

**Dismantling the Deep Earth: Geochemical Constraints from Hotspot Lavas for the Origin and Lengthscales of Mantle Heterogeneity**

by

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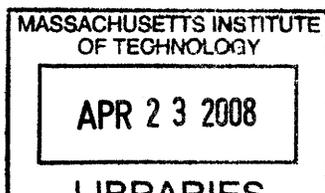
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*Tileinkað Önnu Mariu*



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# **Dismantling the Deep Earth: Geochemical Constraints from Hotspot Lavas for the Origin and Lengthscales of Mantle Heterogeneity**

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Submitted to the Department of Marine Geology and Geophysics,  
Massachusetts Institute of Technology-Woods Hole Oceanographic Institution  
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## **Abstract**

Chapter 1 presents the first published measurements of Sr-isotope variability in olivine-hosted melt inclusions. Melt inclusions in just two Samoan basalt hand samples exhibit most of the total Sr-isotope variability observed in Samoan lavas. Chapter 3 deals with the largest possible scales of mantle heterogeneity, and presents the highest magmatic  $^3\text{He}/^4\text{He}$  (33.8 times atmospheric) discovered in Samoa and the southern hemisphere. Along with Samoa, the highest  $^3\text{He}/^4\text{He}$  sample from each southern hemisphere high  $^3\text{He}/^4\text{He}$  hotspot exhibits lower  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios than their counterparts in the northern hemisphere. Chapter 2 presents geochemical data for a suite of unusually enriched Samoan lavas. These highly enriched Samoan lavas have the highest  $^{87}\text{Sr}/^{86}\text{Sr}$  values (0.72163) measured in oceanic hotspot lavas to date, and along with trace element ratios (low Ce/Pb and Nb/U ratios), provide a strong case for ancient recycled sediment in the Samoan mantle. Chapter 4 explores whether the eclogitic and peridotitic portions of ancient subducted oceanic plates can explain the anomalous titanium, tantalum and niobium (TITAN) enrichment in high  $^3\text{He}/^4\text{He}$  ocean island basalts (OIBs). The peridotitic portion of ancient subducted plates can contribute high  $^3\text{He}/^4\text{He}$  and, after processing in subduction zones, a refractory, rutile-bearing eclogite may contribute the positive TITAN anomalies.



## Introduction

Unlike the earth's crust, it is difficult to sample the mantle directly. However, there are a few isolated (and geographically limited) instances where we can observe the uppermost regions of the mantle. Ophiolites—portions of oceanic plates that have been uplifted and emplaced (obducted) onto the edge of continents—expose the uppermost regions of the mantle on the surface (Dick, 1976). Similarly, at certain, ultra-slow spreading mid-ocean ridges, mantle peridotites are emplaced on the surface (Dick et al., 2003), where they can be recovered with deep submersibles or dredging operations. Additionally, ultramafic xenoliths, which are pieces of the upper mantle entrained in upwelling magmas, allow direct inspection of the composition of the upper mantle (but unlike ophiolites and abyssal peridotites recovered by submersibles, ultramafic xenoliths provide little spatial context for study of the mantle). Thus, except for these few rare instances, we cannot examine the composition of the mantle directly: The deepest hole ever drilled, only ~12.3 km deep, took 22 years and untold Soviet resources to complete-- and got no where near mantle depths (unfortunately, the drill site was atop thick continental crust). While robots rove the surface of a planet, Mars, that is > 56,000,000 km distant, we have not directly observed our own planet at depths >12 km!

Nonetheless, there are indirect methods for evaluating the composition of the deeper earth. Mantle geochemists often use lavas erupted on the surface as “windows” to the composition of the mantle below. Hotspot lavas erupted in oceanic settings, or oceanic island basalts (OIBs), are formed by partially melting and upwelling, solid mantle. During melting, several radiogenic isotope systems and a number of trace element ratios remain unfractionated (or little fractionated, in the case of trace elements with similar compatibilities, at least when the partition coefficients are much less than the degree of melting) from the original, unmelted mantle. OIBs exhibit a great deal of isotopic and trace element heterogeneity, indicating that the mantle from which they were derived is also quite heterogeneous (Zindler and Hart, 1986). This fundamental observation leads to some of the most important questions in the field of mantle

geochemistry: How did the mantle become heterogeneous? At what scale lengths do the heterogeneities exist? How long do the heterogeneities survive?

One common paradigm in mantle geochemistry assumes that oceanic plates, which form by melting and depleting the upper mantle at mid-ocean ridges, are subducted back into the mantle from which they formed. Covered with a veneer of oceanic/continental sediment, subducted oceanic plates inject sediment, oceanic crust and depleted peridotitic mantle lithosphere into the mantle (Hofmann and White, 1982; White and Hofmann, 1982). In this way, mantle heterogeneities are born. Following storage in the mantle, subducted plates and sediment are thought to be sampled by upwelling plumes that melt and erupt lavas on the surface. However, a counteracting mechanism—chaotic mantle convection—stretches, thins, mixes and stirs, and thus homogenizes (or at least greatly attenuates) mantle heterogeneities on long timescales.

In Chapter 1, isotopic heterogeneities are explored at very short lengthscales in olivine-hosted melt inclusions in oceanic OIBs. Using a laser ablation system coupled to a MC-ICPMS (multi-collector inductively coupled plasma mass spectrometer),  $^{87}\text{Sr}/^{86}\text{Sr}$  was measured in olivine-hosted melt inclusions recovered from Samoan basalts. Complementing the pioneering work on Pb-isotopes in olivine-hosted melt inclusions from two Polynesian hotspots (Saal et al., 1998), significant Sr-isotope heterogeneity was also observed in the melt inclusions from individual Samoan basalt hand samples. Melt inclusions in one Samoan lava exhibit a range of  $^{87}\text{Sr}/^{86}\text{Sr}$  from 0.70686 to 0.70926. The isotopic diversity hosted in the melt inclusions from a single lava indicate that the size of the melting zone beneath a Samoan volcano can be larger than the lengthscales of mantle heterogeneities in the mantle upwelling beneath the hotspot. Furthermore, none of the 41 melt inclusions analyzed exhibit  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios lower than the least radiogenic whole-rock basalts in Samoa ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.7044$ ). This  $^{87}\text{Sr}/^{86}\text{Sr}$  data, combined with trace element data on the same melt inclusions, provide strong evidence against assimilation of oceanic crust as the source of the isotopic diversity in the melt inclusions.

Chapter 3 also deals with lengthscales of mantle heterogeneity, but by comparison to Chapter 1, Chapter 3 considers the largest possible lengthscales of heterogeneity in the Earth's mantle. The high  $^3\text{He}/^4\text{He}$  (or FOZO, Focus Zone; Hart et al., 1992) mantle

reservoir, a domain that is considered to be one of the oldest (and deepest?) reservoirs in the mantle, is the focus of chapter 3. In this chapter, the highest magmatic  $^3\text{He}/^4\text{He}$  ratios ever recorded in a southern hemisphere lava (33.8 times atmospheric) are reported in samples from the Samoan island of Ofu. These new measurements from Ofu Island place Samoa in the same category of high  $^3\text{He}/^4\text{He}$  hotspots as Hawaii, Iceland and the Galapagos. Along with Samoa, the highest  $^3\text{He}/^4\text{He}$  sample from each southern hemisphere high  $^3\text{He}/^4\text{He}$  hotspot exhibits lower  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios than their counterparts in the northern hemisphere (excluding lavas erupted in continental, back-arc, and submarine ridge environments). The observation of a large-scale isotopic enrichment in the FOZO-A (austral) high  $^3\text{He}/^4\text{He}$  mantle compared to the FOZO-B (boreal) high  $^3\text{He}/^4\text{He}$  mantle is similar to the DUPAL anomaly, a globe-encircling feature of isotopic enrichment observed primarily in southern hemisphere ocean island basalts. The possible existence of hemispheric-scale heterogeneity in one of the oldest reservoirs in the mantle has important implications for mantle dynamics. It suggests that regions of the (lower?) mantle have escaped the rapid convection motions that dominate the upper mantle. However, the origin of the hemispheric-scale heterogeneity in the FOZO (and DUPAL) reservoir is unknown.

Nonetheless, having defined the variability that exists in the high  $^3\text{He}/^4\text{He}$  mantle, Chapter 3 also explores whether or not the FOZO reservoirs are truly depleted, as is commonly suggested (Hart et al., 1992), or whether they have been re-enriched. The recent discovery of superchondritic  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios in terrestrial (Boyet and Carlson, 2005), martian and lunar (Caro et al., 2007) suggests that bulk silicate earth (BSE) may have superchondritic Sm/Nd ratios ( $^{147}\text{Sm}/^{144}\text{Nd} > 0.209$ , the minimum ratio necessary to generate the terrestrial mantle  $^{142}\text{Nd}/^{144}\text{Nd}$  anomaly relative to chondrites), and that the earth has a minimum  $^{143}\text{Nd}/^{144}\text{Nd}$  of 0.51304. If this is true, then the FOZO reservoirs are actually enriched relative to BSE.

Chapter 2 explores a common paradigm in mantle geochemistry, that subduction of marine/continental sediments can generate geochemically enriched mantle domains that can be sampled by mantle upwellings. In this chapter, remarkably high  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios are reported in submarine lavas recovered from the flanks of the Samoan island of

Savai'i. These Savai'i lavas exhibit the highest  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios reported for ocean island basalts to date. The isotope and trace element data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Importantly, Pb-isotopes in the most enriched Samoan lavas preclude contamination by modern-marine sediment. The ultra-enriched Samoan lavas have most certainly been "contaminated" by sediment, but the sediment is of an ancient (>> 200 Ma) origin and has been recycled into the Samoan mantle source. In summary, Chapter 3 provides the strongest evidence yet that the sediment that goes down in subduction zones does come back up in OIBs. However, given the large mass of sediment that has been subducted into the mantle over geologic history, it is still a mystery why clear signatures of sediment recycling are so rare in OIBs (Hofmann, 1997).

Chapter 4 reports evidence for radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  and enrichment in Titanium, Tantalum and Niobium (TITAN) in high  $^3\text{He}/^4\text{He}$  lavas globally. To explain these observations, the dominant paradigm for the formation of mantle heterogeneity is applied to the high  $^3\text{He}/^4\text{He}$  reservoir: Can the subduction of oceanic plates (crust and peridotite) generate the geochemical signatures associated with the high  $^3\text{He}/^4\text{He}$  reservoir? Radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  and TITAN enrichment are both geochemical signatures that are associated with recycled eclogite, suggesting that the high  $^3\text{He}/^4\text{He}$  lavas were derived from a mantle source hosting a recycled slab component. However, eclogites are quantitatively degassed in subduction zones and do not have intrinsically high  $^3\text{He}/^4\text{He}$ . None-the-less, the peridotitic portion of recycled slabs has been suggested to preserve high  $^3\text{He}/^4\text{He}$  over time (e.g., Parman et al., 2005). The eclogitic and peridotitic portions of subducted plates are intimately associated in space and time, a geometry that is conducive to later mixing in the mantle. Thus, together, the two lithologies can provide the "raw materials" for the formation of the high  $^3\text{He}/^4\text{He}$  mantle.

Importantly, the TITAN enrichment in high  $^3\text{He}/^4\text{He}$  mantle sampled by oceanic hotspot lavas may provide a clue about the location of the "missing" TITAN in the earth. Shallow geochemical reservoirs in the earth—continental crust and the depleted mid-ocean ridge basalt mantle (DMM)—have a shortage of the element Ti, Ta and Nb (TITAN) (McDonough, 1991; Rudnick et al., 2000). The observation of TITAN

enrichment in high  $^3\text{He}/^4\text{He}$  OIB lavas suggests that the mantle domain hosting the Earth's "missing" TITAN is sampled by deep, high  $^3\text{He}/^4\text{He}$  mantle plumes.

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## Chapter 1

# Strontium isotopes in melt inclusions from Samoan basalts: Implications for heterogeneity in the Samoan plume\*

### Abstract

We measured  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios on 41 olivine-hosted melt inclusions from nine Samoan basalts using laser ablation multi-collector (LA-MC) ICPMS.  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios are corrected for mass bias after eliminating major isobaric interferences from Rb and Kr. The external precision averages  $\pm 320$  ppm ( $2\sigma$ ) for the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios on natural Samoan basalt glass standards of a similar composition to the melt inclusions.

All of the Sr-isotope ratios measured by LA-MC-ICPMS on Samoan melt inclusions fall within the range measured on whole rocks using conventional methods. However, melt inclusions from two Samoan basalt bulk rock samples are extremely heterogeneous in  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.70459–0.70926), covering 70% of the variability observed in ocean island basalts worldwide and nearly all of the variability observed in the Samoan island chain (0.7044–0.7089). Seven melt inclusions from a third high  $^3\text{He}/^4\text{He}$  Samoan basalt are isotopically homogeneous and exhibit  $^{87}\text{Sr}/^{86}\text{Sr}$  values from 0.70434 to 0.70469.

Several melt inclusions yield  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios higher than their host rock, indicating that assimilation of oceanic crust and lithosphere is not the likely mechanism contributing to the isotopic variability in these melt inclusions. Additionally, none of the 41 melt inclusions analyzed exhibit  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios lower than the least radiogenic basalts in Samoa ( $^{87}\text{Sr}/^{86}\text{Sr}=0.7044$ ), within the quoted external precision. This provides an additional argument against assimilation of oceanic crust and lithosphere as the source of the isotopic diversity in the melt inclusions.

The trace element and isotopic diversity in Samoan melt inclusions can be modeled by aggregated fractional melting of two sources: A high  $^3\text{He}/^4\text{He}$  source and an EM2 (enriched mantle 2) source. Melts of these two sources mix to generate the isotopic diversity in the Samoan melt inclusions. However, the melt inclusions from a basalt with the highest  $^3\text{He}/^4\text{He}$  ratios in Samoa exhibit no evidence of an enriched component, but can be modeled as melts of a pure high  $^3\text{He}/^4\text{He}$  mantle source.

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## Strontium isotopes in melt inclusions from Samoan basalts: Implications for heterogeneity in the Samoan plume

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### Abstract

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*Keywords:*  $^{87}\text{Sr}/^{86}\text{Sr}$ ; laser ablation; MC-ICPMS; melt inclusion; Samoa; EM2; PHEM; FOZO

### 1. Introduction

Ocean island basalts (OIBs) erupted at hotspots are thought to be the surface expression of buoyantly upwelling mantle plumes that sample the mantle's compositional heterogeneities at various depths and times

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[1,2]. The Samoan islands and seamounts, formed by a mantle plume impinging on the Pacific plate just north of the Tonga Trench, form a time-progressive hotspot track [3,4] which conforms reasonably well to Morgan's hotspot model [5]. Samoan lavas exhibit the highest  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios and the largest  $^{87}\text{Sr}/^{86}\text{Sr}$  variation (0.7044–0.7089) measured in fresh OIBs [4,6], making them ideal for prospecting for diverse Sr-isotope compositions in melt inclusions.

Olivine-hosted melt inclusions in Samoan lavas provide snapshots of diverse magma chemistry before complete melt aggregation, providing an opportunity to see more of the isotopic heterogeneity which exists in the melt source but that is not detectable in whole rocks. However, the chemical variability in melt inclusions may be generated by a number of processes that obscure source variation, including pre-entrapment fractional crystallization, post-entrapment diffusive re-equilibration, crustal assimilation, and degree, type and depth of melting [7–16].

Studies delineating Pb-isotope diversity in melt inclusions have demonstrated that heterogeneous melt source compositions are an important factor in generating compositional variability [17–21]. A landmark Pb-isotope study of melt inclusions hosted in basalts from Mangaia Island in the Cook Islands revealed significantly more isotopic heterogeneity than is found in whole rocks from the island [17]. The results indicate the presence of an unradiogenic Pb-isotope endmember in the melt inclusions not discernable in whole-rock basalts. Problematically, this unradiogenic Pb endmember has been poorly characterized, owing partly to the large uncertainties associated with in situ Pb-isotope measurements: Pb-isotope data from melt inclusions generally are limited to  $^{208}\text{Pb}/^{206}\text{Pb}$  versus  $^{207}\text{Pb}/^{206}\text{Pb}$  isotope projections (due to the inability to collect precise  $^{204}\text{Pb}$  data on silicate melt inclusions), which place DMM (depleted MORB mantle, low  $^3\text{He}/^4\text{He}$ , low  $^{87}\text{Sr}/^{86}\text{Sr}$ ), FOZO (*F*ocus *Z*one, high  $^3\text{He}/^4\text{He}$ , low  $^{87}\text{Sr}/^{86}\text{Sr}$ ), PHEM (*P*rimitive *H*elium *M*antle, high  $^3\text{He}/^4\text{He}$ , middle-range  $^{87}\text{Sr}/^{86}\text{Sr}$ ) and EM2 (enriched mantle 2, low  $^3\text{He}/^4\text{He}$ , high  $^{87}\text{Sr}/^{86}\text{Sr}$ ) in such close graphical proximity that they cannot be unequivocally resolved. The true pedigree of the unradiogenic Pb endmember in Mangaia is still unknown, and could be similar to any of these four endmembers.

An advantage to the Sr-isotope system is that the EM2 endmember has dramatically higher  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios ( $\sim 0.7089$ ) [4] than the DMM (0.7026) [22], PHEM (0.7045) [23] and FOZO (0.7030) [24] mantle endmembers, and can be readily differentiated from the three less radiogenic endmembers. PHEM hosts signif-

icantly more radiogenic Sr than DMM and FOZO and is easily resolved from these two components. Unfortunately, DMM and FOZO exhibit similar  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios and it will be difficult to differentiate between these two components as potential sources of the isotopic diversity in melt inclusions.

We present Sr-isotope data from olivine-hosted melt inclusions recovered from Samoan basalts, some of which lie near the EM2 mantle endmember, with the goal of better understanding the puzzling unradiogenic component sampled by melt inclusions. To this end, we also contribute Sr-isotope data measured on melt inclusions from a recently discovered high  $^3\text{He}/^4\text{He}$  basalt from Samoa [25]. Our strategy is to analyze Sr isotopes in melt inclusions from EM2 and high  $^3\text{He}/^4\text{He}$  endmember basalts from Samoa, to constrain the role of the various components—EM2, PHEM, DMM and FOZO—that may be contributing to the Sr-isotope diversity in the Samoan plume.

## 2. Methods

A detailed description of the protocol used for in situ measurement of Sr isotopes in basaltic glasses (and melt inclusions) by LA-MC-ICPMS is provided in the Supplementary data. In order to measure Sr-isotope ratios in situ, we use a 213 nm NewWave laser ablation system coupled to a Thermo-Finnigan Neptune MC-ICPMS, located in the Plasma Facility at Woods Hole Oceanographic Institution (WHOI). During analytical runs, the laser is run in aperture mode with 100% power, a pulse rate of 20 Hz and a spot size of 120  $\mu\text{m}$ . The raster pattern varies depending on the size and shape of the melt inclusion, and the line speed is 4  $\mu\text{m}/\text{s}$ . Surface contamination is removed by pre-ablation using the same raster and spot size, but with a pulse rate of 5 Hz, 45% power and a raster speed of 30  $\mu\text{m}/\text{s}$ .

During each analytical session, we measure intensities on masses 82 through 88. Raw data are exported to an offline data correction program (TweaKr) for correcting the Rb and Kr isobaric interferences. Runs with low intensities (i.e., < 1 V on mass 88, due to small size or low Sr content) were discarded as they are prone to large systematic errors [26]. Masses 85 and 88 are pure Rb and Sr, respectively, with no significant known interferences, and require correction only for mass fractionation. We correct for Kr interferences on masses 84 and 86 so that the mass fractionation-corrected  $^{84}\text{Sr}/^{86}\text{Sr}$  value is canonical (0.0565725). The protocol for correcting mass 87 for the Rb interference is the following: A Samoan basalt glass with known  $^{87}\text{Sr}/^{86}\text{Sr}$ , from analysis by conventional Thermal Ionization Mass Spectrometry

Table 1  
Corrected and normalized  $^{87}\text{Sr}/^{86}\text{Sr}$  melt inclusion and whole-rock data, including Rb/Sr ratios and raw exported Neptune data (in volts)

Sample name and grain #	Disc #	Mass 82 meas'd-V	Mass 83 meas'd-V	Mass 84 meas'd-V	Mass 85 meas'd-V	Mass 86 meas'd-V	Mass 87 meas'd-V	Mass 88 meas'd-V	$^{85}\text{Rb}/^{87}\text{Rb}$ required	$^{87}\text{Sr}/^{86}\text{Sr}$ corr'd <sup>a</sup>	$^{87}\text{Sr}/^{86}\text{Sr}$ corr'd <sup>b</sup>	Int. prec.	Rb/Sr	Int. prec.	# Cycles	Type	Oliv Fo
AVON3-78-1											0.70890	0.1255					
(whole rock):																	
78-1Z	Year1	0.0013	0.0011	0.0193693	0.158083	0.263023	0.250903	2.25759	2.58966	0.70791	0.70784	87	0.0746	1.82	14 <sup>c</sup>	C	84.6
78-1#7 <sup>d</sup>	Disc1	0.0020	0.0022	0.0193323	0.164332	0.188415	0.199670	1.60867	2.58811	0.70927	0.70926	123	0.1081	2.47	9	H	84.6
78-1#9 <sup>d</sup>	Disc1	0.0019	0.0021	0.0178511	0.111482	0.163039	0.159876	1.38845	2.58811	0.70733	0.70732	103	0.0850	1.90	6	H	84.4
78-1#6 <sup>d</sup>	Disc2	0.0031	0.0034	0.0297219	0.341900	0.308400	0.354647	2.62064	2.58593	0.70846	0.70843	99	0.1397	2.63	32	H	83.6
78-1#9 <sup>d</sup>	Disc2	0.0031	0.0034	0.0316443	0.384820	0.340724	0.395010	2.89882	2.58593	0.70873	0.70870	94	0.1421	3.08	30	H	82.8
78-1#10 <sup>d</sup>	Disc2	0.0030	0.0031	0.0241696	0.094343	0.202069	0.179689	1.70410	2.58663	0.70845	0.70842	104	0.0593	3.21	16	H	83.6
78-1#4 <sup>d</sup>	Disc3	0.0049	0.0057	0.0638080	0.704962	0.756968	0.817352	6.43584	2.58694	0.70869	0.70866	44	0.1179	0.60	40	G	85.2
78-1inside	Disc3	0.0048	0.0049	0.0311385	0.155536	0.152840	0.165959	1.25071	2.58564	0.70689	0.70686	139	0.1337	2.20	16	C	84.3
78-1B	Disc3	0.0044	0.0043	0.0327454	0.189114	0.222802	0.230356	1.86119	2.58675	0.70791	0.70788	127	0.1089	5.76	10	C	84.2
78-1C	Disc3	0.0041	0.0043	0.0408646	0.414636	0.401444	0.448986	3.40453	2.58675	0.70866	0.70863	38	0.1304	1.96	24	C	83.9
78-1D	Disc3	0.0048	0.0048	0.0313504	0.130366	0.156852	0.158811	1.28485	2.58675	0.70699	0.70696	119	0.1091	3.79	14	C	85.3
AVON3-71-2											0.70594	0.0754					
(whole rock):																	
71-2C <sup>d</sup>	Year1	0.0015	0.0012	0.0241555	0.131881	0.348997	0.301171	2.99950	2.58977	0.70488	0.70481	50	0.0469	4.99	10 <sup>c</sup>	G	88.8
71-2C <sup>o</sup>	Disc3	0.0049	0.0051	0.0423813	0.151006	0.357484	0.309961	3.00928	2.58772	0.70487	0.70484	60	0.0539	2.09	19	G	88.8
71-2Da	Year1	0.0014	0.0012	0.0190549	0.143364	0.258958	0.241512	2.22256	2.58977	0.70555	0.70549	53	0.0687	3.92	12 <sup>c</sup>	C	90.1
71-2i_mil <sup>d</sup>	Disc1	0.0041	0.0042	0.0293387	0.151118	0.188888	0.190950	1.57393	2.58696	0.70605	0.70602	42	0.1028	0.89	19	G	90.7
71-2#6_sao1	Disc1	0.0040	0.0044	0.0392836	0.237282	0.399731	0.376067	3.39132	2.58675	0.70532	0.70529	63	0.0750	2.45	32	C	90.3
71-2#6_sao2	Disc1	0.0038	0.0039	0.0311684	0.151330	0.259691	0.241924	2.18859	2.58675	0.70511	0.70508	90	0.0741	3.15	15	C	90.5
71-2pentagon	Disc1	0.0035	0.0036	0.0274768	0.164455	0.221630	0.220515	1.86556	2.58653	0.70603	0.70600	69	0.0944	1.40	35	C	91.1
71-2#5 <sup>d</sup>	Disc3	0.0049	0.0055	0.0608667	0.440578	0.704362	0.672560	5.98503	2.58694	0.70537	0.70534	34	0.0792	1.23	40	G	91.0
71-2#5 <sup>o</sup>	Disc3	0.0049	0.0057	0.0600272	0.456791	0.688053	0.667289	5.84446	2.58694	0.70537	0.70534	34	0.0841	1.21	40	G	91.0
71-2#6 <sup>d</sup>	Disc3	0.0050	0.0053	0.0461581	0.261970	0.420138	0.399032	3.54513	2.58674	0.70602	0.70600	53	0.0794	1.35	40	G	85.7
71-2#6 <sup>o</sup>	Disc3	0.0050	0.0054	0.0511361	0.320998	0.514222	0.489758	4.35389	2.58674	0.70600	0.70598	32	0.0793	1.19	40	G	85.7
71-2Z	Disc3	0.0041	0.0042	0.0286629	0.120538	0.176062	0.169532	1.46336	2.58809	0.70576	0.70573	97	0.0882	2.60	21	C	86.1
71-2N	Disc3	0.0041	0.0043	0.0326106	0.171904	0.249420	0.242412	2.09517	2.58778	0.70554	0.70551	48	0.0879	2.74	26	C	90.5
71-2imi_run2	Disc3	0.0046	0.0048	0.0392641	0.162916	0.326152	0.292704	2.74335	2.58778	0.70520	0.70518	112	0.0638	2.80	12	C	91.0
71-2V	Disc3	0.0034	0.0037	0.0358138	0.196791	0.370738	0.339729	3.15166	2.58754	0.70461	0.70459	122	0.0668	4.32	13	C	87.4
OFU-04-06											0.70458	0.0458					
(whole rock):																	
OFU-04-06-13 <sup>d</sup>	Ofu	0.0012	0.0004	0.0123598	0.087599	0.202878	0.179279	1.74009	2.58774	0.70458	0.70457	226	0.0541	6.04	6	H	79.8
OFU-04-06-15 <sup>d</sup>	Ofu	0.0008	0.0006	0.0166800	0.116272	0.260130	0.231639	2.23575	2.58791	0.70443	0.70443	128	0.0556	1.62	13	H	77.3

OFU-04-06-25 <sup>d</sup>	Ofu	0.0009	0.0008	0.0272537	0.215485	0.450579	0.407078	3.87445	2.58797	0.70455	0.70454	60	0.0595	3.24	20	H	77.2
OFU-04-06-26 <sup>d</sup>	Ofu	0.0007	0.0006	0.0135577	0.076760	0.200599	0.173323	1.72195	2.58771	0.70443	0.70443	86	0.0476	2.73	16	H	76.5
OFU-04-06-28 <sup>d</sup>	Ofu	0.0007	0.0006	0.0138764	0.082936	0.207139	0.180426	1.77796	2.58777	0.70435	0.70434	164	0.0499	1.97	11	H	78.4
OFU-04-06-03 <sup>d</sup>	Ofu	0.0004	0.0005	0.0090473	0.057761	0.131245	0.115793	1.11700	2.58747	0.70443	0.70443	92	0.0560	2.67	15	H	83.8
OFU-04-06-02 <sup>d</sup>	Ofu	0.0007	0.0006	0.0126725	0.082453	0.187441	0.166696	1.61548	2.58588	0.70469	0.70469	184	0.0542	0.96	7	H	76.5
AVON3-71-11											0.70550		0.0700				
(whole rock):																	
71-11a <sup>d</sup>	Disc1	0.0035	0.0035	0.0250240	0.083017	0.152405	0.138888	1.27387	2.58654	0.70653	0.70653	83	0.0691	1.22	16	G	83.0
71-11a <sup>e</sup>	Disc1	0.0019	0.0021	0.0204900	0.125897	0.216562	0.204083	1.85257	2.58857	0.70692	0.70692	109	0.0720	0.92	15	G	83.0
71-11#1 <sup>d</sup>	Disc3	0.0048	0.0048	0.0316250	0.103649	0.160366	0.150752	1.31592	2.58772	0.70683	0.70680	144	0.0846	5.70	8	G	81.4
AVON3-68-3											0.70539		0.0688				
(whole rock):																	
68-3#6mi1 <sup>d</sup>	Disc1	0.0020	0.0021	0.0186133	0.117638	0.165627	0.163878	1.40909	2.58798	0.70625	0.70624	75	0.0884	3.01	20	G	87.9
68-3#6mi2	Disc1	0.0038	0.0040	0.0305189	0.209211	0.258789	0.264567	2.18215	2.58653	0.70577	0.70575	87	0.1027	1.92	12	C	87.9
68-3#6mi2 <sup>e</sup>	Disc1	0.0037	0.0041	0.0371899	0.294848	0.389761	0.392493	3.31142	2.58653	0.70577	0.70574	93	0.0954	1.73	9	C	87.9
68-3Ami	Disc3	0.0041	0.0042	0.0302365	0.092078	0.197293	0.173102	1.64414	2.58754	0.70482	0.70480	144	0.0600	6.45	10	C	90.1
AVON3-63-2											0.70540		0.0679				
(whole rock):																	
63-2mi	Disc3	0.0049	0.0051	0.0641890	0.514757	0.759082	0.741869	6.46111	2.58815	0.70599	0.70596	58	0.0856	2.68	31	C	-
63-2mi <sup>e</sup>	Disc3	0.0047	0.0056	0.0601701	0.544474	0.708353	0.718034	6.03171	2.58815	0.70616	0.70613	69	0.0969	1.22	36	C	-
63-2ami	Disc3	0.0044	0.0045	0.0283814	0.086324	0.138106	0.128251	1.13104	2.58815	0.70522	0.70520	181	0.0818	4.84	11	C	87.1
63-2bmi1	Disc3	0.0044	0.0045	0.0334163	0.133112	0.234231	0.215648	1.95929	2.58815	0.70528	0.70525	55	0.0728	3.15	23	C	86.6
AVON3-74-1											0.70469		0.0484				
(whole rock):																	
74-1mi	Year1	0.0015	0.0012	0.0174643	0.076749	0.224099	0.189719	1.91969	2.58895	0.70483	0.70477	157	0.0426	1.38	9 <sup>o</sup>	C	85.1
S11 (whole rock):											0.70620		0.0597				
S11mi	Disc3	0.0048	0.0048	0.0330920	0.078499	0.189324	0.161268	1.56507	2.58564	0.70520	0.70518	125	0.0539	6.13	11	C	76.5
AVON3-78-3											0.70889		0.1108				
(whole rock):																	
78-3 mi	Year1	0.0014	0.0012	0.0184433	0.207247	0.241204	0.254752	2.06790	2.58966	0.70814	0.70807	61	0.1068	3.51	18 <sup>o</sup>	C	84.1

All samples, except for OFU-04-06 and S11, are dredge samples from the AVON3 dredging cruise. Melt inclusion type refers to whether it was homogenized (H), glassy (G) or crystalline (C). Host olivine forsterite contents are given in the last column. Internal precision (Int. prec.) is standard error ( $2\sigma$ ) and is in ppm ( $^{87}\text{Sr}/^{86}\text{Sr}$ ) or % (Rb/Sr). Melt inclusion  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios are corrected for the isobaric interference from Rb by use of the bracketing method (see text for details). All melt inclusion analyses were performed on 01/27/04, 01/10/05, 01/11/05, 08/09/05 and 08/10/06 and corrected for the isobaric Rb interference using the Samoan glass standard analyses from those 5 days (see Supplementary data). Melt inclusion Sr-isotope and Rb/Sr data from a particular sample are listed immediately below the equivalent whole-rock data.

<sup>a</sup> Sr-isotopes are corrected for Kr and Rb interferences and fractionation-corrected.

<sup>b</sup> After correction for Rb and Sr interferences and fractionation, melt inclusions are corrected for 987 standard solution runs for each analytical session.

<sup>c</sup> Cycle length for melt inclusions from disc "Year1" are 16 s integrations during analysis. Cycles for all other measurements are 8 s integration.

<sup>d</sup> Has major and/or trace element data in Table 2.

<sup>e</sup> Replicate analysis of a melt inclusion is listed immediately below first analysis. Six melt inclusions were large enough to permit replicate analyses. Melt inclusion sample Year1\_71-2 was replicated over a 1-year period, and was renamed Disc3\_71-2C for the replicate analysis. All other replicate analyses were made during the same analytical session.

(TIMS), is measured by laser ablation, Kr-corrected, and the  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio is adjusted until the known TIMS  $^{87}\text{Sr}/^{86}\text{Sr}$  value of the glass is achieved. We adopted the strategy of using the required  $^{85}\text{Rb}/^{87}\text{Rb}$  of the Samoan basalt glass standards to bracket the  $^{85}\text{Rb}/^{87}\text{Rb}$  of the basalt glass unknowns. In order to estimate the overall accuracy associated with this technique, we apply contiguous bracketing of the glass standard runs, and we are able to reproduce the known TIMS values to within an average of  $\pm 320$  ppm ( $2\sigma$  standard deviation). However, due to uncertainty associated with the  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio ( $2.5875 \pm 0.00275$ ,  $2\sigma$ ), the final, corrected  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio exhibits an error magnification that is directly proportional to the Rb/Sr ratio of the sample. Samples with low Rb/Sr will exhibit less error from the Rb correction (145 ppm, Rb/Sr=0.04) than samples with high Rb/Sr (505 ppm, Rb/Sr=0.14) (see Supplementary data). However, over the range of Rb/Sr in the Samoan basalt glass standards (0.045–0.126), we find no relationship between the internal (in-run) precision of  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (which average 45 ppm,  $2\sigma$  standard deviation) and Rb/Sr during analyses of basalt glasses by laser ablation. Similarly, there does not appear to be a relationship between Kr/Sr and the internal precision or reproducibility (external precision) of the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in Samoan glass standards over the range of ratios that we have observed during melt inclusion analysis ( $^{82}\text{Kr}/^{88}\text{Sr}$  from 0.00013 to 0.004). Finally, the reproducibility of  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement does not appear to be related to Sr intensity over the range of Sr intensities observed in lasered Samoan glasses, a range that encompasses the melt inclusion analyses.

Six melt inclusions were large enough for replicate analysis (one melt inclusion, 71-2C, was replicated over a one-year period), and five of their  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios were reproducible within the quoted precision. However, the replicate analysis of melt inclusion 71-11a was different by 550 ppm (see Table 1), while error resulting from the Rb correction is only 260 ppm ( $2\sigma$ ) on Samoan glass standards with similar Rb/Sr ratios. The internal precision of the replicate analysis of this melt inclusion was  $\sim 100$  ppm ( $2\sigma$ ). Data from this melt inclusion indicates that larger-than-usual  $^{85}\text{Rb}/^{87}\text{Rb}$  variations over time can occasionally generate uncertainties (above the  $2\sigma$  level) in  $^{87}\text{Sr}/^{86}\text{Sr}$  that are somewhat larger than error predicted by the data from Samoan glass standards.

An upper limit for the  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement precision on Samoan melt inclusions with low Rb/Sr can be inferred from the near-uniform ratios obtained on melt inclusions from the high  $^3\text{He}/^4\text{He}$  Ofu basalt. The Rb/Sr values were among the lowest during analy-

sis of Ofu melt inclusions, and the tight clustering of the Ofu melt inclusions may be partially explained by decreased error of  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement for these samples compared to other, higher Rb/Sr Samoan glasses and melt inclusions from Vailulu'u and Malu-malu. If we assume that the Ofu melt inclusions are isotopically homogeneous, then the external precision on these 7 melt inclusions is  $\pm 335$  ppm ( $2\sigma$ ). Some of the apparent variability may be a result of error from the Rb correction, which is  $\pm 190$  ppm ( $2\sigma$ ) at Rb/Sr ratios of 0.053, and may not reflect true variability. Additionally, internal precision varied from 60 to 226 ppm ( $2\sigma$  standard error) on the seven Ofu melt inclusions.

Masses 85 and 88 represent pure Rb and Sr, respectively, so that fairly precise measurement of Rb/Sr ratios can be generated. After correcting for mass fractionation ( $\approx 1.5\%$ /amu), Rb/Sr ratios on Samoan basalt glasses measured by laser ablation are reproducible to 17% ( $2\sigma$ , compared to ratios obtained by XRF/ICP techniques on the same samples), and precise (1.7%,  $2\sigma$ ) during multiple runs on the same glass (see Supplementary data).

Major element compositions of glassy and homogenized melt inclusions were obtained with a JEOL-733 automated electron microprobe at the Massachusetts Institute of Technology using an electron beam with current of 10 nA and accelerating potential of 15 kV focused to a spot of 1–2  $\mu\text{m}$  in diameter for olivine analyses, and defocused to 10  $\mu\text{m}$  for glass analyses. Trace element contents were determined with a Cameca IMS 3f ion microprobe following the techniques described in [27,28]. A small beam (5  $\mu\text{m}$  diameter spot), combined with a high-energy filtering technique (80–100 eV window), was used to determine trace element concentrations. Precision for Sr, La, Zr, Y is estimated to be  $\pm 15\%$ , and  $\pm 20$ – $30\%$  for Ba, Nb and Rb. Homogenization of olivine-hosted melt inclusions was performed in a furnace at 1187–1220 C (depending on olivine composition) at 1 atm pressure for 5 min in a graphite capsule.

To correct for the effects of crystallization of olivine in the glassy and homogenized melt inclusions, we add equilibrium olivine to the melt inclusions in 0.1% increments until equilibrium with mantle olivine ( $\text{Fo}_{90}$ ) is achieved, assuming olivine-melt partitioning of Fe and Mg from [29]. Instead of correcting the melt inclusions to be in equilibrium with the host olivine, this correction scheme is chosen so that we can compare them to similarly corrected Samoan whole-rock lavas (after discarding data from the most evolved–MgO < 6.5 wt.%–whole-rock samples).

### 3. Results

#### 3.1. Sr-isotope variability in melt inclusions

Sr-isotopes were measured in melt inclusions from nine geochemically well-characterized basalt samples [4,25] from five islands and seamounts located along the Samoan hotspot track. Olivines ( $F_{0.76-91}$ ) with large ellipsoidal melt inclusions (50–250  $\mu\text{m}$  diameter) were separated from the basalt samples for melt inclusion exposure and isotopic analysis. Most of the melt inclusions were crystalline and usually contained dendritic clinopyroxene in a glassy matrix, with spinel and rare sulfide globules; some ( $\approx 5\%$ ) of the melt inclusions were glassy. The high  $^3\text{He}/^4\text{He}$  basalt sample OFU-04-06 was unique in that amphibole and apatite were common melt inclusion phases, and carbonate was also observed.

Olivines ( $F_{0.82-85}$ ) hosting melt inclusions were separated from Malumalu seamount dredge sample 78-1, a picrite with the highest  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio (0.7089) of any OIB [4]. Melt inclusion-rich olivines were also recovered from dredge sample 71-2 ( $F_{0.84-91}$ ) from Vailulu'u, a seamount that displays intermediate enrichment relative to the other Samoan islands and seamounts (bulk rock  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope ratios from 0.7052 to 0.7067,  $n=20$ ). From Ofu Island, olivines ( $F_{0.76-84}$ ) were recovered from ankaramite dike sample OFU-04-06, which exhibits the highest  $^3\text{He}/^4\text{He}$  measured in a Samoan basalt [25]. Lavas from Ofu are the most isotopically homogeneous of the volcanoes in the Samoan chain ( $^{87}\text{Sr}/^{86}\text{Sr}$  from 0.70444 to 0.70480,  $n=12$ ). Finally, a smaller number of olivines were separated for melt inclusion analysis from Vailulu'u dredge samples 71-11 ( $F_{0.81-83}$ ), 63-2 ( $F_{0.86-87}$ ) and 68-3 ( $F_{0.87-90}$ ), Malumalu dredge sample 78-3 ( $F_{0.84}$ ), Ta'u Island dredge sample 74-1 ( $F_{0.85}$ ), and Savai'i subaerial post-erosional sample S11 ( $F_{0.76}$ ).

The Sr-isotope data from melt inclusions in just three Samoan whole-rock samples (OFU-04-06, 78-1 and 71-2) define a broad array that encompasses the entire spectrum of Sr-isotope variability (0.70434–0.70926) recorded in Samoan basalts (Fig. 1 and Table 1). Eleven melt inclusions from 78-1 display  $^{87}\text{Sr}/^{86}\text{Sr}$  values of 0.70686–0.70926, and encompass over 30% of the isotope variability observed in the OIB mantle. Vailulu'u melt inclusions from dredge samples 71-2 ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70459\text{--}0.70602$ ,  $n=12$  melt inclusions) and 68-3 ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70480\text{--}0.70624$ ,  $n=3$ ) exhibit a smaller range of Sr-isotope values than 78-1. However, the magnitude of the Sr-isotope heterogeneity in the melt inclusions from these two samples is approxi-

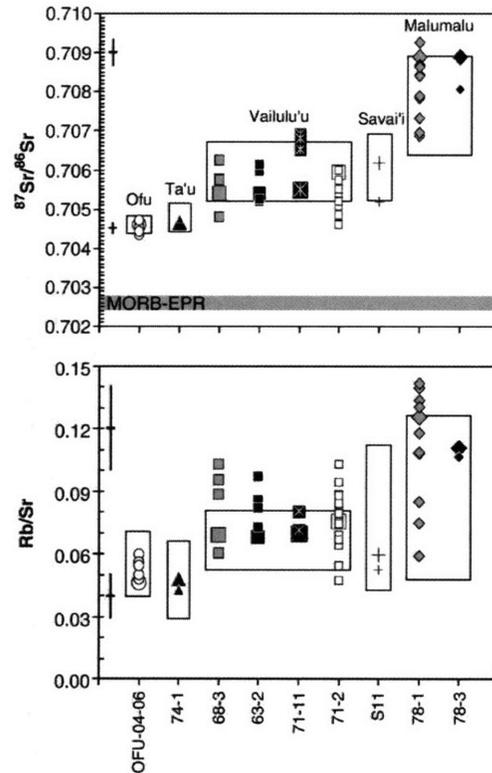


Fig. 1.  $^{87}\text{Sr}/^{86}\text{Sr}$  (upper panel) and Rb/Sr (lower panel) in melt inclusions from nine basalt samples compared to the total variation in each island/seamount (shown by boxes). References for whole-rock data can be found in [4]. The nine whole rocks from which the melt inclusions were separated are represented by larger symbols. No melt inclusions are more isotopically depleted than the most depleted lava measured in the island chain ( $^{87}\text{Sr}/^{86}\text{Sr} > 0.7044$ ). The range for EPR N-MORB [22] (averaged by segment) plots significantly below the Samoan whole rocks and melt inclusions. Replicate analyses on the same melt inclusion are averaged. The Rb/Sr for Savai'i whole-rock sample SAV1-25 is not included because of alteration. Internal precision ( $2\sigma$ , standard error) for  $^{87}\text{Sr}/^{86}\text{Sr}$  is approximately the size of the data symbols.  $^{87}\text{Sr}/^{86}\text{Sr}$  error bars represent error propagated from the Rb correction, as determined on Samoan glass standards, and dominates the error associated with  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement by LA-ICP-MS; maximum and minimum (500 and 150 ppm) errors are shown for reference. Error bars for Rb/Sr are  $\pm 17\%$  ( $2\sigma$ , standard deviation), and are based on the reproducibility of Rb/Sr measurement on Samoan glass standards; internal precision for Rb/Sr averages an order of magnitude better. Samples are listed in order of increasing whole-rock  $^{87}\text{Sr}/^{86}\text{Sr}$ , from left to right.

mately equal to the variability observed in Vailulu'u seamount lavas, although the melt inclusions sample a more depleted component than observed in whole rocks from this seamount. Among the Vailulu'u samples, only sample 71-11 ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70653\text{--}0.70692$ ,  $n=2$ )

hosts inclusions that sample compositions similar to the most enriched isotopic compositions observed in Vailulu'u volcano. Like sample 71-11, isotopic analyses of melt inclusions from sample 63-2 ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70520\text{--}0.70613$ ,  $n=3$ ) lie completely within the isotopic range found in the whole rocks from Vailulu'u seamount. OFU-04-06 melt inclusions exhibit the least isotopic variability ( $^{87}\text{Sr}/^{86}\text{Sr}=0.70434\text{--}0.70469$ ,  $n=7$ ) among the samples with isotopic analyses on more than two different inclusions, and this variability is similar to the variability sampled by whole-rock lavas from the island. There is only a single melt inclusion analysis from each of basalt samples 78-3 (0.70807), 74-1 (0.70477) and S11 (0.70518); the Sr-isotope ratios lie within the range observed in the respective host volcanoes.

Although melt inclusions show more depleted isotopic compositions than the island or seamount from which they were recovered, they are not observed to sample compositions more depleted than whole rocks from the Samoan hotspot ( $^{87}\text{Sr}/^{86}\text{Sr}\geq 0.7044$ ). Therefore, the least radiogenic Sr-isotope ratios in Samoan basalts and melt inclusions are significantly more enriched than the depleted upper mantle sampled by MORB along the EPR (East Pacific Rise; 0.70228–0.70287, N-MORB segment averages from [22]); the old oceanic crust upon which the Samoan island chain is constructed is probably isotopically similar to these modern EPR basalts.

Rb/Sr ratios measured by LA-ICP MS on the same set of melt inclusions tell a story similar to that of the isotopes. Melt inclusions in sample 78-1 exhibit the largest variation in Rb/Sr ratios (0.0593–0.1421), and the variability is similar to that observed in the whole rocks measured from Malumalu seamount. The Rb/Sr ratios in OFU-04-06 melt inclusions show some heterogeneity (Rb/Sr from 0.0476 to 0.0595,  $n=7$ ), but this variability is smaller than the variability sampled by whole-rock lavas from the island. Vailulu'u melt inclusions from dredge samples 71-2 (Rb/Sr from 0.0469 to 0.1028,  $n=12$  melt inclusions) and 68-3 (Rb/Sr from 0.060 to 0.1027,  $n=3$ ) exhibit a range of Rb/Sr values that falls between 78-1 and OFU-04-06. Unlike the isotopes, however, the magnitude of the Rb/Sr heterogeneity in the melt inclusions from samples 71-2 and 68-3 is greater than the variability observed in Vailulu'u seamount lavas, and, within error of measurement, do not sample a component with lower Rb/Sr than observed in whole rocks from this seamount.

The present dataset suggests that the isotopic variability exhibited by the melt inclusions in a basalt sample may be a function of the whole-rock isotopic

composition. In Fig. 2, the Sr-isotope variability of melt inclusions in three basalt samples—those with the largest number of melt inclusion analyses—is plotted against the bulk  $^{87}\text{Sr}/^{86}\text{Sr}$  composition of the respective whole rocks. The  $^{87}\text{Sr}/^{86}\text{Sr}$  variability, determined by the difference between the highest and the lowest  $^{87}\text{Sr}/^{86}\text{Sr}$

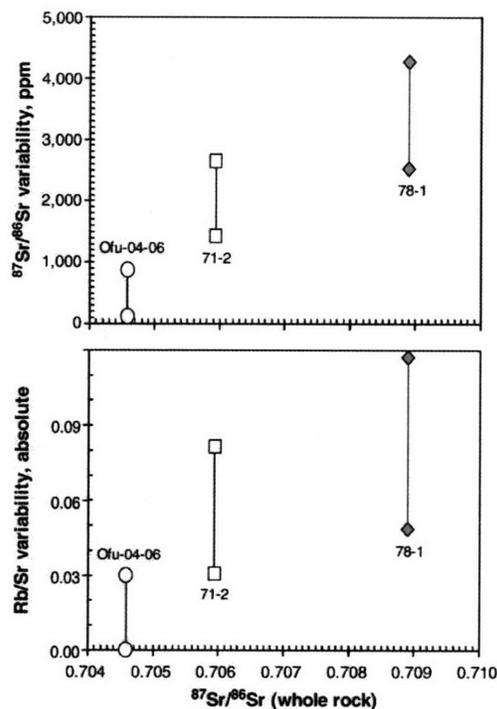


Fig. 2. Upper panel: Melt inclusion isotopic variability (in ppm) as a function of whole-rock  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio for the three Samoan lavas with the greatest number of melt inclusion analyses. Melt inclusion isotopic variability increases with increasing  $^{87}\text{Sr}/^{86}\text{Sr}$  (and decreasing  $^3\text{He}/^4\text{He}$ , not shown) in the three Samoan basalts. Variability is determined by the difference between the most isotopically extreme melt inclusions in a basalt sample. The maximum and minimum values for this variability are a result of the uncertainty introduced by the Rb correction, which is directly related to the Rb/Sr of the melt inclusion. For example, the maximum variability in basalt sample 78-1 is determined by the difference between the highest (0.7092 plus 391 ppm uncertainty) and lowest plausible  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7072 minus 393 ppm uncertainty). The minimum variability is determined by the difference between the lowest plausible  $^{87}\text{Sr}/^{86}\text{Sr}$  in the most enriched melt inclusion (i.e., 0.7089, or 0.7092 minus 389 ppm uncertainty) and the highest plausible  $^{87}\text{Sr}/^{86}\text{Sr}$  in the most depleted melt inclusion (0.7074, or 0.7072 plus 393 ppm uncertainty). Lower panel: Melt inclusion Rb/Sr variability (absolute) as a function of whole-rock  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio. Maximum and minimum variability is determined the same way, but assumes an uncertainty for Rb/Sr of 17%. The magnitude of variability due to internal precision is approximated by the size of the data symbols in both panels.

ratios from melt inclusions in a given basalt sample, exhibits a maximum and a minimum due to the uncertainty introduced by the Rb correction. Although the number of data points is limited, the data are consistent with melt inclusion isotopic diversity increasing with increasing  $^{87}\text{Sr}/^{86}\text{Sr}$  (increasing EM2 component) and decreasing  $^3\text{He}/^4\text{He}$  (not shown). The range of variability for Malumalu sample 78-1 is larger than, but overlaps with, the range of values from Vailulu'u sample 71-2. The melt inclusions from the high  $^3\text{He}/^4\text{He}$  Ofu basalt exhibit the smallest range of probable  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios, and they do not overlap with the range from samples 71-2 and 78-1. A similar observation can be made for the variability of Rb/Sr ratios in the melt inclusions, where Rb/Sr tends toward greater melt inclusion variability in samples 78-1 and 71-2; the Ofu sample has the smallest range of variability, and overlaps slightly with the lowest probable variability in sample 71-2. It is notable that OFU-04-06 exhibits the highest  $^3\text{He}/^4\text{He}$  ratio (and low  $^{87}\text{Sr}/^{86}\text{Sr}$ ) found in Samoa, an observation that may be linked to the small

degree of isotopic and trace element variability in its melt inclusions.

