the deuteron calculation except that $E > 0$.

$$r < r_0 \quad u_0 = C \sin K'r, \quad K' = [2\mu(V_0 + T)]^{1/2} / \hbar$$

$$r > r_0 \quad u_0 = A \sin(kr + \delta_0), \quad k = \sqrt{2\mu T} / \hbar, \quad \mu = \text{reduced mass} = m/2$$

Matching B.C. at $r=r_0$,

$$K' \cot K'r_0 = k \cot(kr_0 + \delta_0)$$

this gives δ_0 in terms of V_0 , r_0 , and T .

To simplify the calculation we will assume

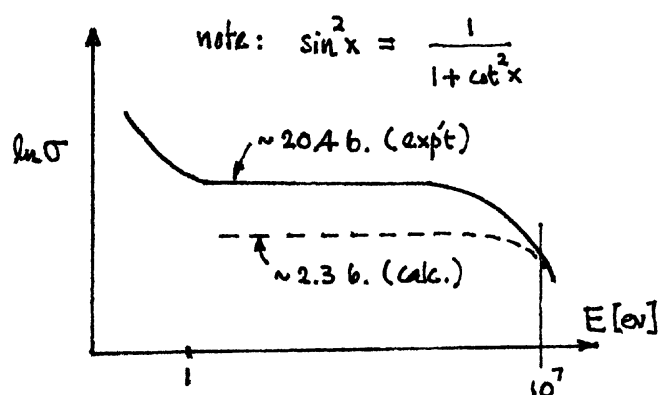
- (i) $a \gg r_0$ so that $kr_0 \ll \delta_0$
- (ii) $T \ll V_0$ so $K' \sim K$

then $k \cot \delta_0 \approx K \cot Kr_0 = -K_V \leftarrow$ from the deuteron problem, $K_V = \sqrt{mE_B}/\hbar$

$$\frac{d\sigma}{d\Omega} = \frac{\sin^2 \delta_0}{k^2} = \frac{1}{k^2} \frac{1}{1 + K_V^2/k^2}$$

$$\frac{d\sigma}{d\Omega} \approx \frac{1}{k^2 + K_V^2} = \frac{\hbar^2}{m} \frac{1}{T + E_B}$$

$$\sigma \approx \frac{4\pi\hbar^2}{m} \frac{1}{T + E_B}$$



Since $T < E_B$ (2.23 Mev), $\sigma \sim 2.3$ barns

This does not agree with the experimental value which is ~ 20.4 barns.

Wigner (1933) has suggested that n-p scattering depends on whether the neutron and proton spins are parallel (triplet state, $I=1$) or anti-parallel (singlet state, $I=0$); for these two cases the potential (and $\therefore \delta_0$) is different. Following this idea one can write

$$\frac{d\sigma}{d\Omega} = \frac{1}{k^2} \left[\frac{1}{4} \sin^2 \delta_{os} + \frac{3}{4} \sin^2 \delta_{ot} \right]$$

↑ singlet ↑ triplet

We already know that the triplet state gives rise to the ground state of the deuteron at $E = -E_B$. If the singlet state produces a state of energy $E = E^*$ (E^* could be positive or negative), then

$$\sigma \approx \frac{\pi\hbar^2}{m} \left(\frac{3}{T + E_B} + \frac{1}{T + |E^*|} \right)$$

With $|E^*| \approx 70$ kev, σ is now in good agreement with experiment.

Since $|E^*| \ll V_0$, just like $E_B \ll V_0$ in the triplet state, the singlet wave function inside the nuclear potential is also $\sim \frac{1}{4}$ of a sine wave,

$$Kr_0 \approx \frac{\pi}{2}, \text{ or}$$

$$V_{os} r_{os}^2 \approx V_{ot} r_{ot}^2 \sim 1.0 \text{ Mev-barn}$$

However, using information from neutron scattering from parahydrogen ($I=0$) one finds a scattering length $a_s = -24 \text{ F}$, thus indicating that the singlet interaction gives rise to a virtual state ($a_s < 0$).

In summary, one finds from such considerations the following results [cf. Preston, Physics of the Nucleus]

<u>Interaction</u>	<u>Scattering lengths a[F]</u>	<u>r_0 [F]</u>	<u>V_0 [Mev]</u>
n-p(triplet)	5.4	2	36
n-p(singlet)	-23.7 (n-p)	~ 2.5	18
	-17 (p-p)		

Final remark:

Experimentally it is known that the total angular momentum (nuclear spin) of the deuteron ground state is $I=1$, where

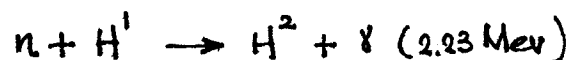
$$\underline{I} = \underline{L} + \underline{S}$$

where \underline{L} is the orbital angular momentum and \underline{S} the intrinsic spin. It is also known that the ground state is mostly $1s$ ($\underline{l}=0$), therefore for this state $S=1$ (n and p spins parallel). Now we have seen from the discussions of part A that the ground state is barely bound ($E_B=2.23 \text{ Mev}$), so all the higher energy states are not bound states. Example, the $1s$ state with $S=0$ (n and p spins antiparallel), is a virtual state; it is unbound by $\sim 60 \text{ kev}$. The significance of this is that nuclear interaction is different for different S states, i.e., nuclear forces are spin-dependent.

From 22.111 Lecture Notes by S. Yip, Chap. III.

A. Bound State of the Deuteron

Experimentally it is known that the binding energy E_B of the deuteron is 2.23 Mev. The deuteron is the only stable bound system of two nucleons. We will see later that the di-neutron and di-proton are not stable. The energy E_B is known from the γ energy in the reaction



The inverse reaction of using electrons of known energy to produce external bremsstrahlung for (γ, n) reaction on H^2 also has been used. Besides the ground state no stable excited states of H^2 have been found. (There is a virtual state at ~ 2.30 Mev.)

Suppose we assume a square well potential (in 3-D),

$$V(r) = \begin{cases} -V_0 & r < r_0 \\ 0 & r > r_0 \end{cases}$$

Then we ask what is the level structure and what values should V_0 and r_0 take in order to be consistent with a bound state at energy $E_B = 2.23$ Mev?

*For a general discussion of nuclear forces, see Chapter 6 of Meyerhof.