### 3.2. Major and trace element characteristics of melt inclusions

Rb/Sr and  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios from the melt inclusions define a broad array that encompasses the entire spectrum of Sr-isotope and Rb/Sr variability recorded in Samoan basalts (Fig. 3). Curiously, the melt inclusions from Malumalu do not form an array by themselves, but plot over a broad region. The Rb/Sr and Sr-isotope data array form a crude mantle isochron of 1.1 Ga.

Major and trace elements were measured on melt inclusions from Vailulu'u, Malumalu and Ofu basalts, and they reveal a large range of compositions (Table 2). Although the trace element compositions of Vailulu'u melt inclusions are similar to whole-rock analyses from this seamount, melt inclusions from EM2 endmember basalt 78-1 record a greater degree of trace element variability than all of the whole rocks measured from Malumalu. One

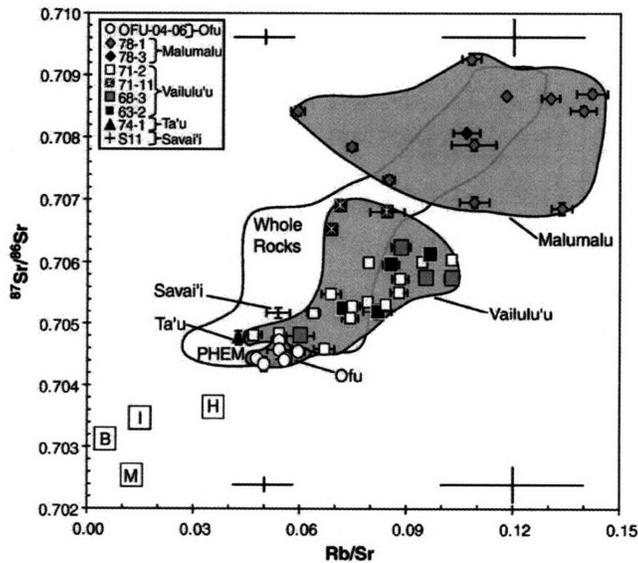


Fig. 3.  $^{87}\text{Sr}/^{86}\text{Sr}$  versus Rb/Sr ratios for Samoan whole rocks and melt inclusions, determined by laser ablation. The shaded regions represent the extent of melt inclusion variability for the volcanoes (Vailulu'u, Malumalu and Ofu). The open region encompasses the least evolved ( $\text{MgO} \geq 6.5$  wt.%) whole-rock measurements from these three volcanoes and Ta'u, thus encompassing islands and seamounts only on the eastern half of the Samoan hotspot track (where ~98% of the analyzed melt inclusions were recovered). Whole-rock data from Ofu is unpublished, and for the other volcanoes is from [4]. MORB range is limited to EPR N-MORB (M) segment averages [22]. High  $^3\text{He}/^4\text{He}$  basalts from Hawaii (H) [37], Iceland (I) [38] and Baffin Island (B) [39] provide an approximation for FOZO (higher  $^3\text{He}/^4\text{He}$  ratios exist for Hawaiian basalts [40], but  $^{87}\text{Sr}/^{86}\text{Sr}$  data are not available for these samples; Rb/Sr data for high  $^3\text{He}/^4\text{He}$  samples from Baffin Island are found in [41]). Error bars on symbols are internal precision of measurement ( $2\sigma$ , mean deviation). Error bars on periphery of figure denote external precision of measurements, as determined by Samoan glass standards with compositions similar to the melt inclusions: Rb/Sr error is 17% ( $2\sigma$ , standard deviation) and  $^{87}\text{Sr}/^{86}\text{Sr}$  error is based on error associated with the Rb correction ( $2\sigma$ , standard deviation). Replicate analyses of the same melt inclusion are shown. The melt inclusion data form a crude mantle isochron age of ~1.1 Ga.

Table 2  
Melt inclusion major and trace element data (uncorrected for olivine fractionation), and host olivine composition

Sample #	Disc #	Grain #	SiO <sub>2</sub>	TiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	FeO	MnO	MgO	CaO	Na <sub>2</sub> O	K <sub>2</sub> O	P <sub>2</sub> O <sub>5</sub>	Total	Mg #	OlFo	Rb	Sr	Y	Zr	Nb	Ba	La	<sup>87</sup> Sr/ <sup>86</sup> Sr <sup>a</sup>	Type
78-1	Disc1	#7	45.64	3.77	12.97	14.11	0.14	6.85	9.53	3.46	2.16	0.46	99.19	49.0	84.6	181	1176	83.6	979	–	–	–	0.70926	H
78-1	Disc1	#9	46.57	2.99	13.06	10.75	0.16	8.02	13.09	2.94	1.57	0.44	99.74	59.6	84.4	49	443	33.0	245	–	–	45.6	0.70732	H
78-1	Disc1	#10–11	44.98	3.22	12.84	11.18	0.14	9.26	11.29	2.62	1.76	0.39	98.10	62.1	84.0	75	507	28.1	269	–	–	–	–	H
78-1	Disc1	#12	46.37	3.76	13.09	11.51	0.17	8.03	11.38	2.93	1.91	0.40	99.65	58.0	84.4	59	545	33.2	339	73.1	390	–	–	H
78-1	Disc2	#5	45.16	3.42	12.62	11.09	0.17	9.15	12.42	2.46	1.42	0.42	98.68	62.0	84.3	130	1127	75.5	568	132.9	796	96.1	–	H
78-1	Disc2	#6	47.22	3.45	13.60	11.80	0.23	7.10	10.17	3.13	2.10	0.42	99.38	54.4	83.6	115	926	54.1	543	122.4	666	93.8	0.70843	H
78-1	Disc2	#7	47.20	2.92	13.65	9.94	0.14	7.83	11.36	2.73	2.00	0.40	98.49	60.9	83.9	–	–	–	–	–	–	–	–	H
78-1	Disc2	#8	45.57	3.20	12.26	11.20	0.17	9.26	11.33	2.71	1.33	0.47	97.88	62.1	83.6	71	679	40.6	360	76.0	409	57.7	–	H
78-1	Disc2	#9	46.81	3.46	14.03	11.81	0.16	7.34	10.54	2.93	1.91	0.43	99.59	55.2	82.8	76	655	38.7	400	84.8	482	72.3	0.70870	H
78-1	Disc2	#10	46.32	3.46	13.16	11.01	0.15	7.65	11.60	4.80	1.27	0.46	100.01	57.9	83.6	33	631	37.3	346	73.4	450	65.1	0.70842	H
78-1	Disc2	#11Le	46.32	3.24	12.74	11.04	0.14	8.18	12.72	3.18	1.33	0.39	99.67	59.5	83.6	55	598	36.7	281	56.1	401	45.5	–	H
78-1	Disc2	#11RtS	–	–	–	–	–	–	–	–	–	–	–	–	83.2	194	1694	100.5	871	193.5	1258	173.6	–	H
78-1	Disc2	#11RtL	44.61	3.77	12.99	11.76	0.19	6.23	13.20	2.70	1.57	0.44	97.85	51.2	83.2	299	2480	158.7	1310	293.5	1887	246.6	–	H
78-1	Disc2	#3	46.09	3.23	12.50	10.68	0.15	8.85	11.47	2.59	1.59	0.43	97.91	62.1	84.1	135	1038	62.5	579	126.8	719	100.6	–	H
78-1	Disc2	#4	46.45	2.99	12.17	10.32	0.15	9.22	12.92	2.37	1.37	0.40	98.80	63.9	83.9	104	901	55.1	465	110.9	686	78.9	–	H
78-1	Disc3	#4	46.07	3.81	13.22	10.47	0.12	6.00	12.52	2.93	1.80	0.42	97.81	53.1	85.2	69	555	26.4	297	64.8	370	55.4	0.70866	G
71-2	Disc1	g#3–4	48.50	2.66	12.55	9.41	0.14	7.38	14.63	2.33	0.84	0.35	99.13	60.8	87.4	26	384	23.0	185	30.4	124	22.4	–	G
71-2	Disc1	i	48.13	2.57	12.54	9.21	0.14	8.07	13.85	2.48	0.78	0.27	98.40	63.4	90.7	–	–	–	–	–	–	–	0.70602	G
71-2	Disc3	#3	46.12	3.48	13.06	10.94	0.15	5.81	14.09	2.55	1.29	0.32	98.23	51.3	84.3	50	517	29.5	237	52.7	347	40.9	–	G
71-2	Disc3	#5	46.71	3.10	12.72	9.66	0.15	7.90	13.10	2.45	1.24	0.32	97.72	61.8	91.0	45	504	27.1	243	58.8	314	35.3	0.70534	G
71-2	Disc3	#6	46.46	3.23	12.31	10.52	0.17	7.54	13.64	2.64	1.14	0.36	98.50	58.7	85.7	50	579	33.3	281	60.1	342	49.9	0.70599	G
71-2	Disc3	C	48.87	3.33	13.16	8.77	0.15	7.63	12.81	2.68	0.83	0.39	98.94	63.3	88.8	27	437	22.2	232	42.6	177	33.9	0.70482	G
71-11	Disc3	#1	49.01	3.13	12.53	11.05	0.18	6.15	12.01	2.85	0.94	0.40	98.51	52.4	81.4	34	419	27.1	232	41.4	205	35.6	0.70680	G
71-11	Disc1	a	49.28	2.82	12.70	10.94	0.14	6.89	11.63	2.60	1.01	0.37	98.73	55.5	83.0	36	409	30.3	223	40.9	214	33.0	0.70672	G
68-3	Disc1	#6mil	48.77	2.88	14.30	8.75	0.13	4.68	14.74	2.64	1.35	0.37	98.98	51.4	87.9	54	533	27.0	260	62.5	378	46.3	0.70624	G
OFU-04-06	Ofu	13	44.25	3.31	12.60	10.10	0.11	7.99	13.38	3.40	0.94	0.50	96.75	61.4	79.8	52	856	41.6	464	76.8	343	106.2	0.70457	H
OFU-04-06	Ofu	14	46.33	3.07	14.03	8.64	0.07	8.03	11.89	3.52	1.08	0.53	97.24	54.9	80.4	61	613	33.3	352	55.6	311	77.8	–	H
OFU-04-06	Ofu	15	46.47	2.81	13.67	10.58	0.14	7.20	10.64	3.96	1.09	0.33	96.96	60.5	77.3	60	700	30.0	374	54.8	303	82.1	0.70443	H
OFU-04-06	Ofu	16	46.74	3.13	14.12	7.00	0.09	8.26	13.11	3.76	1.19	0.45	97.97	49.5	81.9	68	760	38.6	393	72.5	363	91.1	–	H
OFU-04-06	Ofu	18	50.44	1.16	14.17	7.56	0.07	7.33	11.22	4.69	1.38	0.39	98.67	51.4	79.3	63	926	33.8	460	84.1	399	113.9	–	H
OFU-04-06	Ofu	19	47.15	2.14	11.15	17.06	0.21	13.83	2.72	1.56	1.71	0.46	98.00	75.2	76.6	–	–	–	–	–	–	–	–	H
OFU-04-06	Ofu	20	47.36	2.90	14.50	9.00	0.13	7.59	10.84	4.02	1.06	0.39	97.84	55.1	78.6	81	909	39.0	438	70.3	402	32.5	–	H
OFU-04-06	Ofu	24rep	46.35	3.58	13.47	8.58	0.09	8.11	13.70	3.23	0.72	0.29	98.13	55.8	83.0	–	–	–	–	–	–	–	–	H
OFU-04-06	Ofu	25	45.74	3.16	13.55	11.98	0.12	7.71	11.15	3.54	0.89	0.42	98.33	63.6	77.2	59	659	38.1	383	48.7	243	–	0.70454	H
OFU-04-06	Ofu	26	46.14	2.51	12.70	10.60	0.10	8.04	13.61	3.01	0.74	0.34	97.93	62.3	76.5	45	573	37.7	338	42.3	191	–	0.70443	H

OFU-04-06	Ofu	27	44.62	3.36	13.38	10.12	0.10	7.99	12.36	3.35	0.91	0.49	96.82	60.0	79.5	55	807	40.2	457	75.8	321	111.4	-	H
OFU-04-06	Ofu	28	47.60	2.70	13.87	9.87	0.08	7.49	10.92	4.18	0.94	0.33	98.13	58.5	78.4	58	744	34.6	382	71.1	322	94.3	0.70434	H
OFU-04-06	Ofu	30	47.23	3.03	13.82	7.23	0.09	7.78	12.39	3.74	1.15	0.24	96.79	50.9	82.3	63	634	32.8	324	54.0	283	76.6	-	H
OFU-04-06	Ofu	31	42.22	2.01	13.18	14.43	0.12	7.24	12.41	4.17	1.14	0.43	97.95	68.4	80.3	87	1006	46.6	515	79.4	445	136.2	-	H
OFU-04-06	Ofu	33	41.65	2.83	13.21	16.38	0.17	6.96	10.68	3.77	1.06	0.34	97.17	71.1	77.7	66	961	39.1	479	81.9	391	105.9	-	H
OFU-04-06	Ofu	2	46.18	3.00	13.70	9.93	0.13	7.53	11.26	3.70	0.99	0.32	96.80	58.9	76.5	50	800	35.5	409	81.5	326	80.9	0.70469	H
OFU-04-06	Ofu	3	45.75	3.12	13.20	7.91	0.10	8.20	13.67	3.01	0.83	0.37	96.24	54.3	83.8	41	714	36.8	398	72.2	301	72.4	0.70443	H
OFU-04-06	Ofu	4	45.71	3.05	13.40	10.95	0.06	7.62	10.88	3.93	0.93	0.33	96.89	61.8	77.2	56	772	33.6	420	73.2	326	70.2	-	H
OFU-04-06	Ofu	5	47.03	3.01	14.37	8.75	0.12	7.66	11.77	3.70	1.21	0.19	97.92	54.7	81.4	79	661	31.6	333	63.5	393	78.9	-	H
OFU-04-06	Ofu	8	45.83	3.21	13.50	10.96	0.13	7.74	11.19	4.06	0.98	0.36	98.04	61.6	77.7	63	861	40.0	462	81.2	363	102.5	-	H
OFU-04-06	Ofu	9	44.84	3.37	13.52	12.21	0.15	6.89	10.62	4.25	1.04	0.32	97.31	64.1	77.4	77	1006	42.7	528	98.4	444	104.7	-	H
OFU-04-06	Ofu	22	38.04	4.65	10.75	12.83	0.18	9.84	17.18	1.15	0.12	0.59	95.55	70.3	80.9	18	1161	72.4	821	121.9	255	38.9	-	H
OFU-04-06	Ofu	21	45.30	4.63	14.07	10.83	0.20	6.89	9.58	4.35	1.57	0.74	98.40	60.4	78.1	67	779	34.8	417	51.7	271	51.4	-	H
OFU-04-06	Ofu	23	45.58	5.87	12.35	8.58	0.12	8.17	13.00	3.71	1.32	0.84	99.47	57.9	80.6	85	1042	48.6	551	78.9	374	38.4	-	H
OFU-04-06	Ofu	7(air)	-	-	-	-	-	-	-	-	-	-	-	-	-	110	894	21.2	289	112.8	588	97.6	-	H
OFU-04-06	Ofu	7(sec)	-	-	-	-	-	-	-	-	-	-	-	-	-	65	809	37.5	428	74.1	365	104.7	-	H

Rb/Sr ratios by LA-ICP-MS are considered to be more accurate than those obtained by ion probe. Any discrepancy between the Rb/Sr ratios measured by the two methods is probably due to the difficulty of measuring Rb by ion probe techniques. Samples 78-1, 71-2, 71-11 and 68-3 are from the Samoan AVON3 dredging cruise, and their sample names are preceded by the prefix "AVON3." Major element data are in wt.%, and trace elements are in ppm. Forsterite contents of the host olivines (Ol Fo) are also given. Glassy (G) or Homogenized (H).

<sup>a</sup> See data in Table 1 for Sr-isotope data.

melt inclusion from sample 78-1 displays high Sr and Ba concentrations (2500 and 1900 ppm, respectively) that are 6 times more enriched than any whole-rock lavas examined from Malumalu. This ultra-enriched composition is significantly more enriched than observed in any Samoan basalts, and is derived from either an extremely enriched source or from very small degrees of melting.

Melt inclusions from OFU-04-06 also exhibit some unusual major and trace element compositions. One Ofu melt inclusion exhibits unusually low K<sub>2</sub>O concentrations. Additionally, the OFU-04-06 melt inclusions exhibit a large range in SiO<sub>2</sub>, including one sample with unusually low SiO<sub>2</sub> (38.0%) and high CaO (17.2%). These uncommon major element compositions may be due to the combination of unusual phases present in the OFU-04-06 melt inclusions prior to homogenization in the lab. Low glass totals in the OFU-04-06 inclusions are likely due to high volatile contents in the homogenized glasses, a hypothesis consistent with the volatile-rich phases in melt inclusions from this sample.

In addition to Rb/Sr ratios, several other major and trace element parameters correlate with <sup>87</sup>Sr/<sup>86</sup>Sr in the melt inclusions. Melt inclusions from Ofu, Vailulu'u and Malumalu exhibit negative Ba anomalies that also are observed in Samoan shield-stage lavas [4] (Fig. 4). An approximation for this Ba anomaly is (Ba/Nb)<sub>N</sub> (normalized to PUM, primitive upper mantle [30]), a ratio which correlates with <sup>87</sup>Sr/<sup>86</sup>Sr in whole rocks and melt inclusions (Fig. 5). The (Ba/Nb)<sub>N</sub> values are lowest (largest Ba-anomaly) in the basalts and melt inclusions with low <sup>87</sup>Sr/<sup>86</sup>Sr, and highest in the basalts and melt inclusions with elevated <sup>87</sup>Sr/<sup>86</sup>Sr ratios. It is notable that basalts and melt inclusions associated with elevated <sup>3</sup>He/<sup>4</sup>He ratios, and not the high <sup>87</sup>Sr/<sup>86</sup>Sr EM2 end-member basalts, have the largest Ba anomalies. The mechanism that generates this anomaly is unknown [4], and the anomaly also exists in MORB and HIMU basalts, but not in basalts with EM1 characteristics.

The negative K<sub>2</sub>O anomaly in the Samoan melt inclusions (Fig. 4) is commonly observed in OIBs. Curiously, however, PUM-normalized K exhibits little variability regardless of the degree of enrichment of the other trace elements (with the exception of a single Ofu melt inclusion). Despite the limited variability, olivine fractionation-corrected K<sub>2</sub>O concentrations exhibit a relationship with <sup>87</sup>Sr/<sup>86</sup>Sr in the melt inclusions (Fig. 5).

Nb/Zr also correlates with <sup>87</sup>Sr/<sup>86</sup>Sr in Samoan basalts and melt inclusions (Fig. 5). The lowest Nb/Zr ratios are associated with unradiogenic <sup>87</sup>Sr/<sup>86</sup>Sr values. Nb/Zr ratios correlate with Pb-isotopes [4] and inversely with <sup>3</sup>He/

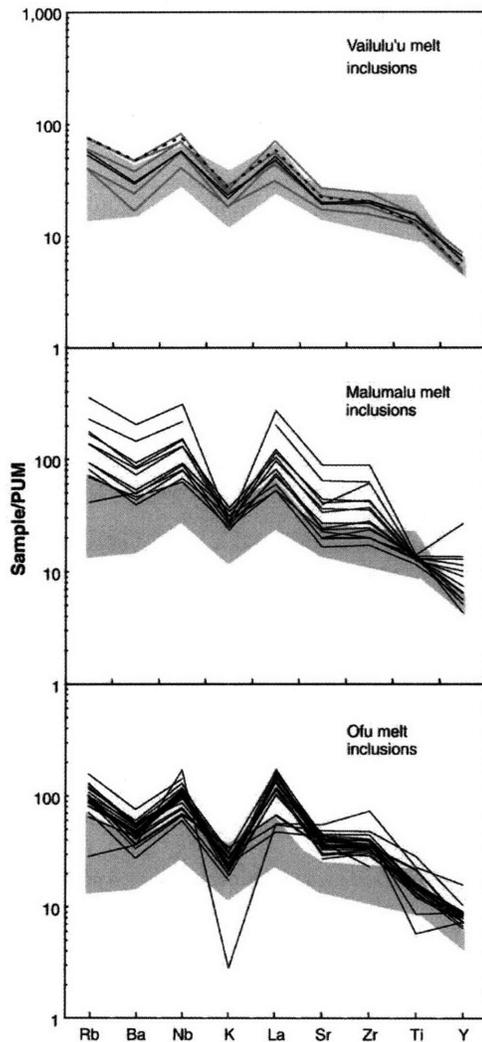


Fig. 4. PUM-normalized trace element patterns of Samoan basalts and melt inclusions; melt inclusion data is from ion probe analysis, except for K and Ti, which were measured by electron probe. Vailulu'u melt inclusions are from three samples: 71-2 (grey lines), 68-3 (dashed line), 71-11 (solid black lines). Malumalu and Ofu melt inclusions are from samples 78-1 and OFU-04-06, respectively. All melt inclusions (except for one melt inclusion from 78-1 and two from OFU-04-06 that have no major element compositions, see Table 2) are corrected for olivine fractionation to be in equilibrium with mantle olivine of  $F_{0.90}$  (see text for correction scheme). The grey field encloses the range of whole-rock patterns from the least evolved basalts (>6.5 wt.% MgO) from Vailulu'u, Ofu and Malumalu; whole-rock compositions have been corrected for olivine addition/fractionation. PUM values from [30].

$^4\text{He}$  (not shown) in Samoan basalts. Nb/Zr may serve as a proxy for  $^3\text{He}/^4\text{He}$  isotopes in Samoan melt inclusions; therefore, it may be important that the Nb/Zr ratios in several of the Ofu-04-06 melt inclusions exhibit values lower than the ratios observed in whole rocks.

### 3.3. Melting models of the EM2 source

It is important to place constraints on the trace element variability introduced by melting processes. A plot of Sr versus Ti/Zr shows that the majority of the melt inclusions form an array that extends outside of the whole-rock field to enriched Sr and low Ti/Zr values (Fig. 6). The low Ti/Zr ratios (15–88) observed in the melt inclusions extend to lower values than observed in whole-rock basalts from the eastern Samoa islands and seamounts (63–130). Such low values cannot be produced by crystal fractionation of melts before olivine entrapment, and assimilation of MORB (Ti/Zr=88) also fails to produce low Ti/Zr values [31]. A model of the Samoan EM2 source composition [4] provides a robust estimate for the mantle source sampled by the extreme EM2 basalt sample 78-1, and variable degrees of aggregated modal fractional melts of this source (Sr=20.0 ppm, Ti/Zr=101.9) in the garnet and spinel stability fields can describe much of the melt inclusion array in Fig. 6. Consistent with the model of the EM2 source as a metasomatized harzburgite [4], we adopt a harzburgite source lithology (1% spinel, 3.6% clinopyroxene [cpx], 20.6% orthopyroxene [opx] and 74.8% olivine), with mineral modes from [32] and mineral/melt partition coefficients from [33]. We assume the mineral modes of a similar bulk composition for melting in the garnet stability field (3.8% garnet, 2.7% cpx, 17.7% opx and 75.8% olivine) using the spinel to garnet conversion from [34]. The two melting models follow similar trajectories, but melting in the spinel stability field is required to generate the exceptionally high Sr and low Ti/Zr observed in the ultra-enriched Malumalu melt inclusions. Interestingly, if the EM2 source [4] has a more cpx-rich lithology than the harzburgite in our melting model, it will not produce melts with the Sr concentrations observed in the most enriched melt inclusion at reasonable degrees (>1%) of melting (Fig. 7).

Other geochemical indicators, including Y/Zr ratios, more clearly resolve the relative roles of melting in the garnet and spinel stability fields. Due to the relative compatibility of Y in residual garnet, low Y/Zr ratios are consistent with melting in the presence of garnet. A role for melting in the garnet stability field is suggested in a plot of Y/Zr against Nb/Zr in Fig. 6. This is particularly true for the Ofu melt inclusions, which trend to the lowest Y/Zr ratios

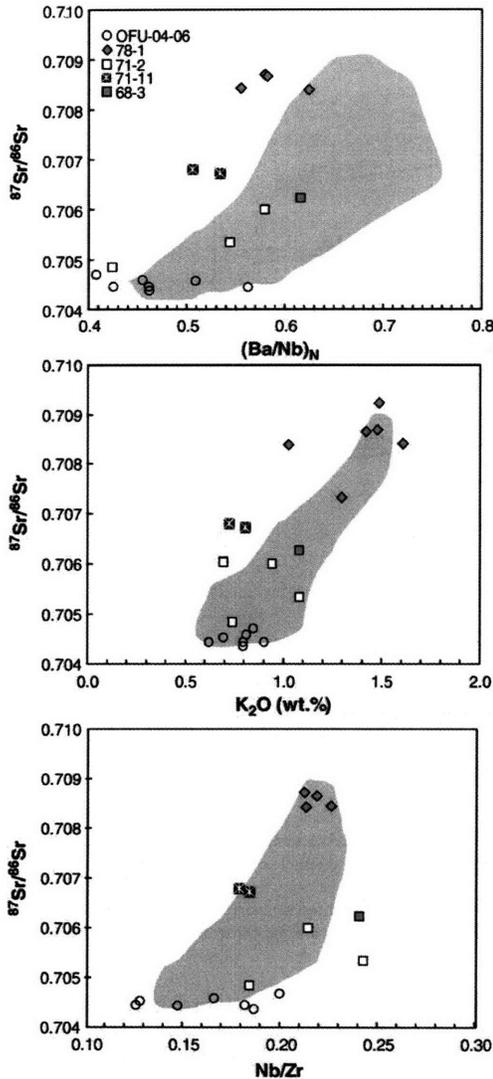


Fig. 5. Nb/Zr,  $K_2O$  and  $Ba/Nb_N$  versus  $^{87}Sr/^{86}Sr$ .  $K_2O$  measured by electron probe, and Nb, Zr and Ba by ion probe.  $(Ba/Nb)_N$  (normalized to PUM) is used as a proxy for the pervasive negative Ba anomaly in Samoan shield-stage basalts. The EM2 basalts and melt inclusions are associated with elevated Nb/Zr,  $K_2O$  and  $(Ba/Nb)_N$  (smaller negative Ba anomalies), and the low  $^{87}Sr/^{86}Sr$  basalts exhibit lower Nb/Zr,  $K_2O$  and  $(Ba/Nb)_N$  (larger negative Ba anomalies). All melt inclusions are corrected for olivine fractionation. The grey fields enclose the range of whole-rock patterns from the least evolved basalts (>6.5 wt.% MgO) from Vailulu'u, Ofu and Malumalu (whole-rock compositions have been corrected for olivine addition/fractionation). Symbols are the same as Fig. 3.  $^{87}Sr/^{86}Sr$  internal precision ( $2\sigma$ , standard error) is approximately the size of the symbol.

observed in the melt inclusions suite and straddle the garnet melting trend at 5% melt. This is similar to the Ti/Zr versus Sr plot, where the garnet melting curve trends through the Ofu melt inclusion field at ~4% melt. Malumalu and Vailulu'u melt inclusions (and Samoan whole rocks) are offset to higher Y/Zr ratios, perhaps suggesting a larger role for melting in the spinel stability field, an observation that is consistent with the same subset of melt inclusions in the Ti/Zr versus Sr melt model.

No single melting model of the EM2 source perfectly describes the melt inclusion fields for all three volcanoes, but we find that a combination of melting and mixing satisfactorily reproduces the melt inclusion geochemical variability. The relative roles of melting and mixing of different components can be partially deconvolved in a plot of  $^{87}Sr/^{86}Sr$  versus  $1/Sr$  (Fig. 8), where two-component mixing trajectories are linear and variable degrees of melting result in horizontal trajectories. The Ofu melt inclusions lie on a horizontal trend, which can be described by various degrees of melting of a single source that exhibits a  $^{87}Sr/^{86}Sr$  ratio of ~0.7045, and the Malumalu and Vailulu'u melt inclusions form a diagonal array that suggests a role for two-component mixing.

#### 4. Discussion

##### 4.1. A homogeneous source for PHEM basalts

Compelling evidence that the Samoan melt inclusions sample a heterogeneous source comes from Sr-isotope analysis of the melt inclusions from Vailulu'u and Malumalu basalts. However, the uniformly unradiogenic character of the Ofu melt inclusions precludes a significant contribution from an enriched, radiogenic (EM2) component. This suggests that Samoan melts with high  $^3He/^4He$  sample a homogeneous source and do not mix with melts of an enriched component. By comparison, the melt inclusions from Malumalu and Vailulu'u span a large range of Sr concentrations and isotopic compositions, indicating that both variable degrees of melting and mixing have occurred. An aggregated fractional melt trajectory for the garnet stability field is plotted (Fig. 8) for the EM2 source [4], and the Malumalu and Vailulu'u melt inclusions form a broad array that trends diagonally away from the horizontal Ofu melting trajectory toward low degree (~1%) melts of the EM2 source; the Vailulu'u and Malumalu melt inclusions can be produced by aggregated melts of an Ofu source that then mix with aggregated fractional garnet melts of the EM2 source.

Unlike the EM2 source, the trace element source composition of the high  $^3He/^4He$ , lower  $^{87}Sr/^{86}Sr$  Ofu source component is less clear. Called PHEM [23], this

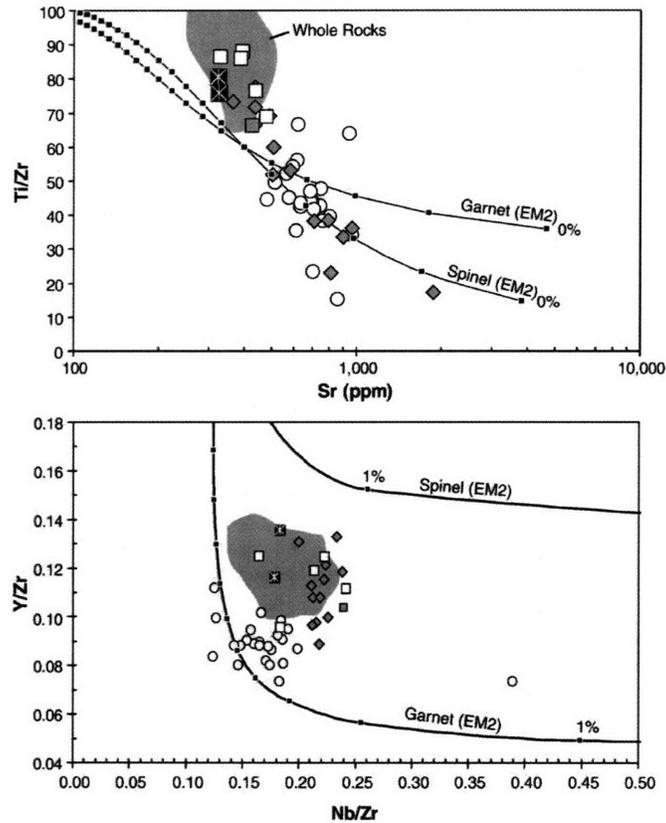


Fig. 6. Upper panel: Ti/Zr versus Sr concentration in melt inclusions and whole rocks with melting models. All melt inclusions and whole rocks are corrected for olivine fractionation to be in equilibrium with mantle olivine. Grey field is for whole-rock samples from Vailulu'u, Malumahu and Ofu. EM2 source [4] in the garnet and spinel stability fields, using partition coefficients from [33]. Harzburgite mineral modes are from [32]. All melting is non-modal. Tick marks are every 1% melting, beginning at 0% and increasing to the left. Lower panel: Y/Zr versus Nb/Zr in melt inclusions and whole rock, including melting models in the garnet and spinel stability fields. Tick marks are every 1% and the degree of melting increases to the left. Melting parameters and grey field are the same as upper panel.

component melts to form basalts and melt inclusions from Ofu. However, it is possible to bracket the source composition of this component and estimate the degree of melting captured in the Ofu melt inclusions. Although the Ofu basalts are more isotopically depleted than the EM2 basalts, an isotopically and trace element depleted DMM lherzolite source (7 ppm Sr) [35] fails to produce the high Sr concentrations observed in the Ofu melt inclusions (Fig. 7). This would suggest that the Ofu source is either more refractory or more trace element enriched, or both, than the lherzolitic [35] DMM source. The first option can be explored by invoking a more refractory, harzburgitic DMM source. However, only unreasonably low ( $F < 1\%$ ) degrees of melting can produce the most enriched Sr concentration observed an

Ofu melt inclusion. The second scenario can be tested by invoking the trace element enriched EM2 source, and assigning it a lherzolitic lithology that is similar to DMM. At reasonable degrees of melting ( $F = 1.5\%$ ), such a source can generate melts with sufficiently high Sr contents to match the range observed in Ofu. Finally, a harzburgitic EM2 source, which is both more refractory and trace element enriched than DMM, can generate the most enriched melt inclusions from Malumahu and Ofu between 1% and 2% melting. The harzburgitic EM2 source can serve as a probable upper limit for the trace element enrichment of the PHEM source because we consider it unlikely for the less isotopically enriched PHEM component to exhibit greater trace element enrichment than EM2. However,

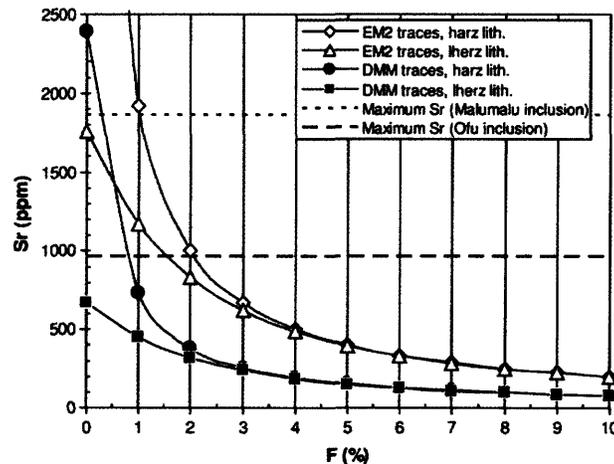


Fig. 7. Sr in melt versus degree of melting ( $F$ ) for various combinations of two lithologies and two trace element source abundances. Two dashed lines represent olivine fractionation-corrected Sr concentrations in the most enriched melt inclusions from Makumalu sample 78-1 (upper line) and Vailulu'u sample 71-2 (lower line). DMM trace element source and lherzolite lithology from [35], converted to equivalent garnet facies lithology using relationship from [34]. EM2 source from [4]. Harzburgite lithology from [32]. Note that only the EM2 source with a harzburgite lithology can produce the highest Sr concentration observed in the melt inclusion from sample 78-1 at  $F \geq 1\%$ . All melting is non-modal in the garnet stability field. Neither modal melting nor melting in the spinel stability field significantly changes our conclusions at  $F \geq 1\%$ .

a PHEM source with a lithology (and/or Sr concentrations) intermediate between DMM and EM2 cannot be ruled out. If PHEM has the same Sr content and lithology as the EM2 source, the Vailulu'u and Makumalu melt inclusion arrays can be explained as mixtures of  $6.5 \pm 1.5\%$  melts of PHEM and  $\sim 1\%$  melts of EM2. However, the degree of melting of the PHEM source that contributes to the Makumalu–Vailulu'u mixing array should be taken as a maximum.

Unlike the Ofu melt inclusions, which sample a pure PHEM source, no melt inclusions sample a pure EM2 melt (as calculated by [4], using ultra-enriched Sr-isotope compositions from Samoan xenoliths [36]). Mixing lines in Fig. 8 between a 1% EM2 melt and  $6.5 \pm 1.5\%$  PHEM melts indicate that the subset of Vailulu'u and Makumalu melt inclusions with both  $^{87}\text{Sr}/^{86}\text{Sr}$  and Sr concentration data are dominated by a PHEM component, and exhibit less than a  $\sim 30\%$  contribution from the EM2 component. However, one Makumalu melt inclusion exhibits a much larger contribution from an EM2 component, as indicated by its high ( $\sim 1865$  ppm) Sr content and low ( $\sim 17$ ) Ti/Zr ratio (Fig. 6). The PHEM–EM2 melt mixing lines in Fig. 8 suggest that this melt inclusion contains more than a 70% contribution from the EM2 melt component. The mixing model suggests that a 70% contribution from an EM2 melt would produce an extrapolated  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of  $\sim 0.712$ , which is a significantly higher ratio than has been observed in a Samoan basalt, but close to the  $^{87}\text{Sr}/$

$^{86}\text{Sr}$  of cpx in metasomatized xenoliths from Savai'i in western Samoa. Unfortunately, this ultra-enriched melt inclusion was too small for isotopic analysis.

#### 4.2. Isotopic variability in Samoan melt inclusions: MORB or FOZO?

Correlations between trace elements and Sr isotopes suggest that trace element variability in Samoan basalts and melt inclusions may reflect heterogeneity in the Samoan mantle. Rb/Sr,  $\text{K}_2\text{O}$ ,  $(\text{Ba}/\text{Nb})_N$  and Nb/Zr ratios correlate with Sr isotopes in Samoan lavas and melt inclusions, suggesting that these trace elements are heterogeneous in the Samoan mantle source. However, a combination of melting processes and variable source lithology may drive the observed correlations, limiting the role for source heterogeneity. For example, the relatively constant  $\text{K}_2\text{O}$  concentrations and negative (PUM-normalized) anomalies in Samoan lavas may be a result of residual phlogopite, which may cause  $\text{K}_2\text{O}$  to behave more compatibly in the Samoan source during melting. However,  $\text{K}_2\text{O}$  correlates with  $^{87}\text{Sr}/^{86}\text{Sr}$  in the whole rocks and melt inclusions (Fig. 5), suggesting a role for  $\text{K}_2\text{O}$  heterogeneity in the Samoan source, and that  $\text{K}_2\text{O}$  concentrations in Samoan basalts and melt inclusions may be controlled only partially by melting processes.

Assuming that Rb/Sr variability in Samoan melts reflects source variability, the array formed by the  $^{87}\text{Sr}/$

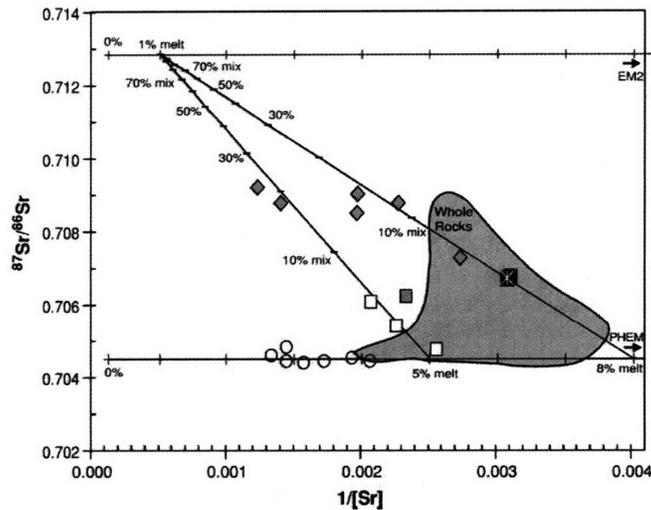


Fig. 8.  $^{87}\text{Sr}/^{86}\text{Sr}$  versus  $1/[\text{Sr}]$  for Samoan melt inclusions and basalts. Malumalu and Vailulu'u melt inclusions form a quasi-linear array extending away from the whole-rock basalts from Malumalu, Ofu and Vailulu'u (grey field, same as Fig. 6). Ofu melt inclusions fall off this array, plotting at lower  $^{87}\text{Sr}/^{86}\text{Sr}$  at a given Sr concentration. Aggregated non-modal fractional melting trends of the EM2 and PHEM sources in the garnet stability field are shown as horizontal lines (tick marks are every 1% melting, beginning at 0% and increasing to the right). Two mixing lines, both extending diagonally from a 1% EM2 melt to 5% and 8% PHEM melts, are shown (each tick represents 10% mixture, starting at 0% EM2 and increasing to the upper left). Ofu melt inclusions lie on the melting trend for PHEM, and require no mixing with an EM2 component. Malumalu and Vailulu'u melt inclusions can be described by mixing of melts from the EM2 and PHEM sources. Sr concentrations in melt inclusions and whole rocks are corrected for olivine fractionation. PHEM and EM2 source compositions plot outside of the figure as indicated, at  $1/\text{Sr}$  values of  $\sim 0.05$  (Sr concentrations of 20.0 ppm). The published source for EM2 [4] is used in this figure. As a limit, PHEM is given the trace element source abundances and lithology of EM2 in the melt model. Symbols same as in Fig. 3.

$^{86}\text{Sr}$  and Rb/Sr data can be modeled as binary mixing between a PHEM and an EM2 component, a model originally proposed by Farley et al. [23]. Such a model is consistent with the trace element melting/mixing model proposed above, which suggests that the Ofu melt inclusions sample only a high  $^3\text{He}/^4\text{He}$  PHEM mantle reservoir and the Vailulu'u and Malumalu melt inclusions result from mixing melts from both the PHEM and EM2 mantle reservoirs.

Previous work on the Pb-isotopic variability in melt inclusions [17] suggested that the unradiogenic end-member in EM2 basalts from Tahaa may be depleted (MORB or FOZO?), and thus lie at even lower  $^{87}\text{Sr}/^{86}\text{Sr}$  values than observed in the high  $^3\text{He}/^4\text{He}$  PHEM lavas from Ofu (Fig. 3). In fact, an extrapolation of the Samoa melt inclusion and whole-rock basalt array in Fig. 3 does indeed trend toward one of two depleted components that are significantly less radiogenic than PHEM: MORB [22] or a common high  $^3\text{He}/^4\text{He}$  mantle component, called FOZO [24] (Fig. 3). The FOZO component is represented by basalts with the highest  $^3\text{He}/^4\text{He}$  from Hawaii [37], Iceland [38] and Baffin Island [39]. Both MORB and FOZO lie on a similar extension of the

Samoa melt inclusion  $^{87}\text{Sr}/^{86}\text{Sr}$ -Rb/Sr array, so it is difficult to distinguish which, if either, of these two components is sampled by the Samoan melt inclusions. If the depleted component is MORB, it may be entrained in melt inclusions by shallow anatexis due to preferential cooling and olivine crystallization near magma chamber and conduit walls [7,8,21]. However, the presence of FOZO (or any other high  $^3\text{He}/^4\text{He}$  component) in the Samoan melt inclusions would require that the isotopic variability in melt inclusions reflect true source heterogeneity, assuming that a high  $^3\text{He}/^4\text{He}$  component does not exist in the oceanic crust or lithosphere.

#### 4.3. The case against MORB

It may be possible to look at other lines of geochemical evidence to discern whether MORB or FOZO play a role in augmenting the isotopic diversity in Samoan melt inclusions. Models suggesting that melt inclusion isotopic variability is caused only by contamination from unradiogenic oceanic crust and lithosphere at shallow levels do not explain how several Samoan melt inclusions have higher  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios than their host

bulk rock compositions (see Fig. 1). For example, Samoan whole-rock basalt sample 71-11 hosts several melt inclusions that have  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (up to 0.70692) that are significantly more enriched than its host rock (0.70550). The presence of  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in melt inclusions that are higher than the bulk rock require that at least some of the isotopic variability present in melt inclusions is derived from the mantle source, because the enriched component in melt inclusion 71-11a is too enriched to be found in the oceanic crust and lithosphere. Therefore, if assimilation of oceanic crust and lithosphere contributes heterogeneity to the Samoan melt inclusions, it cannot be the only means by which isotopic heterogeneity is produced in Samoan melt inclusions, and some contribution from the melt source must be involved as well.

On a different tack, the case for the less radiogenic melt inclusions sampling the depleted oceanic crust and lithosphere by assimilation is limited severely by the observation that, within analytical uncertainty, not a single melt inclusion has an  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio that is lower than the least radiogenic ( $^{87}\text{Sr}/^{86}\text{Sr}=0.7044$ ) whole-rock basalt measured in the Samoan islands (Fig. 1). The Samoan melt inclusions trace out a range of Sr-isotope variability that is confined to the region of Sr-isotope space defined by the Samoan whole-rock data (Fig. 3). On an island-by-island basis, the interpretation is more complicated, as melt inclusions from two Vailulu'u whole-rocks sample a component more depleted than found in whole rocks measured from the seamount. However, the least radiogenic component found in Vailulu'u melt inclusions is also found in lavas from nearby Samoan islands (e.g., Ta'u and Ofu), exhibits elevated  $^3\text{He}/^4\text{He}$  ratios, and is thus known to exist in the Samoan plume. Many of the downstream Samoan seamounts also are dominated by  $^{87}\text{Sr}/^{86}\text{Sr}$  between 0.7044 and 0.7049 [3]. It seems unnecessary, therefore, to invoke contamination from the oceanic crust and lithosphere to explain the presence of the less radiogenic component when it already exists inside the plume! Although the argument can be made that an insufficient number of melt inclusions have been analyzed to detect a component more depleted than what is found in whole rocks, the number of melt inclusions analyzed for Sr isotopes ( $n=41$ ) is already significant, and is equal to ~30% of the number of published  $^{87}\text{Sr}/^{86}\text{Sr}$  whole-rock analyses from the Samoan hotspot.

The high  $^3\text{He}/^4\text{He}$ , unradiogenic Sr component in Samoan basalts (PHEM) is unique in that it exhibits  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios more enriched than in the high  $^3\text{He}/^4\text{He}$  Hawaii, Iceland or Baffin Island basalts,

suggesting that the high  $^3\text{He}/^4\text{He}$  reservoir in the mantle is at least mildly heterogeneous in  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios. The same line of reasoning that precludes the presence of entrained melts from oceanic lithosphere in the Samoan melt inclusions also minimizes the possibility that a traditional, depleted FOZO-like component ( $^{87}\text{Sr}/^{86}\text{Sr}=0.7030$  [24]) serves as the unradiogenic Sr component: If melt inclusion diversity were a result of entrainment of a component (MORB or FOZO) more depleted than found in Samoan basalts, then the melt inclusions would extend to  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios lower than found in whole rocks (0.7044). However, the high  $^3\text{He}/^4\text{He}$  Samoan sample OFU-04-06 defines the lowest  $^{87}\text{Sr}/^{86}\text{Sr}$  portion of the Samoan whole-rock mixing array (see Fig. 3) and the melt inclusions are identical to the whole rock, suggesting that the least radiogenic Sr composition sampled by the Ofu basalts is the same component found in the melt inclusions. Therefore, we maintain that the unradiogenic Sr component in Samoan melt inclusions is more enriched than MORB or FOZO, and is likely the same PHEM component sampled by the high  $^3\text{He}/^4\text{He}$  Samoan basalts, suggesting that a two-component EM2–PHEM mixing model may be the most appropriate for melt inclusions originating in the enriched Samoan mantle. Scatter around such a mixing model (Fig. 8) may be due to minor contributions from other components [4] that may exist in the Samoan mantle.

#### 4.4. Implications for source heterogeneity (or lack thereof)

The results for Sr-isotope measurements in Samoan melt inclusions support an argument for an origin of the isotopic variability in the melt source, not contamination by oceanic crust and lithosphere. We assume that the isotopic variability (or lack thereof) in Samoan melt inclusions is not a product of variable degrees of homogenization in magma conduits and chambers before olivine entrapment, but rather that the isotopic variability in melt inclusions reflects the heterogeneity of the melt source: When the melt source is heterogeneous, melt inclusions capture the range of heterogeneity while the isotopic composition of the bulk rock lava represents an average of the heterogeneity sampled in the melt. By extension, we infer that the high  $^3\text{He}/^4\text{He}$  whole-rock sample OFU-04-06 tends to sample a more homogeneous source, as the melt inclusions are nearly isotopically homogeneous and identical to the bulk rock. Perhaps, then, only melting of a pure PHEM source allows the high  $^3\text{He}/^4\text{He}$  composition to persist in the

Ofu whole-rock lavas. Vailulu'u and Malumalu melt inclusions are more isotopically heterogeneous, and thus are inferred to sample a heterogeneous source that captures much of the mixing spectrum between the EM2 and PHEM components. The contribution of an EM2 component may explain the diminished  $^3\text{He}/^4\text{He}$  composition in the lavas from these two volcanoes. Although  $^{87}\text{Sr}/^{86}\text{Sr}$  analyses of the melt inclusions from high  $^3\text{He}/^4\text{He}$  basalts from other localities are not yet available, perhaps the high  $^3\text{He}/^4\text{He}$  mantle that these basalts sample is homogeneous and devoid of enriched domains. This hypothesis is consistent with the melting–mixing model above (see Fig. 8), which suggests that the isotopically homogeneous melt inclusions from the high  $^3\text{He}/^4\text{He}$  basalt from Ofu exhibit no evidence of mixing with an EM2 component.

It is notable that while near-pure PHEM melts are observed in Samoan melt inclusions, pure EM2 melts ( $^{87}\text{Sr}/^{86}\text{Sr}=0.7128$  [4,36]) were not unequivocally detected in this study (i.e., by measurement of Sr isotopes). Several melt inclusions with ultra-enriched trace element patterns were observed, but most of the melt inclusions are composed of <30% EM2 component. Perhaps this indicates that, compared to the PHEM component, the EM2 component in the Samoan plume is rare. Alternatively, the EM2 component may not be rare in the Samoan plume, but is more refractory and produces less melt than the PHEM component. In this way, perhaps, EM2 melts are less frequently sampled by melt inclusions. Future work on melt inclusions will help resolve the relative contributions of the enriched and high  $^3\text{He}/^4\text{He}$  sources to OIB lavas.

## 5. Summary

The following conclusions can be drawn from this study:

- 1.) The Sr-isotopic diversity in melt inclusions from Samoan basalts does not extend significantly above or below the range defined by whole rocks from the Samoan hotspot.
- 2.) A few melt inclusions exhibit  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios significantly higher than their host whole rock. This is taken as evidence that assimilation of MORB lithosphere cannot be the only mechanism that contributes isotopic diversity to the melt inclusions.
- 3.) The  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of the high  $^3\text{He}/^4\text{He}$  basalt is essentially indistinguishable from the Sr-isotope ratios (0.7044) measured in its melt inclusions, and the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in melt inclusions from

other Samoan basalts do not exhibit ratios lower than 0.7044. This observation is consistent with the hypothesis that a high  $^3\text{He}/^4\text{He}$  component, not MORB, is the unradiogenic Sr endmember in Samoan melt inclusions.

- 4.) Melt inclusions from a high  $^3\text{He}/^4\text{He}$  Samoan basalt (with less radiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$ ) are isotopically more homogeneous than the melt inclusions from basalts with higher  $^{87}\text{Sr}/^{86}\text{Sr}$  (more contribution from an EM2 component). This may indicate that the Samoan high  $^3\text{He}/^4\text{He}$  basalts sample a source that is more isotopically homogeneous than the source that produces basalts with an EM2 component.

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## Appendix A. Supplementary data

Supplementary data associated with this article can be found, in the online version, at doi:10.1016/j.epsl.2006.02.040.

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**Appendix A. Supplementary data** (as it appears, published, on the Elsevier website)

**Background Datasets Text,Figs&Tables**

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## Supplementary Data (Jackson and Hart, 2006)

### Sr isotopes by laser ablation PIMMS: Application to Samoan basaltic melt inclusions

#### 1. Introduction

We have thoroughly, but not exhaustively, investigated the parameters necessary for precise and accurate Sr isotope analysis of basalt glasses by *in situ* laser ablation PIMMS. The lower limit of external precision attained under optimal conditions for the SRM987 Sr solution standard is  $\pm 5$  ppm ( $>25$  volt data,  $1\sigma$ ), and is fairly comparable to good TIMS data. For lower intensity data ( $\sim 1$ -5 volts), more comparable to our typical laser ablation work, the external solution precision is in the 20-80 ppm range. We have instituted a new method for Kr correction that is based on  $^{84}\text{Kr}$ , and a method for Rb correction that utilizes basalt glass standards with significant Rb/Sr ratios and known  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios. Even in basalts requiring very large Rb corrections, we are able to obtain  $^{87}\text{Sr}/^{86}\text{Sr}$  data with  $\sim 320$  ppm external precision and 45 ppm internal (in-run) precision ( $2\sigma$  standard deviation) on a suite of Samoan glasses with known  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios.

We developed an analytical protocol for the measurement of Sr isotope ratios by Laser Ablation PIMMS for analysis of melt inclusions in olivine phenocrysts in OIBs (ocean island basalts) from the EM2 (Samoa), EM1 (Pitcairn, Christmas) and HIMU (Mangaia) mantle end-members. We acknowledge a significant existing body of work on Sr isotope measurement by laser ablation PIMMS, but will not attempt to review this literature or compare it with the protocols we have installed on the NEPTUNE at the Woods Hole Oceanographic Institution [1-6].

#### 2. Techniques

##### 2.1. Instrumental Description

The data discussed here were obtained with a ThermoFinnigan NEPTUNE multicollector ICP-MS, coupled to a NEW Wave UP213 laser, housed in the Plasma Facility at the Woods Hole Oceanographic Institution. This NEPTUNE was installed in June 2003, replacing an earlier one that had been installed in February 2002, but damaged by the October 2002 fire in the Ion Probe Facility.

The typical operating parameters of the NEPTUNE and laser are given in Table 1 (note that our techniques have evolved over time, and some earlier data may have utilized somewhat different protocols). Currently, we aspirate clean dilute (5%) nitric acid during the laser firing, to allow solution standards to be interspersed with laser runs when needed. The arrangement of Faraday cups, and the relevant isotopic masses that are collected, is given in Table 2. Amplifiers are not rotated, as this rotates the Faraday (off-mass) baselines as well. The mean raw Faraday intensities are transferred after all analyses to an offline data reduction program (TweaKr), for the various corrections.

We have opted to run the laser at 100% power and in apertured mode. While we could likely emplace higher energy density in focused mode, this tends to throw off larger chips and particles, particularly from basalt glass samples. We have not made any

investigations of the effects of particle size on mass bias, interferences, and the like. Each analysis takes about 6.5 minutes of lasering, including offpeak baseline integration for 64 seconds and 20 cycles of 16 second integrations (and, more recently, 40 cycles of 8 second integrations). With a 120  $\mu\text{m}$  spot and 200  $\mu\text{m}$  long raster lines (see Table 1), spaced at only 20  $\mu\text{m}$ , the final ablation pit is pyramidal, approximately 300 x 300  $\mu\text{m}$ , and 250  $\mu\text{m}$  deep. Typically, there will be little or no drift of  $^{88}\text{Sr}$  intensity with time during ablation of homogeneous basalt glass standards, and the  $1\sigma$  standard deviation of intensity will fall in the range 5-15%.

### 2.2. Standards and Canonical Isotope Abundances

For Rb and Sr isotope standards, we have used NIST standards SRM984 and SRM987. The certified values for these standards are listed in Table 3, along with the quoted uncertainties. Note that  $^{85}\text{Rb}/^{87}\text{Rb}$  in SRM 984 is only certified to  $\pm 423$  ppm,  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{88}\text{Sr}/^{86}\text{Sr}$  to only 380 ppm, and  $^{84}\text{Sr}/^{86}\text{Sr}$  to 0.25%. It is likely that the uncertainties in  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{88}\text{Sr}/^{86}\text{Sr}$  are not independent, but possibly related by some fractionation-dependence. Note also that the certified  $^{86}\text{Sr}/^{88}\text{Sr}$  value is 0.1193515, not the consent value of 0.119400 adopted by the community. We have corrected all of the SRM 987 certificate ratios, by exponential law, to be consistent with the consent value of 0.1194. Ironically, this results in a “corrected”  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of 0.7101938, which is lower than (but marginally within errors of) the value commonly adopted by the community of 0.710240 (which is itself different from the “uncorrected” certificate value of 0.710339). It will obviously be important for published papers to be very clear as to their usage of SRM 987 standard values. For Kr, we adopt the isotope abundances compiled by Ozima and Podosek (2001).

### 2.3. Backgrounds and Baselines

The NEPTUNE software allows a choice of baseline protocols. PIMMS users frequently use “on-peak” baselines, after sample wash-out [1, 3-6]; the NEPTUNE also allows a “defocused” beam baseline. Our experience is that baselines can be measurably ( $> 50 \mu\text{V}$ ) elevated across a broad mass region during actual sample analysis and thus different from those measured in a defocused “beam-off” state, or while running gas or acid blanks. We have chosen therefore to adopt the common TIMS procedure of running “off-peak” baselines while sample analysis is underway; while this consumes part of the sample, we see no other way to ensure reliability of baselines. (However, baselines on Samoan basalt standards and basaltic melt inclusions do not change significantly over the course of an analytical session, and we have adopted the practice of applying baselines from larger melt inclusions—run during the same analytical session—to exceptionally small melt inclusions to save precious material). We have chosen to measure baselines at a position 0.70 amu below each mass; this avoids potential baseline interferences from doubly-charged half-mass REE peaks, allows the Faradays that are intentionally set at 83.5 and 85.5 in order to monitor these  $\text{REE}^{+2}$  to offset 0.2 amu below mass 83 and 85 during baseline measurement, and places the  $^{88}\text{Sr}$  cup near the minimum in the valley between  $^{87}\text{Sr}$  and  $^{88}\text{Sr}$  during baseline measurement. With this protocol, the only baseline that will have a significant tail contribution is  $^{87}\text{Sr}$ . With the measured abundance sensitivity at Sr mass of  $\sim 1$  ppm at 1 amu, the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio will be elevated by  $< 12$

ppm; this error is small compared to our external precision goal, and will also be normalized, to first order, through the use of the SRM987 standard.

#### 2.4. Kr Correction

As is well known, there are interferences at  $^{84}\text{Sr}$  and  $^{86}\text{Sr}$  masses from Kr. Conventionally, the Kr is considered to be a contaminant in the Ar plasma gas, and is corrected by measurement of  $^{82}\text{Kr}$  or  $^{83}\text{Kr}$ . Over the past two years of measurement here, our  $^{83}\text{Kr}$  intensity has varied from  $\sim 0.05$  mV to  $\sim 20$  mV, and we believe that not all of it comes in with the argon, nor is all of mass 83 actually Kr. Typically, after changing samples in the laser cell, there is a component of air Kr which slowly decreases as the chamber is purged. Ultimately, the Kr intensity with He flowing from the cell into the machine will decrease substantially below that in aspirated solutions (because the He has less Kr than the argon). It appears to us that there is a component of atmospheric Kr dissolved in aspirated solutions and this contributes to the Kr background. This could be possibly moderated by hermetically isolating the solutions from the atmosphere. Air entrainment by the plasma at the torch is also probable, but has proven difficult to constrain.

We have typically measured both  $^{82}\text{Kr}$  and  $^{83}\text{Kr}$  during all Sr analyses. Deviations of up to a factor of two from the canonical  $^{83}\text{Kr}/^{82}\text{Kr}$  ratio of  $\sim 1$  are very common in laser runs, with values most often above one, but also frequently below one (Fig. 1). Even solution runs on the 987 standard will frequently show 10-20% deviations from a ratio of 1, typically with the ratios being too low (interferences on mass 82). Clearly there exist isobaric interferences that seriously hamper efforts to use either of these masses to correct for Kr. We have adopted an alternative scheme which basically uses the most abundant Kr isotope at mass 84 ( $\sim 57\%$ ) to make the Kr correction on mass 86. This mass has a Sr “interference” on it, but for  $^{88}\text{Sr}$  intensities in the 1-2 volt range, the Kr makes up 35-75% of the mass 84 peak (i.e. the Kr and Sr are approximately equal in intensity). By “subtracting” Kr until the  $^{84}\text{Sr}/^{88}\text{Sr}$  ratio equals the canonical value of 0.00675476 (while iterating the mass-bias correction), this allows a robust correction to be made on mass 86 (there is a large error demagnification in this process, due to the fact that  $^{86}\text{Kr}/^{84}\text{Kr}$  is  $\sim 0.30$ , while the  $^{86}\text{Sr}/^{84}\text{Sr}$  ratio is  $\sim 17.7$ ). Obviously, this technique relies on the absence of any other significant isobaric interferences at mass 84; in any event, these are likely to be less fatal than those at mass 83, because of the  $^{84}\text{Kr}/^{83}\text{Kr}$  ratio of  $\sim 5$ . One indication that this calculation method is helpful is the observation that the external precision of samples and standards run at low-intensity (1-2 volts on mass 88) is typically improved by 30-50%, compared to the same data corrected with  $^{83}\text{Kr}$ . Further discussion of isobaric interference issues may be found in section III F, below.

We are able to correct for Kr interferences so that elevated Kr/Sr ratios do not noticeably diminish the precision or accuracy of glass analysis while using our protocol, provided that the  $^{88}\text{Sr}$  intensity is  $> 1$  V and  $^{82}\text{Kr}/^{88}\text{Sr}$  is  $< 0.004$  (Table 4). At exceptionally high Kr/Sr ratios, there may be a tendency for measurement precision to degrade. Therefore, analyses of unknowns that exhibit high Kr/Sr ratios ( $^{82}\text{Kr}/^{88}\text{Sr} > 0.004$ ) are discarded.

Uncertainty in the isotopic ratios of Kr used in our correction scheme does not significantly contribute to diminishing the accuracy of  $^{87}\text{Sr}/^{86}\text{Sr}$  analyses by laser ablation. At elevated Kr/Sr ( $^{82}\text{Kr}/^{88}\text{Sr} = 0.004$ ) — where such uncertainties will play the

largest role in affecting the final Kr-corrected  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio—a 1% uncertainty in any or all of the Kr-isotope ratios will change the final corrected  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio by <15 ppm. Additionally, allowing the Kr-isotope ratios to fractionate by up to 2% (before or during entry into the plasma) changes the  $^{87}\text{Sr}/^{86}\text{Sr}$  by <15 ppm.

### 2.5. Rb Correction

The analysis of basaltic melt inclusions typically involves samples in which 20-50% of intensity on mass 87 is due to Rb, and large Rb corrections are required. Our goal was to devise a protocol where the Rb-corrected  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios are accurate to within a few hundred ppm. We adopted the strategy of using natural basalt glasses (Table 5) with precisely known  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (obtained by TIMS) to bracket the basalt glass unknowns; for each “standard” glass, the  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio required to give the TIMS  $^{87}\text{Sr}/^{86}\text{Sr}$  is calculated, and these bracketing  $^{85}\text{Rb}/^{87}\text{Rb}$  ratios are then used for the Rb corrections in the unknown glasses. This method depends only on the Rb mass bias being relatively invariant on short time scales (20 minutes), and not being a function of the particular major element composition of the various glasses analyzed.

To assess these issues, we performed laser ablation analyses of a suite of 14 Samoan basalt glasses of known  $^{87}\text{Sr}/^{86}\text{Sr}$ , with significant variability in major and trace element composition. The results of nine one-day analysis campaigns are shown in Fig. 2; interspersed with these laser ablation analyses during the first two days were a series of runs on mixed 984-987 Rb-Sr solutions. While the total variability of the “Required  $^{85}\text{Rb}/^{87}\text{Rb}$ ” (basically the variability of the Rb mass bias factor) is substantial (2317 ppm, excluding the 987-984 Rb-Sr solution runs), the variation with time is relatively smooth, such that the bracketing technique will be fairly effective (note that the error on  $^{87}\text{Sr}$  due to the Rb effect is de-magnified by a factor of ~2.59 due to the  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio).

Several things may be noted. First, the “empirically” determined Rb ratio in lasered basalts has an average value of 2.58745, which is some 2100 ppm lower than the canonical value certified for the SRM984 Rb standard (the stated uncertainty in SRM984 is  $\pm 770$  ppm). Whether this reflects the existence of natural isotope variations in Rb, or simply an underestimate of the SRM certified value is unclear. If the former, then it will perhaps be important to use basalt glass standards that are petrogenetically related to the unknowns. Secondly, there is a clear tendency for the Rb-Sr solution runs to lie at the high side of the data series, averaging 2.5900 (1200 ppm lower than the canonical value for the SRM984 Rb standard). This would suggest a slight difference in Rb mass bias for solution runs versus laser ablation runs (and this would not be surprising, given the much larger plasma loading from the laser runs).

To assess the overall accuracy of this correction scheme, we have “corrected” each run in two ways: first, using the contiguous bracketing runs, and secondly, correcting each run with the global average “required”  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio of 2.58745. The resulting “corrected”  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios are then compared with the known values (note that the internal precision of the solution runs averaged about  $\pm 10$  ppm ( $1\sigma$ , standard error); that of the laser runs averaged  $\pm 17$  ppm, with 88 intensities ranging from 1-12 volts). For the “bracketing technique”, the mean deviation (measured against TIMS) for the nine analytical sessions is 127 ppm (320 ppm,  $2\sigma$  standard deviation, assuming a Gaussian distribution); for the “global average” technique, the mean deviation is 155 ppm (Fig. 3). We adopt the first scheme to correct unknowns and when discussing the accuracy of

standard glass (and melt inclusion) runs; this technique has the advantage of capturing the downward drift of the  $^{85}\text{Rb}/^{87}\text{Rb}$  required values over time (see Fig. 2). While obviously not as precise as TIMS analyses, this level of reproducibility is excellent for an *in situ* technique, and will allow us to embark on a realistic program of melt-inclusion analysis.

Due to the uncertainty of the  $^{85}\text{Rb}/^{87}\text{Rb}$  required value, there will be an error magnification on the corrected  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios of basaltic unknowns as Rb/Sr increases. In order to model this effect, we selected several glass standard runs with different Rb/Sr ratios and similar, low Kr/Sr ratios, and varied the  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio by 2 standard deviations about the mean global value of 2.58745 ( $\pm 0.00278$ ,  $2\sigma$  standard deviation). The propagated error increases linearly to 630 ppm when Rb/Sr is 0.14, and the accuracy appears to scale with the Rb/Sr such that the uncertainty in  $^{85}\text{Rb}/^{87}\text{Rb}$  magnifies the error on the final  $^{87}\text{Sr}/^{86}\text{Sr}$  (Fig. 4). However, the “bracketing method” for the Rb correction reduces the uncertainty in the  $^{85}\text{Rb}/^{87}\text{Rb}$  required value by  $\sim 20\%$ , so that the propagated error from the reduced uncertainty translates to an improvement in the accuracy of the  $^{87}\text{Sr}/^{86}\text{Sr}$  (i.e., to 505 ppm when Rb/Sr is 0.14). The internal precision for  $^{87}\text{Sr}/^{86}\text{Sr}$  measurements on lasered glasses is generally an order of magnitude better than the accuracy (this is also true for the Rb/Sr measurements), and does not appear to worsen with increasing Rb/Sr ratios. However, the internal precision of the  $^{87}\text{Sr}/^{86}\text{Sr}$ , but not the accuracy, does vary with Sr intensities (Fig. 5), and, to a lesser extent, the number and length (8 or 16 second integrations) of cycles of analysis. The precision on the melt inclusion analyses also varies as a function of Sr intensity and the number and length of cycles. However, because the accuracy of the  $^{87}\text{Sr}/^{86}\text{Sr}$  of Samoan glasses is not related to Sr intensity over the range of 1-12 Volts on mass 88 (Table 4), somewhat lower Sr intensities during the melt inclusion runs (1-6.5 Volts) should not affect the overall accuracy of  $^{87}\text{Sr}/^{86}\text{Sr}$ .

## 2.6. Other Interferences

We have directly measured possible isobaric mass interferences resulting from Ca dimers and argides,  $\text{FeO}_2$ , doubly-charged Er and Yb and KrH. The Ca dimers and argides invoke errors of less than  $<10$  ppm, for Ca/Sr ratios typical of alkali basalts; there is no straightforward way to monitor or correct for these. Similarly, in a typical basalt laser run (with 500 ppm Sr), the total propagated error in  $^{87}\text{Sr}/^{86}\text{Sr}$  from  $\text{FeO}_2$  will then be less than 40 ppm (and could be nil). Additionally, the REE $^{++}$  invoke errors of  $<30$  ppm for typical Sr/REE ratios in alkali basalts; these can in principal be corrected for by monitoring the 83.5 and 85.5 half-mass peaks due to  $^{167}\text{Er}^{++}$  and  $^{171}\text{Yb}^{++}$ . Finally, there seems to be some evidence for the formation of KrH in the plasma, with Kr/KrH exhibiting ratios from 35-70. Kr hydrides create isobaric interferences on masses 83, 84, 85 and 87, and they may play an important role when Kr/Sr ratios are elevated. In practice, however, it appears that there are frequently other unexplained interferences at these masses, as well as at the Kr and Rb masses, so that correction for these isobaric interferences is not always successful. These problems not only limit the attainable precision of laser ablation analyses, but can limit the precision of straightforward Sr solution analyses as well, even of the SRM987 standard. A continuing investigation of these issues (and a fuller discussion of interferences from Ca dimers, argides,  $\text{FeO}_2$ , KrH and doubly charged Er and Yb) is underway (Hart *et al.*, in prep).

### 2.7. Measurement of Rb/Sr by LA-ICP-MS

Masses 85 and 88 represent pure Rb and Sr, respectively, so that fairly precise measurement of Rb/Sr ratios can be generated. After correcting for mass fractionation during each run, Rb/Sr ratios on Samoan basalt glasses measured by laser ablation are accurate to 7% (1 mean deviation, compared to ratios obtained by XRF/ICP techniques on the same glasses), and precise (0.67%, 1 mean deviation) during multiple runs on a suite of Samoan glasses (Fig. 6). We note that this technique does not appear to work when running a “dry” plasma; Rb and Sr are strongly fractionated from each other, so that the measured Rb/Sr ratios are up to 80% higher than in the standard glass (see last 4 analyses in Table 4).

### 2.8. Replicability

We have not yet made a comprehensive study of reproducibility of sample analyses. We have done an analysis of a group of 50 runs on the SRM987 standard, spread over a 17 month time period, using solutions varying in concentration from 10 ppb to 600 ppb. All errors discussed here will be given at the  $1\sigma$  level. Considering first the 200 and 600 ppb solutions, the average internal precision for these was 7 ppm (for  $^{88}\text{Sr}$  intensities varying from 8-28 volts). The average external precision, calculated from the variations within a single day’s analysis session, was 12 ppm (and the number of standards run during each of these sessions varied from 2-6). While the overall ratio of external to internal precision was  $\sim 1.8$ , there was not a significant correlation between external and internal precision on a session-to-session basis (i.e. the internal precision on individual runs is not a good guide to the expected external precision). For the daily means of 11 sessions over the 17 month time period, the average  $^{87}\text{Sr}/^{86}\text{Sr}$  was 0.710255, with a  $1\sigma$  standard deviation of a single analysis of 15 ppm. We should note that this time period involved a number of different baseline protocols and cup configurations, so the statistics may not be representative of our current procedures. There is one obvious conclusion, however, and that is that the precision on a daily basis is similar to the long-term precision; in other words, the variability in standard runs has almost as much “daily” scatter as it does “yearly” scatter.

Within this data set, there is a fair correlation between Sr intensity and the precision of the data; external precision for a given session is  $\pm 5$  ppm for  $>25$  volt data,  $\pm 7-20$  ppm for 9-15 volt data and  $\pm 50-80$  ppm for 0.3-1 volt data. Overall, the  $1\sigma$  external precision in ppm, as a function of  $^{88}\text{Sr}$  intensity in volts, may be empirically expressed as:  $\sim [80/(\text{volts})^{0.8}]$ .

## 3. Summary

The Finnigan NEPTUNE multi-collector ICP-MS has proven to be an excellent instrument for developing robust Sr isotope analysis protocols. Among its advantages are the stability of the Faraday-amplifier system, the ability to run with intensities up to 50 volts, and the stability of the mass bias for Sr and Rb. While the external precision of solution analyses are not yet comparable to the best TIMS techniques, the *in situ* laser ablation technique appears to be a reliable and very promising tool for the study of small scale-length isotopic heterogeneities, even in samples with a significant Rb component.

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## Figure Captions

Fig. 1. Variation of intensity ratio of mass 82 to mass 83 over time. Series of Samoan basalt glasses analyzed by laser ablation (colored symbols) and mixed Sr-Rb standard solutions (SRM987-SRM984, black diamonds). The certificate value for  $^{82}\text{Kr}/^{83}\text{Kr}$  is 1.004. We observe no relationship between mass 82/mass 83 and the precision or accuracy of  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement by LA-MC-ICPMS.

Fig. 2.  $^{85}\text{Rb}/^{87}\text{Rb}$  ratio required to give correct  $^{87}\text{Sr}/^{86}\text{Sr}$  of a series of Samoan basalt glasses (analyzed by TIMS, open diamonds) and mixed Sr-Rb standard solutions (SRM987-SRM984, filled diamonds). The certificate value for  $^{85}\text{Rb}/^{87}\text{Rb}$  in SRM984 is  $2.593 \pm 0.002$ ; all of the “calculated” Rb ratios are lower than the certificate value, and outside quoted error limits. Runs using a dry plasma not shown.

Fig. 3. Reproducibility of  $^{87}\text{Sr}/^{86}\text{Sr}$  for 12 Samoan basalt glass standards (with known  $^{87}\text{Sr}/^{86}\text{Sr}$  by TIMS) by LA-MC-ICPMS. Two data trends represent reproducibility (external precision) using the “global average” and the “bracketing” correction schemes for the isobaric interference of Rb on mass 87. The bracketing method gives better overall external precision, and is adopted as the correction scheme in the manuscript. Mixed NBS987 and NBS984 solution runs not included in the figure. The downward drift of  $^{85}\text{Rb}/^{87}\text{Rb}$  (see Fig. 2) is noted in the downward drift of error using the “global average” Rb correction scheme. Error bars on symbols are internal (in-run) precision (2 mean deviations). Dashed lines mark external precision of  $\pm 320$  ppm ( $2\sigma$ , standard deviation). Larger scatter in later runs is due to higher Rb/Sr ratios of the glass standards analyzed. Better external precision in early runs is due to smaller variation in  $^{85}\text{Rb}/^{87}\text{Rb}$  over time. The glass standard runs at the beginning and the end of each analytical session are not correctible using the bracketing technique and are not plotted. However, all glass standards are correctible using the “global average” Rb correction scheme, so more data points are plotted for this latter correction scheme.

Fig. 4. Rb/Sr vs. propagated error on  $^{87}\text{Sr}/^{86}\text{Sr}$  measurement due to uncertainty of the  $^{85}\text{Rb}/^{87}\text{Rb}$  required value. The average  $^{85}\text{Rb}/^{87}\text{Rb}$  required values (used for correcting the  $^{87}\text{Rb}$  isobaric interference on  $^{87}\text{Sr}$ ) during all laser runs on Samoan glass standards is  $2.58745 (\pm 0.00278, 2\sigma)$ , and this uncertainty generates larger errors on the final  $^{87}\text{Sr}/^{86}\text{Sr}$  at higher Rb/Sr ratios. The upper model line shows the error ( $2\sigma$ , standard deviation) on final  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio when using the “global average” technique for the Rb correction (see text for description). Symbols on line represent error propagation on (from the uncertainty on the  $^{85}\text{Rb}/^{87}\text{Rb}$ ) using actual basalt analyses by laser ablation. Alternatively, error using the “bracketing” method for the Rb correction is 20% lower, and is described by the lower line. The Rb correction contributes the vast majority of the error on the final  $^{87}\text{Sr}/^{86}\text{Sr}$  in our protocol for measuring basal glasses by LA-MC-ICPMS. Therefore, we consider the lower model curve to be a good approximation of the reproducibility of our method.

Fig. 5. Relationship between intensity on mass 88 and internal precision during laser ablation analysis of Samoan glass standards. Various Neptune and laser operating conditions and variations in basalt and plasma chemistry, including Rb/Sr and Kr/Sr, show little or no relationship with measurement precision. The curve models the relationship between intensity and internal precision (1 mean deviation), and is described by the following relationship:  $\text{Precision} = 45 \cdot (\sqrt{\text{Intensity}} / \text{Intensity})$ .

Fig. 6. Samoan glass Rb/Sr measured by LA-MC-ICPMS Neptune plotted against Rb/Sr measured by standard techniques (XRF or ICP) on powders of the same glasses. Error bars ( $2\sigma$  standard deviation from the mean of measured values) are shown on samples with at least 9 laser analyses. All other glasses had one or two analyses. One-to-one line is solid, and weighted least squares regression (including equation) is dashed.

Table 1. Typical NEPTUNE and Laser Operating Parameters

RF Power	1200 W
Argon cooling gas, flow rate	15L/min
Ar auxiliary gas, flow rate	0.8 L/min
He sample gas, flow rate	0.5 L/min
Interface cones	X-cones
Mass analyzer pressure	$5 \times 10^{-9}$ mbar
Resolution mode	Low (400)
Abundance sensitivity	~ 1 ppm at -1 amu, mass 88 (RPQ off)
Detection system	Faraday cups (9)
Signal analysis set-up	8 sec or 16 sec integration/cycle, 40 or 20 cycles
Background-baseline protocol	1 min integration pre-analysis, -0.7 amu all Faradays
Nebuliser and uptake rate	Elemental Scientific, Inc., PFA 20-50 microliters per minute
Spray chamber	Elemental Scientific, Inc., Stable Introduction System (SIS)
Sensitivity, solution Sr	75 V $^{88}\text{Sr}$ per ppm Sr in 5% $\text{HNO}_3$
Sensitivity, laser Sr	~ 1 volt $^{88}\text{Sr}$ per 100 ppm Sr in basalt at 1 mJ
Laser type	New Wave UP213, quad Nd YAG 213 nm
Carrier gas	Helium
Beam optics	Apertured mode
Spot size	120 $\mu\text{m}$ (60 $\mu\text{m}$ or less in high-Sr carbonates)
Raster pattern	200 x 200 $\mu\text{m}$ , 20 $\mu\text{m}$ line spacing, 4 $\mu\text{m}/\text{sec}$
Pulse rate	20 Hz
Power setting	100% (lower for high-Sr carbonates)
Power delivered	~ 1.0-1.3 mJ
Pre-ablation	same raster and spot size, 5 Hz, 45% power, 30 $\mu\text{m}/\text{sec}$

<b>Table 2. Faraday Cup Configuration Laser Ablation Sr Isotope Analysis</b>	
<u>Cup</u>	<u>Mass</u>
H4	88
H3	87
H2	86
H1	85
C	84
L1	83
L2	82.5
L3	82
L4	81.5

Centered for Sr amu; C ~ 83.913

**Table 3. Adopted Isotopic Abundances.**

<u>Krypton</u>	<u>Isotopic Abundance</u>
78	0.003469
80	0.02257
82	0.11523
83	0.11477
84	0.56998
86	0.17398

From Ozima and Podosek, 2001

<u>Strontium SRM987</u>	<u>Isotopic Abundance**</u>	<u>Isotopic Abundance*</u>		
84	0.005578	0.005574		
86	0.098600	0.098566		
87	0.070029	0.070015		
88	0.825793	0.825845		
	<u>Isotopic Ratios**</u>	<u>Isotopic Ratios*</u>		
84/86	0.056573	0.056549	±2529	ppm
87/86	0.710194	0.710339	±367	ppm
88/86	8.375209	8.37861	±388	ppm
84/88	0.006755	0.006749	±2529	ppm
86/88	0.119400	0.119352	±388	ppm

\*NIST Certificate Values for SRM987 with quoted uncertainties.

\*\*Renormalized to 0.1194 using exponential law; this value has been used by community consensus for decades. The corresponding  $^{87}\text{Sr}/^{86}\text{Sr}$  value adopted by most of the community is 0.71024.

<u>Rubidium SRM984</u>				
85	0.72168			
87	0.27832			
85/87*	2.593	±771		ppm

\*NIST Certificate Values for SRM984, with quoted uncertainties.

Table 4. Corrected and normalized  $^{87}\text{Sr}/^{86}\text{Sr}$  glass standard data, including Rb/Sr ratios and raw exported Neptune data (in Volts) on masses 82 to 88

Date	Glass	Mass 82	Mass 83	Mass 84	Mass 85	Mass 86	Mass 87	Mass 88	$^{86}\text{Rb}/^{87}\text{Rb}$ Mass	$^{87}\text{Sr}/^{86}\text{Sr}$	Rb/Sr	Rb/Sr	Rb/Sr	Rb/Sr	Rb/Sr	mass	# of							
Name <sup>a</sup>	Meas'd	Meas'd	Meas'd	Meas'd	Meas'd	Meas'd	Meas'd	Meas'd	Required <sup>b</sup>	Bias <sup>c</sup>	Laser <sup>d</sup>	Laser <sup>d</sup>	Laser <sup>d</sup>	I. Prec. <sup>e</sup>	E. Prec. <sup>f</sup>	E. Prec. <sup>g</sup>	TIMS <sup>h</sup>	Laser <sup>d</sup>	I. Prec. <sup>e</sup>	E. Prec. <sup>f</sup>	XRF <sup>i</sup>	82/88 <sup>j</sup>	Cycles <sup>k</sup>	
01/28/04	75-10	0.0019	0.0015	0.070160	0.363076	1.161550	0.981164	9.99332	2.58799	1.45	0.70451	0.70450	--	10	41	--	0.704533	0.041	1.18	8.4	0.046	0.00019	20	
01/28/04	75-10	0.0019	0.0015	0.068130	0.410905	1.130263	0.970173	9.72817	2.58762	1.47	0.70452	0.70452	0.70454	11	19	12	0.704533	0.045	0.42	0.9	0.046	0.00020	20	
01/28/04	75-10	0.0019	0.0015	0.066872	0.316660	0.929578	0.788622	7.99056	2.58761	1.48	0.70452	0.70452	0.70456	10	16	38	0.704533	0.042	0.97	5.4	0.046	0.00023	20	
01/28/04	75-10	0.0018	0.0014	0.038589	0.210880	0.891571	0.505435	5.08531	2.58867	1.49	0.70447	0.70448	0.70450	11	97	41	0.704533	0.044	1.00	1.1	0.045	0.00035	20	
01/28/04	75-10	0.0019	0.0015	0.059566	0.310559	0.978671	0.821938	8.42300	2.58865	1.48	0.70448	0.70447	0.70456	9	85	34	0.704533	0.038	3.32	12.0	0.046	0.00022	20	
01/28/04	76-13	0.0017	0.0014	0.055907	0.544459	0.915110	0.871425	7.97805	2.58963	1.50	0.70620	0.70619	0.70633	13	285	89	0.706395	0.074	0.30	4.5	0.077	0.00022	20	
01/28/04	76-8	0.0017	0.0014	0.054455	0.528363	0.889211	0.845639	7.85417	2.58922	1.49	0.70621	0.70621	0.70639	11	231	22	0.706374	0.073	0.16	9.8	0.081	0.00023	20	
01/28/04	83-13	0.0017	0.0013	0.045848	0.441395	0.731734	0.898360	6.29880	2.58914	1.50	0.70536	0.70536	0.70547	11	226	71	0.706520	0.075	0.90	4.7	0.071	0.00028	20	
01/28/04	88-3	0.0018	0.0013	0.045375	0.373755	0.727534	0.868448	6.26098	2.58798	1.50	0.70535	0.70534	0.70542	15	62	51	0.705388	0.064	0.60	7.3	0.069	0.00029	20	
01/28/04	72-2	0.0018	0.0013	0.048745	0.438044	0.789158	0.737371	6.79224	2.58771	1.50	0.70537	0.70537	0.70538	19	37	16	0.705395	0.069	0.37	15.1	0.060	0.00028	20	
01/28/04	73-12	0.0019	0.0013	0.045837	0.450865	0.733186	0.704204	6.30680	2.58719	1.50	0.70688	0.70687	0.70670	14	28	63	0.706853	0.076	0.58	0.1	0.078	0.00030	18	
01/28/04	71-2	0.0019	0.0013	0.046247	0.412128	0.726657	0.883580	6.25381	2.58762	1.50	0.70583	0.70582	0.70587	14	27	37	0.706843	0.070	0.22	6.8	0.076	0.00030	20	
01/28/04	73-1	0.0018	0.0013	0.060270	0.533282	0.818134	0.797957	7.04356	2.58866	1.51	0.70680	0.70680	0.70688	10	175	86	0.706720	0.081	0.49	0.3	0.081	0.00026	20	
01/28/04	70-1	0.0018	0.0010	0.035357	0.270189	0.577103	0.518443	4.86111	2.58848	1.43	0.70530	0.70530	0.70540	14	107	41	0.706371	0.058	0.70	2.1	0.060	0.00033	20	
01/28/04	71-11	0.0016	0.0010	0.036483	0.316405	0.580123	0.547012	5.07213	2.58808	1.43	0.70526	0.70526	0.70537	12	194	27	0.706394	0.067	0.27	4.7	0.070	0.00031	20	
01/28/04	71-22	0.0017	0.0013	0.041431	0.382145	0.873708	0.825087	5.79313	2.58923	1.44	0.70533	0.70532	--	14	211	--	0.705473	0.067	0.55	6.1	0.071	0.00029	20	
01/27/04	73-1	0.0015	0.0010	0.042357	0.501491	0.715500	0.713457	6.18177	2.58963	1.85	0.70855	0.70848	--	14	338	--	0.706720	0.088	0.54	7.0	0.081	0.00024	20	
01/27/04	76-13	0.0012	0.0008	0.041502	0.427055	0.703095	0.874782	6.07876	2.58938	1.67	0.70628	0.70621	0.70841	14	259	25	0.706395	0.075	0.45	3.4	0.077	0.00020	17	
01/27/04	83-13	0.0011	0.0009	0.035882	0.355703	0.588686	0.571649	5.18386	2.58951	1.69	0.70539	0.70533	0.70553	17	270	8	0.706520	0.073	1.43	1.8	0.071	0.00021	20	
01/27/04	78-1	0.0011	0.0008	0.041195	0.638387	0.702572	0.761301	6.07862	2.58977	1.70	0.70862	0.70856	0.70875	16	482	185	0.708886	0.112	1.28	11.1	0.128	0.00018	20	
01/27/04	75-10	0.0011	0.0008	0.042356	0.280778	0.725189	0.631281	6.27295	2.58913	1.70	0.70455	0.70449	0.70483	12	81	137	0.704533	0.047	0.41	6.1	0.046	0.00017	20	
01/27/04	76-13	0.0018	0.0013	0.055752	0.552520	0.923794	0.881225	7.95726	2.58977	1.51	0.70624	0.70618	0.70632	12	304	111	0.706395	0.074	0.27	4.1	0.077	0.00020	10	
01/27/04	76-13	0.0016	0.0013	0.055487	0.554683	0.920418	0.879848	7.83062	2.58968	1.53	0.70625	0.70618	0.70641	10	285	17	0.706395	0.075	0.28	3.5	0.077	0.00020	20	
01/27/04	76-13	0.0016	0.0013	0.058108	0.578887	0.970484	0.925582	8.38648	2.58986	1.55	0.70624	0.70617	0.70639	11	315	24	0.706395	0.074	0.41	0.8	0.077	0.00019	20	
01/27/04	78-13	0.0015	0.0013	0.049800	0.508317	0.817493	0.786881	7.04381	2.58966	1.54	0.70625	0.70618	--	12	301	--	0.706395	0.077	0.42	4.6	0.077	0.00022	20	
02/13/04	71-11	0.0020	0.0020	0.052127	0.412318	0.785318	0.723859	6.73813	2.58794	1.41	0.70542	0.70535	--	13	62	--	0.705394	0.065	0.21	6.5	0.070	0.00030	20	
02/13/04	78-8	0.0020	0.0019	0.049567	0.579772	0.740226	0.759702	6.35140	2.58835	1.42	0.70757	0.70750	--	13	159	--	0.707614	0.089	0.40	0.5	0.087	0.00032	20	
03/28/04	70-1	0.0019	0.0016	0.048811	0.327874	0.704302	0.633370	6.08027	2.58880	1.56	0.70645	0.70641	--	17	60	--	0.705371	0.058	0.27	3.2	0.060	0.00032	20	
12/08/04	75-10	0.0027	0.0026	0.062513	0.321199	0.902189	0.770882	7.75825	2.58956	1.56	0.70485	0.70458	--	9	64	--	0.704533	0.044	0.34	1.4	0.045	0.00035	49	
12/08/04	75-10	0.0025	0.0025	0.082778	0.327432	0.917974	0.785039	7.90096	2.58872	1.61	0.70484	0.70457	0.70459	10	52	92	0.704533	0.044	0.17	1.5	0.045	0.00032	48	
12/08/04	76-13	0.0025	0.0024	0.083843	0.557882	0.950839	0.802824	8.19880	2.58901	1.67	0.70632	0.70625	0.70629	10	203	144	0.706395	0.072	0.13	6.5	0.077	0.00030	48	

12/09/04	76-13	0.0024	0.0024	0.062189	0.541662	0.926303	0.879131	7.98773	2.58901	1.69	0.70632	0.70625	0.70624	9	203	215	0.706395	0.072	0.13	6.9	0.077	0.00030	48
12/09/04	75-10	0.0021	0.0022	0.041088	0.190853	0.563339	0.478042	4.85665	2.58660	1.77	0.70469	0.70463	0.70466	16	133	175	0.704533	0.042	1.64	7.0	0.045	0.00043	50
12/09/04	75-10	0.0020	0.0019	0.044719	0.210583	0.648179	0.547229	5.59980	2.58695	1.82	0.70462	0.70456	--	12	32	--	0.704533	0.040	2.09	11.2	0.045	0.00035	50
01/10/05	76-13	0.0029	0.0032	0.062582	0.558318	0.908212	0.873053	7.83316	2.58638	1.73	0.70650	0.70649	--	12	136	--	0.706395	0.076	0.17	2.2	0.077	0.00037	37
01/10/05	76-13	0.0028	0.0031	0.062148	0.561306	0.908499	0.874702	7.83946	2.58683	1.74	0.70645	0.70645	0.70634	11	76	76	0.706395	0.076	0.18	1.8	0.077	0.00036	38
01/10/05	71-2	0.0033	0.0035	0.045848	0.326355	0.560609	0.529545	4.81303	2.58615	1.70	0.70606	0.70605	0.70601	15	158	95	0.705943	0.072	0.24	4.7	0.075	0.00069	38
01/10/05	71-2	0.0035	0.0038	0.053888	0.403808	0.691899	0.654374	5.94536	2.58697	1.69	0.70599	0.70598	0.70585	12	54	132	0.705943	0.072	0.30	4.5	0.075	0.00059	40
01/10/05	76-13	0.0036	0.0044	0.065799	0.643348	0.910299	0.907215	7.83305	2.58571	1.66	0.70658	0.70658	0.70655	13	258	223	0.706395	0.087	0.36	12.9	0.077	0.00046	40
01/10/05	71-2	0.0033	0.0034	0.031707	0.179591	0.292804	0.278502	2.49353	2.58737	1.73	0.70595	0.70595	0.70589	19	5	81	0.705943	0.076	0.51	1.2	0.075	0.00133	38
01/10/05	71-2	0.0019	0.0019	0.019831	0.128680	0.205503	0.197576	1.75832	2.58781	1.78	0.70581	0.70580	0.70600	26	56	75	0.705943	0.078	0.34	3.4	0.075	0.00106	40
01/10/05	71-2	0.0019	0.0020	0.018789	0.111868	0.176592	0.169506	1.50584	2.58933	1.77	0.70576	0.70575	0.70575	43	268	278	0.705943	0.079	0.68	4.2	0.075	0.00129	27
01/10/05	76-13	0.0021	0.0027	0.040442	0.401984	0.569223	0.567925	4.90895	2.58686	1.76	0.70646	0.70645	0.70665	19	83	360	0.706395	0.087	0.38	12.2	0.077	0.00043	40
01/10/05	76-13	0.0021	0.0026	0.027671	0.212827	0.335122	0.323659	2.87683	2.58910	1.71	0.70624	0.70623	0.70619	47	233	290	0.706395	0.078	0.98	1.5	0.077	0.00073	30
01/10/05	76-13	0.0018	0.0025	0.034968	0.346399	0.485332	0.485705	4.18601	2.58712	1.78	0.70643	0.70643	0.70663	19	44	338	0.706395	0.088	0.30	13.3	0.077	0.00046	40
01/10/05	71-2	0.0018	0.0018	0.016824	0.097919	0.154975	0.148658	1.32174	2.58951	1.82	0.70574	0.70574	--	43	291	--	0.705943	0.078	0.84	3.8	0.075	0.00133	31
01/11/05	71-2	0.0046	0.0051	0.069247	0.570788	0.881891	0.852357	7.52080	2.58694	1.29	0.70599	0.70596	--	18	26	--	0.705943	0.082	0.66	8.1	0.075	0.00061	40
01/11/05	71-2	0.0045	0.0051	0.058418	0.470999	0.717218	0.695719	6.11468	2.58698	1.33	0.70601	0.70599	0.70592	17	60	36	0.705943	0.083	0.58	9.6	0.075	0.00073	40
01/11/05	71-2	0.0040	0.0048	0.061762	0.582067	0.830511	0.822118	7.10423	2.58651	1.38	0.70607	0.70604	0.70597	16	135	39	0.705943	0.088	0.63	16.3	0.075	0.00056	40
01/11/05	76-13	0.0031	0.0035	0.042900	0.404831	0.547640	0.550584	4.88347	2.58655	1.44	0.70652	0.70649	0.70636	19	135	53	0.706395	0.092	0.85	19.6	0.077	0.00066	40
01/11/05	76-13	0.0034	0.0041	0.054000	0.586498	0.753766	0.770557	6.45857	2.58594	1.43	0.70659	0.70657	0.70643	16	245	48	0.706395	0.097	0.66	25.8	0.077	0.00052	40
01/11/05	76-13	0.0035	0.0044	0.053429	0.584883	0.737877	0.758293	6.31971	2.58591	1.42	0.70680	0.70658	0.70648	16	255	126	0.706395	0.099	0.47	28.2	0.077	0.00055	40
01/11/05	71-2	0.0046	0.0049	0.064039	0.471377	0.775869	0.736914	6.60756	2.58735	1.29	0.70597	0.70595	0.70584	17	7	149	0.705943	0.077	0.54	1.6	0.075	0.00070	40
01/11/05	76-13	0.0049	0.0053	0.064899	0.528307	0.775994	0.759615	6.60372	2.58653	1.27	0.70651	0.70649	0.70650	16	128	148	0.706395	0.086	0.60	11.3	0.077	0.00074	40
01/11/05	71-2	0.0048	0.0050	0.063217	0.452777	0.748571	0.708570	6.35639	2.58772	1.31	0.70594	0.70591	0.70589	20	43	79	0.705943	0.076	0.49	1.4	0.075	0.00075	40
01/11/05	71-2	0.0050	0.0052	0.059835	0.419486	0.671236	0.641146	5.70452	2.58771	1.29	0.70594	0.70591	0.70585	17	42	131	0.705943	0.079	0.47	4.7	0.075	0.00087	40
01/11/05	76-13	0.0052	0.0066	0.099886	1.041701	1.410914	1.417626	12.0573	2.58577	1.27	0.70661	0.70658	0.70649	10	280	135	0.706395	0.093	0.23	20.2	0.077	0.00043	40
01/11/05	76-13	0.0050	0.0065	0.096918	1.069271	1.373301	1.402463	11.7439	2.58552	1.31	0.70664	0.70662	0.70656	11	315	227	0.706395	0.098	0.46	26.6	0.077	0.00043	40
01/11/05	71-2	0.0048	0.0050	0.053745	0.351423	0.586159	0.539079	4.80289	2.58798	1.29	0.70592	0.70589	0.70584	12	78	142	0.705943	0.079	0.39	4.2	0.075	0.00101	40
01/11/05	71-2	0.0042	0.0044	0.045297	0.297450	0.470912	0.450880	4.00124	2.58832	1.40	0.70588	0.70585	0.70590	26	126	55	0.705943	0.080	0.71	5.5	0.075	0.00106	26
01/11/05	71-2	0.0042	0.0045	0.058885	0.471070	0.730507	0.708057	6.24085	2.58785	1.42	0.70593	0.70590	0.70596	15	63	22	0.705943	0.081	0.58	7.0	0.075	0.00067	40
01/11/05	71-2	0.0039	0.0044	0.054667	0.450610	0.678735	0.661178	5.79870	2.58770	1.42	0.70594	0.70591	0.70584	14	43	11	0.705943	0.083	0.24	10.2	0.075	0.00067	40
01/11/05	71-2	0.0048	0.0055	0.079954	0.671802	1.083082	1.022383	9.08473	2.58739	1.35	0.70597	0.70594	0.70601	15	1	94	0.705943	0.079	0.35	5.1	0.075	0.00053	40
01/11/05	71-2	0.0047	0.0050	0.052864	0.347356	0.586080	0.537973	4.80994	2.58846	1.35	0.70587	0.70584	--	28	141	--	0.705943	0.077	0.89	2.6	0.075	0.00098	40
08/09/05	71-2	0.0014	0.0011	0.022728	0.258610	0.388432	0.379989	3.33312	2.58724	1.28	0.70596	0.70596	--	19	23	--	0.705943	0.083	1.38	10.5	0.075	0.00041	37

08/09/05 71-2	0.0010	0.0012	0.039887	0.410998	0.878025	0.848988	5.81840	2.58823	1.44	0.70587	0.70587	0.70583	21	108	20	0.705843	0.078	0.77	0.1	0.075	0.00017	40
08/09/05 75-10	0.0010	0.0009	0.041822	0.267759	0.717313	0.818485	6.17228	2.58758	1.43	0.70453	0.70452	0.70480	13	15	96	0.704533	0.048	1.00	3.8	0.045	0.00016	40
08/09/05 71-2	0.0009	0.0009	0.034384	0.350906	0.581418	0.555389	5.00488	2.58832	1.46	0.70577	0.70577	0.70574	16	248	288	0.705843	0.075	0.42	0.7	0.075	0.00019	40
08/09/05 75-10	0.0010	0.0012	0.044398	0.305583	0.781872	0.886350	8.55471	2.58863	1.41	0.70458	0.70458	0.70488	14	87	209	0.704533	0.050	0.81	11.8	0.045	0.00015	40
08/09/05 76-13	0.0009	0.0011	0.040285	0.445631	0.888601	0.888344	5.90586	2.58878	1.41	0.70828	0.70828	0.70619	18	192	281	0.706395	0.081	0.51	4.5	0.077	0.00016	40
08/09/05 76-13	0.0011	0.0021	0.058288	0.720332	1.014084	1.011825	8.72088	2.58876	1.38	0.70847	0.70848	0.70683	16	97	331	0.706395	0.088	0.79	14.5	0.077	0.00013	40
08/09/05 71-2	0.0005	0.0005	0.017882	0.187819	0.288184	0.271203	2.45882	2.58809	1.03	0.70580	0.70580	0.70570	22	210	347	0.705843	0.074	0.79	2.2	0.075	0.00020	40
08/09/05 75-10	0.0008	0.0008	0.029739	0.191258	0.502920	0.433088	4.28411	2.58584	1.03	0.70483	0.70482	--	18	127	--	0.704533	0.048	1.71	7.8	0.045	0.00015	40
08/10/05 71-2	0.0008	0.0008	0.030282	0.288853	0.507454	0.483010	4.38372	2.58702	1.85	0.70588	0.70588	--	16	48	--	0.705843	0.072	0.63	4.0	0.075	0.00018	40
08/10/05 75-10	0.0008	0.0010	0.038555	0.248734	0.860744	0.572473	5.70862	2.58474	1.85	0.70488	0.70488	0.70482	17	217	123	0.704533	0.048	0.83	3.9	0.045	0.00015	40
08/10/05 76-13	0.0010	0.0017	0.048809	0.547845	0.808282	0.798810	8.98314	2.58548	1.83	0.70858	0.70858	0.70848	17	280	92	0.706395	0.083	0.77	8.0	0.077	0.00014	40
08/10/05 71-2	0.0008	0.0009	0.031222	0.318881	0.530073	0.508828	4.58580	2.58748	1.72	0.70584	0.70584	0.70574	22	10	280	0.705843	0.073	0.28	2.9	0.075	0.00018	40
08/10/05 75-10	0.0008	0.0009	0.038888	0.201178	0.868617	0.558368	5.78558	2.58498	1.84	0.70484	0.70484	0.70488	12	157	177	0.704533	0.037	4.72	17.4	0.045	0.00013	40
08/10/05 71-2	0.0007	0.0008	0.018731	0.177420	0.288853	0.284878	2.58205	2.58782	1.70	0.70580	0.70580	0.70570	28	67	348	0.705843	0.073	0.85	3.4	0.075	0.00029	38
08/10/05 76-13	0.0010	0.0021	0.053202	0.858724	0.930221	0.931071	8.04115	2.58547	1.85	0.70880	0.70880	--	20	285	--	0.706395	0.087	0.80	12.8	0.077	0.00013	40
08/11/05 71-2	0.0008	0.0008	0.028474	0.280326	0.481040	0.457445	4.18744	2.58755	1.80	0.70583	0.70583	--	19	19	--	0.705843	0.071	0.54	5.7	0.075	0.00018	40
08/11/05 71-2	0.0008	0.0008	0.025828	0.248742	0.428075	0.407073	3.70873	2.58701	1.78	0.70588	0.70588	0.70587	21	49	43	0.705843	0.071	0.58	5.5	0.075	0.00021	40
08/11/05 71-2	0.0008	0.0008	0.027774	0.287384	0.463111	0.438103	4.00773	2.58715	1.75	0.70588	0.70588	0.70582	20	31	33	0.705843	0.071	1.25	6.4	0.075	0.00021	33
08/11/05 71-2	0.0008	0.0008	0.025284	0.255547	0.433576	0.413754	3.75880	2.58877	1.83	0.70800	0.70800	0.70587	23	80	38	0.705843	0.072	0.48	4.8	0.075	0.00016	40
08/11/05 71-2	0.0007	0.0008	0.028528	0.274040	0.455528	0.438804	3.94888	2.58887	1.83	0.70588	0.70588	0.70587	26	58	42	0.705843	0.073	1.15	2.9	0.075	0.00017	28
08/11/05 71-2	0.0007	0.0008	0.022004	0.218339	0.387855	0.350877	3.18822	2.58782	1.83	0.70580	0.70581	0.70588	19	53	93	0.705843	0.072	0.40	5.0	0.075	0.00021	40
08/11/05 71-2	0.0007	0.0008	0.024842	0.251123	0.417278	0.400127	3.81888	2.58720	1.82	0.70588	0.70588	0.70587	19	25	40	0.705843	0.073	0.51	2.8	0.075	0.00019	40
08/11/05 71-2	0.0007	0.0008	0.025228	0.258815	0.427220	0.410275	3.70300	2.58721	1.82	0.70588	0.70588	0.70588	21	25	54	0.705843	0.074	0.34	2.2	0.075	0.00019	40
08/11/05 71-2	0.0007	0.0007	0.023358	0.232289	0.385225	0.378807	3.42880	2.58804	1.84	0.70588	0.70588	--	17	81	--	0.705843	0.072	0.31	5.1	0.075	0.00019	40
01/18/08 78-1	0.0082	0.0081	0.070811	0.888532	0.782508	0.818940	8.52057	2.58430	1.89	0.70847	0.70832	--	13	810	--	0.708888	0.112	0.85	10.6	0.128	0.00088	45
01/18/08 78-1	0.0058	0.0058	0.050415	0.385045	0.405885	0.432382	3.43848	2.58480	1.72	0.70838	0.70824	0.70881	22	484	38	0.708888	0.113	1.03	10.3	0.128	0.00172	45
01/18/08 78-1	0.0058	0.0057	0.088038	1.008880	1.140008	1.218188	9.78527	2.58588	1.88	0.70825	0.70808	0.70881	16	284	106	0.708888	0.108	0.72	12.9	0.128	0.00060	45
01/18/08 78-1	0.0058	0.0055	0.082132	0.584585	0.881023	0.704447	5.85188	2.58572	1.70	0.70827	0.70811	0.70878	28	323	148	0.708888	0.110	0.90	12.6	0.128	0.00088	39
01/18/08 78-1	0.0053	0.0053	0.042805	0.278774	0.311453	0.328480	2.82848	2.58408	1.71	0.70851	0.70835	0.70813	21	854	340	0.708888	0.112	0.32	11.0	0.128	0.00201	45
01/18/08 78-1	0.0050	0.0050	0.033418	0.140870	0.184178	0.188158	1.35428	2.58587	1.72	0.70825	0.70910	0.70873	33	288	214	0.708888	0.110	0.58	12.1	0.128	0.00371	45
01/18/08 78-1	0.0050	0.0050	0.036845	0.208480	0.230855	0.243558	1.93288	2.58548	1.70	0.70831	0.70816	0.70910	35	382	308	0.708888	0.113	0.85	9.7	0.128	0.00258	27
01/18/08 78-1	0.0085	0.0085	0.048816	0.121730	0.148878	0.145002	1.15711	2.58816	1.77	0.70884	0.70878	0.70884	52	148	489	0.708888	0.111	0.56	11.3	0.128	0.00732	45
01/18/08 71-2	0.0155	0.0157	0.081273	0.138273	0.263105	0.227057	2.08873	2.58584	1.78	0.70824	0.70808	0.70805	28	187	147	0.705843	0.071	0.37	5.4	0.075	0.00752	45
01/18/08 71-2	0.0183	0.0185	0.123513	0.333251	0.811080	0.551484	5.04318	2.58588	1.72	0.70823	0.70808	--	13	180	--	0.705843	0.070	0.37	7.2	0.075	0.00384	45

01/16/06 78-1 <sup>j</sup>	0.0105	0.0098	0.071513	0.683739	0.480786	0.592918	3.84615	2.58924	1.48	0.70862	0.70846	--	37	602	--	0.708886	0.190	0.19	51.3	0.126	0.00273	41
01/16/06 78-1 <sup>j</sup>	0.0111	0.0104	0.090872	1.314651	0.775036	1.069266	6.54519	2.58742	1.44	0.70804	0.70888	--	24	8	--	0.708886	0.215	0.82	71.1	0.126	0.00169	45
01/16/06 71-2 <sup>k</sup>	0.0119	0.0112	0.086823	0.853507	0.631526	0.700407	5.29883	2.58759	1.45	0.70607	0.70591	--	27	43	--	0.705843	0.132	0.55	74.8	0.075	0.00225	45
01/16/06 71-2 <sup>k</sup>	0.0126	0.0119	0.093511	0.738091	0.694135	0.778190	5.82947	2.58734	1.45	0.70611	0.70595	--	20	14	--	0.705843	0.135	0.38	79.4	0.075	0.00217	45

Analyses are listed in chronological order. Cycles are 16 second integrations of measurement during analysis until 3/26/04. Cycles are 8 second integrations thereafter.

<sup>a</sup> All samples are dredge samples from the AVON3 dredging cruise (Workman *et al.*, 2004) [7]

<sup>b</sup>  $^{86}\text{Rb}/^{87}\text{Rb}$  required values are obtained after correcting for Kr and Rb interferences, fractionation correction, and normalization to NBS 987 solution run during each analytical session. The average  $^{86}\text{Rb}/^{87}\text{Rb}$  required value is 2.58745, which is used to obtain the Rb correction using the "global average" technique

<sup>c</sup> Sr-isotopes are corrected for Kr and Rb interferences and fractionation corrected

<sup>d</sup> After correction for Rb and Sr interferences and fractionation, glass standard runs are corrected for 987 standard solution runs for each analytical session

<sup>e</sup> Internal (in-run) precision (I. Prec.) is quoted as 1 $\sigma$  standard error, and is calculated offline using the Tweekr program. In the above dataset, the average internal precision is 17 ppm for  $^{87}\text{Sr}/^{86}\text{Sr}$ , and 0.67% for Rb/Sr.

<sup>f</sup> External precision (E. Prec.) is calculated assuming that the TIMS  $^{87}\text{Sr}/^{86}\text{Sr}$  (ppm) and XRF Rb/Sr (%) values obtained on bulk powders are correct. The average external precision of  $^{87}\text{Sr}/^{86}\text{Sr}$  measurements on the above glass standards is 155 ppm mean deviation (1 $\sigma$ ), which translates to 388 ppm standard deviation (2 $\sigma$ ) if a Gaussian distribution is assumed. The average external precision on Rb/Sr measurements on glass standards is 6.7% (mean deviation, 1 $\sigma$ ), which translates to a standard deviation of 16.7% (2 $\sigma$ ). Uses global average  $^{86}\text{Rb}/^{87}\text{Rb}$  value (2.58745) for isobaric Rb interference.

<sup>g</sup>  $^{87}\text{Sr}/^{86}\text{Sr}$  and Rb/Sr values obtained by conventional methods (XRF or ICP) are reported in Workman *et al.*, 2004 [7].

<sup>h</sup> Raw voltages on mass 82 divided by mass 88. This is a proxy for Kr/Sr during analysis

<sup>i</sup> Samples are run at less than 100% laser power (between 45-90% power)

<sup>j</sup> Unlike all other Samoan glass standard and melt inclusion analyses, these four analyses were performed with a "dry" plasma (i.e., no aspiration of the 5% nitric solution during laser analysis).

These dry runs are not used in calculations or in the figures, and were performed for comparative purposes only. We find that the Rb is strongly fractionated from Sr, and that the  $^{86}\text{Rb}/^{87}\text{Rb}$  required and mass bias tend to be lower than similar runs made with a wet plasma during the same analytical session.

<sup>k</sup> The amount of mass bias correction required to obtain the canonical value for  $^{86}\text{Sr}/^{88}\text{Sr}$ . Units are %/amu.

<sup>l</sup> External precision is calculated assuming that the TIMS  $^{87}\text{Sr}/^{86}\text{Sr}$  is correct. External precision is calculated using the "bracketing" technique for the correction of the isobaric interference from Rb on mass 87. The external precision is somewhat improved (127 ppm, 1 mean deviation) compared to the global average technique for the Rb correction (155 ppm) when comparing the same set of glass analyses.

Bracketing is a useful technique for individual, 1-day analytical sessions only, as  $^{86}\text{Rb}/^{87}\text{Rb}$  values can change significantly between analytical sessions. Therefore, the first and last runs of the Samoan glass standards of each analytical session cannot be corrected with the bracketing technique. Runs with a dry plasma are not examined.

<sup>m</sup> Same as previous column, but  $^{87}\text{Sr}/^{86}\text{Sr}$  values calculated by correcting for the Rb interference using the "bracketing method", our preferred correction scheme and the one adopted in the manuscript for correcting unknowns.

Table 5. Compositions of Samoan glasses ("standards") used in determining <sup>87</sup>Rb/<sup>87</sup>Rb required values in this study.

Mass Nam	76-8	78-1	78-8	76-13	75-10	63-13	70-1	71-11	71-22	72-2	73-1	73-12	71-2	68-3
Volcano	Malumalu	Malumalu	Malumalu	Malumalu	Ta'u	Vailulu'u								
<sup>87</sup> Sr/ <sup>87</sup> Sr	0.706374	0.706896	0.707614	0.706395	0.704533	0.705520	0.705371	0.705394	0.705475	0.705395	0.706720	0.706653	0.705943	0.705388
SiO <sub>2</sub>	47.04	45.54	47.33	46.35	47.10	47.16	47.57	47.57	45.21	45.53	47.14	47.10	43.85	47.84
Al <sub>2</sub> O <sub>3</sub>	13.89	8.94	14.19	13.55	16.28	14.76	13.63	13.71	8.13	14.10	13.97	12.07	6.37	11.86
TiO <sub>2</sub>	3.88	1.96	3.29	3.86	4.02	3.01	2.81	2.75	1.68	3.31	2.80	2.46	1.44	2.55
FeO*	11.93	11.89	11.45	12.78	12.56	11.59	11.44	10.64	10.77	12.13	11.26	10.25	10.80	10.51
MnO	0.18	0.17	0.17	0.20	0.20	0.18	0.18	0.18	0.17	0.19	0.18	0.18	0.17	0.17
CaO	11.86	8.23	11.47	11.80	8.11	12.37	13.05	12.72	8.99	11.88	12.83	12.80	8.07	12.75
MgO	6.27	19.68	6.93	6.42	5.16	6.57	7.58	8.60	21.79	8.56	7.52	11.70	27.44	10.68
K <sub>2</sub> O	1.92	1.41	2.05	1.84	1.46	1.28	0.98	1.15	0.87	1.40	1.37	1.15	0.59	0.99
Na <sub>2</sub> O	2.75	1.92	2.74	2.61	3.47	2.72	2.35	2.35	1.39	2.54	2.50	2.02	1.11	2.36
P <sub>2</sub> O <sub>5</sub>	0.49	0.27	0.40	0.49	0.63	0.37	0.34	0.33	0.20	0.37	0.34	0.28	0.17	0.29
Pre-total	99.27	100.27	99.74	99.12	98.58	99.01	98.81	98.74	98.71	98.39	99.10	99.13	98.50	98.49
Ni	85	616	97	71	5	59	57	135	778	110	79	220	1080	285
Cr	137	1289	106	130	15	110	98	316	1687	227	202	768	2141	822
V	366	194	342	384	282	339	338	324	212	367	325	295	177	300
Ga	22	14	21	26	28	20	21	19	14	23	14	17	9	20
Cu	72	49	79	95	21	58	82	91	156	79	70	59	138	82
Zn	121	102	104	125	140	95	92	92	78	102	86	77	77	89
Cs	0.51	0.39	0.80	0.5	0.39	0.28	0.31	0.34	0.21	0.39	0.38	0.31	0.19	0.30
Rb	44.3	41.8	51.4	42.8	27.6	33.6	25.0	30.4	19.0	29.3	38.1	30.4	16.9	26.6
Ba	355	255	372	341	264	279	203	247	149	287	347	282	128	218
Th	6.4	6.0	6.3	5.3	4.9	5.2	4.2	4.4	2.6	5.0	6.1	4.8	2.3	4.0
U	1.30	1.05	1.36	1.29	1.10	1.07	0.83	0.85	0.56	1.11	1.15	0.87	0.49	0.87
Nb	58.7	39.0	59.0	57.5	51.3	45.1	37.1	40.5	24.7	47.5	48.7	38.2	20.7	36.4
Ta	3.9	2.6	3.8	3.8	3.6	3.1	2.5	2.8	1.6	3.2	3.1	2.5	1.4	2.5
La	48.4	39.4	52.0	48.6	43.1	37.4	31.0	32.9	20.2	38.3	44.2	36.1	17.1	28.3
Ce	95.9	73.0	98.8	98.3	90.5	74.4	65.0	65.6	40.4	78.3	83.6	68.8	34.4	58.7
Pb	4.2	4.2	4.6	5.2	3.3	2.8	2.2	2.3	1.6	2.5	4.8	6.2	1.3	2.5
Pr	11.2	7.9	11.5	11.8	11.1	8.7	7.6	7.9	4.7	9.1	9.4	7.8	4.0	6.9
Nd	46.3	31.0	44.2	47.2	48.0	35.4	31.5	31.6	18.0	37.2	36.0	30.2	16.5	28.5
Sr	544	333	529	554	617	470	420	434	267	492	473	398	224	387
Zr	276	183	258	281	327	218	198	198	117	221	212	177	98	168
Hf	7.2	4.7	6.4	7.1	6.3	5.8	5.4	5.2	3.1	5.9	5.9	4.9	2.8	4.4
Sm	10.5	6.5	9.1	10.2	11.6	7.8	7.5	7.3	4.4	8.4	7.8	6.7	3.7	6.5
Eu	3.2	1.9	2.5	2.9	3.8	2.5	2.3	2.3	1.4	2.6	2.4	2.1	1.1	2.0
Gd	9.5	5.8	8.0	9.1	10.7	7.1	6.8	6.4	3.9	7.3	6.8	6.0	3.2	5.8
Tb	1.4	0.8	1.1	1.3	1.6	1.1	1.0	1.0	0.6	1.1	1.1	0.9	0.5	0.9
Dy	7.4	4.4	5.8	6.6	8.8	5.9	5.7	5.4	3.3	6.0	5.6	5.0	2.7	4.8
Ho	1.3	0.8	1.0	1.2	1.6	1.1	1.0	1.0	0.6	1.1	1.0	0.9	0.5	0.9
Y	32.2	19.8	30.5	33.1	40.2	27.5	26.4	25.9	15.6	29.3	27.3	23.5	12.6	22.5
Er	3.0	1.8	2.5	2.8	3.8	2.7	2.5	2.4	1.4	2.7	2.5	2.2	1.2	2.1
Tm	0.38	0.23	0.34	0.37	0.50	0.36	0.33	0.32	0.19	0.36	0.33	0.28	0.16	0.28
Yb	2.05	1.24	1.97	2.13	2.80	2.02	1.87	1.83	1.11	2.02	1.90	1.64	0.87	1.63
Lu	0.29	0.19	0.29	0.31	0.40	0.30	0.27	0.27	0.16	0.29	0.28	0.23	0.13	0.24
Sc	31.0	25.2	28.5	25.0	20.5	33.4	41.9	39.2	36.0	38.0	39.1	42.7	28.8	40.9

Most glass data can be found in [7]. Major element data normalized to 100%. All glass standards names have the prefix "AVONS"

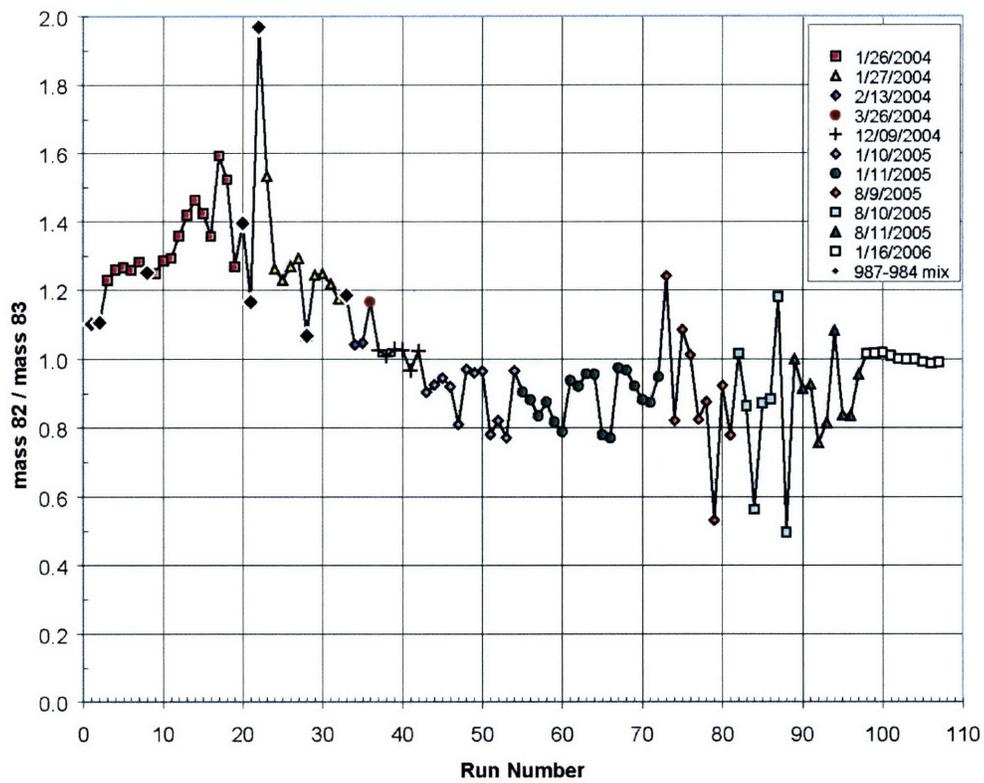


Fig. 1

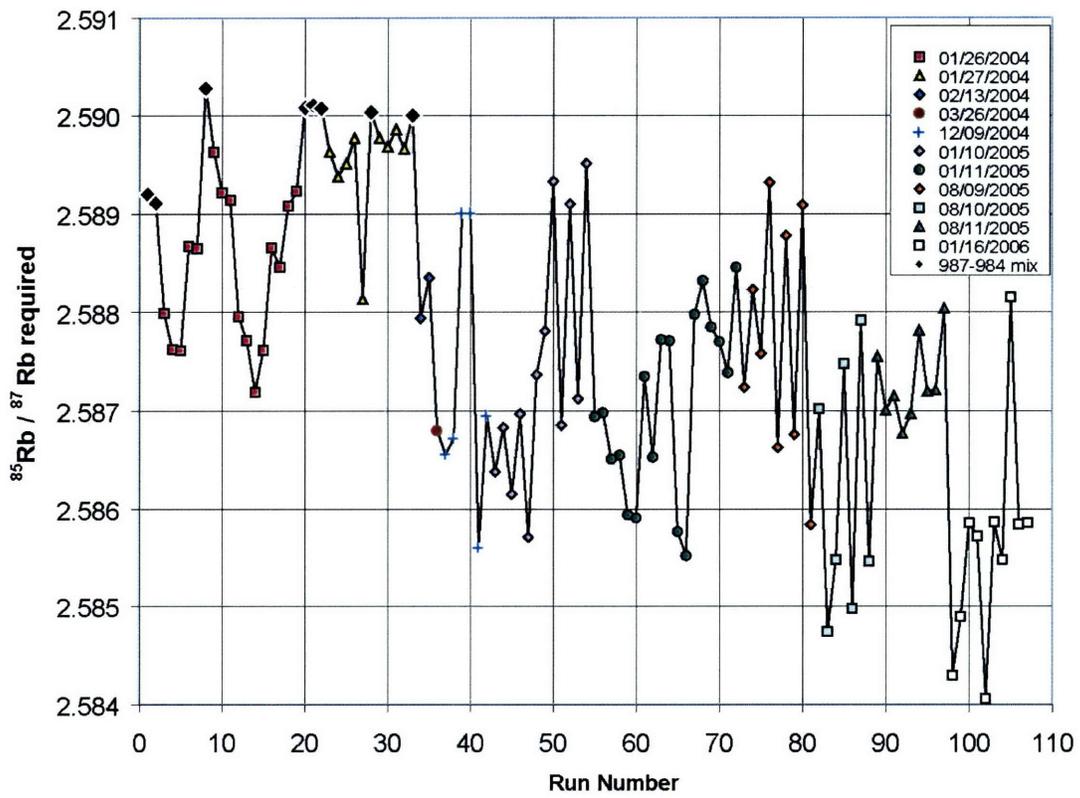


Fig. 2

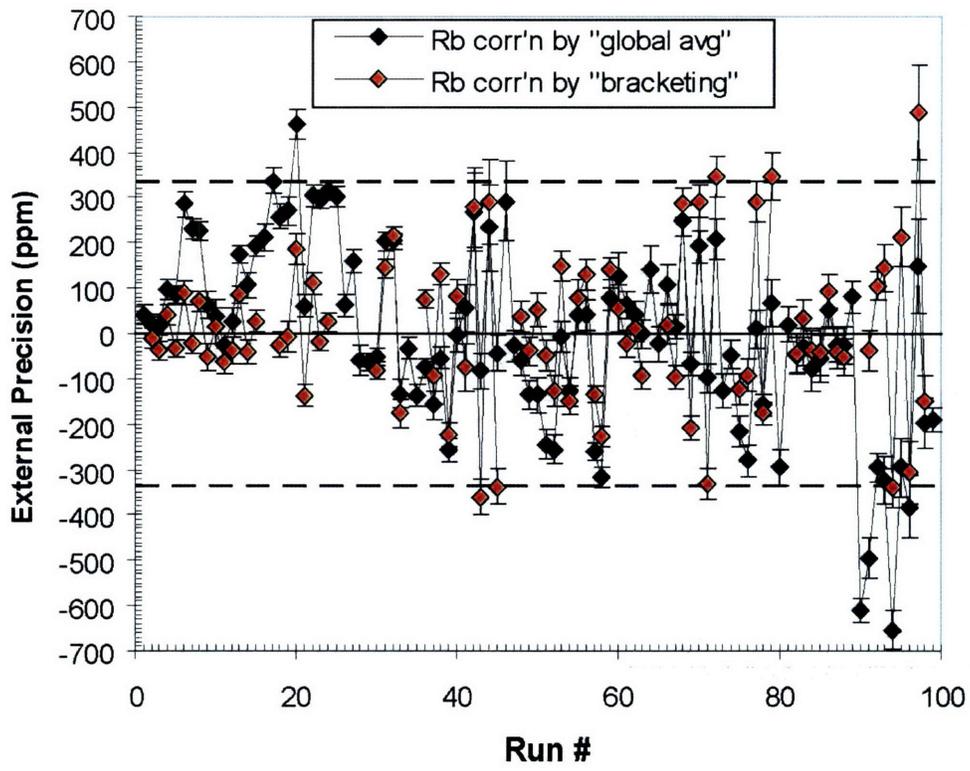
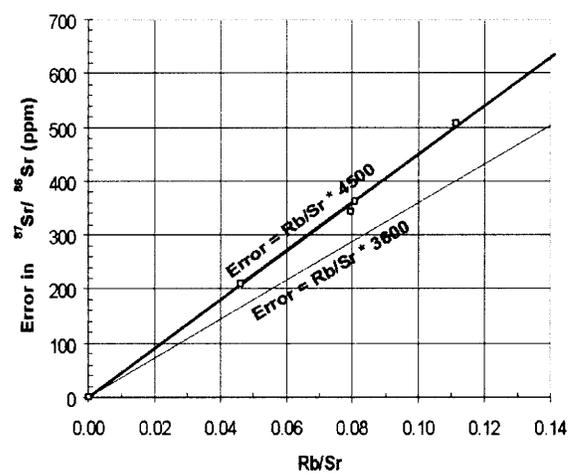


Fig.3



Supp Data Fig. 4

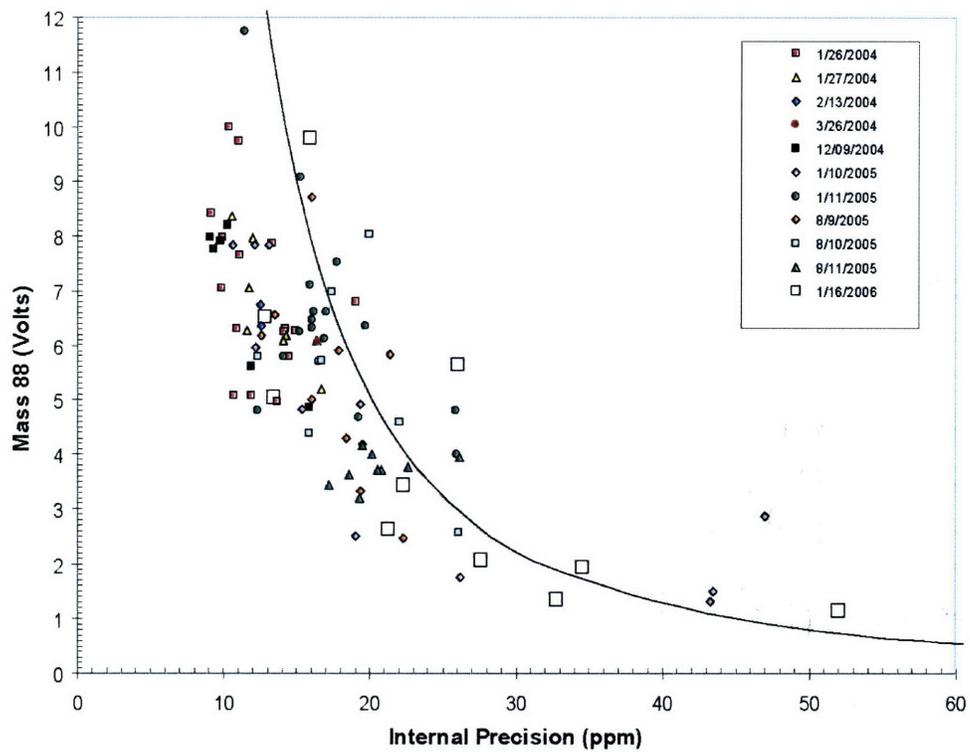


Fig. 5

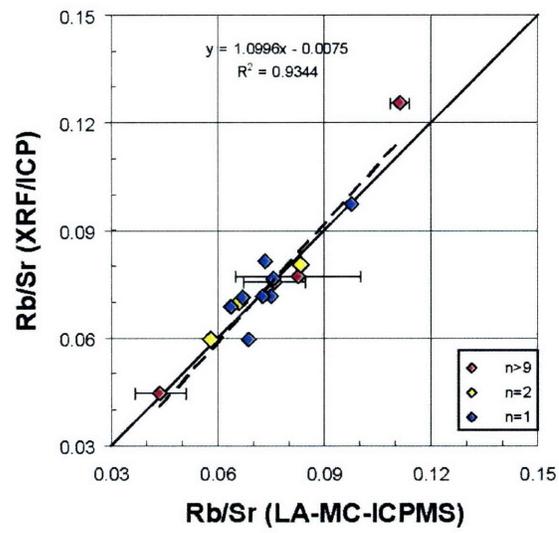


Fig. 6



## Chapter 2

# The return of subducted continental crust in Samoan lavas\*

### Abstract

Substantial quantities of terrigenous sediments are known to enter the mantle at subduction zones, but little is known about their fate in the mantle<sup>1</sup>. Subducted sediment may be entrained in buoyantly upwelling plumes and returned to the Earth's surface at hotspots<sup>2-5</sup>, but the proportion of recycled sediment in the mantle is small, and clear examples of recycled sediment in hotspot lavas are rare<sup>6,7</sup>. Here we report remarkably enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  isotope signatures in Samoan lavas from three dredge locations on the underwater flanks of Savai'i island, Western Samoa. The submarine Savai'i lavas represent the most extreme  $^{87}\text{Sr}/^{86}\text{Sr}$  isotope compositions reported for ocean island basalts to date. The data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Trace-element data show affinities similar to those of the upper continental crust—including exceptionally low Ce/Pb and Nb/U ratios<sup>8</sup>—that complement the enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  isotope signatures. The geochemical evidence from these Samoan lavas significantly redefines the composition of the EM2 (enriched mantle 2; ref. 9) mantle endmember, and points to the presence of an ancient recycled upper continental crust component in the Samoan mantle plume.

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## LETTERS

## The return of subducted continental crust in Samoan lavas

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Substantial quantities of terrigenous sediments are known to enter the mantle at subduction zones, but little is known about their fate in the mantle<sup>1</sup>. Subducted sediment may be entrained in buoyantly upwelling plumes and returned to the Earth's surface at hotspots<sup>2,3</sup>, but the proportion of recycled sediment in the mantle is small, and clear examples of recycled sediment in hotspot lavas are rare<sup>4,5</sup>. Here we report remarkably enriched <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd isotope signatures in Samoan lavas from three dredge locations on the underwater flanks of Savai'i island, Western Samoa. The submarine Savai'i lavas represent the most extreme <sup>87</sup>Sr/<sup>86</sup>Sr isotope compositions reported for ocean island basalts to date. The data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Trace-element data show affinities similar to those of the upper continental crust—including exceptionally low Ce/Pb and Nb/U ratios<sup>6</sup>—that complement the enriched <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd isotope signatures. The geochemical evidence from these Samoan lavas significantly redefines the composition of the EM2 (enriched mantle 2; ref. 9) mantle endmember, and points to the presence of an ancient recycled upper continental crust component in the Samoan mantle plume.

The Earth's mantle, as sampled by ocean island basalts erupted at hotspots, is chemically and isotopically heterogeneous. However, the origin of the geochemical heterogeneity of the mantle is not well understood. One model for the geochemical evolution of the mantle assumes that much of the chemical diversity is a result of subduction, a tectonic process that introduces enriched oceanic crust and compositionally heterogeneous sediment into a largely primitive (or slightly depleted) mantle<sup>5,10,11</sup>. Following subduction, these surface materials mix with a peridotitic mantle, thus imprinting their enriched chemical and isotopic signatures on its various domains. A number of isotopically distinct geochemical reservoirs, as sampled by ocean island basalts, have resulted from this process. The isotopic endmembers are often referred to as HIMU (high  $\mu = ^{238}\text{U}/^{204}\text{Pb}$ ), EM1 (enriched mantle 1) and EM2 (enriched mantle 2) and DMM (depleted mid-ocean-ridge basalt mantle)<sup>9</sup>. Although the most radiogenic Pb isotope ratios observed in the HIMU component have been proposed to result from a contribution of recycled oceanic crust<sup>5,12</sup>, most models for the creation of the EM1 and EM2 mantle reservoirs invoke a small portion of lithologically distinct sediments that have been recycled into the mantle<sup>5,13</sup>.

The volcanically active Samoan islands and seamounts define a hotspot track with a classical EM2 pedigree<sup>7,14,15</sup>. The first high-precision <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd measurements from Samoan lavas were interpreted as evidence of sediment recycling<sup>5</sup>. Recently,

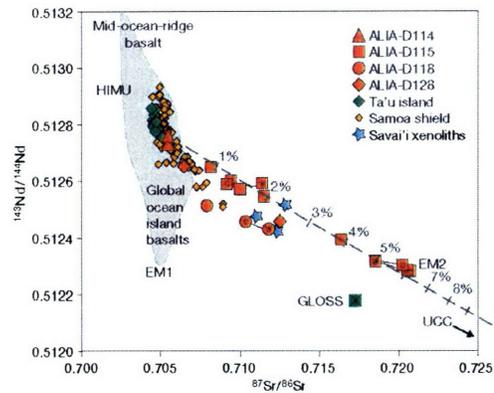
however, the proposed recycled sediment origin of the enriched Samoan basalts has been questioned (see Supplementary Discussion), and an alternative model favouring source enrichment by metasomatic processes was proposed<sup>7</sup>. The extreme isotopic and chemical enrichment in the new Samoan EM2 lavas exhibit distinctly continental fingerprints, and argue for a role for a component similar to ancient recycled upper continental crust (UCC) in the Samoan plume (see Supplementary Discussion for the ALIA 2005 cruise dredge locations and geochemical data).

The most isotopically enriched Samoan whole-rock <sup>87</sup>Sr/<sup>86</sup>Sr signature (0.720469, Mg# = 57.2) is recorded in a trachyandesite, dredge sample D115-21, which was taken from the southwestern flank of Savai'i. Clinopyroxene mineral separates from the same sample yielded an even higher <sup>87</sup>Sr/<sup>86</sup>Sr ratio (0.721630). A trachybasalt (D115-18) hosts the second-most-enriched <sup>87</sup>Sr/<sup>86</sup>Sr (0.718592, Mg# = 58.7), and clinopyroxene mineral separates from the sample also gave more enriched ratios (0.720232–0.720830). Six other lavas recovered in the same dredge also exhibit enriched <sup>87</sup>Sr/<sup>86</sup>Sr ratios (0.708175–0.716394, Mg# = 52.0–65.1). Dredge D118, located on the far western end of the Savai'i lineament, contained an alkali basalt with enriched <sup>87</sup>Sr/<sup>86</sup>Sr (0.710337, measured on fresh clinopyroxene). Dredge D128, taken on the northeastern flanks of Savai'i, yielded a transitional basalt with a high <sup>87</sup>Sr/<sup>86</sup>Sr ratio (0.712500, Mg# = 70.5) and several other basalts with less enriched <sup>87</sup>Sr/<sup>86</sup>Sr (0.706397–0.708170, Mg# = 61.2–63.9). Dredge D114, taken on the southwestern flanks of Savai'i, provided younger shield basalts of transitional chemistry and normal <sup>87</sup>Sr/<sup>86</sup>Sr (0.705422–0.705435, Mg# = 67.2 and 76.3).

The <sup>87</sup>Sr/<sup>86</sup>Sr isotopes in the basalts from all three ultra-enriched sampling localities are complemented by enriched (low) <sup>143</sup>Nd/<sup>144</sup>Nd and the lowest <sup>3</sup>He/<sup>4</sup>He ratios (4.31–4.93 Ra, or ratio to atmosphere) observed in Samoan basalts. Together, the new data extend the Samoan isotope array to a region outside the global ocean island basalt field (Fig. 1). Highly enriched EM2 signatures have previously been observed only in metasomatized xenoliths from Savai'i (<sup>87</sup>Sr/<sup>86</sup>Sr up to 0.712838; ref. 16), and the Samoan EM2 basalts provide the first evidence that the enriched component hosted in these xenoliths also occurs as erupted basalts. The enriched <sup>87</sup>Sr/<sup>86</sup>Sr and <sup>143</sup>Nd/<sup>144</sup>Nd isotope ratios, coupled with the low <sup>3</sup>He/<sup>4</sup>He, are consistent with a recycled UCC component in the mantle source of the Samoan EM2 basalts.

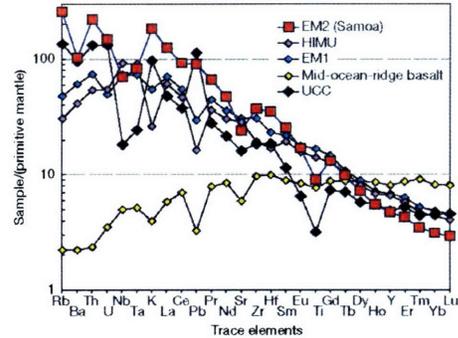
The UCC reservoir exhibits several diagnostic trace-element characteristics that can be useful for detecting its presence in Samoan EM2 lavas. Compared to ocean island basalt and mid-ocean-ridge basalt lavas, UCC displays exceptional depletion in Nb (and Ta), Ti and Eu, and enrichment in Pb (Fig. 2). Samoan basalts have

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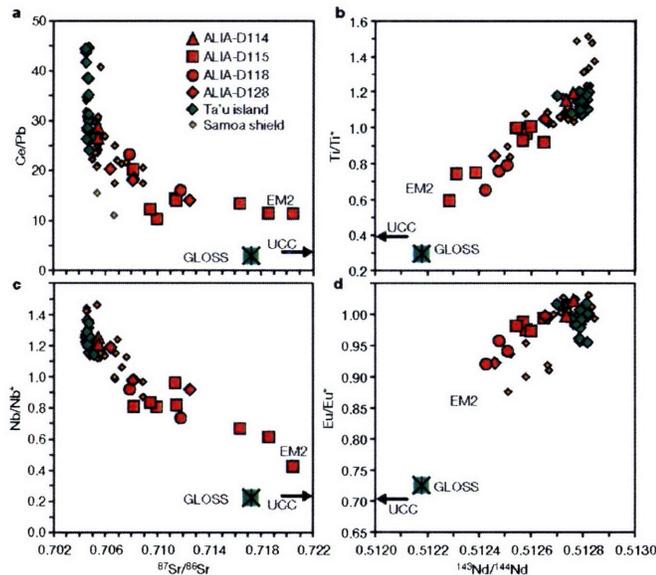
**Figure 1** |  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  isotope ratios of new enriched Samoan lavas. The values are compared with other Samoan shield basalts<sup>7</sup>, global ocean island basalt compositions and GLOSS (global subducting sediment)<sup>1</sup>. Analyses in which whole-rock (w) powders and clinopyroxene (c) analyses are performed on the same sample are connected by a tie-line. A model mixing line between depleted Ta'u peridotite and UCC is marked at 1% intervals, with increasing contribution from the latter component. The hypothetical UCC mixing endmember lies outside the figure. Approximately 5% UCC is required to produce the spidergram of sample D115-18 (see Supplementary Discussion), and ~6% is required to generate the  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  in D115-21.

trace-element characteristics that are increasingly similar to UCC with more enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  values (Fig. 3). Although the most isotopically depleted basalts from Samoa show slight positive anomalies in Nb and Ti, the magnitude of these anomalies decreases monotonically towards the most enriched Samoan EM2 basalts. Similarly, a correlation exists between greater



**Figure 2** | Primitive-mantle-normalized<sup>27</sup> trace-element patterns for the Samoan EM2 endmember. The EM2 spidergram is the most isotopically enriched Samoan lava, sample D115-21 ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.720469$ ,  $\text{Mg}\# = 57$ ). The other mantle endmembers (corrected to  $\text{Mg}\#$  numbers of 60–62) and UCC<sup>17</sup> are plotted for comparison. Similar to UCC, the Samoan EM2 lava exhibits large negative Ti and Nb (and Ta) anomalies and an excess of Pb (and K).

Pb enrichment and increasing isotopic enrichment in Samoan basalts. Importantly, the Eu anomaly is increasingly negative in the most isotopically enriched Samoan EM2 lavas (excluding basalts with  $\text{MgO} < 6.5$  wt%), and the Rb/Sr and U/Pb are too low in the lower (or middle) continental crust<sup>17</sup> to be consistent with the new Samoan Sr and Pb isotope data; these observations rule out the involvement of lower (or middle) continental crust. Furthermore, rare xenoliths with enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  from the subcontinental lithospheric mantle<sup>9</sup> suggest that this mantle domain can be isotopically enriched. However, the subcontinental lithospheric mantle does not appear to exhibit the trace-element anomalies observed in the most isotopically enriched Samoan lavas<sup>18</sup>. Instead, isotope ratios



**Figure 3** |  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios suggest the presence of a UCC component in Samoan EM2 lavas. The  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios are plotted against diagnostic UCC trace-element indicators in Samoan basalts: the more isotopically enriched Samoan basalts exhibit trace-element characteristics that are increasingly similar to UCC. UCC<sup>17</sup> plots outside the panels, and its trace-element and isotopic composition is indicated by the level (and direction) of the arrows. **a–c**, All submarine Savai'i samples are plotted, as are other Samoan shield lavas with  $\text{MgO} > 6.5$  wt%. **d**, Submarine Savai'i samples with  $\text{MgO} < 6.5$  wt% are excluded (avoiding possible effects of plagioclase fractionation). All trace-element data shown are by ICP-MS (inductively coupled plasma mass spectrometry). Element anomalies are calculated as follows (where subscript N means normalized to primitive mantle<sup>27</sup>):  $\text{Ti}/\text{Ti}^* = \text{Ti}_N / (\text{Nd}_N^{0.0558} \times \text{Sm}_N^{0.333} \times \text{Gd}_N^{0.722})$ ;  $\text{Nb}/\text{Nb}^* = \text{Nb}_N / (\text{Th}_N \times \text{La}_N)$ ;  $\text{Eu}/\text{Eu}^* = \text{Eu}_N / (\text{Sm}_N \times \text{Gd}_N)$ .

and trace-element anomalies (Nb, Ti, Eu and Pb) in Samoan basalts generate arrays that trend towards a composition similar to UCC.

We can exclude a shallow origin for the anomalous enrichment observed in the Samoan EM2 lavas. Owing to the close proximity of the Tonga trench, located only 120 km south of Savai'i, rapid cycling of sediment from the subduction zone into the Samoan plume was proposed as a mechanism for generating the extreme isotopic enrichment in Samoan lavas<sup>14</sup>. However, at the time that the submarine Savai'i lavas were erupted 5 Myr ago<sup>31</sup>, plate reconstructions indicate that the northern terminus of the Tonga trench was located 1,300 km to the west of Savai'i (ref. 19), and sediment input from the Tonga trench can be ruled out as a source of enrichment in these lavas.

Evidence from Pb isotopes suggests that it is unlikely that shallow-level contamination by modern marine sediments is responsible for the isotopic enrichment in the Samoan EM2 basalts. In  $\Delta^{207}\text{Pb}/^{204}\text{Pb} - \Delta^{208}\text{Pb}/^{204}\text{Pb}$  isotope space, Samoan basalts and global marine sediments exhibit non-overlapping fields with diverging trends (Fig. 4). Moreover, three composite cores taken from the Samoan region, and a single ferromanganese crust from the flanks of Savai'i, plot in the global marine sediment field and exhibit no geochemical relationship with the extremely enriched Samoan lavas. It is also unlikely that the Samoan plume has been contaminated by stranded continental crust, such as was found beneath the Kerguelen plateau<sup>20</sup> and the southern Mid-Atlantic Ridge<sup>21</sup>, or by ancient limestone blocks like those discovered in the Romanche fracture zone<sup>22</sup>. The tectonic history of the Samoan region places it neither at the locus of continental rifting, which was responsible for the marooned Kerguelen and southern Atlantic continental blocks, nor in proximity to any Pacific fracture zones<sup>23</sup>.

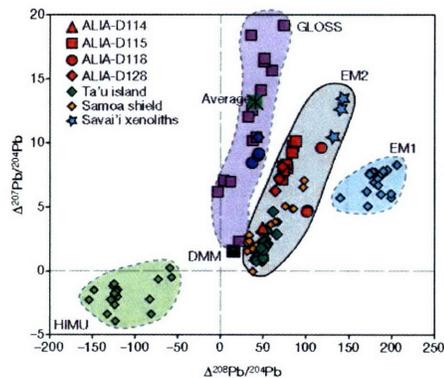
Large quantities of sediment derived from UCC have entered the mantle at subduction zones over geologic time<sup>1</sup>, and such a reservoir is ideally suited as an enriched source for the Samoan plume. The array formed by the Samoan EM2 basalts in  $^{143}\text{Nd}/^{144}\text{Nd}$ - $^{87}\text{Sr}/^{86}\text{Sr}$  isotope space is anchored on the depleted end by basalts from

Ta'u, one of the youngest, easternmost Samoan islands. The  $^{143}\text{Nd}/^{144}\text{Nd}$ - $^{87}\text{Sr}/^{86}\text{Sr}$  array suggests mixing between this dominant, slightly depleted Ta'u component and a rare, enriched component that exhibits isotope and trace-element characteristics similar to UCC. The proportion of the enriched component in the Samoan EM2 lavas can be estimated by calculating trace-element concentrations in the depleted Ta'u mantle and mixing this composition with UCC (see Supplementary Discussion). A contribution of 5% UCC to the depleted Ta'u mantle generates a composition that, after mixing and melting, produces a trace-element pattern similar to that observed in Samoan EM2 sample D115-18 (with  $^{87}\text{Sr}/^{86}\text{Sr}$  of 0.718592). Fixing the proportions of the depleted and UCC components in the Samoan EM2 source in this way then defines the  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  isotopic composition of this material as 0.7421 and 0.5117, respectively. The most isotopically enriched Samoan lavas have higher  $^{87}\text{Sr}/^{86}\text{Sr}$  than the average continental crust inferred from suspended river sediments ( $^{87}\text{Sr}/^{86}\text{Sr} \approx 0.716$ ; ref. 24) and global marine sediments ( $\sim 0.717$ ; ref. 1), values that are biased towards younger continental crust. However, composites of directly sampled ancient continental shield rocks show isotopically enriched compositions<sup>25</sup> that bracket the calculated composition of the recycled UCC sediment in the Samoan mantle. The new ultra-enriched EM2 lavas suggest an unusually enriched recycled protolith in the Samoan mantle.

Despite the large volumes of sediment entering the mantle at subduction zones (estimated at 0.5–0.7 km<sup>3</sup> yr<sup>-1</sup>; ref. 1), isotopic signatures associated with recycled UCC are rare in ocean island basalts<sup>26</sup>. This enriched component is also uncommon in the Samoan plume, where the highly enriched Samoan EM2 lava D115-18 is calculated to have only 5% recycled UCC (and 95% depleted Ta'u source), and 90% of the remaining Samoan basalts exhibit depleted  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios (>0.512638). In addition to being rare in other ocean island basalts, recycled UCC may exist in low abundance in the Samoan plume. The reason for this may be that most subducted sediment melts and is rapidly returned to the surface in subduction zone volcanoes, or is simply scraped off onto the forearc and is never subducted. Alternatively, if a significant portion of UCC has been subducted over the past 4 Gyr (0.5–0.7 km<sup>3</sup> yr<sup>-1</sup>) and has survived subduction zone melting, the resulting accumulated reservoir in the mantle will constitute only  $\sim 0.15\%$  of its mass. Such a small reservoir may be diluted by the ambient mantle after convective stirring, a mechanism that efficiently attenuates mantle heterogeneities. Therefore, recycled crustal signatures can be greatly diluted and difficult to detect. By contrast, the recycled UCC component in the Samoan plume is an anomalous survivor in a chaotic mantle.

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**Figure 4** |  $\Delta\text{Pb}$  isotope compositions of Samoan lavas and marine sediment samples indicate that the Samoan EM2 lavas are not contaminated with modern marine sediment. Samoan basalts show no overlap with oceanic sediments contributing to GLOSS<sup>1</sup> (purple squares, where the GLOSS average composition is represented by a green crossed 'average' square), composite sections from three sediment cores taken in the Samoan region (blue circles), and a Samoan ferromanganese rind (blue diamond). Pb-isotope data for mantle endmembers DMM (using data from mid-ocean ridge basalt normal segments, as defined and catalogued by ref. 28), EM1 (Pitcairn), and HIMU (Mangaia and Tubuai) are from the literature (see Supplementary Information for reference citations). The use of  $\Delta\text{Pb}$  isotope notation<sup>29</sup> to identify sediment components in ocean island basalts is discussed elsewhere<sup>30</sup>. All Samoan data shown are determined using a high-precision Tl-spike protocol<sup>1</sup>.

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- Author Contributions** M.G.J. performed most of the experimental work, developed the model and wrote the paper. S.R.H. and H.S. conceived the project, and were co-chiefs of the ALIA expedition. A.A.P.K. and J.K. were responsible for the cruise bathymetry, and A.A.P.K. greatly improved the figures. J.B. and M.K. provided analytical assistance and access to facilities. A.A.P.K., J.B., J.K. and J.A.R. helped with sample preparation. All authors participated in the discussion and interpretation of results, and commented on the manuscript.
- Author Information** Reprints and permissions information is available at [www.nature.com/reprints](http://www.nature.com/reprints). The authors declare no competing financial interests. Correspondence and requests for materials should be addressed to M.G.J. ([mjackson@whoi.edu](mailto:mjackson@whoi.edu)).

## SUPPLEMENTARY INFORMATION

Contents of Supplementary Information (in a merged PDF titled jacksonsuppinfo.pdf;  
Size: 0.7 MB total):

1. Supplementary Discussion
2. Supplementary Figure 1 and caption
3. Supplementary Figure 2 and caption
4. Supplementary Figure 3 and caption
5. Supplementary Figure 4 and caption
6. Supplementary Table 1
7. Supplementary Table 2
8. Supplementary Table 3

The file contains a Supplementary Discussion that gives a model of the Samoan EM2 source, including calculations and assumptions. The file also contains a Supplementary Table 1 and a Supplementary Table 2 that together provide the model parameters used to calculate the depleted Ta'u and EM2 Samoan sources, respectively. The file also contains a Supplementary Table 3 that provides the new Samoan geochemical data. Four supplementary figures also are included in the file: Supplementary Figure 1 shows new helium isotope data from the remarkably enriched Samoan lavas; Supplementary Figure 2 is an expanded view of main text Figure 1, and shows the composition of the Samoan lavas in the context of individual samples from the upper continental crust; Supplementary Figure 3 compares the spidergram of a model melt of the EM2 source with the spidergram of an isotopically-enriched Samoan EM2 lava. Supplementary Figure 4 provides a model that explains the correlation between SiO<sub>2</sub> and <sup>87</sup>Sr/<sup>86</sup>Sr in lavas from dredge ALIA D115.

## 1. Supplementary Discussion

**Recycled UCC in the Samoan EM2 mantle.** The standard model for the genesis of the EM2 reservoir suggests a role for recycled sediment, probably terrigenous in origin<sup>1-7</sup>. However, a number of problems with the standard model were outlined in ref. 8, and an alternative origin of the Samoan EM2 source (recycled, metasomatized lithosphere) was proposed. In light of the remarkably enriched Samoan lavas presented in this paper, we reconsider the arguments against a recycled terrigenous component in the EM2 plume. We demonstrate that recycled upper continental crust (UCC) can have an important role in the origin of the EM2 source. However, we maintain that it is (most likely) not “upper continental crust” (UCC) that is being recycled, but sediment with a composition resembling that of UCC. While the sediment recycling model presented here does not invalidate the metasomatic model for the lavas presented earlier<sup>8</sup>, the new lavas presented in this study are consistent with a recycled sediment component.

The problems with recycled marine sediment in the Samoan EM2 mantle source (as outlined in ref. 8) are as follows. The variation in  $^{187}\text{Os}/^{188}\text{Os}$  in Samoan lavas is not consistent with mixing between recycled marine sediment and depleted MORB mantle (DMM). Such a scenario would require 35% marine sediment in the EM2 source, a quantity not observed in the trace element patterns of Samoan basalts. Additionally, the smooth trace element spidergrams observed in Samoan EM2 lavas were suggested to be inconsistent with a sediment component in the Samoan EM2 source. This is because, unlike Samoan EM2 lavas presented earlier<sup>8</sup>, marine sediments exhibit “jagged” spidergrams, marked by large negative anomalies for Nb (and Ta), Ti and Eu and large positive anomalies for Pb (and K). It was also noted that EM2 lavas exhibit negative Ba-

anomalies, a feature not characteristic of marine sediments<sup>8</sup>. Furthermore, because EM2 lavas exhibit high  $^3\text{He}/^4\text{He}$  ratios ( $\sim 8$  Ra, ratio to atmosphere, or  $\sim 1.4 \times 10^{-6}$ ), a trait not shared with the low  $^3\text{He}/^4\text{He}$  ratios in sediments (0.05-1 Ra)<sup>9</sup> and continental crust (0.007 Ra)<sup>10</sup>, it was suggested that the Samoan EM2 source does not host a sediment component<sup>8</sup>. Finally, Pb-isotope composition of modern marine sediment was observed to be unsuitable as an endmember for the enriched Samoan basalts<sup>8</sup>. Modern marine sediments exhibit  $^{208}\text{Pb}/^{204}\text{Pb}$  ratios that are too low at a given  $^{206}\text{Pb}/^{204}\text{Pb}$  to serve as mixing endmembers for the Samoan EM2 lavas.

Below, we suggest that, instead of a marine sediment composition such as GLOSS (Global Subducting Sediment<sup>11</sup>), sediment with a composition like UCC (a composition here approximated by UCC from ref. 12) is more suitable for generating a source sampled by the most isotopically-enriched Samoan lavas. A small portion ( $\sim 5\%$ ) of (sediment with the composition of) UCC mixed with a depleted Samoan plume component generates a peridotite that, when melted, produces a spidergram similar to that observed in the most isotopically-enriched Samoan lavas. Like Samoan EM2 lavas, UCC exhibits a negative Ba anomaly, a feature not shared with GLOSS<sup>11</sup>. Furthermore, published Os-isotopes<sup>8</sup> in enriched Samoan EM2 lavas are not inconsistent with a UCC component in the plume. The Os-isotope signature in Samoan basalts is not likely a result of mixing between DMM ( $^{187}\text{Os}/^{188}\text{Os} = 0.125$ ) and marine sediment; however, a mixture of UCC ( $^{187}\text{Os}/^{188}\text{Os} = 1.05$ ,  $[\text{Os}] = 30 \text{ ppt}^{13}$ ) and a depleted Samoan plume component generates a mixing trend that describes the Os-isotope data for Samoan shield basalts (excluding samples with  $< 100 \text{ ppt Os}$ , and assuming measurement precision of  $\pm 1.5\%$  in data from ref. 8). This mixing scenario uses measured Ta<sup>u</sup>  $^{187}\text{Os}/^{188}\text{Os}$  ratios

of 0.129<sup>8</sup> and assumes that the Ta'u peridotite source also has Os concentrations that are depleted relative to primitive mantle (>3,900 ppt<sup>14</sup>). Moreover, new helium isotope data for Samoan EM2 lavas exhibit low ratios (<5 Ra) that plot on a trajectory that trends to UCC, and are consistent with the existence of a recycled UCC component in the Samoan plume (Supplementary Figure 1). Finally, our model for the formation of the EM2 mantle agrees with Pb-isotope constraints. UCC exhibits a large range of Pb-isotope values, and the Samoan EM2 lavas exhibit Pb-isotope values that plot in the field previously defined for UCC<sup>15</sup>.

**Model for the EM2 source.** In order to model the generation of the mantle source sampled by the new Samoan EM2 lavas, we take advantage of the array formed by Samoan basalts in <sup>87</sup>Sr/<sup>86</sup>Sr–<sup>143</sup>Nd/<sup>144</sup>Nd isotope space. The array suggests that the most enriched Samoan basalts were formed as products of mantle-mixing between a depleted component (here represented by the mantle source sampled by the isotopically-depleted and remarkably homogeneous lavas from Ta'u island) and a component similar to UCC (Supplementary Figure 2). In order to determine the proportion of UCC in the Samoan plume, we first calculate a trace element peridotite source for the depleted Ta'u endmember that is consistent with the radiogenic isotopes of Nd, Hf, Sr and Pb. We then determine the amount of UCC that must be added to the depleted Ta'u peridotite source so that the final mixture, an enriched peridotite, can be melted to generate a spidergram similar to the new Samoan EM2 lavas.

In the following modeling exercise, we generate a model for the EM2 source that describes the array formed by the most isotopically-enriched submarine Savai'i lavas from dredge D115, and we make no attempt to model the other components previously

identified<sup>8</sup> in lavas from the Samoan hotspot. We emphasize that the model is just one possible model that is consistent with the isotopes and trace elements in Samoan EM2 lavas. The model is presented only to demonstrate that recycling sediment into Samoan EM2 lavas is possible.

**Composition of the depleted Ta'u peridotite source.** Following ref. 8, we generate an average olivine fractionation corrected trace element budget for Ta'u lavas (Supplementary Figure 3 and Supplementary Table 1). In order to determine a trace element mantle source sampled by the average Ta'u lava composition, we first assume a peridotite source lithology, an aggregated fractional melting model, and we adopt mineral-melt partition coefficients from ref. 16 (Supplementary Table 1). We also assume a two-stage isotope model for the evolution of the depleted Ta'u source, and that this differentiation event of a primitive mantle composition occurred at 1.8 Ga. This age is commonly quoted as the average mantle differentiation age<sup>17</sup>, and is an age that is consistent with the array formed by Ta'u lavas in  $^{207}\text{Pb}/^{204}\text{Pb}$  vs.  $^{206}\text{Pb}/^{204}\text{Pb}$  isotope space (see Fig. 7 in ref. 8). Given these assumptions, a Ta'u source is calculated so that parent-daughter ratios—Sm/Nd and Lu/Hf—will generate the present-day depleted  $^{143}\text{Nd}/^{144}\text{Nd}$  (Ta'u average is 0.512789<sup>8</sup>) and  $^{176}\text{Hf}/^{177}\text{Hf}$  (Ta'u average is 0.282987<sup>18</sup>) isotopes measured in Ta'u lavas given the two-stage isotope model. The Ta'u source is not very sensitive to its age of formation: ages of 1.0-2.5 Ga require only small variations in melting—4.5 to 5.5%—to generate the average Ta'u lava. However, assuming a formation age of 1.8 Ga, a 5.1% melt (with ~50.6% garnet melting) of the hypothetical Ta'u source will generate a model melt with a spidergram that is both identical to the average measured Ta'u lava and consistent with isotopic constraints. These highly

specific melting parameters are presented only to generate an average Ta'U source that is consistent with isotopic constraints, and by presenting them we are not suggesting such precise knowledge of the actual mantle "plumbing" beneath Ta'U island.

The  $^{87}\text{Sr}/^{86}\text{Sr}$  of primitive mantle is unconstrained, so it is not possible to evaluate whether the  $^{87}\text{Sr}/^{86}\text{Sr}$  measured in Ta'U lavas is enriched or depleted relative to primitive mantle. However, the average  $^{87}\text{Sr}/^{86}\text{Sr}$  in Ta'U lavas (0.704650) and the Rb/Sr of the Ta'U source (0.0268) can be modeled as having evolved from primitive mantle at 1.8 Ga if the present-day primitive mantle  $^{87}\text{Sr}/^{86}\text{Sr}$  is 0.70508, a value that is in the range typically assigned to primitive mantle. In order for the two-stage isotope model to produce the observed average Ta'U  $^{206}\text{Pb}/^{204}\text{Pb}$  (19.271) and  $^{206}\text{Pb}/^{204}\text{Pb}$  (15.597) (excluding T14 and considering only T1-spiked data<sup>8</sup>), the proportion of sulfide<sup>19</sup> in the Ta'U source mineralogy is adjusted to obtain an appropriate parent-daughter U/Pb source ratio. The resulting calculated Th/Pb source ratio is within error of the Th/Pb ratio required to produce the average Ta'U  $^{208}\text{Pb}/^{204}\text{Pb}$  composition (39.424) in 1.8 Ga. While non-unique (a different melt model could be chosen or the Ta'U source formation age may be different, etc.), the trace element source calculated from Ta'U lavas is consistent (within the uncertainties of the data) with constraints from radiogenic isotopes.

**Generation of a source for the Samoan EM2 lavas.** Determining the precise nature of the recycled component contributing to the enrichment in the Samoan plume is not straightforward. UCC rocks and individual marine sediment cores show a large degree of trace element heterogeneity that varies considerably with geography and provenance<sup>11,12</sup>, and the composition of a recycled sediment may depend on the geography of the subduction zone. More problematic is the issue of temporal variability<sup>20,21</sup>: the trace

element budgets of sediments encountered in modern oceans may be poor analogues for the sediments subducted in the past. Finally, the poorly constrained processes operating in subduction zones, including fluid loss and/or melting of sediments, may modify the composition of the subducted sediment component<sup>22-25</sup>.

With these caveats aside, we take a simple approach. We assume that the compositions of subducted material are conserved in the subduction zone. Perhaps fluids and melts are removed from the slab and inoculated into the mantle wedge; but the resulting slab residue and the fertilized mantle wedge may in some cases stay together as a package. This package is what enters the general circulation, to eventually be remixed to end up looking like "closed system" slab recycling. While the composition of the original material sent into the subduction zone is underconstrained, we explore whether or not a modern UCC composition, when mixed with the depleted Ta'u source, can generate a suitable source for the Samoan EM2 lava D115-18. We assume that after the depleted Ta'u source mixed completely with the enriched UCC component, the resulting peridotite (treated here as a single lithology) was melted in a modal, aggregated fractional melting system. The modal abundances of the mantle phases, the contribution of UCC to the depleted Ta'u source, and the degree of melting (and the proportion of garnet and spinel melting) of the resulting mixture (the EM2 source) are all adjusted to generate a trace element spidergram that is similar to the enriched Samoan EM2 basalts. The combination of these parameters that generates a "best-fit" spidergram to the Samoan EM2 lava composition is given in Supplementary Table 2. The agreement between the model spidergram and the spidergram for Samoan EM2 lava D115-18 is optimized (the

fit is within 14% for all the trace elements considered, or ~11% if U is excluded) when the contribution of the UCC component is ~5%.

Having fixed the proportions of UCC (5%) and the depleted Ta'u source (95%) in the EM2 source sampled by lava D115-18 (with  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  of 0.718592 and 0.512314, respectively), we use trace element budgets and isotopic constraints to calculate the  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  of the UCC component. Assuming a measured isotopic ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.704650$  and  $^{143}\text{Nd}/^{144}\text{Nd} = 0.512789$ ) and calculated trace element composition (Supplementary Table 1) of the Ta'u source and a trace element composition of UCC<sup>12</sup>, we calculate the Sr and Nd isotopic composition of the UCC component in the Samoan plume to be 0.7421 and 0.5117, respectively. The  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  values calculated for the UCC endmember in the Samoan plume are within the range of values measured in ancient UCC shield rocks<sup>26</sup> (Supplementary Figure 2).

While the addition of UCC to the depleted Ta'u source generates a spidergram that is a close match to Samoan EM2 lavas, the fit is not perfect. In particular, the element that exhibits the least perfect fit to the data is U. Samoan EM2 lavas have lower U concentrations than the model result, and the disagreement may be a result of U-loss during weathering (note the high Th/U in sample D115-18; Supplementary Table 3).

**Addressing contamination by marine sediment.** Implicit in the model of the EM2 source is the assumption that the sediment signature in the Samoan EM2 lavas is a primary mantle signal and not a result of shallow-level sediment contamination. A plot of  $\Delta^{207}\text{Pb}/^{204}\text{Pb} - \Delta^{208}\text{Pb}/^{204}\text{Pb}$  (see Main Text Figure 4) indicates that sediment contamination is not an issue. Further evidence comes from  $^{87}\text{Sr}/^{86}\text{Sr}$  measurements in sediments from the Samoan region, which exhibit  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (0.70614 to 0.70824)

that are much lower than the most enriched Samoan EM2 lavas (0.72047 in whole-rock sample D115-21). It is very difficult to generate the high  $^{87}\text{Sr}/^{86}\text{Sr}$  observed in the new Samoan EM2 lavas by sediment contamination. For example, the GLOSS average  $^{87}\text{Sr}/^{86}\text{Sr}$  composition is only 0.7173 (327 ppm Sr)<sup>11</sup>, a value that is much lower than the most isotopically-enriched lavas from Samoa. Even if the most isotopically-enriched marine sediment ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.73493$ , 251 ppm Sr) in the compilation from ref. 11 were added to the least isotopically-enriched submarine Savai'i lava ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.705435$ , 374 ppm Sr), over 60% sediment assimilation would be required to generate the most radiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$  observed in the Samoan lavas. Such large quantities of sediment are not visible in the trace element spidergrams of the most enriched Samoan lavas.

In general, Samoan lavas do not show a correlation between  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $\text{SiO}_2$  (Supplementary Figure 4). However,  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in whole-rock lavas from dredge D115 do correlate with  $\text{SiO}_2$ . Having ruled out sediment assimilation using Sr and Pb isotope data, the  $^{87}\text{Sr}/^{86}\text{Sr} - \text{SiO}_2$  array formed by ALIA D115 samples is interpreted to be a result of magma mixing between an evolved (high  $\text{SiO}_2$ ), isotopically-enriched magma and a less evolved, less isotopically-enriched magma. Other major and trace element data are consistent with this scenario (see Supplementary Figure 4). Perhaps the isotopically-enriched magmatic endmember evolved by crystal fractionation in a magma chamber and, just before eruption, mixed with a later pulse of a less evolved, less isotopically-enriched magma. Magma mixing is not an uncommon phenomenon; heterogeneous  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios recorded in olivine-hosted melt inclusions from individual Samoan basalt samples suggest that mixing of magmas from isotopically-distinct sources is not uncommon in Samoa<sup>27</sup>.

**Data sources for HIMU and EM1 lavas.** In Fig. 2 and Fig. 4 of the main text, trace element and Pb-isotope data for the HIMU endmember are from samples collected at Mangaia and Tubuai islands and are reported in refs. 41 and 42. Trace element and Pb-isotope data from the EM1 endmember are from samples collected at Pitcairn and include data from ref. 43 and unpublished data (S.R. Hart and E.H. Hauri). Only data from the freshest, most isotopically-extreme samples were used.

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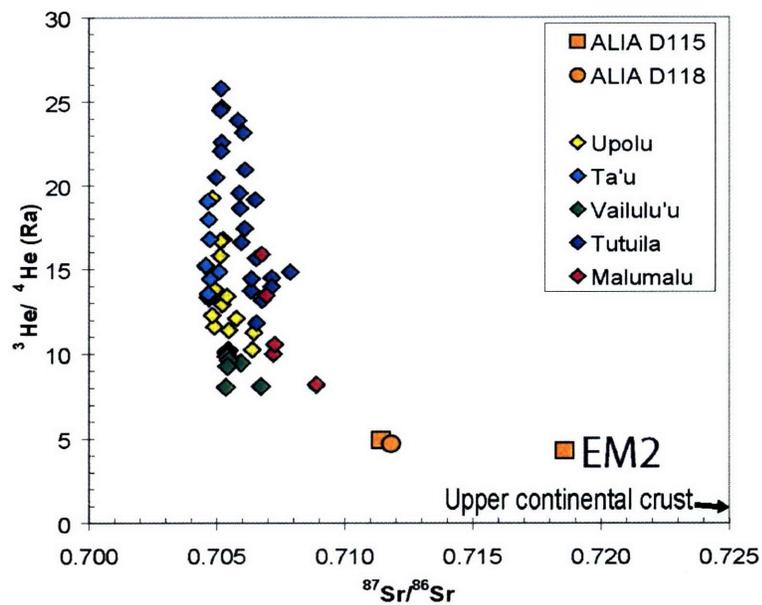
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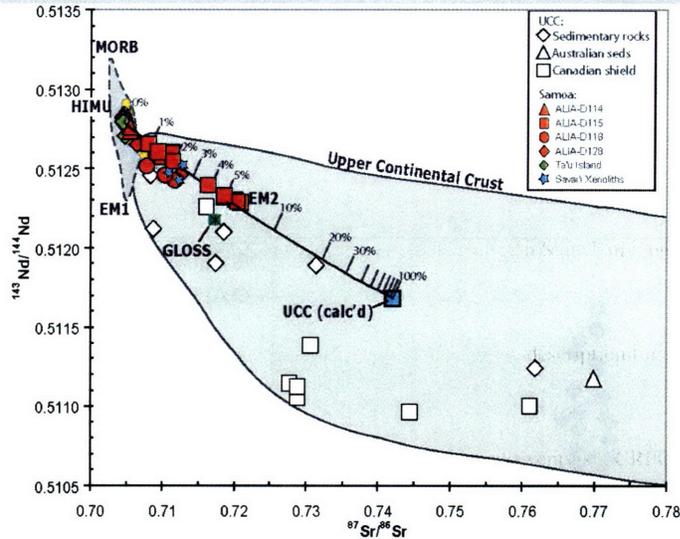
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## 2. Supplementary Figures

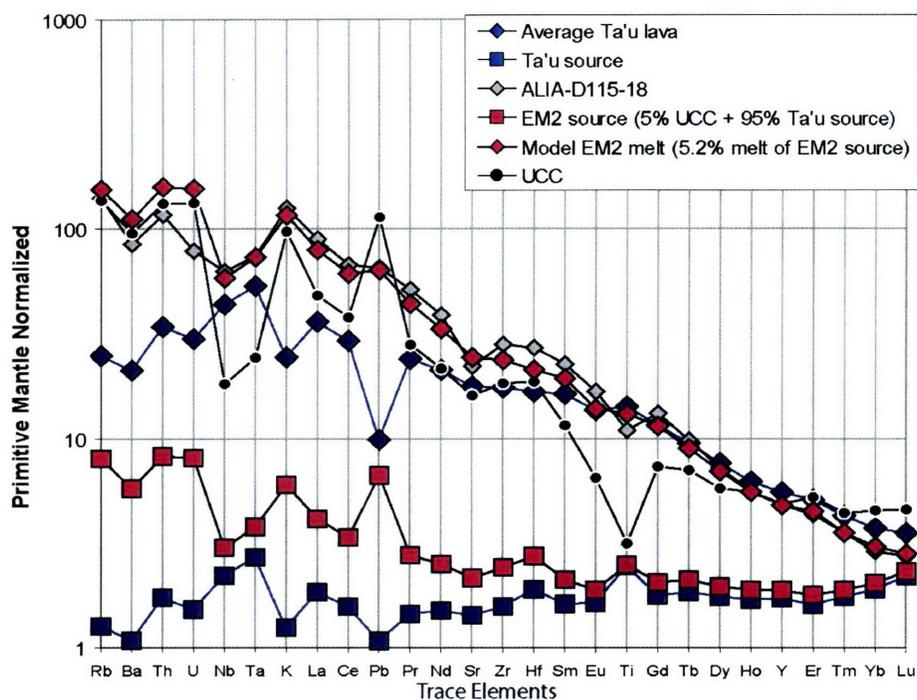


**Supplementary Figure 1**  $^{87}\text{Sr}/^{86}\text{Sr}$  vs.  $^3\text{He}/^4\text{He}$  isotopes in Samoan basalts.

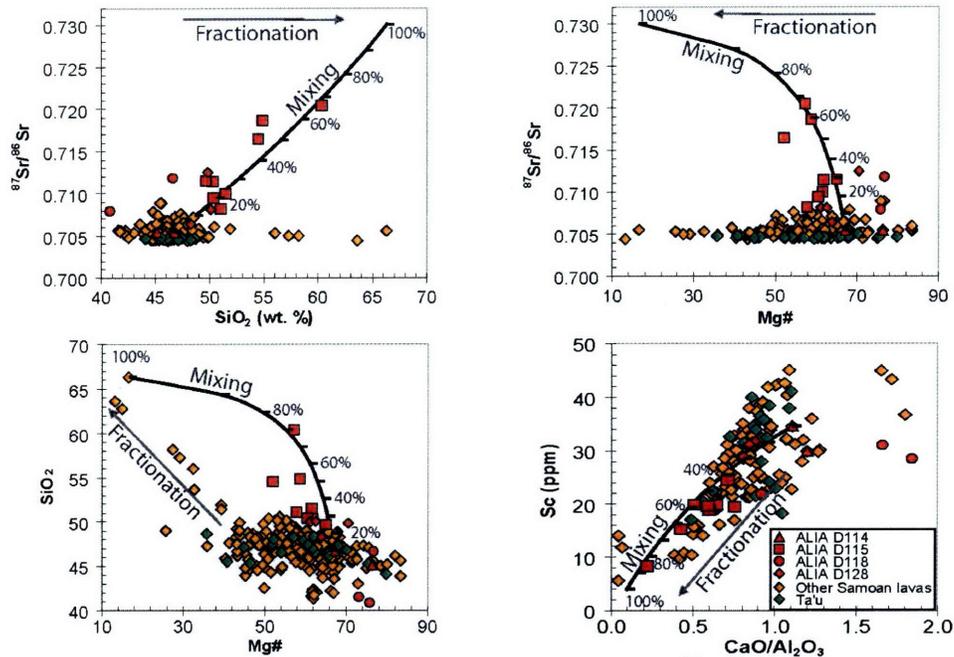
New  $^3\text{He}/^4\text{He}$  ratios in the enriched Samoan EM2 basalts are low (<5 Ra), and plot on a trajectory that extrapolates to a UCC component with low  $^3\text{He}/^4\text{He}$  (0.007 Ra)<sup>10</sup> and high  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7421). Data for the other Samoan islands and seamounts are from refs 8 and 40.



**Supplementary Figure 2**  $^{87}\text{Sr}/^{86}\text{Sr}$  vs.  $^{143}\text{Nd}/^{144}\text{Nd}$  of Samoan shield basalts compared with samples from upper continental crust (UCC) from ref. 26. The grey field (with solid outline) encompasses Sr and Nd isotopic analyses on UCC samples<sup>26</sup>, and extends outside of the figure to values of 1.1862 and 0.51023, respectively. The grey field (with dashed outline) encompasses the non-Samoan OIB field. The orange field describes Samoan shield basalts observed in previous studies; the new Samoan EM2 basalts (red) extend well into the field for UCC. Tic marks represent addition of the UCC component to the depleted peridotite Ta'u source. The isotopic composition of the UCC endmember in the Samoan plume is calculated to have  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  of 0.7421 and 0.5117, respectively (light blue box, see Supplementary Discussion), and lies in the range of values previously measured on UCC rocks<sup>26</sup>. 5% of this hypothetical UCC composition is required to generate the spidergram of Samoan sample D115-18 (and an estimated 6% UCC is required to generate the  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios in the most enriched cpx from sample D115-21).



**Supplementary Figure 3** Primitive mantle<sup>28</sup> normalized trace element patterns for Samoan basalts and peridotite sources. These compositions are used in a sediment recycling model for generating a Samoan EM2 mantle source. Ta'u lavas are assumed to be melts of a depleted plume component, and the average, olivine-fractionation corrected Ta'u lava composition is plotted. A Ta'u source for the average Ta'u lava is calculated to satisfy isotopic constraints. 5% of UCC<sup>12</sup> is added to 95% of the Ta'u source to make the EM2 source sampled by the Samoan lava D115-18, and the model melt of the EM2 peridotite source is plotted. The spidergram of the model melt is similar to sample D115-18 (the second most isotopically-enriched Samoan sample). The most isotopically-enriched lava, D115-21, is too evolved to reliably reconstruct its trace element composition. Plotted compositions can be found in Supplementary Tables 1 & 2.



**Supplementary Figure 4** A model for the co-variation of  $\text{SiO}_2$  and  $^{87}\text{Sr}/^{86}\text{Sr}$  in ALIA D115 samples.

Samoan lavas do not exhibit a one-to-one relationship between  $\text{SiO}_2$  (or  $\text{Mg}\#$ ) and  $^{87}\text{Sr}/^{86}\text{Sr}$ ; evolved Samoan lavas exhibit both high (up to 0.7205) and low (0.7044)  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios. However, lavas from dredge ALIA D115 exhibit a correlation between  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $\text{SiO}_2$ . We explain this correlation in ALIA D115 lavas as a result of mixing between a highly evolved, isotopically-enriched magma and a less evolved, less isotopically-enriched magma. In this scenario, the mixing endmembers are a Savai'i submarine lava sample (ALIA D114-01; Supplementary Table 3) and a highly evolved lava from Tutuila (91TP-128;  $\text{SiO}_2 = 66.30$  wt.%,  $\text{MgO} = 0.41$  wt.%,  $\text{FeO} = 4.06$  wt.%,  $\text{CaO} = 1.62$  wt.%,  $\text{Al}_2\text{O}_3 = 17.03$  wt.%,  $\text{Sr} = 395$  ppm; Natland unpubl. data). The measured  $^{87}\text{Sr}/^{86}\text{Sr}$  of the Tutuila lava is 0.705535<sup>8</sup>. However, for the sake of argument, this sample is given an  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio of 0.730 in the mixing model in order to determine whether mixing between an evolved magma and a less evolved magma can generate a mixing array consistent with the data from dredge ALIA D115. Sc was not measured on the Tutuila lava, but its concentration is assumed to be 4 ppm. The reduction of Sc and  $\text{CaO}/\text{Al}_2\text{O}_3$  (including reduced  $\text{Mg}\#$ 's and increased  $\text{SiO}_2$ ) in the most isotopically-enriched lavas from ALIA D115 is consistent with cpx fractionation. Crystal fractionation of an isotopically-enriched magma followed by mixing with a less evolved, less isotopically-enriched magma describes the geochemical data from dredge ALIA D115.

## 3. Supplementary Tables

Supplementary Table 1. Calculation of the depleted component in the Samoan plume, using lavas from Ta'u island

	Primitive mantle <sup>1</sup>	Avg. Ta'u lava <sup>2</sup> (olivine corr'd to Fo <sub>93</sub> )	Variability in 18 Ta'u lavas (1σ, std. dev., %)	Bulk partition coeffs <sup>3</sup> (garnet field)	Bulk partition coeffs <sup>3</sup> (spinel field)	Depleted Ta'u source <sup>4</sup> (calculated from avg. Ta'u lava)
Rb	0.6	14.9	16.5	0.000144	0.000124	0.759
Ba	6.6	140	11.9	0.000144	0.000124	7.14
Th	0.0795	2.71	13.4	0.00219	0.00178	0.138 (0.117) <sup>5</sup>
U	0.0203	0.61	12.7	0.00074	0.00018	0.0309
Nb	0.658	28.53	11.5	0.00393	0.00302	1.45
Ta	0.037	1.97	11.7	0.00393	0.00302	0.10
K	240	5860	9.8	0.00150	0.00128	299
La	0.648	23.27	11.0	0.0114	0.00983	1.20
Ce	1.675	48.93	10.3	0.0191	0.0161	2.64
Pb	0.15	1.49	18.4	0.0834	0.0816	0.161
Pr	0.254	6.09	10.0	0.0314	0.0247	0.368
Nd	1.25	26.58	9.1	0.0463	0.0357	1.88
Sr	19.9	354.1	9.9	0.0665	0.0364	28.3
Zr	10.5	184	9.9	0.0867	0.0371	16.5
Hf	0.283	4.74	9.0	0.114	0.0607	0.538
Sm	0.406	6.64	7.8	0.0868	0.0574	0.657
Eu	0.154	2.10	7.9	0.121	0.0671	0.253
Ti	1205	17190	7.4	0.176	0.122	2960
Gd	0.544	6.32	7.8	0.179	0.0769	0.965
Tb	0.099	0.94	8.1	0.254	0.0876	0.18
Dy	0.674	5.17	8.2	0.316	0.0964	1.18
Ho	0.149	0.93	8.1	0.402	0.0972	0.25
Y	4.30	23.86	7.3	0.475	0.102	7.40
Er	0.438	2.24	7.3	0.481	0.105	0.70
Tm	0.068	0.29	7.2	0.673	0.110	0.12
Yb	0.441	1.85	7.1	0.865	0.115	0.84
Lu	0.0675	0.24	7.5	1.09	0.124	0.148

<sup>1</sup>Primitive mantle is from ref 28.

<sup>2</sup>Ta'u basalts (with MgO > 6.5 wt.%) used in the average lava composition include: T10, T16, T19, T22, T23, T25, T30, T32, T33, T44, T45, T46, T47, T48, T51, T54, T55, and 74-1. T14 is suspected of Pb contamination and is not included. The 18 Ta'u basalts were individually corrected for olivine fractionation to be in equilibrium with a mantle olivine composition of Fo<sub>93</sub> (assuming Fe<sup>2+</sup>/Fe<sup>total</sup> is 0.90), trace element corrected, and then averaged.

<sup>3</sup>Bulk partition coefficients are calculated by assuming the following modal abundances: spinel stability field: 3.72% spinel, 17.8% clinopyroxene, 26% orthopyroxene, 52.39% olivine, 0.162% sulfide; garnet stability field: 10.7% garnet, 20.8% clinopyroxene, 13.4% orthopyroxene, 55.1% olivine, 0.161% sulfide. Mineral-melt partition coefficients are from ref. 16, with the following additions and modifications: Ba and Ta are assumed to have the same bulk partition coefficients as Rb and Nb, respectively; Tm and Sr are assumed to have a bulk partition coefficient that is the average of the elements that bracket them on the spidergram; the sulfide-melt partition coefficient for Pb is assumed to be 42. The mineral modes in the garnet and spinel stability fields are calculated to agree with a primitive mantle major element composition. Although the Ta'u source is slightly depleted, and therefore has suffered a small amount of melt extraction during its history, the low degree of melt extraction will not significantly change the calculated mineral modes.

<sup>4</sup>The Ta'u source assumes that the average, olivine-fractionation corrected Ta'u lava resulted from a 5.1% modal aggregated fractional melt with a 50.6% melt contribution from the garnet stability field (and the remainder from the spinel stability field). This precise set of melting parameters allows the parent-daughter ratios (Lu/Hf and Sm/Nd) in the Ta'u source to be consistent with the <sup>176</sup>Hf/<sup>177</sup>Hf and <sup>143</sup>Nd/<sup>144</sup>Nd ratios measured in the Ta'u lavas (given a two-stage evolution model, with differentiation from primitive mantle at 1.8 Ga).

<sup>5</sup>The Th concentration in parenthesis, which yields a Th/Pb ratio of 0.727, provides a value required for the Ta'u source to be consistent with constraints from <sup>208</sup>Pb/<sup>204</sup>Pb isotopes measured in Ta'u lavas. The other Th value for the Ta'u source is calculated by adjusting the modal sulfide abundance to give U/Pb ratios that are consistent with <sup>206</sup>Pb/<sup>204</sup>Pb and <sup>207</sup>Pb/<sup>204</sup>Pb. This Th value (0.138) generates a Th/Pb ratio (0.855) that is within error (18%) of the measured Th/Pb variability (±23%) in Ta'u lavas, and is therefore (within uncertainty of the data available on Ta'u lavas) consistent with measured <sup>209</sup>Pb/<sup>204</sup>Pb ratios.

## 3. Supplementary Tables

Supplementary Table 1. Calculation of the depleted component in the Samoan plume, using lavas from Ta'u island

	Primitive mantle <sup>1</sup>	Avg. Ta'u lava <sup>2</sup> (olivine corr'd to Fo <sub>90</sub> )	Variability in 18 Ta'u lavas (1σ, std. dev., %)	Bulk partition coeffs <sup>3</sup> (garnet field)	Bulk partition coeffs <sup>3</sup> (spinel field)	Depleted Ta'u source <sup>4</sup> (calculated from avg. Ta'u lava)
Rb	0.6	14.9	16.5	0.000144	0.000124	0.759
Ba	6.6	140	11.9	0.000144	0.000124	7.14
Th	0.0795	2.71	13.4	0.00219	0.00178	0.138 (0.117) <sup>5</sup>
U	0.0203	0.61	12.7	0.00074	0.00018	0.0309
Nb	0.658	28.53	11.5	0.00393	0.00302	1.45
Ta	0.037	1.97	11.7	0.00393	0.00302	0.10
K	240	5860	9.8	0.00150	0.00128	299
La	0.648	23.27	11.0	0.0114	0.00983	1.20
Ce	1.675	48.93	10.3	0.0191	0.0161	2.64
Pb	0.15	1.49	18.4	0.0834	0.0816	0.161
Pr	0.254	6.09	10.0	0.0314	0.0247	0.368
Nd	1.25	26.58	9.1	0.0463	0.0357	1.88
Sr	19.9	354.1	9.9	0.0665	0.0364	28.3
Zr	10.5	184	9.9	0.0867	0.0371	16.5
Hf	0.283	4.74	9.0	0.114	0.0607	0.538
Sm	0.406	6.64	7.8	0.0868	0.0574	0.657
Eu	0.154	2.10	7.9	0.121	0.0671	0.253
Ti	1205	17190	7.4	0.176	0.122	2960
Gd	0.544	6.32	7.8	0.179	0.0769	0.965
Tb	0.099	0.94	8.1	0.254	0.0876	0.18
Dy	0.674	5.17	8.2	0.316	0.0964	1.18
Ho	0.149	0.93	8.1	0.402	0.0972	0.25
Y	4.30	23.86	7.3	0.475	0.102	7.40
Er	0.438	2.24	7.3	0.481	0.105	0.70
Tm	0.068	0.29	7.2	0.673	0.110	0.12
Yb	0.441	1.85	7.1	0.865	0.115	0.84
Lu	0.0675	0.24	7.5	1.09	0.124	0.148

<sup>1</sup>Primitive mantle is from ref 28.

<sup>2</sup>Ta'u basalts (with MgO > 6.5 wt.%) used in the average lava composition include: T10, T16, T19, T22, T23, T25, T30, T32, T33, T44, T45, T46, T47, T48, T51, T54, T55, and 74-1. T14 is suspected of Pb contamination and is not included. The 18 Ta'u basalts were individually corrected for olivine fractionation to be in equilibrium with a mantle olivine composition of Fo<sub>90</sub> (assuming Fe<sup>2+</sup>/Fe<sub>total</sub> is 0.90), trace element corrected, and then averaged.

<sup>3</sup>Bulk partition coefficients are calculated by assuming the following modal abundances: spinel stability field: 3.72% spinel, 17.8% clinopyroxene, 26% orthopyroxene, 52.39% olivine, 0.162% sulfide; garnet stability field: 10.7% garnet, 20.8% clinopyroxene, 13.4% orthopyroxene, 55.1% olivine, 0.161% sulfide. Mineral-melt partition coefficients are from ref. 16, with the following additions and modifications: Ba and Ta are assumed to have the same bulk partition coefficients as Rb and Nb, respectively; Tm and Sr are assumed to have a bulk partition coefficient that is the average of the elements that bracket them on the spidergram; the sulfide-melt partition coefficient for Pb is assumed to be 42. The mineral modes in the garnet and spinel stability fields are calculated to agree with a primitive mantle major element composition. Although the Ta'u source is slightly depleted, and therefore has suffered a small amount of melt extraction during its history, the low degree of melt extraction will not significantly change the calculated mineral modes.

<sup>4</sup>The Ta'u source assumes that the average, olivine-fractionation corrected Ta'u lava resulted from a 5.1% modal aggregated fractional melt with a 50.6% melt contribution from the garnet stability field (and the remainder from the spinel stability field). This precise set of melting parameters allows the parent-daughter ratios (Lu/Hf and Sm/Nd) in the Ta'u source to be consistent with the <sup>176</sup>Hf/<sup>177</sup>Hf and <sup>143</sup>Nd/<sup>144</sup>Nd ratios measured in the Ta'u lavas (given a two-stage evolution model, with differentiation from primitive mantle at 1.8 Ga).

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**Supplementary Table 3. Isotope, major and trace element analyses for Samoan (submarine Sava'i) lavas and mineral separates, including sediment cores and a ferromanganese crust from the Samoan region.**

Sample ID	D114-01	D114-03	D115-03	D115-07	D115-18	D115-18 rep	D115-20	D115-21	D115-22	D115-26	D115-28	D118-13	D118-22	D118-23	D128-01	D128-17	D128-19	D128-21	RC 13-43 <sup>a</sup>	VM-18-253 <sup>a</sup>	RC-12-207 <sup>a</sup>	D116-22crust <sup>a</sup>	W-2	BHV0-1	
Latitude (°S)	13.875	13.875	14.090	14.090	14.090	14.090	14.090	14.090	14.090	14.090	14.090	13.325	13.325	13.325	13.213	13.213	13.213	13.213	12.253	12.233	14.617	13.225			
Longitude (°W)	172.724	172.724	172.898	172.898	172.898	172.898	172.898	172.898	172.898	172.898	172.898	173.430	173.430	173.430	172.083	172.083	172.083	172.083	172.083	172.083	172.083	171.683	168.5	173.430	
Water Depth (m)	2490	2490	3220	3220																					
Lithology	Basalt	Picrite	Hawaite	Basalt	Trachybasalt		Hawaite	Trachyandesite	Basalt	Trachybasalt	Hawaite	basalt	Basalt	Basalt	sediment	sediment	sediment	Fe-Mn crust							
SiO <sub>2</sub>	46.30	44.82	49.18	46.30	52.29	46.85	57.20	47.38	51.80	50.00	39.82	39.71	42.23	48.94	47.73	48.36	49.85	35.62							
Al <sub>2</sub> O <sub>3</sub>	10.00	7.42	12.59	11.53	14.20	13.16	14.63	12.64	14.26	13.44	7.86	7.42	7.81	12.68	12.07	12.61	12.11	9.85							
TiO <sub>2</sub>	4.00	3.00	3.50	3.13	2.45	3.55	1.72	3.43	2.71	3.30	3.46	3.20	2.21	2.80	3.09	2.85	2.28	2.23							
FeO <sup>a</sup>	11.78	12.33	9.74	9.19	7.09	9.13	5.36	8.65	7.46	7.88	12.06	10.82	10.08	9.73	11.31	9.91	9.67	8.28							
MnO	0.18	0.19	0.13	0.13	0.07	0.13	0.05	0.12	0.09	0.12	0.19	0.15	0.11	0.17	0.17	0.16	0.17	0.23							
CaO	11.08	8.89	8.17	8.75	6.08	7.89	3.21	7.47	7.27	8.62	12.70	13.68	7.26	10.68	10.60	10.54	9.81	12.81							
MgO	12.16	20.07	7.97	8.63	5.09	6.34	3.61	7.47	4.08	6.37	16.63	17.13	16.74	8.22	10.10	7.88	11.65	6.86							
K <sub>2</sub> O	1.13	0.83	3.09	2.69	4.01	2.57	5.05	2.31	3.52	2.73	1.01	1.32	0.89	1.58	1.33	1.56	1.23	1.26							
Na <sub>2</sub> O	2.09	1.55	2.85	2.41	3.64	3.32	3.68	3.11	3.50	3.15	1.58	1.78	2.61	2.41	2.15	2.40	2.12	3.78							
P <sub>2</sub> O <sub>5</sub>	0.48	0.36	0.51	0.45	0.36	0.61	0.31	0.52	0.42	0.52	0.81	0.61	0.48	0.37	0.35	0.36	0.25	0.31							
Pre-total	99.28	99.65	97.73	93.21	95.28	95.35	94.82	94.10	95.20	97.10	95.91	94.81	90.61	97.88	96.92	96.84	98.24	81.22							
Mg# <sup>b</sup>	87.15	76.32	61.84	65.06	59.73	57.89	67.24	60.52	62.02	61.63	73.20	75.81	78.68	62.60	63.89	61.17	70.46	62.11							
LOI	0.29	nr	1.80	6.25	4.40	4.27	4.68	5.56	4.42	2.51	3.63	4.72	8.97	2.00	0.75	3.05	1.43	16.48							
Ni	377	820	222	385	116	191	53	145	106	180	478	472	528	142	207	108	285	168							
Cr	1001	1410	370	484	244	292	78	298	484	484	872	830	855	438	567	393	713	225							
V	328	258	237	228	154	274	89	255	221	270	280	283	189	284	312	291	261	203							
Ga	20	13	24	21	27	25	27	22	27	24	12	14	14	20	21	19	18	15							
Cu	84	76	41	52	23	56	17	52	53	52	70	47	45	56	58	62	41	85							
Zn	125	127	120	119	111	122	78	142	150	103	80	98	92	108	116	114	115	106							
Cs	0.30	0.22	0.70	0.88	0.59	0.81	1.07	0.29	0.58	0.66	0.33	0.61	4.19	0.59	0.4	0.56	0.41	0.79	1.05	0.92	0.15	0.89	0.10		
Rb	22.5	14.2	93.1	80.0	97.4	74.0	155.9	57.4	99.8	81.4	33.2	44.5	28.5	44.2	32.4	28.4	26.1	32.6	27.1	32.6	27.1	32.6	27.1	32.6	27.1
Ba	244	177	569	1736	646	486	683	466	586	510	216	477	701	345	257	328	240	369	350	276	1761	171	130		
Th	4.44	3.28	7.13	7.84	10.74	10.13	17.46	8.26	10.31	8.89	15.45	14.04	12.83	6.53	5.21	8.28	4.29	4.00	4.64	3.61	9.95	2.15	1.27		
U	0.96	0.83	1.83	2.89	1.84	2.04	2.98	1.87	2.10	1.77	1.48	1.97	1.58	1.38	1.10	1.31	0.72	0.89	1.07	0.74	8.08	0.51	0.43		
Nb	47.81	33.94	54.12	49.84	47.48	59.73	49.14	52.26	47.73	54.31	107.52	99.75	74.53	49.87	50.79	48.45	31.26	34.75	40.93	24.84	157.4	6.84	17.2		
Ta	3.34	2.31	3.68	3.39	3.17	3.95	3.08	3.50	3.23	3.61	7.18	6.64	4.55	3.43	3.47	3.33	2.17	2.56	2.89	1.89	9.39	0.46	1.17		
La	39.13	28.05	53.17	55.63	66.92	64.01	81.33	56.82	58.58	61.04	113.07	98.37	95.85	47.92	41.65	46.44	32.25	29.16	34.38	24.95	110.7	10.9	15.6		
Ce	78.51	60.89	104.2	110.9	130.1	129.3	154.5	114.2	118.6	116.4	208.2	194.8	172.8	96.98	81.65	94.30	66.79	56.84	65.72	50.89	680.2	23.3	38.0		
Pb	2.78	2.28	7.27	7.95	11.29	6.42	13.56	9.31	8.84	11.61	8.03	8.42	10.90	5.37	4.02	5.13	4.77	8.29	9.55	7.87	694.4	7.71	2.03		
Pr	9.59	7.75	11.97	12.99	15.02	15.17	16.97	13.43	14.39	13.76	24.87	22.12	19.42	11.60	9.32	11.32	8.55	6.87	8.26	6.25	16.24	3.06	5.44		
Nd	40.69	32.01	48.31	48.91	55.85	58.29	58.94	51.35	55.43	52.43	82.59	82.00	67.46	44.89	37.85	44.23	34.67	29.08	34.92	26.72	69.48	13.0	24.6		
Sr	499	374	617	714	511	662	479	589	528	645	515	627	317	496	466	481	353	709	690	392	1092	192	385		
Zr	253	196	305	289	341	358	390	302	320	313	284	260	243	248	227	242	179	168	208	157	642	92	167		
Hf	6.60	4.80	8.02	7.35	8.86	8.81	9.87	7.90	8.26	7.88	7.34	6.94	6.37	5.39	5.88	6.36	5.02	4.22	5.11	4.20	16.41	2.56	4.44		
Sm	9.56	7.08	10.68	10.02	10.86	11.75	10.24	10.42	11.24	10.48	15.38	13.87	11.86	8.98	8.27	8.93	7.50	6.88	8.26	6.72	14.10	3.31	6.11		
Eu	2.97	2.22	3.14	2.91	3.00	3.52	2.62	3.07	3.20	3.14	4.16	3.85	3.12	2.74	2.55	2.89	2.17	2.11	2.54	2.09	4.14	1.06	2.06		
Gd	8.61	6.20	9.02	8.16	8.34	9.92	7.19	8.89	9.20	9.01	11.40	10.15	8.94	7.81	7.35	7.76	6.86	6.55	7.82	6.33	17.03	3.86	6.19		
Tb	1.23	0.90	1.25	1.16	1.13	1.41	0.87	1.24	1.28	1.24	1.45	1.27	1.18	1.14	1.10	1.14	1.02	0.98	1.16	0.98	2.89	0.61	0.93		
Dy	6.52	4.89	6.44	5.93	5.67	7.38	4.88	6.58	6.86	6.85	7.28	6.27	5.83	6.27	6.17	6.28	5.49	5.43	6.35	5.58	20.42	3.86	5.33		
Ho	1.14	0.86	1.06	1.01	0.97	1.26	0.82	1.12	1.15	1.15	1.24	1.05	0.98	1.14	1.11	1.13	1.03	0.98	1.16	1.03	4.74	0.82	1.00		
Y	29.14	20.58	27.26	24.47	24.05	31.03	20.29	27.44	28.78	28.16	36.16	26.42	25.83	27.93	28.52	27.68	25.78	26.63	31.32	25.92	113.70	21.6	26.4		
Er	2.86	2.00	2.36	2.31	2.20	2.92	1.87	2.58	2.65	2.68	2.89	2.29	2.20	2.73	2.70	2.75	2.47	2.47	2.82	2.53	13.82	2.22	2.55		
Tm	0.33	0.25	0.30	0.28	0.28	0.36	0.24	0.33	0.33	0.33	0.36	0.28	0.28	0.37	0.37	0.36	0.34	0.32	0.36	0.34	2.08	0.32	0.33		
Yb	1.86	1.41	1.59	1.54	1.47	1.99	1.38	1.82	1.87	1.89	1.95	1.55	1.47	2.14	2.09	2.11	1.98	1.87	2.10	1.98	13.18	2.01	1.90		
Lu	0.27	0.20	0.22	0.23	0.21	0.27	0.20	0.26	0.27	0.27	0.30	0.22													

in house standards (BCR-P, BMP-01 and MON-01) that are dissolved and processed with each batch of unknowns. The calibration values for these three standards were determined in 2001 using 17 international reference materials, including: BCR-1 (Basalt), W-2 (Diabase), G-2 (Granite), DNC-1 (Diabase), AGV-1 (Andesite), BIR-1 (Basalt), BHVC-1 (Basalt), AC-E (Granite), QLC-1 (Qtz Latite), GSP-1 (Granodiorite), DNC-1 (Dolerite), FK-N (K-feldspar), JG-1 (Granodiorite), JB-1a (Basalt), etc. The accepted values for these standards are taken from refs 32 and 33 and the GEOREM database. Single analyses of the international standards W-2 and BHVO-1, run as an "unknowns", are reported in the table above.

Basalt Sr, Nd and Pb isotope analyses were performed on the same powder following 3 hours of leaching in 6.2 N HCl at 100 °C. Sr and Nd isotopes were also measured on leached cpx separates (following ref. 34) from selected samples (two separate cpx populations distinguished by colour—black [bl] and green [gr]—were extracted from sample D115-18), and the data are reported along with the whole rock data above. Signal intensities during Sr and Nd isotope measurements of cpx separates and whole rock basalt powders were similar, as were the associated errors. Sr and Nd chemistry was done with conventional ion chromatography using DOWEX 50 cation resin and HDEHP-treated teflon for Nd separation (following ref. 34); the HBr-HNO<sub>3</sub> procedure of refs. 35 and 36 were used for Pb chemistry, with a single column pass. Sr, Nd, and Pb isotope analyses were done on the NEPTUNE multi-collector ICP-MS at WHOI. The protocol and associated errors and standard normalizations can be found in ref. 30. Internal precision is 5–10 ppm (2σ) for Sr and Nd isotope measurements, and after adjusting to 0.710240 (SRM987 Sr standard) and 0.511847 (La Jolla Nd standard), the corresponding external precision for Sr and Nd is estimated at 15–25 ppm (2σ). The internal precision for Pb isotope measurements on XXXX204 ratios is 15–30 ppm. SRM97 Ti was used as an internal standard, and external reproducibility following full chemistry ranges from 17 ppm (2σ) for <sup>207</sup>Pb/<sup>206</sup>Pb to 117 ppm (2σ) for <sup>208</sup>Pb/<sup>206</sup>Pb. Pb isotope ratios are adjusted to the SRM 981 values of ref. 37. Further documentation of the precision for Pb isotope measurement at WHOI (on a standard solution and a basalt) can be found in ref. 38.

Helium isotopes were measured at WHOI (see ref. 39) on cpx (in samples D118-23 and D115-18) and olivine (in sample D115-03) mineral separates. Gas contents were 2.5x10<sup>-8</sup>, 1.9x10<sup>-8</sup> and 1.7x10<sup>-8</sup> cc/g, respectively, with 1σ internal precision of ±0.10, ±0.1 and ±0.064 Ra (Ra, or ratio to atmosphere, is 1.384x10<sup>-5</sup>). All helium data were obtained by crushing of cpx and olivine in vacuum. It is assumed here that radiogenic contributions are negligible. However, without total fusion helium data, we cannot exclude the possibility of implanted radiogenic <sup>3</sup>He from the basalt matrix.

<sup>1</sup>Replicate measurement was made on a leached powder prepared from a different batch of rock chips from the same rock sample.

<sup>2</sup>Cores available at Lamont-Doherty Earth Observatory core repository, and all cores were taken in the region of the Samoan islands and seamounts. Care was taken to sample all lithologies and, when making a composite of each core, the lithologies sampled were combined in approximately the same proportions as found in the respective cores. Composites were not subject to leaching prior to isotopic and major and trace element analyses.

<sup>3</sup>The ferromanganese rind was taken from the rim of basalt sample D118-22. The Mn-rind was powdered before analysis and not subjected to leaching.

<sup>4</sup>All Fe reported as FeO.

<sup>5</sup>Mg# = molar ratio of MgO/(MgO + 0.9FeO)

## Chapter 3

# **New Samoan lavas from Ofu Island reveal a hemispherically heterogeneous high $^3\text{He}/^4\text{He}$ mantle \***

### Abstract

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## New Samoan lavas from Ofu Island reveal a hemispherically heterogeneous high $^3\text{He}/^4\text{He}$ mantle

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### Abstract

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**Keywords:** Samoa; FOZO; PHEM; C;  $^3\text{He}/^4\text{He}$ ; mantle; basalt; geochemistry; DUPAL; hotspot

### 1. Introduction

Oceanic lavas with high  $^3\text{He}/^4\text{He}$  signatures are rare, and derive from ancient reservoirs in the earth's mantle. Volcanically active hotspots with high  $^3\text{He}/^4\text{He}$  lavas, such as Samoa and Hawaii, sample melts of buoyantly upwelling regions of the deep mantle where the high

$^3\text{He}/^4\text{He}$  reservoir is thought to reside (Kurz et al., 1982; Hart et al., 1992; Class and Goldstein, 2005). Consequently, ocean island basalts (OIBs) erupted at hotspots provide a unique tool for probing the composition and history of the deep mantle. Radiogenic isotopes in OIBs are commonly used as tracers for the various mantle components revealed at hotspots, and show a diverse range of compositions, or endmembers, including DMM (depleted mid-ocean ridge basalt [MORB] mantle), HIMU (high  $\mu$ , or  $^{238}\text{U}/^{204}\text{Pb}$  mantle), EM1 and EM2 (enriched mantle 1 and 2) (Zindler and Hart, 1986).

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44 Emerging from this taxonomic diversity, a unifying  
 45 theory in mantle geochemistry maintains that a single  
 46 mantle component exists that is common to all hotspots  
 47 (Hart et al., 1992): Mixing arrays from individual ocean  
 48 islands originate near the mantle endmembers in various  
 49 radiogenic isotope spaces and converge on a region  
 50 characterized by depleted isotope ratios that is distinct  
 51 from normal MORB (Hart et al., 1992). Lavas plotting  
 52 in this region of convergence often exhibit elevated  
 53  $^3\text{He}/^4\text{He}$  ratios, and are suggested to sample a com-  
 54 ponent in the mantle common to all hotspots. Variably  
 55 called FOZO (Hart et al., 1992), PHEM (*Primitive*  
 56 *Helium Mantle*; Farley et al., 1992), or C (Common;  
 57 Hanan and Graham, 1996), the high  $^3\text{He}/^4\text{He}$  common  
 58 component is thought to be a relatively less degassed  
 59 region of the (lower?) mantle (Kurz et al., 1982; Class  
 60 and Goldstein, 2005).

61 New geochemical data from the Samoan hotspot are  
 62 not entirely consistent with this view of the mantle.  
 63 The new  $^3\text{He}/^4\text{He}$  ratios (up to  $33.8 \pm 0.2$  Ra, ratio to  
 64 atmosphere) from the Samoan Island of Ofu are the  
 65 highest yet recorded in the southern hemisphere, and are  
 66 significantly higher than  $^3\text{He}/^4\text{He}$  ratios previously  
 67 measured (25.8 Ra) in Samoan basalts and xenoliths  
 68 (Farley et al., 1992; Poreda and Farley, 1992; Workman  
 69 et al., 2004). The new helium isotope data from Samoa  
 70 extend the range of observed  $^3\text{He}/^4\text{He}$  up to values  
 71 comparable to those found in Hawaii (32.3 Ra; Kurz  
 72 et al., 1982), Iceland (37.7 Ra; Hilton et al., 1999) and the  
 73 Galapagos (30.3 Ra; Kurz and Geist, 1999; Saal et al.,  
 74 2007), referred to here as HIG. The Ofu lavas are  
 75 isotopically more enriched (higher  $^{87}\text{Sr}/^{86}\text{Sr}$  and lower  
 76  $^{143}\text{Nd}/^{144}\text{Nd}$ ) than the high  $^3\text{He}/^4\text{He}$  samples from HIG  
 77 and exhibit elevated incompatible trace element con-  
 78 centrations. Due to this isotopic and trace element  
 79 enrichment relative to HIG lavas, the new data from  
 80 Samoan high  $^3\text{He}/^4\text{He}$  lavas are inconsistent with recent  
 81 models that describe the evolution of the high  $^3\text{He}/^4\text{He}$   
 82 mantle, and the Ofu data suggest that the high  $^3\text{He}/^4\text{He}$   
 83 mantle domain is isotopically heterogeneous.

## 84 2. Methods and results

### 85 2.1. Sample location and state of preservation

86 Ofu Island is located in the eastern province of the  
 87 Samoan archipelago, an age-progressive hotspot track  
 88 (Hart et al., 2004; Koppers et al., submitted for  
 89 publication) located just north of the northern terminus  
 90 of the Tonga subduction zone. The samples were col-  
 91 lected at various locations on the perimeter of Ofu and  
 92 Olosega islands (Ofu hereafter; sample location map is

available in the supplementary data in the Appendix).  
 93 Tholeiitic lavas in Samoa are rare (Natland, 1980; 94  
 Workman et al., 2004), and with the exception of a 95  
 cumulate (OFU-04-14) and a gabbro (OFU-04-17), 96  
 the Ofu lavas presented in this study are alkali basalts 97  
 (Table 1). The Ofu samples are generally quite fresh. 98  
 With the exception of sample OFU-04-12, which has a 99  
 Th/U ratio of 4.8 (and may indicate U-loss during sub- 100  
 aerial weathering), the range of Th/U ratios in the Ofu 101  
 sample suite is 4.0–4.4. The Ba/Rb ratios for Ofu sam- 102  
 ples ( $9.1 \pm 1.1$  at  $1\sigma$ ) are similar to the values for young 103  
 Samoan basalts reported previously (Workman et al., 104  
 2004), and somewhat lower than the canonical value of 105  
 $\sim 12$  for fresh OIB lavas (Hofmann and White, 1983). 106  
 Excluding the cumulate sample OFU-04-14, which has a 107  
 Rb/Cs ratio of 280, the range of Rb/Cs values from the 108  
 Ofu samples (from 73 to 137) is close to the canonical 109  
 range of 85–95 (Hofmann and White, 1983). These 110  
 weathering proxies indicate that elements equally or less 111  
 mobile than U, Rb and Cs yield useful petrogenetic 112  
 information. 113

### 114 2.2. He, Sr, Nd and Pb isotopes in Ofu lavas

New  $^3\text{He}/^4\text{He}$  values (19.5 to 33.8 Ra) were measured 115  
 at Woods Hole Oceanographic Institution on olivine 116  
 and clinopyroxene (cpx) phenocrysts separated from 12  
 117 hand samples (Table 1). Measurements were made by 118  
 crushing and fusion *in vacuo*, following the protocol 119  
 reported in (Kurz et al., 2004). The sample with the 120  
 highest  $^3\text{He}/^4\text{He}$  value, OFU-04-06, was taken from an 121  
 ankaramite dike exposed at 2 m depth in a recent road 122  
 cut. Olivines from this sample are relatively gas rich 123  
 ( $67.5 \times 10^{-9} \text{ cm}^3 \text{ STP g}^{-1}$ , the sum of crushing and 124  
 fusion) and yielded similar  $^3\text{He}/^4\text{He}$  ratios on two sepa- 125  
 rate crushes of the same olivine separate (OFU-04-06cr1 126  
 followed by OFU-04-06cr2). Following the crushing 127  
 experiments, a fusion extraction of the resulting olivine 128  
 powder (OFU-04-06fus) yielded lower  $^3\text{He}/^4\text{He}$ , indi- 129  
 cating the presence of ingrown radiogenic helium, prob- 130  
 ably implanted from the Th and U-rich matrix. Two 131  
 different olivine populations (lighter and darker olivines) 132  
 separated from sample OFU-04-06 yielded similar 133  
 $^3\text{He}/^4\text{He}$  ratios (33.4 and 33.6 Ra). These fusion and 134  
 crushing experiments, coupled with sampling depths, 135  
 preclude the influence of cosmogenic helium for 136  
 this important sample. The lava with the second highest 137  
 $^3\text{He}/^4\text{He}$  value measured by crushing—sample OFU-04- 138  
 15—also yielded lower  $^3\text{He}/^4\text{He}$  ratios upon fusion 139  
 (OFU-04-15fus) of the crushed olivine powder. Addi- 140  
 tionally, crushing experiments of sample OFU-04-03 141  
 (OFU-04-03cr1 followed by OFU-04-03cr2) yielded 142

Table 1  
Isotopic and (selected) trace element analyses of Ofu and Moorea lavas

Sample	Phase analyzed for helium	Rock type	Geological formation	$^4\text{He}$ ( $10^{-9}\text{cm}^3\text{STP g}^{-1}$ )	$^3\text{He}/^4\text{He}$ ( $R/R_a$ )	$1\sigma$	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{143}\text{Nd}/^{144}\text{Nd}$	$^{206}\text{Pb}/^{204}\text{Pb}$	$^{207}\text{Pb}/^{204}\text{Pb}$	$^{208}\text{Pb}/^{204}\text{Pb}$	Th	Sr	Th/U	Ba/Rb	Rb/Cs	Ba/Nb	Pb/Pb*	
OFU-04-02	Olivine	Ankaramite	Dike	38.6	24.6	0.1	NA	NA	NA	NA	NA	NA	NA	NA	NA	NA	NA	NA	NA
OFU-04-03cr1	Olivine	Ankaramite	Dike	119	24.0	0.1	0.704756	0.512811	19.257	15.591	39.418	3.26	390	4.29	7.19	117.1	5.3	0.44	
OFU-04-03cr2	Olivine	Ankaramite	Dike	11.5	23.9	0.2	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-05	Olivine	Ankaramite	Dike	70.8	24.4	0.1	0.704793	0.512814	19.254	15.589	39.412	4.24	546	4.26	7.81	136.7	5.6	0.43	
OFU-04-06cr1	Olivine	Ankaramite	Dike	44.0	33.7	0.2	0.704584	0.512827	19.189	15.571	39.202	3.96	599	3.99	9.71	77.7	4.9	0.42	
OFU-04-06cr2	Olivine	Ankaramite	Dike	5.0	33.8	0.2	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-06 fus	Olivine	Ankaramite	Dike	18.5	27.0	0.1	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-06 (light olivines)	Olivine	Ankaramite	Dike	27.3	33.4	0.2	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-06 (dark olivines)	Olivine	Ankaramite	Dike	40.0	33.6	0.1	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-06 <sup>a</sup>	cpx	Ankaramite	Dike	0.2	26.1	3.4	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-07	none	Trachybasalt	Dike	NA	NA	NA	0.704623	0.512814	19.177	15.582	39.268	5.24	767	4.04	9.59	122.0	5.0	0.39	
OFU-04-08	Olivine	Ankaramite	Dike	4.5	21.3	0.3	0.704703	0.512802	19.207	15.586	39.291	4.06	573	4.11	8.78	77.5	4.7	0.42	
OFU-04-09	Olivine	Ankaramite	Flow	3.2	25.6	0.5	0.704538	0.512835	19.165	15.577	39.208	3.25	503	4.03	10.20	131.6	5.4	0.46	
OFU-04-10	Olivine	Alkali Basalt	Flow	8.0	22.0	0.2	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-10	cpx	Alkali Basalt	Flow	1.0	19.7	0.6	0.704648	0.512815	19.187	15.585	39.261	3.65	495	4.39	10.79	73.0	5.8	0.43	
OFU-04-11	None	Trachybasalt	Flow	NA	NA	NA	0.704635	0.512804	19.184	15.594	39.299	6.60	929	4.30	9.09	108.0	5.3	0.45	
OFU-04-12	Olivine	Alkali Basalt	Flow	3.5	21.2	0.3	0.704795	0.512800	19.202	15.594	39.403	3.29	331	4.78	10.36	118.3	6.2	0.42	
OFU-04-13	None	Hawaiite	Flow	NA	NA	NA	0.704625	0.512816	19.176	15.581	39.264	5.02	723	4.07	9.09	99.2	4.8	0.40	
OFU-04-14	Olivine	Cumulate	Boulder	4.5	25.0	0.2	0.704517	0.512819	19.126	15.584	39.164	2.95	292	4.39	9.88	280.2	4.9	0.37	
OFU-04-15	Olivine	Ankaramite	Boulder	14.3	29.6	0.2	0.704559	0.512822	19.141	15.580	39.169	3.89	601	4.02	9.28	88.0	4.8	0.43	
OFU-04-15 fus	Olivine	Ankaramite	Boulder	3.2	21.2	0.2	"	"	"	"	"	"	"	"	"	"	"	"	"
OFU-04-16	cpx	Alkali Basalt	Boulder	0.2	19	1	0.704438	0.512834	19.216	15.587	39.276	3.54	639	3.96	8.60	129.6	5.3	0.46	
OFU-04-17	Olivine	Gabbro	Boulder	1.0	26.4	0.4	0.704498	0.512844	19.160	15.576	39.173	2.18	310	4.14	7.30	118.2	5.1	0.41	
MO01-01 (Moorea) <sup>b</sup>	Olivine	Ankaramite	NA	1.3	17.0	1.6	0.704621	0.512774	19.216	15.603	38.933	2.11	381	3.75	11.14	270.4	8.5	0.50	

Samples were collected at Ofu Island, American Samoa in August of 2004. All  $^3\text{He}/^4\text{He}$  ratios are reported relative to atmospheric ( $R/R_a$ ) using an atmospheric value of  $1.384 \times 10^{-6}$ . The Sr, Nd, and Pb chemistry, mass spectrometry and associated measurement precision are reported in Hart and Bhattacharya (2006). Sr, Nd and Pb-isotope analyses were performed on leached whole-rock powders. Trace elements analyses were performed on unleached whole-rock powders by ICP at the Washington State University GeoAnalytical Lab.

Errors associated with trace element analyses can be found in Jackson et al. (2007). Only two samples (OFU-04-06 fus and OFU-04-15 fus) were fused for helium isotope measurement. All other helium measurements were made by crushing. When an olivine sample was crushed and measured more than once, the crush experiment is labeled as "cr1" or "cr2".

<sup>a</sup> Required a blank correction of 46% due to low gas concentrations. The  $^3\text{He}/^4\text{He}$  ratio on this clinopyroxene measurement is similar to the olivine crush values at the  $2\sigma$  level.

<sup>b</sup> The sample is from Moorea in the Societies island chain, a hotspot that exhibits EM2 affinities. Helium isotope data were previously published on Societies sample MO01-01 in Hanyu and Kaneoka (1997). Sr, Nd and Pb isotopes by same method as Ofu rocks.

143 reproducible results. The reproducibility of the measure-  
 144 ments for Ofu samples, the relatively high helium  
 145 concentrations, the absence of higher  $^3\text{He}/^4\text{He}$  ratios  
 146 on melting, and shielding of many of the samples all  
 147 indicate that cosmogenic  $^3\text{He}$  is not a factor in generating  
 148 the remarkably high  $^3\text{He}/^4\text{He}$  ratios.

149 The Sr, Nd, and Pb chemistry, mass spectrometry,  
 150 associated measurement precision and standard normal-  
 151 izations are reported in Hart and Blusztajn (2006) and  
 152 references therein. Basalt Sr, Nd and Pb-isotope analyses  
 153 were performed on the same powder following 1 h of  
 154 leaching in 6.2 N HCl at 100 °C. Sr and Nd chemistry

was done with conventional ion chromatography using 184  
 DOWEX 50 cation resin and HDEHP-treated teflon for 185  
 Nd separation (Taras and Hart, 1987). An HBr–HNO<sub>3</sub> 186  
 procedure (Galer, in press; Abouchami et al., 1999) was 187  
 used for Pb chemistry, with a single column pass. Sr, Nd, 188  
 and Pb-isotope analyses were done on the NEPTUNE 189  
 multi-collector ICP-MS at WHOI. Internal precision is 190  
 5–10 ppm ( $2\sigma$ ) for Sr and Nd-isotope measurements. 191  
 Adjusting to 0.710240 (SRM987 Sr standard) and 192  
 0.511847 (La Jolla Nd standard) gives an estimated 193  
 external precision for Sr and Nd of 15–25 ppm ( $2\sigma$ ). 194  
 The internal precision for Pb-isotope ratios is better 195  
 than 15–30 ppm, and using SRM997 Tl as an internal 196  
 standard, the external reproducibility following full 197  
 chemistry ranges from ~20 ppm ( $2\sigma$ ) for  $^{207}\text{Pb}/^{206}\text{Pb}$  198  
 to ~120 ppm ( $2\sigma$ ) for  $^{208}\text{Pb}/^{204}\text{Pb}$  (Hart and Blusztajn, 199  
 2006). Pb-isotope ratios are adjusted to the SRM 981 200  
 values of Todt et al. (1996). 201

Ofu lavas exhibit slight variations in the radiogenic 202  
 isotope ratios of  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.704438–0.704795), 203  
 $^{143}\text{Nd}/^{144}\text{Nd}$  (0.512800 to 0.512844) and  $^{206}\text{Pb}/^{204}\text{Pb}$  204  
 (19.126 to 19.257) (Fig. 1 and Table 1). Together with 205  
 Ofu, data from Samoan lavas in general define a wedge- 206  
 like shape in  $^3\text{He}/^4\text{He}$ – $^{87}\text{Sr}/^{86}\text{Sr}$  isotope space: lavas with 207  
 the highest (more enriched)  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios exhibit low 208  
 $^3\text{He}/^4\text{He}$  ratios and samples with the least radiogenic 209  
 $^{87}\text{Sr}/^{86}\text{Sr}$  ratios, found primarily at Ofu, are associated 210  
 with the highest  $^3\text{He}/^4\text{He}$  ratios (Fig. 1). Similarly, 211  
 Samoan lavas with the most enriched (lowest)  $^{143}\text{Nd}/^{144}\text{Nd}$  212  
 $^{143}\text{Nd}/^{144}\text{Nd}$  have low  $^3\text{He}/^4\text{He}$  ratios. However, Nd and Pb 213  
 isotopes measured in Ofu lavas do not define end-member 214  
 values for Samoa, but instead fall toward the upper end 215  
 found in the hotspot. 216

In order to further constrain the nature of elevated 217  
 $^3\text{He}/^4\text{He}$  lavas associated with EM2 hotspots, we present 218  
 Sr, Nd and Pb isotopes for a high  $^3\text{He}/^4\text{He}$  sample from 219

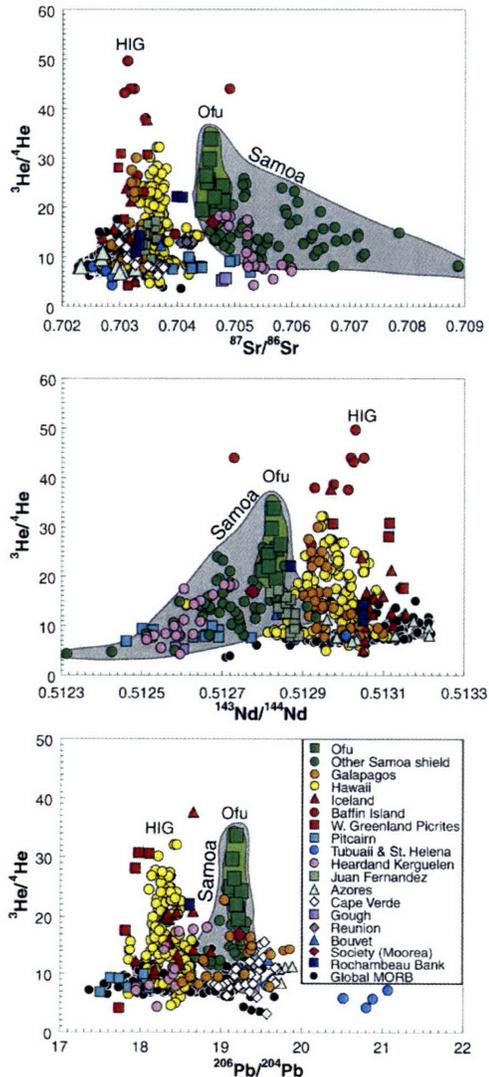


Fig. 1. The new Ofu data indicate that the Samoan high  $^3\text{He}/^4\text{He}$  lavas have more enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  than the highest  $^3\text{He}/^4\text{He}$  lavas from Hawaii, Iceland and Galapagos (abbreviated HIG). The field of Samoan shield basalts is shaded grey, and the field of Samoan lavas from Ofu Island is shaded green (see Table 1). HIG lavas trend to a more isotopically depleted high  $^3\text{He}/^4\text{He}$  component than lavas from Ofu. A single sample from Baffin Island is an exception to the Sr and Nd-isotope separation between the Ofu and HIG lavas, but was indicated by Stuart et al. (2003) to be crustally contaminated. Data sources are summarized in Graham (2002) or noted in Table 2. In the isotope projections shown in this figure, the separation between the highest  $^3\text{He}/^4\text{He}$  samples from the northern and southern hemisphere high  $^3\text{He}/^4\text{He}$  hotspots is not perfect. For example, the Pb-isotopes of the highest  $^3\text{He}/^4\text{He}$  lavas from Samoa and the Galapagos overlap. However, the isotopic and hemispheric separation of FOZO-A and FOZO-B samples (Table 2) is observed in the isotope projections shown in Fig. 2. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

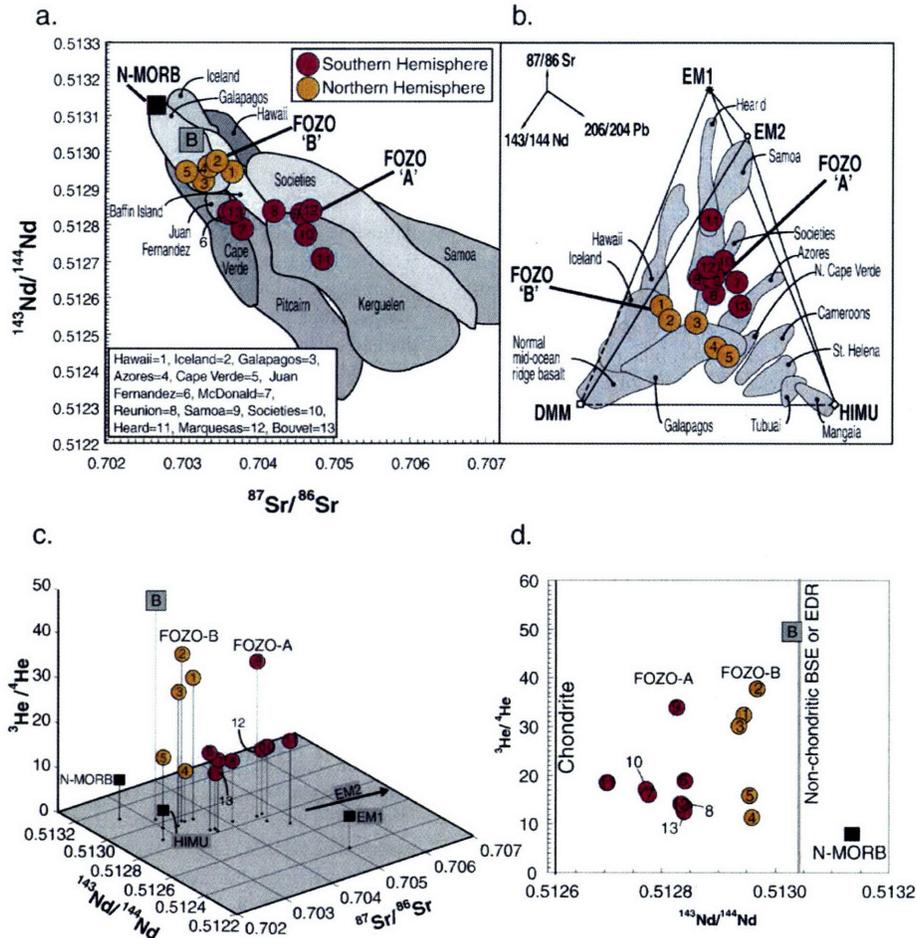


Fig. 2. The highest  $^3\text{He}/^4\text{He}$  lavas from northern hemisphere (FOZO-B) hotspots separate from the highest  $^3\text{He}/^4\text{He}$  samples from southern hemisphere (FOZO-A) hotspots in a.)  $^{87}\text{Sr}/^{86}\text{Sr}$ - $^{143}\text{Nd}/^{144}\text{Nd}$  space, b.) the mantle tetrahedron, c.) a “helium balloon” pinplot showing  $^{87}\text{Sr}/^{86}\text{Sr}$ - $^{143}\text{Nd}/^{144}\text{Nd}$ - $^3\text{He}/^4\text{He}$ , and d.)  $^{143}\text{Nd}/^{144}\text{Nd}$ - $^3\text{He}/^4\text{He}$  isotope spaces. The criteria for selecting high  $^3\text{He}/^4\text{He}$  lavas with Sr–Nd–Pb isotopes most representative of the OIB mantle are described in Table 2 and Sections 3 and 4.1 of the text. The EDR and the non-chondritic BSE are described in Fig. 6. Although the Baffin Island lava (square marked with a “B”) is erupted in a continental setting, and is thus excluded from the compilation of high  $^3\text{He}/^4\text{He}$  lavas (see Table 2), it hosts the highest magmatic  $^3\text{He}/^4\text{He}$  on record and is included in the figure for reference. The highest  $^3\text{He}/^4\text{He}$  lavas from each hotspot (and the high  $^3\text{He}/^4\text{He}$  Baffin Island lava) are significantly more enriched than the N-MORB composition from Su (in press), consistent with earlier observations (Hart et al., 1992). Compared to other southern hemisphere high  $^3\text{He}/^4\text{He}$  lavas, Samoan high  $^3\text{He}/^4\text{He}$  lavas do not exhibit anomalous Sr, Nd and Pb isotopes.

220 Moorea island in the Societies hotspot. The high  $^3\text{He}/^4\text{He}$  He  
 221 sample from Moorea, MO01-01, has a  $^3\text{He}/^4\text{He}$  ratio of  
 222 17.0 Ra (Hanyu and Kaneoka, 1997), and has Sr, Nd and  
 223 Pb-isotope ratios similar to Ofu basalts (Table 1).

224 **3. Ofu in a global context**

225 The  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$  compositions of Ofu  
 226 lavas show that the high  $^3\text{He}/^4\text{He}$  reservoir sampled by the

Samoan hotspot is more enriched than the high  $^3\text{He}/^4\text{He}$  227  
 reservoir sampled by HIG lavas. Data in Fig. 1 show a 228  
 clear “valley” between the highest  $^3\text{He}/^4\text{He}$  Samoan and 229  
 HIG lavas; samples with  $^3\text{He}/^4\text{He}$  greater than  $\sim 22$  Ra 230  
 have not been found in this gap. The Ofu samples have 231  
 $^{87}\text{Sr}/^{86}\text{Sr}$  ratios greater than 0.7044, while the high 232  
 $^3\text{He}/^4\text{He}$  HIG lavas have  $^{87}\text{Sr}/^{86}\text{Sr}$  less than 0.7038; 233  
 a single Baffin Island sample, which is considered by 234  
 Stuart et al. (2003) to be affected by contamination from 235

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t2.1 Table 2

t2.2 Summary of the radiogenic isotope data for the highest  $^3\text{He}/^4\text{He}$  samples used to define FOZO-A and FOZO-B

t2.3	Hotspot	Island/ seamount/ region	ID of sample with highest $^3\text{He}/^4\text{He}$ used	$^3\text{He}/^4\text{He}$ ( $\pm 1\sigma$ )	$^{87}\text{Sr}/^{86}\text{Sr}$	$^{143}\text{Nd}/^{144}\text{Nd}$	$^{206}\text{Pb}/^{204}\text{Pb}$	$^{207}\text{Pb}/^{204}\text{Pb}$	$^{208}\text{Pb}/^{204}\text{Pb}$	Epsilon $^{143}\text{Nd}$	
t2.4	<i>FOZO-A</i>										
t2.5	Samoa <sup>a</sup>	Ofu	Ofu-04-06	33.8±0.2	0.704584	0.512827	19.189	15.571	39.202	4.2	
t2.6	Societies <sup>a</sup>	Moorea	MO-01-01	17.0±1.6	0.704621	0.512774	19.216	15.603	38.933	3.2	
t2.7	Macdonald <sup>a</sup>	Macdonald	SO-47 64DS2	15.8±0.3	0.703755	0.512777	19.46	15.61	39.27	3.2	
t2.8	Juan Fernandez <sup>b,c</sup>	Mas a Tierra	PIN-8	17.2±0.3	0.70363	0.512840	<i>19.13</i>	<i>15.61</i>	<i>38.98</i>	4.5	
t2.9	Reunion <sup>b</sup>	Reunion	VP1931	13.6±0.3	0.704130	<i>0.51284</i>	<i>18.88</i>	<i>15.59</i>	<i>38.96</i>	4.5	
t2.10	Kerguelen <sup>d</sup>	Heard	HB24A	18.3±0.5	0.704862	0.512705	18.776	15.588	39.170	1.8	
t2.11	Bouvet <sup>b</sup>	Bouvet	WJ8B	12.4±1	0.70371	<i>0.51284</i>	19.588	15.653	39.243	4.5	
t2.12	Marquesas <sup>a</sup>	Hiva Oa	HO-AT-1	14.4±1.4	0.70471	0.512834	19.017	15.570	38.823	4.4	
t2.13	<i>FOZO-B</i>										
t2.14	Hawaii <sup>a,c</sup>	Loihi	KK18-8	32.3±0.6	0.703680	0.512945	18.448	15.477	38.189	6.5	
t2.16	Iceland <sup>a</sup>	Westfjords	SEL-97	37.7±2.0	0.703465	0.512969	18.653	15.473	38.453	7.0	
t2.17	Galapagos <sup>a</sup>	Fernandina	NSK 97-214	30.3±0.2	0.703290	0.512937	19.080	15.537	38.711	6.4	
t2.18	Cape Verde <sup>b</sup>	San Nicolau	SN-10	15.7±0.1	0.703050	<i>0.51296</i>	19.538	15.586	38.964	6.7	
t2.19	Azores <sup>a,c</sup>	Terciera	T2	11.3±0.8	0.703520	0.512960	19.883	15.630	39.310	6.8	

If Sr, Nd and Pb isotopes are unavailable for a sample with high  $^3\text{He}/^4\text{He}$ , they are estimated based on the criteria below and the estimated values are italicized in the table.

High  $^3\text{He}/^4\text{He}$  lavas erupted at deep submarine ridge environments are excluded, and include the following locations as summarized in Graham (2002): Southeast Indian ridge near Amsterdam and St. Paul (14.1 Ra), east and west rifts of the Easter microplate (11.7 Ra), Gulf of Tadjoura near Afar (14.7 Ra), southern Mid-Atlantic ridge near Shona and Discovery (12.3 and 15.2 Ra, respectively), Southwest Indian Ridge near Bouvet Island (14.9 Ra) and the Manus Basin back-arc spreading center (15.1 Ra). However, sub-aerially erupted high  $^3\text{He}/^4\text{He}$  lavas sampled at ridge-centered hotspots, including Bouvet Island (Kurz et al., 1998) and Iceland (Hilton et al., 1999), are included in the dataset. Additionally, high  $^3\text{He}/^4\text{He}$  lavas erupted in continental settings are not considered, and include: Baffin Island (Stuart et al., 2003) and West Greenland picrites (Graham et al., 1998), Yellowstone (Graham et al., 2006) and Afar (Scarsi and Craig, 1996). Finally, high  $^3\text{He}/^4\text{He}$  lavas erupted in back-arc environments, such as

Rochambeau Bank (Poreda and Craig, 1992) and Manus Basin (Macpherson et al., 1998), are excluded from our treatment of the high  $^3\text{He}/^4\text{He}$  mantle.

<sup>a</sup> Many of the high  $^3\text{He}/^4\text{He}$  lavas in this table have a full complement of lithophile radiogenic isotope analyses measured on the same samples, and are from the following locations: The Galapagos (Kurz and Geist, 1999; Saal et al., 2007), Hawaii (Kurz et al., 1983; Staudigel et al., 1984), Samoa (Table 1), Iceland (Hilton et al., 1999), Societies (Hanyu and Kaneoka, 1997; Table 1), Marquesas (Castillo et al., 2007), Macdonald (Hémond et al., 1994; Moreira and Allègre, 2004) and Azores (Turner et al., 1997; Moreira et al., 1999).

<sup>b</sup> A full complement of Sr, Nd and Pb-isotope data do not exist for all the high  $^3\text{He}/^4\text{He}$  samples listed in the table above. Values for the “missing” lithophile radiogenic isotopes (data in italics) are generated based on the following criteria. The high  $^3\text{He}/^4\text{He}$  Cape Verde lava has Sr and Pb isotopes but lacks Nd isotopes (Doucelance et al., 2003). Sr and Nd-isotope data correlate well in the Cape Verde islands, and the Nd-isotope ratio for the Cape Verde sample is made by regression through existing data. The high  $^3\text{He}/^4\text{He}$  Bouvet Island sample (Kurz et al., 1998) also has Sr and Pb-isotope data, but lacks a Nd-isotope measurement. The missing Nd isotopic value is estimated by averaging existing Bouvet island data from O’Nions et al. (1977). A high  $^3\text{He}/^4\text{He}$  sample from Reunion with Sr isotope data (Graham et al., 1990) lacks Nd and Pb-isotope data; Reunion lavas are isotopically uniform, and the GEOROC database (<http://georoc.mpch-mainz.gwdg.de/georoc/>) was used to estimate the Nd and Pb isotopic compositions for the Reunion sample. Sr and Nd-isotope data are reported for a sample with the second highest  $^3\text{He}/^4\text{He}$  measured from Juan Fernandez (Farley et al., 1993), but Pb-isotope data are missing; the high  $^3\text{He}/^4\text{He}$  Juan Fernandez sample exhibits one of the least radiogenic Nd-isotope ratios (0.51284) in the Juan Fernandez suite, so Pb-isotope compositions ( $^{206}\text{Pb}/^{204}\text{Pb}$  from 19.045 to 19.214) from three other Juan Fernandez samples (Gerlach et al., 1986) with the least radiogenic Nd isotopes (0.51282–0.51284) were averaged to generate a Pb-isotope composition for the high  $^3\text{He}/^4\text{He}$  lava.

<sup>c</sup> When the  $^3\text{He}/^4\text{He}$  ratios quoted above are lower than the maximum from a hotspot, it is because Sr, Nd and/or Pb-isotope ratios could not be estimated reliably for the sample, or because higher reported  $^3\text{He}/^4\text{He}$  ratios were not measured as precisely (or were suspected of having a cosmogenic  $^3\text{He}$  influence). For example, a  $^3\text{He}/^4\text{He}$  value of 14.8 Ra was reported for an Azores lava (Madureira et al., 2005), but lithophile radiogenic isotopes were not reported and a cosmogenic  $^3\text{He}$  influence cannot be ruled out for this sample. Similarly, a value of up to 35.3 Ra has been reported from Loihi (Valbracht et al., 1997), but analytical uncertainties on this measurement were large and lithophile isotopes were not reported.

<sup>d</sup> The Sr, Nd and Pb isotopes of the Heard sample (Barling and Goldstein, 1990; Hilton et al., 1995) were not measured on the same sample as the  $^3\text{He}/^4\text{He}$ , but they were measured on a sample (69244) from the same flow.

236 continental crust, is an exception to the separation of Ofu  
237 and HIG lavas. Ofu’s trend to a more enriched region  
238 of Nd-isotope space than that sampled by the HIG lavas  
239 is also clear ( $^{143}\text{Nd}/^{144}\text{Nd} < 0.51285$  in Ofu versus  $^{143}\text{Nd}/$   
240  $^{144}\text{Nd} > 0.51290$  in HIG lavas). Previous work suggested

that the Samoan data trends toward a similar high  $^3\text{He}/^4\text{He}$  241  
component as the HIG lavas (Farley et al., 1992; Hart 242  
et al., 1992). The new isotope data argue against this 243  
hypothesis because Ofu has  $^3\text{He}/^4\text{He}$  as high as HIG but 244  
with more enriched  $^{87}\text{Sr}/^{86}\text{Sr}$  and  $^{143}\text{Nd}/^{144}\text{Nd}$ . 245

246 In order to place constraints on the Sr, Nd and Pb-  
 247 isotope heterogeneity of the high  $^3\text{He}/^4\text{He}$  mantle reser-  
 248 voir (Fig. 2), we have compiled a dataset of Sr, Nd and Pb  
 249 isotopes for lavas that are representative of the high  
 250  $^3\text{He}/^4\text{He}$  OIB mantle. We include only the hotspots that  
 251 have  $^3\text{He}/^4\text{He} > 11 \text{ Ra}$ , thereby excluding hotspots such as  
 252 Tristan, Gough, and St. Helena, etc., that do not sample a  
 253 component from the high  $^3\text{He}/^4\text{He}$  mantle. We also limit  
 254 our discussion to just the *single highest  $^3\text{He}/^4\text{He}$  sample*  
 255 *from each high  $^3\text{He}/^4\text{He}$  hotspot*, in an attempt to define  
 256 the Sr–Nd–Pb-isotope composition of the highest  $^3\text{He}/$   
 257  $^4\text{He}$  reservoir sampled by a hotspot. Finally, high  $^3\text{He}/^4\text{He}$   
 258 lavas erupted at continental, back-arc and deep submarine  
 259 ridge environments are excluded (see Table 2 for a list of  
 260 specific samples from these three environments).

261 Using the above criteria, Table 2 lists the high  
 262  $^3\text{He}/^4\text{He}$  samples that best represent the Sr, Nd and Pb  
 263 compositions of the high  $^3\text{He}/^4\text{He}$  OIB mantle reservoir.  
 264 Combined with the new Ofu data, the high  $^3\text{He}/^4\text{He}$  OIB  
 265 dataset in Table 2 suggests the existence of two iso-  
 266 topically distinct high  $^3\text{He}/^4\text{He}$  reservoirs in the mantle.  
 267 These two high  $^3\text{He}/^4\text{He}$  reservoirs separate in the  
 268 northern (boreal) and southern (austral) hemispheres  
 269 (Fig. 2). The highest  $^3\text{He}/^4\text{He}$  samples from each of the  
 270 southern hemisphere hotspots, Macdonald seamount,  
 271 Bouvet, Kerguelen, Juan Fernandez, Societies, Reunion  
 272 and Marquesas appear to sample the more isotopically  
 273 enriched (or less isotopically depleted, see Section 4.5),  
 274 high  $^3\text{He}/^4\text{He}$  mantle component found in Ofu lavas. By  
 275 contrast, HIG lavas and two other northern hemisphere  
 276 hotspots—Cape Verde and Azores—sample a more iso-  
 277 topically depleted, high  $^3\text{He}/^4\text{He}$  component than  
 278 observed in southern hemisphere high  $^3\text{He}/^4\text{He}$  lavas.  
 279 The highest  $^3\text{He}/^4\text{He}$  samples from the boreal and austral  
 280 high  $^3\text{He}/^4\text{He}$  hotspots separate toward the depleted  
 281 and enriched ends, respectively, of the global  $^{87}\text{Sr}/^{86}\text{Sr}$ –  
 282  $^{143}\text{Nd}/^{144}\text{Nd}$  OIB array (Fig. 2). Fig. 2 shows that there  
 283 are two separate  $^3\text{He}/^4\text{He}$  peaks in Sr–Nd–He isotope  
 284 space: One peak is formed by the highest  $^3\text{He}/^4\text{He}$   
 285 northern hemisphere samples and another peak, an-  
 286 chored by the high  $^3\text{He}/^4\text{He}$  Ofu lava, is formed by the  
 287 highest  $^3\text{He}/^4\text{He}$  southern hemisphere samples. Samples  
 288 with the highest  $^3\text{He}/^4\text{He}$  from each of the boreal and  
 289 austral FOZO hotspots also are distinguished isotopi-  
 290 cally inside the mantle tetrahedron formed by the  
 291 isotopes of Sr, Nd and Pb. The boreal (FOZO-B) domain  
 292 plots closer to the DMM–HIMU join while the austral  
 293 (FOZO-A) domain plots closer to the EM1–EM2 join  
 294 (Fig. 2). FOZO-A and FOZO-B also clearly separate in  
 295 a mantle tetrahedron constructed from the isotopes of Pb  
 296 (a figure of the Pb-isotope tetrahedron is available as  
 297 supplementary data in the Appendix).

## 4. Discussion

298

### 4.1. FOZO-A and FOZO-B: some important caveats

299

Unfortunately, there are limited numbers of hotspots  
 300 with moderately high  $^3\text{He}/^4\text{He}$  (from 11.3–18.3 Ra, 9  
 301 hotspots) and very high  $^3\text{He}/^4\text{He}$  (>30 Ra, 4 hotspots);  
 302 lithophile isotope data on lavas from many of these high  
 303  $^3\text{He}/^4\text{He}$  hotspots are also limited (see Table 2). There-  
 304 fore, we cannot exclude the possibility that the apparent  
 305 hemispheric separation is related to sparse data, and that  
 306 existing He, Sr, Nd and Pb-isotope data do not yet  
 307 adequately characterize the terrestrial mantle. Further-  
 308 more, the isotopic and hemispheric separation of the two  
 309 FOZOs is not observed in all isotope projections: In  
 310  $^3\text{He}/^4\text{He}$  vs.  $^{206}\text{Pb}/^{204}\text{Pb}$  and  $^3\text{He}/^4\text{He}$  vs.  $^{87}\text{Sr}/^{86}\text{Sr}$  iso-  
 311 tope spaces (not shown), the FOZO-A and FOZO-B  
 312 fields overlap. However, in  $^3\text{He}/^4\text{He}$  vs.  $^{143}\text{Nd}/^{144}\text{Nd}$ -  
 313 isotope space the two FOZOs are completely resolved  
 314 (Fig. 2).  
 315

An additional characteristic of the “two FOZOs”  
 316 model is that the isotopic and hemispheric distinction is  
 317 not always observed in the highest  $^3\text{He}/^4\text{He}$  lavas erupted  
 318 in continental, back-arc and deep submarine mid-ocean  
 319 ridge settings. One possible explanation is that mantle  
 320 plume interaction with the shallow geochemical reser-  
 321 voirs in these three environments can decouple the deep  
 322 mantle  $^3\text{He}/^4\text{He}$  signatures from the associated Sr, Nd and  
 323 Pb isotopes. If so, then the high  $^3\text{He}/^4\text{He}$  ratios measured  
 324 in lavas from these three settings cannot be traced to the  
 325 same source components as the Sr, Nd and Pb isotopes.  
 326 High  $^3\text{He}/^4\text{He}$  lavas erupted in continental settings may  
 327 have suffered crustal assimilation such that the Sr, Nd and  
 328 Pb isotopes are not representative of the high  $^3\text{He}/^4\text{He}$   
 329 mantle. For example, a high  $^3\text{He}/^4\text{He}$  Baffin Island  
 330 sample (43.9 Ra; Stuart et al., 2003) with the least  
 331 depleted  $^{143}\text{Nd}/^{144}\text{Nd}$  in the suite–0.512730—is inferred  
 332 to be crustally contaminated, and is not interpreted to be a  
 333 northern hemisphere expression of FOZO-A. High  
 334  $^3\text{He}/^4\text{He}$  lavas erupted in back-arc environments may  
 335 not have lithophile isotope ratios that reflect the high  
 336  $^3\text{He}/^4\text{He}$  mantle source, owing to possible involvement of  
 337 slab-derived fluids that may affect the Sr, Nd and Pb  
 338 isotopes (even though such fluids may have a negligible  
 339 effect on the  $^3\text{He}/^4\text{He}$  ratios; Macpherson et al., 1998).  
 340 The high  $^3\text{He}/^4\text{He}$  signatures observed in deeply erupted  
 341 submarine ridge lavas may host Sr, Nd and Pb isotopes  
 342 that reflect entrainment of the shallow depleted mantle.  
 343 However, we presume that the high plume flux neces-  
 344 sary for the generation of a sub-aerially-exposed, ridge-  
 345 centered hotspot volcano enhances the plume contribution  
 346 in the erupted lavas, thus overwhelming the depleted  
 347

348 upper mantle component. Therefore, sub-aerially erupted  
 349 Bouvet Island and Iceland high  $^3\text{He}/^4\text{He}$  lavas are  
 350 included in the database (both lavas host Sr, Nd and Pb  
 351 isotopes that more enriched than nearby ridges, and are  
 352 taken to reflect a plume component).

353 How can high  $^3\text{He}/^4\text{He}$ , deep mantle signatures be  
 354 preserved and become decoupled from the deep mantle  
 355 Sr, Nd and Pb signatures in continental, back-arc and  
 356 deep submarine ridge environments? As a stimulus for  
 357 further investigation, we suggest that the He/(Sr,Nd,Pb)  
 358 ratios may be higher in the high  $^3\text{He}/^4\text{He}$  mantle  
 359 (or mantle melt) than in the shallow contaminating  
 360 reservoirs found in these three geological settings. In this  
 361 way, high  $^3\text{He}/^4\text{He}$  signatures often may be little affected  
 362 by contamination from shallow reservoirs while Sr, Nd  
 363 and Pb isotopes in such lavas can be strongly decoupled  
 364 from the original high  $^3\text{He}/^4\text{He}$  mantle source.

#### 365 4.2. FOZO-A vs. FOZO-B: is the difference sediment?

366 The isotopic enrichment observed in FOZO-A lavas,  
 367 particularly from Ofu, relative to the FOZO-B reservoir  
 368 is not easily reconciled with existing models for the  
 369 evolution of the high  $^3\text{He}/^4\text{He}$  mantle. One hypothesis for  
 370 the generation of the relative lithophile isotope and trace  
 371 element enrichment in high  $^3\text{He}/^4\text{He}$  lavas from Samoa  
 372 maintains that the high  $^3\text{He}/^4\text{He}$  mantle beneath this  
 373 hotspot was recently contaminated by rapidly cycled sedi-  
 374 ment from the nearby Tonga trench (Class and Goldstein,  
 375 2005). However, evidence from Pb-isotopes rules out  
 376 modern marine sediment contamination of the Samoan  
 377 high  $^3\text{He}/^4\text{He}$  mantle. The sediment-OIB discriminating  
 378 properties (Hart, 1988) of  $\Delta^{207}\text{Pb}/^{204}\text{Pb}-\Delta^{208}\text{Pb}/^{204}\text{Pb}$ -  
 379 Pb-isotope space (Hart, 1984) show that Samoan basalts and  
 380 modern global marine sediments (Plank and Langmuir,  
 381 1998) exhibit non-overlapping fields with diverging  
 382 trends (Fig. 3) (Jackson et al., 2007). Pb-isotope data  
 383 thus preclude the presence of modern marine sediment  
 384 (including sediments outboard of the Tonga trench) in the  
 385 FOZO-A mantle sampled by Ofu lavas.

386 Sr, Nd and Pb-isotope compositions for the high  
 387  $^3\text{He}/^4\text{He}$  Moorea lava (Hanyu and Kaneoka, 1997)  
 388 provide further evidence that the more enriched radio-  
 389 genic isotope compositions of Samoan high  $^3\text{He}/^4\text{He}$   
 390 lavas (compared to HIG lavas) are not necessarily a result  
 391 of rapidly cycled, subducted sediment (Table 1). Moorea  
 392 island is not located near a subduction zone, yet the high  
 393  $^3\text{He}/^4\text{He}$  Moorea sample has  $^{87}\text{Sr}/^{86}\text{Sr}$ ,  $^{143}\text{Nd}/^{144}\text{Nd}$  and  
 394  $^{206}\text{Pb}/^{204}\text{Pb}$ -isotope ratios similar to Ofu basalts.

395 The observation of elevated trace element concentra-  
 396 tions in Samoan high  $^3\text{He}/^4\text{He}$  lavas (relative to HIG  
 397 high  $^3\text{He}/^4\text{He}$  lavas) has been used as evidence for recent

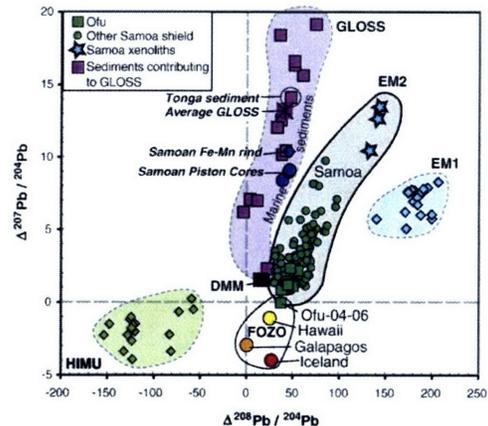


Fig. 3. In  $\Delta^{207}\text{Pb}/^{204}\text{Pb}-\Delta^{208}\text{Pb}/^{204}\text{Pb}$ -isotope space (Hart 1984, 1988), Samoan basalts (including Ofu basalts) and global marine sediments exhibit non-overlapping fields with diverging trends. The Samoan data are plotted with oceanic sediments (the Tonga sediment is circled) contributing to global subducting sediment (GLOSS) from Plank and Langmuir (1998). Samoan lavas from Ofu Island plot close to the highest  $^3\text{He}/^4\text{He}$  lavas from Hawaii, Iceland and Galapagos. Samoan xenoliths from Savai'i island are plotted as stars (Hauri et al., 1993). Other plotted Pb-isotope data include endmember MORB (average of normal ridge segments; Su, in press), EM1 (Pitcairn; Eisele et al., 2002; Hart and Hauri, unpubl. data) and HIMU (Mangaia and Tubuai; Hauri and Hart, 1993; Woodhead, 1996). The figure is adapted from Jackson et al. (2007), and data for sediments from the Samoan region (piston cores) and the ferromanganese rind are from the same source. Pb-isotope data preclude the presence of modern marine sediment (including sediment recently subducted into the Tonga trench) in the FOZO-A mantle sampled by Ofu lavas.

388 sediment contamination of the Samoan high  $^3\text{He}/^4\text{He}$  mantle  
 389 (Class and Goldstein, 2005). However, our favored  
 390 explanation for the elevated trace element concentrations  
 391 in the Ofu (and Moorea) high  $^3\text{He}/^4\text{He}$  lavas is that they are  
 392 products of low degrees of melting, a mechanism that can  
 393 greatly increase the trace element concentrations in lavas  
 394 relative to their mantle source. Ofu and Moorea high  
 395  $^3\text{He}/^4\text{He}$  lavas are alkaline, a petrologic feature that is  
 396 uncommon among high  $^3\text{He}/^4\text{He}$  HIG lavas and one that is  
 397 likely to be a result of low degrees of mantle melting. Therefore,  
 398 the alkaline nature of Samoan high  $^3\text{He}/^4\text{He}$  lavas may help  
 399 resolve the apparent paradox of high U and Th concentrations  
 400 in Samoan lavas that also host high  $^3\text{He}/^4\text{He}$  ratios (Fig. 4).

401 Although the Samoan plume exhibits a (low  $^3\text{He}/^4\text{He}$ ) component  
 402 with recycled sediment, the high  $^3\text{He}/^4\text{He}$  Samoan lavas from  
 403 Ofu exhibit no evidence for an ancient recycled sediment  
 404 component. The most enriched lavas from Samoa display  
 405 evidence for an ancient recycled sediment component (White  
 406 and Hofmann, 1982; Wright 417

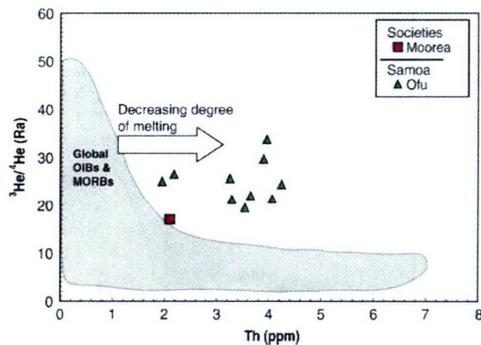


Fig. 4.  $^3\text{He}/^4\text{He}$ –Th for Ofu lavas compared to other OIBs and MORBs. Ofu basalts and the high  $^3\text{He}/^4\text{He}$  basalt from Moorea (MO-01-01) exhibit higher Th concentrations than high  $^3\text{He}/^4\text{He}$  basalts from other hotspots (see Table 1). The field for global OIBs (excluding Samoa) and MORBs is from Class and Goldstein (2005). All Ofu and Moorea Th data shown are by ICP, and the Ofu Th data are not corrected for olivine or cpx fractionation. The high Th concentrations in the Ofu and Moorea high  $^3\text{He}/^4\text{He}$  basalts are interpreted as resulting from lower degrees of melting, a hypothesis that is consistent with the alkaline nature of the lavas from these two islands.

418 and White, 1987; Farley et al., 1992), including high  
 419  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios (up to 0.7216) and low  $^3\text{He}/^4\text{He}$  ratios  
 420 ( $<5$  Ra; Jackson et al., 2007). By contrast,  $^{87}\text{Sr}/^{86}\text{Sr}$  and  
 421  $^{143}\text{Nd}/^{144}\text{Nd}$  isotopes measured in Ofu lavas are among  
 422 the most depleted in Samoa (Table 1). Furthermore, the  
 423 positive Pb anomalies and high Ba/Nb ratios that are  
 424 diagnostic of sediment are observed in the most iso-  
 425 topically enriched Samoan lavas, but are absent in the  
 426 isotopically depleted Ofu lavas (Fig. 5). The Ba/Nb and  
 427 Pb/Pb\* measured in the HIG (FOZO-B) high  $^3\text{He}/^4\text{He}$   
 428 lavas are indistinguishable from the Ofu lavas, providing  
 429 further evidence that the Ofu lavas do not host an ancient  
 430 recycled sediment component (Fig. 5).

#### 431 4.3. Ofu lavas from recycled harzburgite?

432 Some recent models advocate a role for ancient,  
 433 depleted, high  $^3\text{He}/^4\text{He}/(\text{U}+\text{Th})$  harzburgitic mantle litho-  
 434 sphere in the formation of the high  $^3\text{He}/^4\text{He}$  mantle  
 435 domain (Anderson, 1998; Paman et al., 2005; Heber  
 436 et al., 2007). This assumes that cpx-poor melt residues  
 437 may have the property of increased  $^3\text{He}/^4\text{He}/(\text{U}+\text{Th})$  relative  
 438 to the initial, unmelted source (Parman et al., 2005), and  
 439 thus may preserve high  $^3\text{He}/^4\text{He}$  ratios over time. Trace  
 440 element budgets in abyssal peridotites provide a critical  
 441 test of the hypothesis that recycled harzburgite is involved  
 442 in the source of high  $^3\text{He}/^4\text{He}$  OIB lavas. While a harz-  
 443 burgite lithology is consistent with the cpx-poor nature of  
 444 the upper oceanic lithosphere, the abyssal peridotites that

comprise the upper oceanic lithosphere are too depleted in  
 445 incompatible trace elements (Workman and Hart, 2005) to  
 446 act as a source for the high  $^3\text{He}/^4\text{He}$  Ofu basalts. 447  
 448 For example, the maximum Sr concentration that can be  
 449 generated by extremely low degrees of aggregated frac-  
 450 tional melting of modern harzburgitic abyssal peridotites  
 451 (samples with  $<5\%$  cpx, with an average reconstructed  
 452 bulk rock Sr concentration of  $\sim 0.02$  ppm; see figure  
 453 available as supplementary data in the Appendix) is  
 454  $\sim 5$  ppm, which is far below the Sr concentrations  
 455 ( $\sim 600$  ppm) of primary Ofu lavas. Trace element budgets  
 456 in modern abyssal peridotites suggest that Ofu basalts are

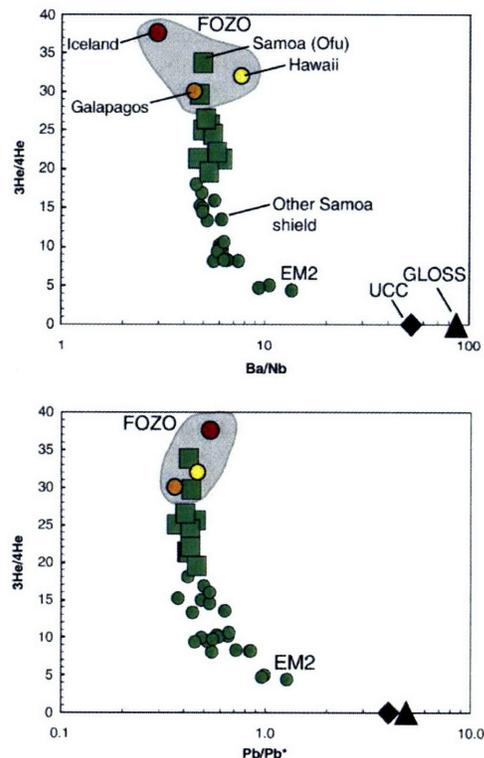


Fig. 5.  $^3\text{He}/^4\text{He}$  variation with Ba/Nb and Pb/Pb\* in Samoan lavas and high  $^3\text{He}/^4\text{He}$  samples from Hawaii (Kurz et al., 1982), Iceland (Hilton et al., 1999) and the Galapagos (Kurz and Geist, 1999; Saal et al., 2007). Ba/Nb ratios in high  $^3\text{He}/^4\text{He}$  lavas from all four hotspots are low, and do not indicate input from marine sediment (GLOSS; Plank and Langmuir, 1998) or upper continental crust (UCC; Rudnick and Gao, 2003). Pb/Pb\* values in high  $^3\text{He}/^4\text{He}$  lavas are also unlike GLOSS and UCC. On the other hand, the enriched (EM2) Samoan lavas have low  $^3\text{He}/^4\text{He}$ , high Ba/Nb and high Pb/Pb\* values that suggest a recycled sediment signature (Jackson et al., 2007). All trace element data were measured by ICP, but trace element data from the Iceland and Hawaii samples are unpublished.  $\text{Pb}/\text{Pb}^* = \text{Pb}_N / \sqrt{(\text{Ce}_N \times \text{Nd}_N)}$ , where N signifies normalization to primitive mantle (McDonough and Sun, 1995).

not the melts of ancient, recycled (depleted) oceanic upper mantle lithosphere.

The base of the oceanic mantle lithosphere, which is thought to have a lherzolitic composition similar to DMM (Workman and Hart, 2005), can produce the Sr concentrations observed in Ofu lavas if it is melted at low degrees. Lherzolitic melt residues also may have the property of increased  $^3\text{He}/(\text{U}+\text{Th})$  relative to the initial, unmelted source (Parman et al., 2005). However, results from recent helium partitioning experiments yield conflicting results (Heber et al., 2007; see Section 4.5 below), and suggest that lherzolitic melt residues will have greatly diminished  $^3\text{He}/(\text{U}+\text{Th})$ . The observation of radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  in high  $^3\text{He}/^4\text{He}$  lavas from Iceland is inconsistent with a purely peridotitic origin of the high  $^3\text{He}/^4\text{He}$  mantle (Brandon et al., in press), and suggest that helium partitioning studies based on melting pure peridotite compositions may not be applicable to melting of the high  $^3\text{He}/^4\text{He}$  reservoirs in the earth's mantle.

#### 4.4. Ofu lavas from chondritic primitive (undegassed) mantle or regassed mantle?

An alternative hypothesis suggests that high  $^3\text{He}/^4\text{He}$  Samoan lavas are derived from chondritic primitive mantle (referred to as PHEM; Farley et al., 1992). However, such a model fails to explain the non-chondritic  $^{143}\text{Nd}/^{144}\text{Nd}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  ratios obtained in Ofu lavas: the  $^{143}\text{Nd}/^{144}\text{Nd}$  and  $^{206}\text{Pb}/^{204}\text{Pb}$  ratios of Ofu lavas, which are 0.51283 and 19.19, respectively, are much higher than the corresponding chondritic primitive mantle values, which are 0.512611 (Boyet and Carlson, 2005) and 17.675, respectively.

Another, more recent model attempts to reconcile the non-chondritic lithophile isotopes in high  $^3\text{He}/^4\text{He}$  lavas assumes that a depleted upper mantle source was “regassed” by mixing with a small proportion of helium-rich, high  $^3\text{He}/^4\text{He}$  chondritic mantle (Stuart et al., 2003; Ellam and Stuart, 2004). This model suggested that regassed, depleted mantle is in the source of the high  $^3\text{He}/^4\text{He}$  ratios found in basalts associated with the proto-Iceland plume (PIP). However, Ofu lavas fall well outside of the array formed by PIP lavas in  $^3\text{He}/^4\text{He}$ – $^{143}\text{Nd}/^{144}\text{Nd}$  isotope space (Ellam and Stuart, 2004), highlighting the need for a model that includes a heterogeneous high  $^3\text{He}/^4\text{He}$  mantle.

#### 4.5. Implications of $^{142}\text{Nd}/^{144}\text{Nd}$ for “re-enrichment” of the high $^3\text{He}/^4\text{He}$ mantle

The discovery of  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios in accessible terrestrial rocks that are higher than chondrite (Boyet and

Carlson, 2005) has important consequences for the origin and evolution of the two FOZOs. The terrestrial  $^{142}\text{Nd}/^{144}\text{Nd}$  anomaly indicates that all measured terrestrial rocks were derived from a reservoir that had superchondritic Sm/Nd during the lifetime of  $^{146}\text{Sm}$ , as imperfect mixing of nucleosynthetic material in the solar nebula (Ranen and Jacobsen, 2006) does not explain the terrestrial excess in  $^{142}\text{Nd}$  (Hidaka et al., 2003; Andreasen and Sharma, 2006; Carlson et al., 2007; Wombacher and Becker, 2007). For example, if Sm/Nd ratios were heterogeneously distributed in the solar nebula at the time of accretion, bulk silicate earth (BSE) may have acquired higher Sm/Nd and thus higher time-integrated  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios than chondrites (Boyet and Carlson, 2006). If the assumptions for a non-chondritic BSE are valid (Boyet and Carlson, 2006), then both FOZO-A and FOZO-B are isotopically enriched relative to BSE (Fig. 6, top panel). Alternatively, if BSE has chondritic Sm/Nd, the terrestrial  $^{142}\text{Nd}/^{144}\text{Nd}$  anomaly in accessible terrestrial mantle rocks could have been generated by an early global terrestrial differentiation event within 30 Myr of accretion (Boyet and Carlson, 2005). In this case, the resulting early depleted reservoir (EDR) has superchondritic Sm/Nd and evolves superchondritic  $^{142}\text{Nd}/^{144}\text{Nd}$  (a complementary hidden early enriched reservoir, or EER, evolved sub-chondritic  $^{142}\text{Nd}/^{144}\text{Nd}$ ; Boyet and Carlson, 2005). The FOZO reservoirs exhibit  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios that are lower (more enriched) than the minimum  $^{143}\text{Nd}/^{144}\text{Nd}$  of the EDR (Fig. 6, bottom panel). If the FOZO reservoirs (like all accessible terrestrial mantle rocks) were ultimately derived from the EDR at some point in earth's history, then the high  $^3\text{He}/^4\text{He}$  FOZO mantle has been re-enriched since the early differentiation event. In summary, if the terrestrial  $^{142}\text{Nd}/^{144}\text{Nd}$  anomaly relative to chondrites is due to the decay of  $^{146}\text{Sm}$ , the observed  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios in FOZO-A and FOZO-B lavas require that they were re-enriched relative to either a non-chondritic BSE or the EDR.

There are a number of mechanisms by which the high  $^3\text{He}/^4\text{He}$  mantle could have been re-enriched. However, the enriched material added to the high  $^3\text{He}/^4\text{He}$  reservoirs must also have the property of preserving a high  $^3\text{He}/^4\text{He}$  signature over time. Thus, re-enrichment by addition of recycled sediments seems unlikely, as lavas exhibiting clear evidence of sediment recycling, such as the high  $^{87}\text{Sr}/^{86}\text{Sr}$  Samoan lavas from Samoa, exhibit low  $^3\text{He}/^4\text{He}$  (Jackson et al., 2007). On the other hand, re-enrichment by addition of recycled (oceanic crust) eclogite plums to a high  $^3\text{He}/^4\text{He}$  mantle source may not necessarily diminish the high  $^3\text{He}/^4\text{He}$  signature (Brandon et al., in press). As an alternative to re-enrichment by eclogite addition, depleted oceanic mantle lithosphere

557 that has been re-enriched (metasomatized) with melt may  
558 serve as a source for high  $^3\text{He}/^4\text{He}$  basalts (White, 2005).  
559 However, if the re-enriching melt is a result of imperfect  
560 melt extraction near a mid-ocean ridge (Workman et al.,  
561 2004), the  $^3\text{He}/(\text{U}+\text{Th})$  of the melt will have to be  
562 similar to or higher than DMM to preserve high  $^3\text{He}/^4\text{He}$   
563 in the recycled upper oceanic mantle lithosphere over  
564 time (i.e., following subduction and isolation in the  
565 lower mantle). This condition requires that the compat-  
566 ibility of helium is similar to or lower than U and Th  
567 during peridotite melting.

568 However, such partitioning behavior is inconsistent  
569 with the helium partitioning results of Parman et al.  
570 (2005). Recent helium partitioning results of Heber et al.  
571 (2007) do suggest that He is less compatible than U and  
572 Th during peridotite melting (assuming a DMM  
573 lherzolite lithology from Workman and Hart (2005)  
574 and U and Th partition coefficients from Kelemen et al.  
575 (2004)). If the partitioning results of Heber et al. (2007)  
576 are correct, then oceanic mantle lithosphere hosting  
577 trapped melt may preserve high  $^3\text{He}/^4\text{He}$  over time.  
578 Thus, long-term isolation of melt-impregnated oceanic

579 mantle lithosphere may provide a viable mechanism for  
580 preserving a high  $^3\text{He}/^4\text{He}$  signature, and may have  
581 potential for describing the radiogenic isotope composi-  
582 tions and trace element budgets observed in global high  
583  $^3\text{He}/^4\text{He}$  lavas.

#### 4.6. Origin of the hemispheric separation of FOZO-A and FOZO-B: hints from the DUPAL anomaly

584 The observation of a southern hemisphere high  
585  $^3\text{He}/^4\text{He}$  domain that is isotopically more enriched  
586 than its northern hemisphere counterpart is reminiscent  
587 of the DUPAL anomaly (Hart, 1984)—a globe-encircling  
588 feature of isotopic enrichment observed primarily in the  
589 southern hemisphere mantle—and indicates a long-term  
590 separation of the earth's northern and southern hemi-  
591 sphere high  $^3\text{He}/^4\text{He}$  mantle. While there are a number of  
592 processes by which the re-enriched high  $^3\text{He}/^4\text{He}$  mantle  
593 can be generated, the mechanisms responsible for the  
594 hemispheric separation, and long-term preservation, of  
595 two isotopically distinct FOZO reservoirs are elusive.

596 Like the high  $^3\text{He}/^4\text{He}$  reservoir (Kurz et al., 1982;  
597 Hart et al., 1992) the DUPAL anomaly was suggested to  
598 be an ancient feature residing in the lower mantle (Hart,  
599 1988; Castillo, 1988). The DUPAL anomaly is the only

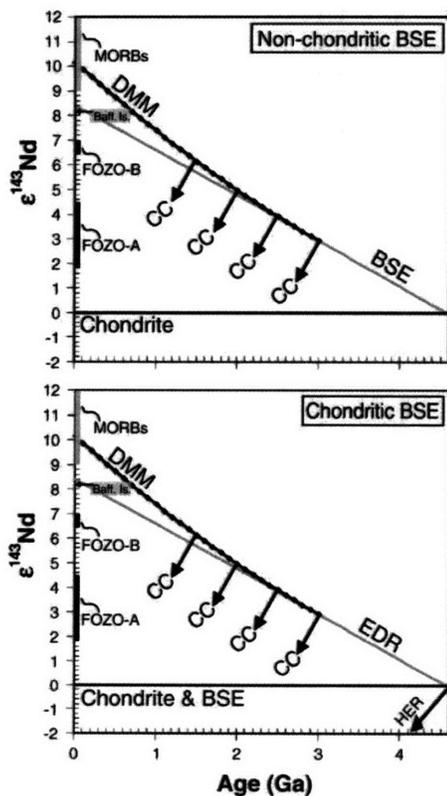


Fig. 6. Implications of terrestrial  $^{142}\text{Nd}/^{144}\text{Nd}$  anomalies for the  $^{143}\text{Nd}/^{144}\text{Nd}$  evolution of BSE and the origin of the two FOZOs. If the observed superchondritic terrestrial mantle  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios (Boyet and Carlson, 2005) were generated by  $^{146}\text{Sm}$  decay, either 1.) (top panel) the earth accreted from non-chondritic material and BSE has superchondritic Sm/Nd, or 2.) (bottom panel) BSE is chondritic but underwent an early depletion event and all available terrestrial mantle rocks derive from an early depleted reservoir (EDR) with superchondritic Sm/Nd. All calculations are after Boyet and Carlson (2005, 2006), and assume a chondrite average  $^{147}\text{Sm}/^{144}\text{Nd}$  of 0.1948,  $^{143}\text{Nd}/^{144}\text{Nd}$  of 0.512611, a solar system initial  $^{146}\text{Sm}/^{144}\text{Sm}$  of 0.008, and an identical age of 4.567 Ga for the accretion time of the earth (top panel) and for the early differentiation event (bottom panel). Given these assumptions, the non-chondritic BSE and the EDR require  $^{147}\text{Sm}/^{144}\text{Nd}$  ratios of  $\sim 0.209$  to generate the  $^{142}\text{Nd}/^{144}\text{Nd}$  anomalies that are 20 ppm higher than chondrite (later formation times for the accretion of the earth, or for the early differentiation event, require even higher Sm/Nd ratios). Thus, the lowest present-day  $^{143}\text{Nd}/^{144}\text{Nd}$  value for a non-chondritic BSE and the EDR is 0.51304 ( $\epsilon^{143}\text{Nd} = +8.4$ ). Both FOZOs exhibit  $^{143}\text{Nd}/^{144}\text{Nd}$  values lower than the non-chondritic BSE (top panel) and the EDR (bottom panel), and both FOZOs thus exhibit evidence for re-enrichment. Starting at 3 Ga, the DMM reservoir evolved by continuous extraction of continental crust (Workman and Hart, 2005) from the non-chondritic BSE (top panel) or from the EDR (bottom panel); DMM calculations are identical to Boyet and Carlson (2006). The trajectories of CC and EER are estimated. CC is not extracted as a single event, but is continuously extracted after 3 Ga. The MORB field is from Boyet and Carlson (2005) and FOZO-A and FOZO-B are from Table 2; the Baffin Island (Baff. Is.) high  $^3\text{He}/^4\text{He}$  lava (Stuart et al., 2003) is plotted for reference.  $\epsilon_{\text{Nd}}(t) = [^{143}\text{Nd}/^{144}\text{Nd}_{\text{Sample}}(t)/^{143}\text{Nd}/^{144}\text{Nd}_{\text{Chondrite}}(t) - 1] \times 10^4$ .

602 other mantle domain suggested to occupy similar,  
603 hemispheric proportions. The DUPAL anomaly shows  
604 that the southern hemisphere exhibits generally more  
605 isotopically enriched mantle domains than the northern  
606 hemisphere, and we find a similar southern hemisphere  
607 enrichment in the high  $^3\text{He}/^4\text{He}$  FOZO-A reservoir; the  
608 depletion in the northern hemisphere DUPAL reference  
609 reservoir is mirrored in the FOZO-B reservoir. The  
610 geographic and geochemical similarities between the  
611 FOZO and DUPAL reservoirs suggest that their origin  
612 may have been linked.

613 One hypothesis for the formation of the DUPAL  
614 domain is that oceanic mantle and crustal lithosphere were  
615 injected into the mantle, thereby enriching this region  
616 of the mantle. This model has the advantage of being  
617 compatible with two of the possible re-enrichment  
618 mechanisms suggested for the FOZO reservoirs in Section  
619 4.5: 1.) oceanic crustal eclogite plums were added to the  
620 depleted high  $^3\text{He}/^4\text{He}$  mantle via subduction, or 2.) re-  
621 enriched (metasomatized by melt near a mid-ocean ridge)  
622 oceanic mantle lithosphere was subducted and isolated in  
623 the (deep?) mantle and became the high  $^3\text{He}/^4\text{He}$  mantle  
624 domain. These injection models may produce a random  
625 pattern of isotopic enrichment throughout the mantle, with  
626 no coherent hemispheric pattern (Hart, 1984). However,  
627 if injection were focused around the perimeter of a  
628 supercontinent during a period of anomalous subduction,  
629 a hemispheric pattern in the FOZO and DUPAL reservoirs  
630 might emerge. This model would suggest that the  
631 hemispheric and isotopic separation may be a surviving  
632 artifact of the paleo-arrangement of the subduction zones  
633 and continents during the formation of the FOZO reser-  
634 voirs. While the timing of the formation of the FOZO  
635 reservoirs is unknown, we note that the Northern  
636 Hemisphere Reference Line (NHRL) separates the two  
637 FOZOs in  $^{207}\text{Pb}/^{204}\text{Pb}$ – $^{206}\text{Pb}/^{204}\text{Pb}$ -isotope space (see  
638 figure included in the supplementary data in the  
639 Appendix), possibly suggesting a similar formation time  
640 ( $\sim 1.8$  Ga) for the DUPAL and FOZO reservoirs (Pb-  
641 isotope ratios in high  $^3\text{He}/^4\text{He}$  lavas are displaced from  
642 the Geochron and preclude coeval formation of the FOZO  
643 and the hypothetical EDR reservoirs).

#### 644 4.7. FOZO-A and FOZO-B: implications for mantle 645 dynamics

646 A clear implication of the isotopic and geographic  
647 separation of the two high  $^3\text{He}/^4\text{He}$  reservoirs is that they  
648 had to be isolated from each other for long timescales. The  
649 long-term separation of the two FOZO domains requires  
650 preservation despite convective stirring, a mechanism that  
651 efficiently attenuates mantle heterogeneities (van Keken

et al., 2002). However, the hemispheric heterogeneity in 652  
the high  $^3\text{He}/^4\text{He}$  reservoir is apparently not a feature 653  
preserved in the convecting upper mantle sampled by 654  
mid-ocean ridges (see Sections 3.4.1). Thus, the shallow 655  
mantle may not be an ideal location for the preservation of 656  
the FOZO-A and FOZO-B mantle domains. However, the 657  
less rapid convective motions of the lower mantle may 658  
make it a more suitable home for the FOZOs (Hart et al., 659  
1992; Macpherson et al., 1998; van Keken et al., 2002; 660  
Class and Goldstein, 2005). 661

662 One mechanism for the preservation of hemispheric-  
scale heterogeneity may be the isolation of the high 663  
 $^3\text{He}/^4\text{He}$  domains in a dense boundary layer at the core- 664  
mantle boundary (CMB). Seismic tomography suggests 665  
that some hotspots, like Samoa and Hawaii, may well 666  
originate as upwellings from this region of the mantle 667  
(Montelli et al., 2006), and many hotspots with high 668  
 $^3\text{He}/^4\text{He}$  have been associated with velocity anomalies in 669  
the deep mantle (Courtillet et al., 2003). The post- 670  
perovskite phase proposed to exist at this depth may 671  
exhibit a sufficiently large density contrast to isolate it 672  
from the overlying convecting mantle (Guignot et al., 673  
2007), and a new seismic technique may allow detection 674  
of this phase at the base of the mantle (van der Hilst et al., 675  
2007). Exploiting this new seismic tool, it may be pos- 676  
sible to better resolve the spatial relationships between 677  
velocity anomalies at the CMB and high  $^3\text{He}/^4\text{He}$  ocean 678  
islands at the earth's surface. Thus, a confluence of 679  
geochemical and geophysical observations may ulti- 680  
mately reveal the mechanism responsible for the long- 681  
term preservation of the hemispheric heterogeneity in the 682  
deep mantle, a feature that hinges on new observations of 683  
high  $^3\text{He}/^4\text{He}$  in Samoan lavas from Ofu. 684

## 685 5. Conclusions

686 From this study we draw the following conclusions: 686

- 687 (1) New high  $^3\text{He}/^4\text{He}$  ratios (19.5–33.8 times 687  
atmospheric) from Ofu island are the highest 688  
from Samoa (and the southern hemisphere), and 689  
place Samoa in the same category of high  $^3\text{He}/^4\text{He}$  690  
hotspots as Hawaii, Iceland and the Galapagos. 691
- 692 (2) The new Ofu data reveal that at least two distinct 692  
high  $^3\text{He}/^4\text{He}$  reservoirs—one more isotopical- 693  
ly enriched than the other—exist in the earth's 694  
mantle. 695
- 696 (3) The two high  $^3\text{He}/^4\text{He}$  reservoirs separate in the 696  
earth's northern (FOZO-B, boreal) and southern 697  
(FOZO-A, austral) hemispheres. 698
- 699 (4) The trace element budgets and isotopic composi- 699  
tions of the new high  $^3\text{He}/^4\text{He}$  samples from Samoa 700

are inconsistent with derivation from recycled harzburgite. Additionally, the Samoan high  $^3\text{He}/^4\text{He}$  mantle does not appear to be contaminated with rapidly cycled sediment from the Tonga subduction zone. Furthermore, the Nd and Pb isotopes of the highest  $^3\text{He}/^4\text{He}$  Samoan basalts demonstrate that they were not derived from a chondritic primitive mantle.

- (5) If the terrestrial  $^{142}\text{Nd}/^{144}\text{Nd}$  anomaly relative to chondrites is due to the decay of  $^{146}\text{Sm}$ , the observed  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios in FOZO-A and FOZO-B lavas require that they were re-enriched relative to either a non-chondritic BSE or the EDR.
- (6) The high  $^3\text{He}/^4\text{He}$  mantle is an ancient reservoir, and the discovery of isotopically distinct northern and southern hemisphere high  $^3\text{He}/^4\text{He}$  mantle domains suggests that these regions of the mantle escaped the convective mixing and stirring that has efficiently attenuated heterogeneities in the upper mantle. This observation provides an important constraint for future dynamic and isotopic models describing the evolution of the earth's mantle.

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#### Appendix A. Supplementary data

Supplementary data associated with this article can be found, in the online version, at doi:10.1016/j.epsl.2007.09.023.

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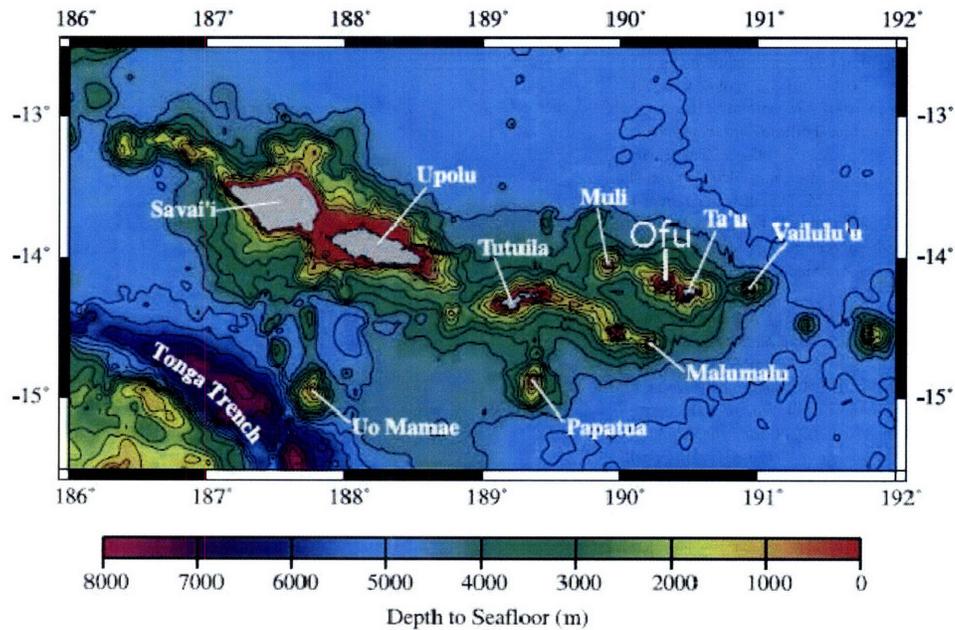
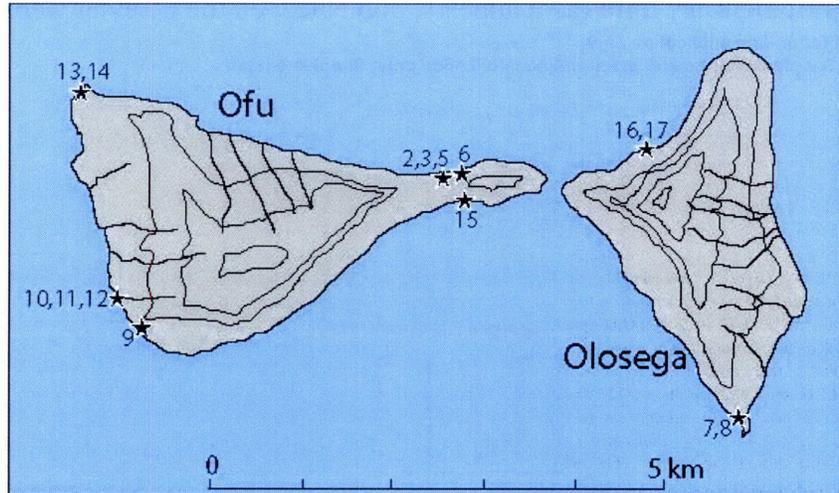
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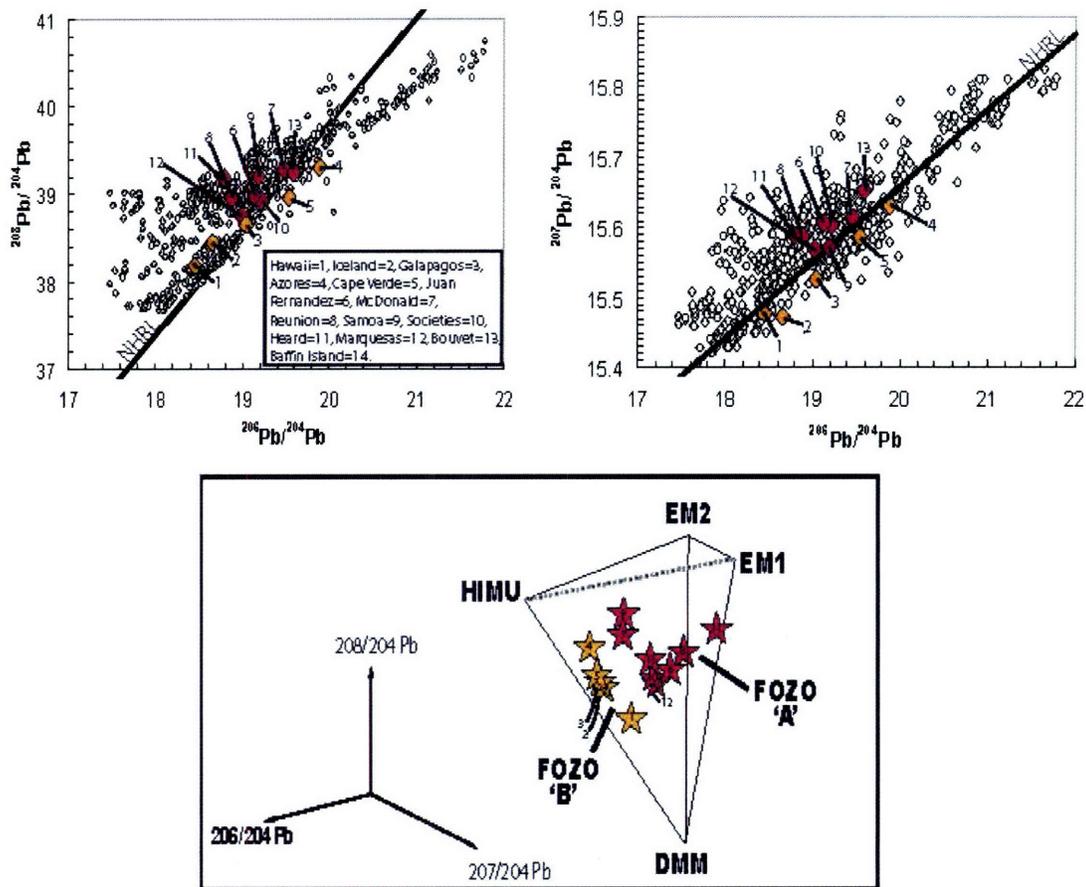
**Appendix A. Supplementary data** (as it appears, published, on the Elsevier website)

**Supplementary material for on-line publication only**

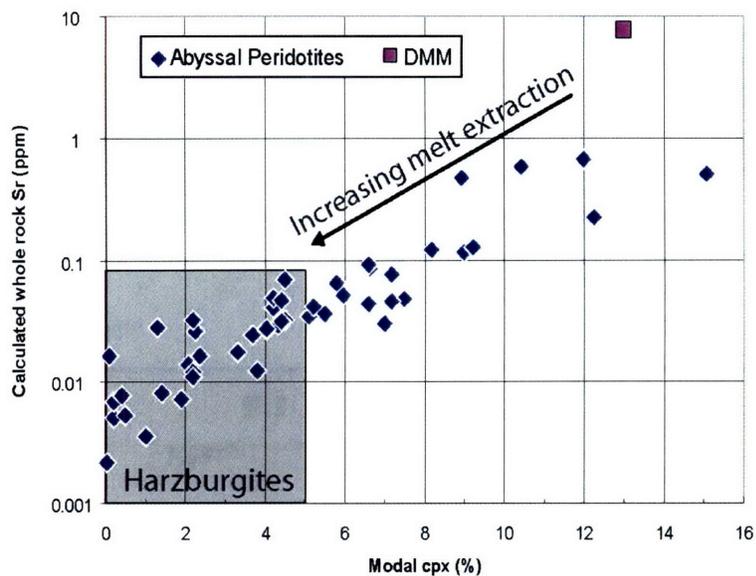
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Supplementary Material Fig 1. Upper panel: Location map for samples taken from Ofu and Olosega islands that are reported in this study. Samples were taken from the major geologic formations recorded by Stice and McCoy (1968). Lower Panel: Ofu Island relative to the subaerial islands (Savai'i, Upolu, Tutuila and Ta'u) and submarine volcanoes (Muli, Malumalu and Vailulu'u) of the Eastern volcanic province (map from Workman et al. (2004)). Papatua and Uo Mamae are isolated, "off-axis" seamounts. Note Samoa's proximity to the northern termination of the Tonga arc/trench system.



Supplementary Material Fig. 2. The separation of FOZO-A and FOZO-B in Pb-isotope space. Top panels: In two dimensional Pb-isotope space ( $^{208}\text{Pb}/^{204}\text{Pb}$  vs.  $^{206}\text{Pb}/^{204}\text{Pb}$  and  $^{207}\text{Pb}/^{204}\text{Pb}$  vs  $^{206}\text{Pb}/^{204}\text{Pb}$  isotope spaces), the two FOZOs can be separated by a line. However, the Northern Hemisphere Reference Line (NHRL) from Hart (1984) does not appear to divide the two FOZO's in  $^{208}\text{Pb}/^{204}\text{Pb}$  –  $^{206}\text{Pb}/^{204}\text{Pb}$  space, but does divide them in  $^{207}\text{Pb}/^{204}\text{Pb}$  –  $^{206}\text{Pb}/^{204}\text{Pb}$  space. Open black diamonds are global OIB data. Lower panel: FOZO-A and FOZO-B separate in 3D Pb-isotope space, indicating that Pb-isotopes are consistent with the hemispheric separation of the two high  $^3\text{He}/^4\text{He}$  reservoirs. Pb-isotope data used to define FOZO-A and FOZO-B are found in Table 2.



Supplementary Material Fig. 3. Modal abundance of cpx compared to reconstructed whole-rock Sr concentrations in abyssal peridotites, a proxy for recycled, depleted, oceanic upper mantle lithosphere. Sr concentrations in abyssal peridotites diminish rapidly with small reductions in cpx modal abundance, a result of melt extraction from DMM in a fractional melting regime. Harzburgites and other cpx-poor lithologies that are produced in the mantle lithosphere at mid-ocean ridges are extremely trace element depleted. These peridotites may not be good candidates for a source that generates the enriched trace-element budgets observed in Ofu lavas. The abyssal peridotite compilation accompanies Workman and Hart (2005), and the estimate for the Sr concentration in DMM concentrations is from the same source. Samples with modal plagioclase have been excluded.

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## Chapter 4

# **High $^3\text{He}/^4\text{He}$ hotspot lavas expose the Earth's "missing" titanium, tantalum and niobium (TITAN): The missing link between continental crust and depleted mantle found?**

### Abstract

A shortage of the elements titanium, tantalum and niobium (TITAN) exists in the Earth's shallow geochemical reservoirs—the depleted MORB (mid-ocean ridge basalt) mantle and continental crust—and the location of these missing elements is unknown. Here we report evidence for a global, TITAN-enriched reservoir sampled by OIBs (ocean island basalts) with high  $^3\text{He}/^4\text{He}$  ratios, an isotopic signature associated with the deep mantle. Excesses of Ti (and to a lesser degree Nb and Ta) correlate remarkably well with  $^3\text{He}/^4\text{He}$  in a dataset of global OIBs. The observation of TITAN enrichment in high  $^3\text{He}/^4\text{He}$  OIB lavas suggests that the mantle domain hosting the Earth's "missing" TITAN is sampled by deep, high  $^3\text{He}/^4\text{He}$  mantle plumes. The TITAN enrichment in the high  $^3\text{He}/^4\text{He}$  reservoir has profound implications for the origin of the high  $^3\text{He}/^4\text{He}$  mantle component, and suggests that, far from being a primitive reservoir, or simply a depleted peridotite reservoir, the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs appears to host a component of recycled, refractory, rutile-bearing oceanic crust that was processed in subduction zones.

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## 1. Introduction

The standard model for the evolution of the silicate earth maintains that the depleted MORB (mid-ocean ridge basalt) mantle (DMM) is the residue of continental crust (CC) extraction from an early primitive mantle (Jacobsen and Wasserburg, 1979; O’Nions et al., 1979; Allegre et al., 1980; Hofmann, 1988, 1997). If the earth has chondritic abundances of the refractory elements, DMM and CC must be geochemically complementary reservoirs within the earth. However, the TITAN trio of elements are prominently depleted in the continents (Rudnick and Gao, 2003), and their absence is not balanced by a corresponding enrichment in DMM (Workman and Hart, 2005). Thus, another deeper reservoir hosting the missing TITAN elements has been proposed to exist in the earth (McDonough, 1991; Rudnick et al., 2000; Kamber and Collerson, 2000).

Oceanic plates are formed by melting and depletion of the upper mantle at mid-ocean ridges. The resulting oceanic lithosphere, composed of mafic oceanic crust and the uppermost region of the depleted peridotite mantle, is subducted back into the mantle at trenches, thereby contributing to its compositional heterogeneity (Hofmann and White, 1980, 1982; Chase, 1981; Zindler and Hart, 1986; van Keken et al., 2002). During subduction, the mafic portion of the plate is dehydrated and may be partially melted, and the resulting lavas erupted at subduction zone volcanoes are depleted in the TITAN elements. Incompatible elements are largely lost to the overlying mantle during dehydration and melting of the eclogite portion of slabs. By contrast, titanium-rich phases, such as rutile, may preferentially sequester the TITAN elements in the mafic portion of the downgoing slab, balancing the depletion observed in subduction zone lavas (Green and Pearson, 1986; Ryerson and Watson, 1987; Brennan et al., 1994; Foley et al., 2000; Schmidt et al., 2004; Kessel et al., 2005). Refractory, rutile-bearing eclogites have been subducted in large quantities over geologic time, and may form a reservoir in the mantle that hosts the Earth’s missing TITAN (McDonough, 1991; Rudnick et al., 2000).

TITAN enrichment in hotspot lavas also has enormous potential as a geochemical tracer for recycled oceanic plates (Rudnick et al., 2000). If TITAN-enriched refractory eclogites are returned to the surface in mantle plumes, their presence would be evident as TITAN enrichment in hotspot lavas. A common mantle dynamics paradigm maintains

that mantle plumes complete the process of recycling by transporting plate remnants from the core-mantle boundary back to the surface, where they melt and erupt as ocean island basalts (OIBs) (Hofmann and White, 1980, 1982; Chase, 1981). However, the reservoir hosting the Earth's missing TITAN has been difficult to detect in hotspot lavas (McDonough, 1991; Rudnick et al., 2000).

Rare, high  $^3\text{He}/^4\text{He}$  ( $>30$  Ra, ratio to atmosphere) ratios in lavas erupted at some hotspots, including Hawaii, Iceland, Galapagos and Samoa, are thought to be tracers of buoyantly upwelling mantle plumes that sample an ancient reservoir residing in the mantle (e.g., Kurz et al., 1982; Hart et al., 1992). Variouslly called FOZO (Focus Zone; Hart et al., 1992; Jackson et al., 2007a), PHEM (Primitive Helium Mantle; Farley et al., 1992) or C (Common; Hanan and Graham, 1996), the precise location (shallow or deep mantle) of the high  $^3\text{He}/^4\text{He}$  reservoir is a source of intense debate (e.g., Anderson, 1998). Nonetheless, there is a growing consensus that this reservoir hosts a significant component of depleted mantle peridotite (e.g., Hart et al., 1992; Anderson, 1998; Parman et al., 2005; Heber et al., 2007). Recent work suggests that the high  $^3\text{He}/^4\text{He}$  mantle may also host a component of recycled eclogite (Dixon et al., 2001; Brandon et al., 2007). In this paper, we argue that the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs hosts a component of recycled, refractory eclogite, and that this component balances Earth's budget for the elements titanium, tantalum and niobium (TITAN).

## 2. New Data and Observations

We report new trace element data by ICP-MS (inductively coupled plasma mass spectrometer) on the highest  $^3\text{He}/^4\text{He}$  lavas from Hawaii (32.3 Ra; Kurz et al., 1982), Iceland (37.7 Ra; Hilton et al., 1999) and Samoa (33.8 Ra; Jackson et al., 2007a) (see Table 1). In Fig. 1, these new data are presented together with previously published ICP-MS trace element data for the highest  $^3\text{He}/^4\text{He}$  Galapagos lava (30.2 Ra; Kurz and Geist, 1999; Saal et al., 2007). These high  $^3\text{He}/^4\text{He}$  lavas exhibit Ti, Ta, and Nb excesses, or positive anomalies, relative to elements of similar compatibility in peridotite (on a primitive mantle normalized basis, see Fig. 1). While the association of positive TITAN anomalies and high  $^3\text{He}/^4\text{He}$  (or plume) signatures were previously observed regionally in

Hawaii (Dixon et al., 2001), Iceland (Fitton, 1997) and the Galapagos (Kurz and Geist, 1999; Saal et al., 2007), we suggest that the large, positive TITAN anomalies are a global phenomenon in high  $^3\text{He}/^4\text{He}$  OIBs. The primitive-mantle normalized trace element patterns (spidergrams) of the highest  $^3\text{He}/^4\text{He}$  lavas from Hawaii, Iceland, Galapagos and Samoa all share prominent, anomalous enrichment in the TITAN elements compared to elements of similar compatibility in peridotite (Fig. 1). In fact, the Nb/U ratios in the high  $^3\text{He}/^4\text{He}$  OIB lavas are all higher than the average Nb/U value of 47 previously proposed for OIBs and MORBs (Hofmann et al., 1986).

By contrast, the mantle endmembers (Zindler and Hart, 1986) with low  $^3\text{He}/^4\text{He}$ , including HIMU (high ‘ $\mu$ ’, or  $^{238}\text{U}/^{204}\text{Pb}$ ; Graham et al., 1992; Hanyu and Kaneoka, 1997), EM1 (enriched mantle 1; Honda and Woodhead, 2005), EM2 (enriched mantle 2; Workman et al., 2004; Jackson et al., 2007b) and DMM exhibit spidergrams (Hart and Gaetani, 2006) that lack such pronounced TITAN anomalies (Fig. 1). While HIMU basalts can have positive Nb and Ta anomalies (Weaver et al., 1987; Weaver et al., 1991; Chauvel et al., 1992), they generally have lower anomalies than high  $^3\text{He}/^4\text{He}$  lavas, and can even have negative Nb anomalies (Sun and McDonough, 1989). Importantly, HIMU lavas exhibit flat or negative Ti-anomalies (McDonough, 1991), and thus lack the positive Ti-anomalies observed in high  $^3\text{He}/^4\text{He}$  lavas.

Available data (see Appendix A) also indicate that the TITAN enrichment is enhanced with increasing  $^3\text{He}/^4\text{He}$  in OIB lavas (Fig. 2): Large, positive Ti (high Ti/Ti\*), Nb (elevated Nb/Nb\*) and Ta (Ta/Ta\*, not shown) anomalies are observed in the highest  $^3\text{He}/^4\text{He}$  basalts. While all high  $^3\text{He}/^4\text{He}$  lavas (>30 Ra) have large, positive TITAN anomalies, not all lavas with positive TITAN anomalies have high  $^3\text{He}/^4\text{He}$ . In the plots of TITAN anomalies vs.  $^3\text{He}/^4\text{He}$ , the OIB data outline a “wedge-shaped” pattern. For example, Cape Verde lavas have large, positive Nb/Nb\* values, but have low  $^3\text{He}/^4\text{He}$  (Fig. 2). However, high  $^3\text{He}/^4\text{He}$  lavas with negative TITAN anomalies are absent in the available dataset.

High  $^3\text{He}/^4\text{He}$  lavas exhibit moderately radiogenic Os-isotopes ( $^{187}\text{Os}/^{188}\text{Os}>0.135$ ), an observation that is contrary to previous suggestions that this reservoir hosts unradiogenic  $^{187}\text{Os}/^{188}\text{Os}$  ratios (Hauri et al., 1996) similar to DMM

(Standish et al., 2002) (Fig. 2). Radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  in the high  $^3\text{He}/^4\text{He}$  reservoir is consistent with the positive correlation observed between  $^3\text{He}/^4\text{He}$  and  $^{187}\text{Os}/^{188}\text{Os}$  in Icelandic hotspot lavas (Brandon et al., 2007). However, radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  is not unique to the high  $^3\text{He}/^4\text{He}$  mantle: the low  $^3\text{He}/^4\text{He}$  mantle endmembers EM1 and HIMU also have radiogenic  $^{187}\text{Os}/^{188}\text{Os}$ . Importantly, unradiogenic  $^{187}\text{Os}/^{188}\text{Os}$  is not observed in high  $^3\text{He}/^4\text{He}$  OIB lavas.

### **3. The case for a refractory, rutile-bearing eclogite component in the high $^3\text{He}/^4\text{He}$ mantle sampled by OIBs**

#### **3.1. TITAN enrichment and high $^{187}\text{Os}/^{188}\text{Os}$ : Evidence for refractory eclogite.**

The lack of large TITAN anomalies in DMM (Fig. 1) demonstrates that phases contained in upper mantle peridotites do not preferentially sequester the TITAN elements relative to other incompatible lithophile elements. By contrast, subduction zone lavas are TITAN-depleted, indicating that processes operating in their mantle source can fractionate TITAN from the other lithophile trace elements. It was suggested that eclogite melting in subduction zones may generate rutile-bearing residues that are residually-enriched in TITAN elements (e.g., McDonough, 1991). Experimental studies indicate that the TITAN elements are strongly partitioned into rutile during eclogite melting (Green and Pearson, 1986; Ryerson and Watson, 1987; Ayers, 1998; Stalder et al., 1998; Foley et al., 2000; Schmidt et al., 2004; Kessel et al., 2005), thereby generating positive TITAN anomalies in refractory, rutile-bearing slab residues that balance the depletion observed in subduction zone lavas.

Like TITAN-enrichment, radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  is not a geochemical signature typically associated with a peridotite reservoir. While peridotites tend to have low Re/Os and  $^{187}\text{Os}/^{188}\text{Os}$ , mafic igneous rocks generally exhibit elevated Re/Os and  $^{187}\text{Os}/^{188}\text{Os}$  ratios, (Walker et al., 1989; Reisberg et al., 1991; Hauri and Hart, 1993; Reisberg et al., 1993; Snow and Reisberg, 1995; Becker, 2000). Oceanic crust enters subduction zones with initially high Re/Os, and owing to moderate compatibility of Re in garnet, oceanic crust may retain high Re/Os (and with time, radiogenic  $^{187}\text{Os}/^{188}\text{Os}$ ) ratios during

subduction zone processing in the garnet stability field (Righter and Hauri, 1998). While it has been suggested that Os and Re are extracted from the slab in the subduction zone (e.g., Brandon et al., 1996; McInnes et al., 1999; Becker et al., 2000), we note that the Re/Os ratios of altered oceanic crust ( $^{187}\text{Re}/^{188}\text{Os}$  averages of 349 and 353 in composites of two separate drill cores; Peuker-Ehrenbrink et al., 2003) and oceanic crust gabbros ( $^{187}\text{Re}/^{188}\text{Os}$  average of 472, Hart et al., [1999]) are similar to the ratios found in metabasalts metamorphosed in paleosubduction zones (median  $^{187}\text{Re}/^{188}\text{Os} = 326$ , including eclogites, blueschists and mafic granulites; Becker, 2000). Thus, there is ample evidence supporting the contention that high Re/Os ratios can be preserved in the slab during subduction zone metamorphism. Like TITAN enrichment, the observation of moderately radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  in high  $^3\text{He}/^4\text{He}$  lavas is consistent with a refractory eclogite component in their mantle sources, and not consistent with the high  $^3\text{He}/^4\text{He}$  reservoir being primitive mantle.

### 3.2. Depletion in the $^4\text{He}$ -producing elements (U and Th) and TITAN enrichment.

While recycled oceanic crust has been suggested to be ubiquitous in the mantle sources beneath hotspots (Sobolev et al., 2007), the close association of a refractory, TITAN-enriched mafic component with the high  $^3\text{He}/^4\text{He}$  mantle may appear contradictory since eclogites are quantitatively degassed in subduction zones (Staudacher and Allegre, 1988; Moreira and Kurz, 2001; Moreira et al., 2003). Far from hosting the high  $^3\text{He}/^4\text{He}$  signature in the mantle source of high  $^3\text{He}/^4\text{He}$  OIB lavas, the eclogite will instead contribute  $^4\text{He}$  (via alpha decay of U and Th) and generate low time-integrated  $^3\text{He}/^4\text{He}$  ratios. However, the long-term  $^4\text{He}$  production of an eclogite can be greatly reduced by melt (or fluid) extraction of highly incompatible elements like U and Th from the slab during subduction processing. By contrast, the TITAN elements will be conserved if melt (or fluid) extraction occurs in the presence of rutile (e.g., McDonough, 1991). In this scenario, a positive TITAN anomaly indicates U and Th (including other incompatible trace elements, like La, Tb and Sm) depletion in a slab. The positive TITAN anomaly is formed by relative enrichment of TITAN due to conservation of Ti,

Ta and Nb and concomitant loss of U and Th. Therefore, the refractory eclogite model of McDonough (1991) is potentially consistent with the preservation of a high  $^3\text{He}/^4\text{He}$  in the mantle. A U- and Th-depleted, TITAN-enriched eclogite will not produce significant post-subduction radiogenic  $^4\text{He}$  ingrowth (see below), and may explain why TITAN enrichment is associated with high  $^3\text{He}/^4\text{He}$  signatures.

By contrast, subducted oceanic crust that conserves much of its original U and Th budget through the subduction zone will not possess positive TITAN anomalies and will produce more  $^4\text{He}$  over time than a U and Th-depleted slab. Consequently, the abundant  $^4\text{He}$  produced by an undepleted slab would infect the surrounding mantle with radiogenic  $^4\text{He}$  due to rapid diffusion of helium in the mantle (Hart, 1984; Trull and Kurz, 1993; Hart et al., 2007), and is not compatible with the preservation of a high  $^3\text{He}/^4\text{He}$  signature in the mantle. This hypothesis is consistent with the observation that high  $^3\text{He}/^4\text{He}$  lavas never have negative (or flat) TITAN anomalies.

### 3.3. Refractory eclogite and high $^3\text{He}/^4\text{He}$ peridotite: the raw materials for the high $^3\text{He}/^4\text{He}$ OIB mantle.

While refractory, rutile-bearing eclogite possesses positive TITAN anomalies, it does not have intrinsically high  $^3\text{He}/^4\text{He}$  ratios: high  $^3\text{He}/^4\text{He}$  signatures in the TITAN-enriched, high  $^3\text{He}/^4\text{He}$  OIB source must be derived from another lithology. Ancient mantle peridotites can potentially preserve elevated  $^3\text{He}/^4\text{He}$  ratios over time (Hart et al., 1992; Anderson, 1998; Parman et al., 2005; Heber et al., 2007; Jackson et al., 2007a) if they were isolated from the convecting mantle early in Earth's history. However, mantle peridotites do not generally exhibit positive TITAN anomalies (McDonough, 1991). Alone, neither eclogite nor peridotite can contribute both TITAN-enrichment and high  $^3\text{He}/^4\text{He}$  to the mantle source sampled by high  $^3\text{He}/^4\text{He}$  OIB lavas. Both refractory eclogite and the high  $^3\text{He}/^4\text{He}$  peridotite are required to generate such a mantle source. Thus, an important question is how the TITAN-enriched eclogites came to be associated (i.e., mixed) with high  $^3\text{He}/^4\text{He}$  peridotites in the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs.

### 3.4. Peridotite and eclogite portions of ancient subducted slabs: A high $^3\text{He}/^4\text{He}$ , TITAN-enriched “package”.

One possible explanation for the generation of a hybrid lithology in the high  $^3\text{He}/^4\text{He}$  OIB mantle is that the crust and mantle components of ancient, subducted oceanic lithosphere contain the raw materials required for the generation of a mantle source sampled by high  $^3\text{He}/^4\text{He}$  basalts. TITAN-enriched refractory eclogite exists at the top of the slab, and ancient high  $^3\text{He}/^4\text{He}$  asthenospheric peridotite (with the same composition as contemporary DMM) is coupled to the underside of the downgoing plate. After processing in subduction zones, the crustal portion of the subducted plate hosts TITAN enrichment, and the asthenospheric DMM coupled to the bottom of the downgoing plate remains unscathed by subduction zone processes. The two components, TITAN-enriched eclogite and high  $^3\text{He}/^4\text{He}$  asthenospheric peridotite, should be intimately associated as a “package” in space and time within a subducted plate, a geometry that is conducive to later mixing in the mantle.

We present a simple model for the generation of the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs, whereby the eclogitic and peridotitic components of the oceanic plate are subducted and isolated from the convecting mantle at 3 Ga and mixed during storage in the lower mantle (Fig 3). We assume that the upper mantle began with a primitive composition (McDonough and Sun, 1995) at 4.4 Ga, and evolved by continuous depletion to the present-day DMM composition of Workman and Hart (2005). Oceanic plates were continuously injected into the mantle over this time period, and the eclogitic and peridotitic portions of these subducted plates were thoroughly mixed in the lower mantle and were later sampled by upwelling mantle plumes and erupted at hotspots.

In order to develop a quantitative geochemical model for the  $^3\text{He}/^4\text{He}$ ,  $^{187}\text{Os}/^{188}\text{Os}$ , and TITAN of the two lithologies (peridotite and eclogite) in a subducted plate, we make the following assumptions about their compositions over time:

**i.  $^3\text{He}/^4\text{He}$  of ancient asthenospheric DMM peridotite and refractory eclogite:** The  $^3\text{He}/^4\text{He}$  of DMM is thought to have decreased significantly over time (Fig. 3). However, portions of asthenospheric DMM peridotite that are isolated with the downgoing plate

(i.e., removed from the upper mantle) are not subject to further depletion. In the model, the subducted asthenospheric DMM peridotite “locked in” the high  $^3\text{He}/^4\text{He}$  and  $^3\text{He}/^{238}\text{U}$  ratios of ancient, less-depleted DMM, and evolved by closed-system decay of U and Th. By contrast, if the subducted portion of ancient asthenospheric DMM had remained as part of the convecting asthenosphere, it would have evolved by continuous melt extraction to become modern DMM with low  $^3\text{He}/^4\text{He}$  and  $^3\text{He}/^{238}\text{U}$  (see Fig. 3 and Appendix B for details of  $^3\text{He}/^4\text{He}$  evolution in the mantle). Consequently, following subduction and isolation, the  $^3\text{He}/^4\text{He}$  of the isolated, ancient DMM rapidly diverged from (and preserved higher  $^3\text{He}/^4\text{He}$  than) its upper mantle counterpart, which continued to be depleted (in He relative to U and Th, see Appendix B) by continental and oceanic crust extraction (Fig. 3).

By comparison, we model the  $^3\text{He}/^4\text{He}$  evolution of the subducted eclogite using the trace element composition of a hypothetical refractory eclogite calculated by McDonough (1991) (see Fig. 4 for a spidergram of the refractory eclogite). The eclogite is assumed to start with a  $^3\text{He}/^4\text{He}$  ratio equal to DMM at its time of isolation, and following over 99% degassing in the subduction zone, the U and Th of the eclogite generates  $^4\text{He}$  by decay, thus rapidly diminishing the  $^3\text{He}/^4\text{He}$  of the eclogite over time (Fig. 3, Table 2). Degassing of the eclogite portion of the subducted slab is simulated by increasing the  $^{238}\text{U}/^3\text{He}$  of the eclogite by a factor of 1,000 relative to contemporary DMM.

**ii.  $^{187}\text{Os}/^{188}\text{Os}$  of ancient asthenospheric DMM peridotite and refractory eclogite:**

The  $^{187}\text{Os}/^{188}\text{Os}$  of DMM and Primitive Mantle are not very different (0.12-0.13; Standish et al., 2002; Meisel et al., 2001), and thus the isolated, ancient asthenospheric peridotite portions of the subducted plates are assumed to have an intermediate present-day composition (0.125). DMM is also assumed to have had an Os concentration of 3000 ppt over geologic time. In contrast to the asthenospheric DMM peridotite, the eclogitic portion of the subducted plate likely evolved extremely radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  over time (Fig. 3). The  $^{187}\text{Os}/^{188}\text{Os}$  evolution of the refractory eclogite is modeled using the median  $^{187}\text{Re}/^{188}\text{Os}$  (325) and Os (6 ppt) from eclogites metamorphosed in paleosubduction zones

reported by Becker (2000): earlier subduction injection of eclogites with this composition yielded higher  $^{187}\text{Os}/^{188}\text{Os}$  in the present day. The high calculated present-day  $^{187}\text{Os}/^{188}\text{Os}$  ratios in the recycled eclogites are similar to the most radiogenic eclogites presented by Becker (2000).

**iii. TITAN anomalies of ancient asthenospheric DMM peridotite and refractory eclogite:** The trace element content of the present-day asthenospheric DMM peridotite is assumed to be the same as the DMM compositions calculated by Workman and Hart (2005). However, DMM has likely become increasingly depleted throughout geologic time. We assume that DMM began with a primitive mantle composition (McDonough and Sun, 1995) and evolved by continuous depletion until the present day. Using the continuous transport equations in Appendix B, the trace element budget of DMM is calculated at various times, and “snapshots” of DMM compositions through time are plotted as spidergrams in Fig. 4 (see Table 2 for compositions). By contrast, the subducted eclogites are assumed to have the same present-day trace element composition as the hypothetical refractory eclogite from McDonough (1991), regardless of the isolation time (see Table 2). This hypothetical eclogite has a trace element composition similar to the eclogites with the largest positive TITAN anomalies presented in Becker et al. (1999) (see Table 2).

### 3.5. Mixing asthenospheric DMM peridotite with refractory eclogite.

The high  $^3\text{He}/^4\text{He}$ , moderately radiogenic  $^{187}\text{Os}/^{188}\text{Os}$ , and positive TITAN anomalies observed in the highest  $^3\text{He}/^4\text{He}$  OIB lavas can be generated by mixing refractory eclogite and the asthenospheric DMM peridotite that were subducted together (in the same plate) at 3 Ga. When the proportion of refractory eclogite in the mixture is between 20 and 25%, the model generates  $^3\text{He}/^4\text{He}$ ,  $^{187}\text{Os}/^{188}\text{Os}$  and positive TITAN anomalies that are similar to the highest  $^3\text{He}/^4\text{He}$  OIB lavas (Fig. 2). If the proportion of refractory eclogite is increased, the present-day  $^3\text{He}/^4\text{He}$  in the resulting mixture is diminished and the positive TITAN anomalies and  $^{187}\text{Os}/^{188}\text{Os}$  ratios are both increased to values above those observed in high  $^3\text{He}/^4\text{He}$  OIB lavas.

Fig. 4 shows that the addition of 20% eclogite to a 3 Ga DMM composition generates a hybrid eclogite-peridotite spidergram that is similar in shape to the spidergram of the highest  $^3\text{He}/^4\text{He}$  lava from Iceland. (Sr is a poor fit, however, due to the positive Sr anomalies in the Icelandic lava, a possible result of interaction with shallow lithospheric gabbros [Gurenko and Sobolev, 2006]). This hybrid spidergram may thus serve as a plausible melt source for the high  $^3\text{He}/^4\text{He}$  lava. It is important to note that the TITAN elements are “bracketed” by elements of similar compatibility (as long as rutile is absent) in the spidergram, and these bracketing elements are used to calculate the TITAN anomalies. Thus, the magnitude of TITAN anomalies are little affected by partial melting beneath a hotspot, and the positive TITAN anomalies in the mantle source of high  $^3\text{He}/^4\text{He}$  lavas are reflected in the erupted hotspot lavas. Clearly, this assumes that the rutile present in the downgoing slab is no longer present in the high  $^3\text{He}/^4\text{He}$  mantle that melts beneath hotspots. One way to destabilize rutile is to completely mix the (smaller proportion of) eclogite and the (larger proportion of) peridotite in the ancient recycled slab. Alternatively, if eclogite is still present in the source of the high  $^3\text{He}/^4\text{He}$  OIB mantle source, it must be melted to a sufficiently high degree to eliminate rutile as a phase in the residue of melting (Gaetani et al., 2007).

This model for the generation of high  $^3\text{He}/^4\text{He}$ , TITAN-enriched mantle can also generate a low  $^3\text{He}/^4\text{He}$  that has positive TITAN anomalies. Subduction is a continuous process that has operated for much of geologic time, and the  $^3\text{He}/^4\text{He}$  ratio of the upper mantle has likely decreased significantly (Fig. 3). Thus, the peridotite portion of more recently subducted oceanic plates will trap and preserve a lower  $^3\text{He}/^4\text{He}$  upper mantle signature than ancient subducted plates, and the refractory eclogite portion of recently subducted plates will also host positive TITAN anomalies. For example, the lower  $^3\text{He}/^4\text{He}$ , TITAN-enriched mantle source sampled by Cape Verde lavas can be generated by mixing refractory eclogite (10-30% by mass) and asthenospheric DMM components of a plate subducted between 1-2 Ga (Fig. 2).

The “wedge-shaped” outline of the OIB data in Fig. 2 highlights a striking absence of high  $^3\text{He}/^4\text{He}$  lavas with negative (or even small positive) TITAN anomalies. Why do all high  $^3\text{He}/^4\text{He}$  lavas exhibit TITAN-enrichment, or rather, why have high

$^3\text{He}/^4\text{He}$ -TITAN depleted (or only mildly TITAN enriched) lavas not been found? Continental crust and arc lavas (and associated sediments) compose the only known reservoir to exhibit TITAN depletion. Due to its extremely high U and Th contents, admixture of CC with high  $^3\text{He}/^4\text{He}$  peridotite may not be conducive to the preservation of high  $^3\text{He}/^4\text{He}$  (Jackson et al., 2007b), and could explain why TITAN-depleted lavas always exhibit low  $^3\text{He}/^4\text{He}$  (Fig. 2). OIB lavas lacking TITAN anomalies may host an eclogite component that was not U and Th-depleted (and thus did not acquire positive TITAN anomalies) in a subduction zone, and may also produce significant  $^4\text{He}$ .

On the other hand, the absence of high  $^3\text{He}/^4\text{He}$  lavas with flat (or even slightly positive) TITAN anomalies may be explained by the intimate spatial and temporal association between the TITAN-enriched eclogites and high  $^3\text{He}/^4\text{He}$  peridotites suggested by this model: The peridotite and refractory eclogite components—the raw materials for the formation of the high  $^3\text{He}/^4\text{He}$ , TITAN-enriched mantle—are always together in subducting plates, and it may not be possible to melt pure high  $^3\text{He}/^4\text{He}$  peridotite without also melting some eclogite. This will be true particularly if subducted slabs are stretched, thinned and folded in the dynamic mantle (Allegre and Turcotte, 1986), such that the diminished thickness of the (eclogite and peridotite) slab is less than the width of melting zones beneath hotspots. Such a process might explain why an eclogite signature (positive TITAN anomalies and radiogenic  $^{187}\text{Os}/^{188}\text{Os}$ ) is invariably present in high  $^3\text{He}/^4\text{He}$  lavas.

### 3.6. Two alternative models for a hybrid high $^3\text{He}/^4\text{He}$ mantle: “Eclogite Injection” and diffusion of “Ghost” primordial helium.

An alternative process for generating the hybrid eclogite-peridotite high  $^3\text{He}/^4\text{He}$  mantle assumes that subducted slab peridotite contributes little to the helium budget of the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs. In this model, the refractory slab eclogite penetrates into the lower mantle and mixes with a hypothetical lower mantle high  $^3\text{He}/^4\text{He}$  peridotite reservoir (i.e., not associated with the peridotite portion of slabs). The hybrid mixture then rises in a plume where it is melted beneath a hotspot. By comparison to the slab peridotite-eclogite “package” model describe in section 3.4 and 3.5 above, this

alternative “eclogite injection” model does not guarantee an intimate spatial and temporal association of the high  $^3\text{He}/^4\text{He}$  peridotite and TITAN-enriched eclogite components. For example, deep mantle high  $^3\text{He}/^4\text{He}$  peridotite could upwell in a plume without first being inoculated with refractory eclogite, in which case the erupted high  $^3\text{He}/^4\text{He}$  lavas would lack positive TITAN anomalies. Such lavas have not been observed. While this model is not explored further here, one possible solution could be that, due to mixing by stretching and thinning of slabs in a convecting mantle over geologic time (Allegre and Turcotte, 1986), recycled eclogite has become pervasive in the mantle and is distributed at lengthscales smaller than the melting zones beneath hotspot volcanoes. In this way, volcanoes fed by upwelling high  $^3\text{He}/^4\text{He}$  mantle plumes will inevitably sample subducted eclogite.

Instead of mechanically mixing the high  $^3\text{He}/^4\text{He}$  peridotite and TITAN-enriched eclogite, as suggested in the “slab package” and “eclogite injection” models above, it may be possible to diffusively mix helium from a high  $^3\text{He}/^4\text{He}$  peridotite into a degassed, U and Th-poor pyroxenite (i.e., refractory eclogite) (Albarede and Kaneoka, 2007). Due to the higher diffusivity of helium compared to non-volatile major and trace elements, helium isotopes may become decoupled from other lithophile isotope tracers (Hart et al., 2007), and primordial helium may become associated with recycled materials like refractory eclogites (Albarede and Kaneoka, 2007). Albarede and Kaneoka (2007) propose that helium from deep (high  $^3\text{He}/^4\text{He}$ ) mantle peridotites can diffuse into embedded, tightly folded layers of stretched and thinned refractory eclogite. They suggest that changing the duration of the diffusion process, as well as the U and Th contents of the refractory eclogite layers, can generate mantle sources for both high and low  $^3\text{He}/^4\text{He}$  hotspots. U and Th-poor refractory eclogite that was processed in subduction zones will have positive TITAN anomalies (McDonough, 1991), and because such eclogites will produce little  $^4\text{He}$  over time, they are perfect “containers” for preserving diffusively acquired high  $^3\text{He}/^4\text{He}$  signatures. If these eclogites become sufficiently thinned (to <1-2 km thickness, by mantle mixing), they could acquire high  $^3\text{He}/^4\text{He}$  signatures by diffusion from the ambient deep mantle peridotite (Hart et al., 2007). Thus, the “ghost” helium model of Albarede and Kaneoka (2007) may offer a

resolution to the paradoxical association of high  $^3\text{He}/^4\text{He}$  signatures in lavas with strong eclogite signatures.

Nonetheless, the “ghost helium” model suffers from the same spatial and temporal issues as the “eclogite injection” model: There is no obvious mechanism preventing ambient lower mantle (eclogite-free) peridotite from upwelling and melting beneath a hotspot, thus generating a high  $^3\text{He}/^4\text{He}$  lava that lacks positive TITAN anomalies. However, one possible resolution is that plume-entrained, high  $^3\text{He}/^4\text{He}$  lower mantle peridotites are too refractory to contribute significantly to melting beneath OIBs (i.e., Albarede and Kaneoka [2007] invoke dunites and harzburgites), so that pure peridotites are never melted. Alternatively, if eclogite layers are pervasive in the mantle at lengthscales smaller than those sampled by melting zones (<100 km), it may be inescapable that eclogites always contribute to mantle melts. An additional problem with the “ghost helium” model is that rutile may still be stable in a pure eclogite mantle component that is upwelling beneath a hotspot; the residual rutile will hold back TITAN in the source, and as a result, positive TITAN anomalies will not be observed in the erupted lavas. However, if the eclogite is melted to sufficiently high degrees beneath a hotspot, the rutile will be completely consumed and not present in the residue (Gaetani et al., 2007).

### 3.7. High $^3\text{He}/^4\text{He}$ lavas without positive TITAN anomalies?

If a high  $^3\text{He}/^4\text{He}$  peridotite could be melted in pure (no eclogite) form, there would be no positive TITAN-anomalies in the erupted lavas. There would also be no contribution of  $^4\text{He}$  ingrowth from the eclogite, and an even higher  $^3\text{He}/^4\text{He}$  might be expected in the melts of such a mantle source. Such lavas have not been identified. However, with the highest magmatic  $^3\text{He}/^4\text{He}$  values on record, Baffin Island lavas (Stuart et al., 2003) may provide an important test case for this hypothesis.

### 3.8. TITAN anomalies due to partitioning between lower mantle phases?

It is difficult to rule out the possibility of TITAN fractionation in lower mantle materials. Experimental studies of high pressure partitioning and mineralogy are in the

early phases. However, Ca-perovskite in peridotitic and basaltic systems shows negative Ti and Nb partitioning patterns compared to Th, U and the rare earth elements (REEs) (Hirose et al., 2004). This means that a Ca-perovskite bearing solid assemblage would have negative anomalies, but melt equilibrated with Ca-perovskite could have positive anomalies. If it is possible to generate Ca-perovskite melts at the appropriate pressures and temperatures (for example, D''), and extract them from the lower mantle (the inferred home of the high  $^3\text{He}/^4\text{He}$  domain), then Ca-perovskite melting in the lower mantle may offer a potential explanation for the origin of TITAN-enriched, high  $^3\text{He}/^4\text{He}$  lavas.

#### **4. The high $^3\text{He}/^4\text{He}$ , TITAN-enriched mantle: A reservoir for the “missing” TITAN elements in the earth?**

In addition to offering insights into the composition and generation of the high  $^3\text{He}/^4\text{He}$  mantle reservoir, TITAN-enrichment in high  $^3\text{He}/^4\text{He}$  OIBs may provide information about the location and composition of the reservoir hosting the earth's missing TITAN elements. If the refractory lithophile trace elements have chondritic abundances in BSE (bulk silicate earth), mass balance constraints require that the TITAN elements missing in the shallow earth reservoirs—DMM and CC—must exist in the deeper earth (McDonough, 1991; Rudnick et al., 2000), perhaps in a deep reservoir composed of subducted oceanic plates (Christensen and Hofmann, 1994). In this section, we calculate the trace element budget of a hypothetical missing TITAN-rich reservoir that, when added together with DMM and CC, generates a BSE (McDonough and Sun, 1995) spidergram. We show that, like high  $^3\text{He}/^4\text{He}$  OIB lavas, the spidergram of the TITAN-rich reservoir has positive TITAN anomalies.

In order to estimate the trace element composition of this deep, TITAN-enriched reservoir, we assume that the composition of BSE can be approximated with just three reservoirs: CC, DMM and subducted plates (here called PLATE, which is composed of oceanic crust, mantle lithosphere, and everything else—like sediment—subducted along with downgoing plates). The trace element budget of the PLATE reservoir is calculated using the following mass balance relationship:

$$[x]_{CC} \times M_{CC} + [x]_{DMM} \times M_{DMM} + [x]_{PLATE} \times M_{PLATE} = [x]_{BSE} \times M_{BSE} \quad (\text{eq. 1})$$

where  $[x]$  represents the concentration of the element  $x$  in the four reservoirs, and  $M$  is the mass of the reservoirs. Assuming the mass of CC is fixed at 0.6% of BSE, and given that the mass proportions of the DMM and PLATE reservoirs are unknown, the following relationships are used to calculate possible trace element budgets for the PLATE reservoir:

$$M_{CC} = 0.006 \times M_{BSE} \quad (\text{eq. 2})$$

$$M_{PLATE} = \beta \quad (\text{eq. 3})$$

$$M_{DMM} = 0.994 - \beta \quad (\text{eq. 4})$$

where the mass of the three reservoirs, PLATE, CC, and DMM sum to the mass of the total silicate mantle. Employing the trace element budgets previously derived for CC (Rudnick and Gao, 2003), DMM (Workman and Hart, 2005) and BSE (McDonough and Sun, 1995), the trace element budget of the subducted PLATE reservoir is calculated for different values of  $M_{PLATE}$  ( $\beta$ ). The results of the mass balance model are plotted in Fig. 5 (assuming that  $\beta \geq 0.4$ ). The calculated PLATE spidergram changes as a function of its mass proportion ( $\beta$ ) of BSE. For large values of  $\beta$ , the PLATE reservoir is more trace element depleted, and for smaller values of  $\beta$ , the PLATE reservoir is increasingly trace element enriched. The most important observation is that, regardless of the value of  $\beta$ , the PLATE reservoir exhibits positive TITAN anomalies.

It is possible to place constraints on the minimum size of the subducted PLATE reservoir and show that its mass proportion of BSE is unlikely to be small. Following Rudnick et al. (2000), we assume that the present mass of the oceanic crust ( $5.3 \times 10^{21}$  kg) has been subducted every 100 Ma for the past 2.5 Ga, and the total mass of the reservoir comprised of subducted oceanic crust is  $1.3 \times 10^{23}$  kg. Assuming that the portion of the subducted oceanic plate that is oceanic mantle lithosphere is 10 times thicker and 10% denser than the oceanic crustal lithosphere, then the total mass of oceanic plates

subducted over the past 2.5 Ga is  $\sim 1.4 \times 10^{24}$  kg ( $\sim 8\%$  of which is oceanic crust), which is  $\sim 40\%$  of the mass of BSE. We consider this a minimum estimate for the size of the subducted PLATE reservoir, as subduction probably operated before 2.5 Ga, and the Archaen mantle was hotter, thereby leading to higher rates of plate formation and subduction (Rudnick et al., 2000). Thus, the PLATE reservoir composes  $\geq 40\%$  of BSE, and all possible spidergrams for the plate reservoir are more depleted than the PLATE spidergram shown in Fig. 5.

Therefore, mass balance considerations (equations 1-4) and estimates of plate recycling budgets suggest that a large mantle reservoir host for the missing TITAN elements. The spidergram of the calculated PLATE reservoir exhibits hints of the positive TITAN-anomalies observed in high  $^3\text{He}/^4\text{He}$  OIBs. Thus, it is not implausible that high  $^3\text{He}/^4\text{He}$  lavas are sampling the “missing” TITAN hosted in the PLATE reservoir.

While there are many similarities, there is some disagreement between the shape of the spidergrams of the calculated PLATE reservoir and the high  $^3\text{He}/^4\text{He}$  OIB lavas. Most significantly, the positive TITAN anomalies in the PLATE reservoir are not as large as those observed in the high  $^3\text{He}/^4\text{He}$  lavas from Hawaii, Iceland, Galapagos and Samoa. One solution to this discrepancy is the following: The PLATE reservoir is likely to be heterogeneous, as it includes everything in BSE that is not DMM or CC. These other components in the PLATE reservoir will dilute the large, positive TITAN-anomalies contributed from recycled, rutile-bearing eclogites. While the PLATE reservoir is not purely a high  $^3\text{He}/^4\text{He}$  reservoir, its spidergram balances the Earth’s budget of TITAN, and recycled, refractory, rutile-bearing eclogite likely contributes the positive TITAN anomalies to the PLATE reservoir (McDonough, 1991). It is this TITAN-enriched domain of the PLATE reservoir that may be associated with the source for high  $^3\text{He}/^4\text{He}$  lavas.

In section 3.5 it was shown that, if mixed with high  $^3\text{He}/^4\text{He}$  peridotite, approximately 20 to 25% refractory, rutile-bearing eclogite can generate the  $^3\text{He}/^4\text{He}$ ,  $^{187}\text{Os}/^{188}\text{Os}$  and TITAN anomalies in high  $^3\text{He}/^4\text{He}$  OIBs. However, this percentage of eclogite greatly exceeds the percentage of eclogite in the PLATE reservoir ( $\sim 8\%$ ).

Perhaps this inconsistency can be explained by the fact that eclogites have lower solidus temperatures than pure, unadulterated peridotites. Thus, when melting the high  $^3\text{He}/^4\text{He}$  portion of the PLATE reservoir beneath a hotspot, the eclogite's (be it pure eclogite or eclogite completely mixed into a peridotite) contribution to the melt will exceed its mass proportion in the mantle source.

## **5. Implications for $^{142}\text{Nd}/^{144}\text{Nd}$ measurements on terrestrial mantle rocks**

A recent discovery demonstrated that the  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios in all measured terrestrial mantle rocks are 20 ppm higher than chondrite (Boyet and Carlson, 2005, 2006), indicating that these rocks were derived from a reservoir that had superchondritic Sm/Nd during the lifetime of  $^{146}\text{Sm}$  (the first few hundred million years following accretion). If the BSE has chondritic abundances of the refractory elements, then the superchondritic  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios observed in the accessible terrestrial rocks suggest that they sample the depleted residue (early depleted reservoir, or EDR) of an early differentiation event. In this model, a complementary hidden “early enriched reservoir” (EER) with subchondritic  $^{142}\text{Nd}/^{144}\text{Nd}$  must exist in the deep earth (Boyet and Carlson, 2005, 2006; Andreasen and Sharma, 2006; Carlson et al., 2007). Alternatively, if BSE accreted from non-chondritic materials, an early differentiation of the silicate earth is not required (Caro et al., 2007).

If the earth accreted from non-chondritic materials, the superchondritic  $^{142}\text{Nd}/^{144}\text{Nd}$  ratios in the terrestrial mantle could be a result of superchondritic  $^{147}\text{Sm}/^{144}\text{Nd}$  ( $\geq 0.209$ ) ratios in BSE (Boyet and Carlson, 2006). However, a non-chondritic earth does not necessarily obviate the need for a deep, TITAN-enriched SLAB reservoir, because we see evidence for its presence in high  $^3\text{He}/^4\text{He}$  OIBs.

Alternatively, if BSE does have chondritic abundances of the refractory elements, neither of the two early formed reservoirs—EDR and EER—have the appropriate isotopic and trace element characteristics to be the mantle source of high  $^3\text{He}/^4\text{He}$  lavas. The EDR was suggested to be the high  $^3\text{He}/^4\text{He}$  reservoir variously called FOZO, PHEM

or C (Boyet and Carlson, 2006). While the EDR does have higher  $^{143}\text{Nd}/^{144}\text{Nd}$  than chondrite, consistent with globally high (superchondritic)  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios in high  $^3\text{He}/^4\text{He}$  lavas (Jackson et al., 2007a), the EDR is calculated to have negative TITAN anomalies (Boyet and Carlson, 2005) and is not consistent with being the mantle source of TITAN-enriched, high  $^3\text{He}/^4\text{He}$  OIB lavas. By contrast, the EER does have positive TITAN anomalies like those observed in high  $^3\text{He}/^4\text{He}$  OIB lavas. An early crust isolated at the bottom of the mantle has been suggested to host high  $^3\text{He}/^4\text{He}$  ratios (Tolstikhin and Hofmann, 2005). However, the EER cannot be the mantle source of the high  $^3\text{He}/^4\text{He}$  as it has Sm/Nd ratios lower (more enriched) than chondrite, and will generate lower  $^{143}\text{Nd}/^{144}\text{Nd}$  ratios than observed in high  $^3\text{He}/^4\text{He}$  OIBs. Thus, neither of the two initial reservoirs suggested by Boyet and Carlson (2005) describe both the isotopic and trace element characteristic of the high  $^3\text{He}/^4\text{He}$  OIB mantle. The depleted peridotite and refractory, rutile-bearing eclogite in ancient plates that were subducted and stored in the mantle are ideally suited to be a mantle source for high  $^3\text{He}/^4\text{He}$  OIB lavas.

## **6. The fate of slabs and the high $^3\text{He}/^4\text{He}$ reservoir**

The presence of recycled eclogites in the mantle source of high  $^3\text{He}/^4\text{He}$  lavas has important implications for the helium isotope evolution of the mantle, since subduction zone processing likely plays an important role in determining the composition of recycled eclogites. Specifically, the uniqueness of the thermal regimes of different subduction zones may affect the composition of eclogites in dramatically different ways. Slab melting and dehydration in ancient, hot subduction zones may residually enrich the slab in TITAN elements while depleting it in U and Th, a process that is conducive to the formation and preservation of a high  $^3\text{He}/^4\text{He}$ , TITAN-enriched mantle reservoir. Alternatively, cooler subduction zones may not generate TITAN anomalies (or deplete the  $^4\text{He}$ -producing elements, U and Th) in the slab, a scenario that may produce mantle reservoirs with small positive TITAN anomalies and low  $^3\text{He}/^4\text{He}$ . Thus, the fate of the  $^3\text{He}/^4\text{He}$  evolution of the various mantle reservoirs may hinge on the processes operating in subduction zones.

## **Appendix A: $^3\text{He}/^4\text{He}$ , $^{187}\text{Os}/^{188}\text{Os}$ and trace element data compilation.**

$^3\text{He}/^4\text{He}$ ,  $^{187}\text{Os}/^{188}\text{Os}$  and trace element data for representative OIB samples (plotted in fig. 2) are from the GEOROC database (<http://georoc.mpch-mainz.gwdg.de/>), from the helium database of Abedini et al. (2006), and from the literature. Some of the  $^3\text{He}/^4\text{He}$  and  $^{187}\text{Os}/^{188}\text{Os}$  data from Iceland are unpublished, as are  $^{187}\text{Os}/^{188}\text{Os}$  data for Samoan samples with  $^3\text{He}/^4\text{He} > 20 \text{ Ra}$  (the protocol's for measuring  $^{187}\text{Os}/^{188}\text{Os}$  are the same as Skovgaard et al. [2001] for the unpublished Icelandic data and Workman et al. [2004] for the unpublished Samoan data). The  $^3\text{He}/^4\text{He}$  data from OIBs (plotted in Fig. 2) were obtained by both crushing and fusion of olivine, clinopyroxene and glass, and are not filtered based on helium concentrations. However, samples suggested to have suffered shallow contamination by crust (e.g., several samples in Macpherson et al., 2005) were not included. Additionally, very evolved rocks ( $\text{MgO} < 5.3 \text{ wt.}\%$ ) were excluded, so as to preclude the effects of extensive fractional crystallization on the various trace element ratios (e.g.,  $\text{Ti}/\text{Ti}^*$  and  $\text{Nb}/\text{Nb}^*$ ). Using  $\text{Ba}/\text{Rb}$  as a filter for alteration, samples with high ratios ( $\text{Ba}/\text{Rb} > 25$ ) were not considered. Only trace element (Th, La, Sm, Tb) data measured by ICP-MS and neutron activation are included, thereby eliminating samples with low-precision trace element measurements.

## **Appendix B: Helium isotope model.**

The precise timing of the genesis of the high  $^3\text{He}/^4\text{He}$  reservoirs cannot be calculated using helium isotopes because the degassing history of the DMM reservoir, the initial  $^4\text{He}/^3\text{He}$  and  $^3\text{He}$  abundance of a (hypothetical) undegassed mantle (and DMM), and the present-day  $^3\text{He}$  abundance in DMM are not well known. Thus, the age of the reservoir (3 Ga), as sampled by the high  $^3\text{He}/^4\text{He}$  lava from Iceland (sample SEL 97, see Table 1), is poorly constrained. For example, in the model presented in Section 3.5 above, a present-day undegassed mantle  $^{238}\text{U}/^3\text{He}$  ratio of 70 was assumed (Table 2). However, if the initial  $^{238}\text{U}/^3\text{He}$  ratio of the mantle is increased to a value of  $\sim 250$ , then our model can generate the TITAN anomalies,  $^{187}\text{Os}/^{188}\text{Os}$  and  $^3\text{He}/^4\text{He}$  of the mantle sampled by the Icelandic lava by subducting and isolating a plate at 4 Ga (in this case, the amount of eclogite required to generate the high  $^3\text{He}/^4\text{He}$ ,  $^{187}\text{Os}/^{188}\text{Os}$  and positive

TITAN anomalies in this older, high  $^3\text{He}/^4\text{He}$  mantle reservoir increases slightly). In order to generate a TITAN-enriched, high  $^3\text{He}/^4\text{He}$  mantle at 2 Ga, unrealistically low  $^{238}\text{U}/^3\text{He}$  ratios ( $^{238}\text{U}/^3\text{He} < 10$ ) are required.

Our preferred model for the time evolution of  $^3\text{He}/^4\text{He}$  in DMM and primitive mantle are as follows. If an initial  $^{238}\text{U}/^3\text{He}$  and  $^3\text{He}/^4\text{He}$  of the undegassed mantle are assumed (see Table 2 for assumed values), and if DMM was formed by continuous depletion by extraction of oceanic and continental crust from a chondritic primitive mantle over Earth's history, the known  $^3\text{He}/^4\text{He}$  of present day DMM (8 Ra) can be used to calculate the present-day  $^{238}\text{U}/^3\text{He}$  of DMM. Thus, assuming that the continuous transport equations (Allegre, 1969; Hart and Brooks, 1970; Workman and Hart, 2005) accurately model the continuously depleting upper mantle (DMM), the  $^{238}\text{U}/^3\text{He}$  and  $^3\text{He}/^4\text{He}$  of DMM can then be calculated at any time from 4.4 Ga to the present day (see Table 2 for list of these values for DMM at different times in earth's history). The helium isotope evolution of the continuously depleting DMM reservoir (shown in Fig. 3) is modeled using the following continuous transport equations:

$$\begin{aligned}
^4\text{He}/^3\text{He}_t = & ^4\text{He}/^3\text{He}_T + \\
& 8\lambda_{238}/(\lambda_{238}+k_{238})(^{238}\text{U}/^3\text{He})_T(1-\exp(-1(T-t)(\lambda_{238}+k_{238}))) + \\
& 7\lambda_{235}/(\lambda_{235}+k_{235})(^{235}\text{U}/^3\text{He})_T(1-\exp(-1(T-t)(\lambda_{235}+k_{235}))) + \\
& 6\lambda_{232}/(\lambda_{232}+k_{232})(^{232}\text{Th}/^3\text{He})_T(1-\exp(-1(T-t)(\lambda_{232}+k_{232}))) \quad (\text{eq. 5})
\end{aligned}$$

where

$$k_{238} = -1 \ln((^{238}\text{U}/^3\text{He})_0 / (^{238}\text{U}/^3\text{He})_T) / (T-t) - \lambda_{238}; \quad (\text{eq. 6})$$

$$k_{235} = -1 \ln((^{235}\text{U}/^3\text{He})_0 / (^{235}\text{U}/^3\text{He})_T) / (T-t) - \lambda_{235}; \quad (\text{eq. 7})$$

$$k_{232} = -1 \ln((^{232}\text{Th}/^3\text{He})_0 / (^{232}\text{Th}/^3\text{He})_T) / (T-t) - \lambda_{232} \quad (\text{eq. 8})$$

where  $\lambda_{238}$  ( $1.55 \times 10^{-10} \text{ y}^{-1}$ ),  $\lambda_{235}$  ( $9.85 \times 10^{-10} \text{ y}^{-1}$ ) and  $\lambda_{232}$  ( $4.95 \times 10^{-11} \text{ y}^{-1}$ ) are the decay constants for  $^{238}\text{U}$ ,  $^{235}\text{U}$ , and  $^{232}\text{Th}$ , respectively, and  $k_{238}$ ,  $k_{235}$  and  $k_{232}$  are the continuous transport coefficients for the U-Th-He system in DMM. The k-value is the difference in transport coefficients for U (or Th) and He, and is related to the difference in bulk

partition coefficients between U (or Th) and He; negative k-values in the model indicate that He is transported from the mantle more efficiently than U and Th.

When  $k_{238,235,232}$  equals zero, equations 5-8 describe closed-system  $^3\text{He}/^4\text{He}$  evolution. It is assumed that the undegassed mantle has been closed to degassing since 4.4 Ga. Thus, the closed-system model starts at time  $T=4.4$  Ga, where “t” is the time before present day (and  $T-t$  equals elapsed time). The  $^4\text{He}/^3\text{He}_T$  (initial ratio) of primitive mantle is unconstrained, but is assumed to be 5,995 (or 120 Ra) and is assumed to increase to 7,224 (or  $\sim 100$  Ra) today. The  $^{232}\text{Th}/^{238}\text{U}_0$  (present-day) ratio of primitive mantle is assumed to be 4.05 (McDonough and Sun, 1995). The primitive mantle  $^{238}\text{U}/^3\text{He}_0$  (present-day) ratio is then 70. From the  $^{238}\text{U}/^3\text{He}_0$  ratio, the  $^3\text{He}$  can be calculated by employing a present-day primitive mantle U concentration of 0.0203 ppm (McDonough and Sun, 1995), and is  $7.3 \times 10^{11}$  atoms/g.

The model for the  $^3\text{He}/^4\text{He}$  and  $^{238}\text{U}/^3\text{He}$  of DMM starts at time  $T=4.4$  Ga. It is assumed that DMM and primitive mantle had the same composition ( $^4\text{He}/^3\text{He}_T$ ,  $^{238}\text{U}/^3\text{He}_T$  and  $^{232}\text{Th}/^{238}\text{U}_T$ ) at 4.4 Ga, and that DMM began forming immediately by melt extraction from primitive mantle starting at 4.4 Ga. It is further assumed that DMM has evolved to exhibit present-day  $^{232}\text{Th}/^{238}\text{U}_0$  and  $^4\text{He}/^3\text{He}_0$  values of 2.55 (similar to the value for average DMM in Workman and Hart [2005]) and  $\sim 89,900$  (8 Ra), respectively. Equations 5 through 8 are then solved for  $^{238}\text{U}/^3\text{He}_0$ , which is calculated to be  $\sim 54,000$ . For this solution,  $k_{238}$  and  $k_{235}$  are both  $-1.51 \times 10^{-9} \text{ y}^{-1}$  and  $k_{232}$  is  $-1.41 \times 10^{-9} \text{ y}^{-1}$ . If a DMM U concentration of 0.0032 ppm is assumed (Workman and Hart, 2005), the  $^3\text{He}$  of present-day DMM is calculated to be  $1.5 \times 10^8$  atoms/g.

The  $^3\text{He}$  abundance in DMM calculated with the continuous depletion model is within a factor of 3 to 20 of the  $^3\text{He}$  abundances inferred from MORB samples and  $^3\text{He}$  flux from ridges (see Ballentine et al. [2002] for summary). Assuming 10% melting of the mantle source,  $^3\text{He}$  concentrations for DMM were derived from  $\text{CO}_2$  concentrations and canonical mantle  $\text{CO}_2/^3\text{He}$  ratios in MORB melt inclusions ( $4.52 \times 10^8 \pm 1.93$  atoms  $^3\text{He}/\text{g}$  (Saal et al., 2002), “popping” rock ( $>2.69 \times 10^9$  atoms  $^3\text{He}/\text{g}$  [Moreira et al., 1998]), and flux of  $^3\text{He}$  out of mid-ocean ridges ( $1.18 \times 10^9$  atoms  $^3\text{He}/\text{g}$  [Farley et al., 1995; Ballentine et al., 2002]).

The increase in  $^{238}\text{U}/^3\text{He}$  in DMM over the age of the earth (Table 2) requires that He is less compatible than U and Th during melting of this mantle reservoir. However, the assumption of increased  $^{238}\text{U}/^3\text{He}$  in DMM over time is realistic given the lherzolitic lithology of DMM, a mantle reservoir that hosts an estimated cpx modal abundance of ~13% (Workman et al., 2005). Results from a recent helium partitioning study (Heber et al., 2007) are consistent with helium being less compatible than U and Th (assuming U and Th partition coefficients from a recent compilation [Kelemen et al., 2003]) during mantle melting of a lherzolite lithology. However, helium partitioning during mantle melting is a controversial subject. Parman et al. (2005) reported olivine-melt partition coefficients for helium suggesting that helium may be more compatible than U and Th during melting of a cpx-poor lherzolite or harzburgite. However, Heber et al. (2007) report values that are over an order of magnitude smaller (less compatible), suggesting the helium may be more compatible than U and Th only when melting cpx-poor harzburgites or dunites. The discrepancy in olivine-melt helium partition coefficients between these two studies is not yet resolved.

Finally, the continuous transport equations can be written to calculate the concentrations of any element in DMM at any time in earth's history, assuming that DMM formed by continuous depletion of BSE starting at 4.4 Ga:

$$X_{\text{DMM},t} = X_{\text{BSE},0}(\exp(-1(\alpha_X)[T-t])) \quad (\text{eq. 9})$$

where

$$\alpha_X = -1 \ln(X_{\text{DMM},0}/X_{\text{BSE},0})/(T) \quad (\text{eq. 10})$$

where  $\alpha_X$  is proportional to the transport of element X out of DMM over time;  $X_{\text{BSE},0}$  and  $X_{\text{DMM},0}$  are the present-day concentrations of element X in BSE and DMM, respectively; "t" is time before the present day and  $T = 4.4$  Ga, and  $X_{\text{DMM},t}$  is the concentration of element X in DMM at any time "t" before the present day. Using equations 9 and 10, concentrations of trace elements that are known in BSE and present-day DMM can be used to calculate their time-dependent concentrations in DMM. For example, the present-day Ti concentrations in DMM (716 ppm) and BSE (1205 ppm) yield an  $\alpha_{\text{Ti}}$  value

of  $1.18 \times 10^{-10} \text{ y}^{-1}$  in equation 10. Thus, using this  $\alpha_{\text{Ti}}$  value and solving for  $\text{Ti}_{\text{DMM},t}$  in equation 10, the concentration of Ti in DMM can be calculated at any time in earth's history. More incompatible elements have larger values for  $\alpha$ . For example,  $\alpha_{\text{Th}} = 5.25 \times 10^{-10} \text{ y}^{-1}$  and  $\alpha_{\text{U}} = 4.20 \times 10^{-10} \text{ y}^{-1}$ , where Th is more incompatible than U. In Table 2, the abundances of several trace elements in DMM are provided at various times in earth's history.

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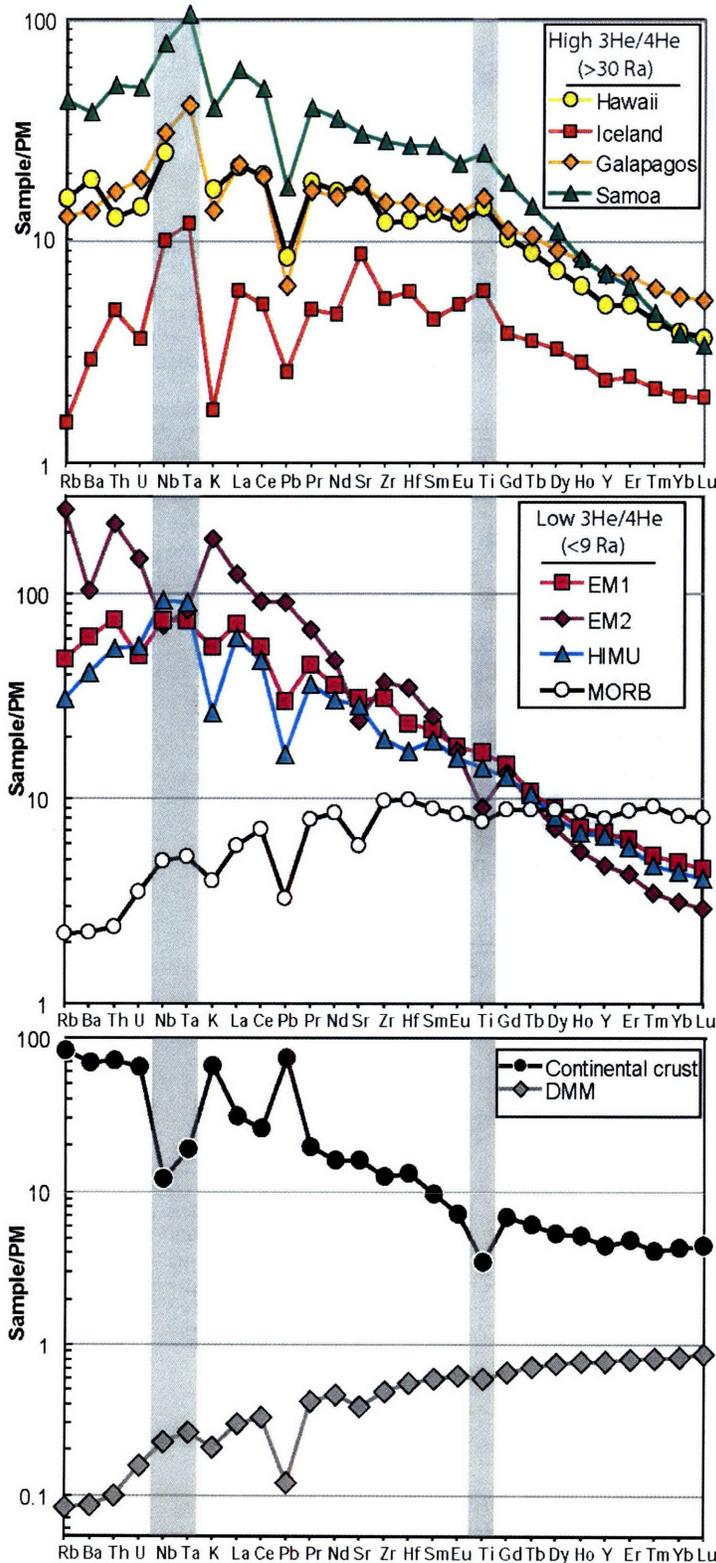
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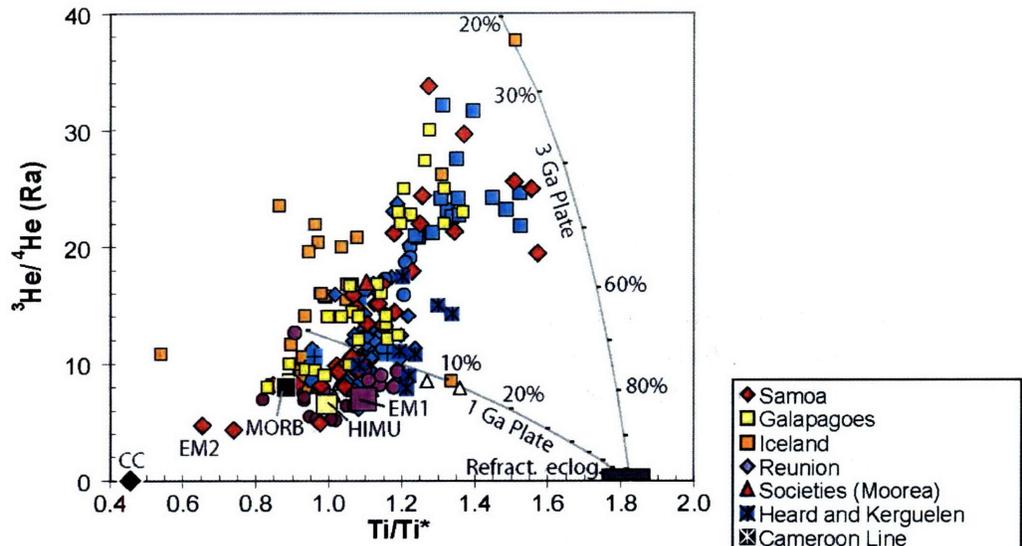
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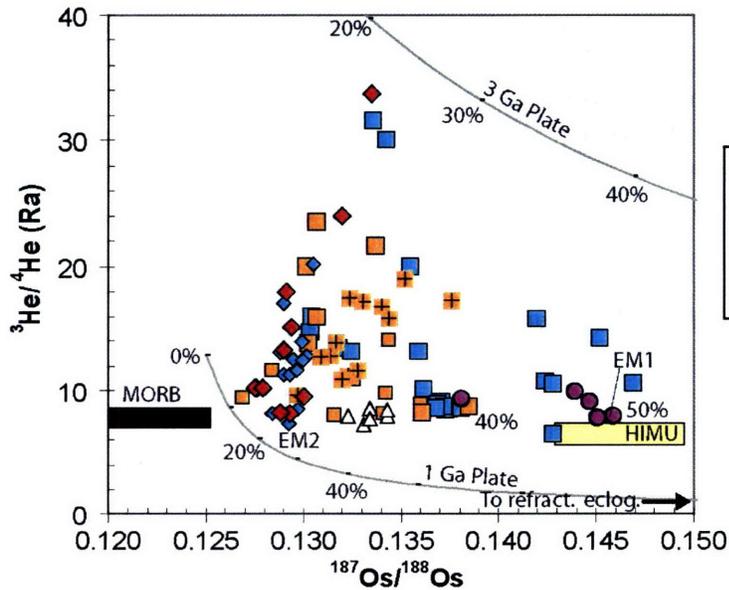
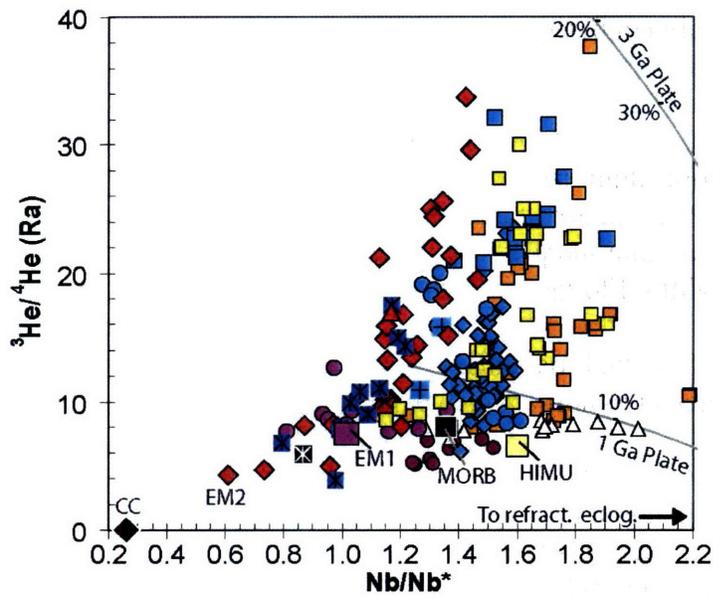
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**Figure 1.** Titanium, tantalum and niobium (TITAN) are highly enriched in ocean island basalts (OIBs) with high  $^3\text{He}/^4\text{He}$ . Top panel: Primitive mantle (PM) normalized trace element data (spidergrams) for basalts with the highest  $^3\text{He}/^4\text{He}$  (>30 Ra) from 4 different hotspots. All high  $^3\text{He}/^4\text{He}$  samples exhibit TITAN enrichment relative to neighboring elements on the spidergrams. Middle panel: Spidergrams for the low  $^3\text{He}/^4\text{He}$  (<9 Ra) mantle endmembers. EM1 lavas are from Pitcairn, EM2 lavas are from Samoa, HIMU lavas are from Mangaia and Tubuaii. Spidergrams for the mantle endmembers and MORB are from Hart and Gaetani (2006) and (for EM2) Jackson et al. (2007b). TITAN enrichment is not as pronounced, or is absent, in the average spidergrams of the low  $^3\text{He}/^4\text{He}$  mantle endmembers. Bottom panel: Spidergrams for DMM (Workman and Hart, 2005) and continental crust (Rudnick and Gao, 2003) indicate that these two shallow earth reservoirs are not entirely geochemically complementary: TITAN depletion in the continents is not balanced by corresponding enrichment in DMM (see Fig. 5 for mass balance details). However, the TITAN-enrichment in high  $^3\text{He}/^4\text{He}$  lavas suggests that these lavas sample a (deeper) mantle reservoir that hosts the TITAN that is “missing” in the shallow geochemical reservoirs, DMM and continental crust. Sources of the  $^3\text{He}/^4\text{He}$  data for the mantle endmembers and the high  $^3\text{He}/^4\text{He}$  lavas can be found in the text and in Table 1.

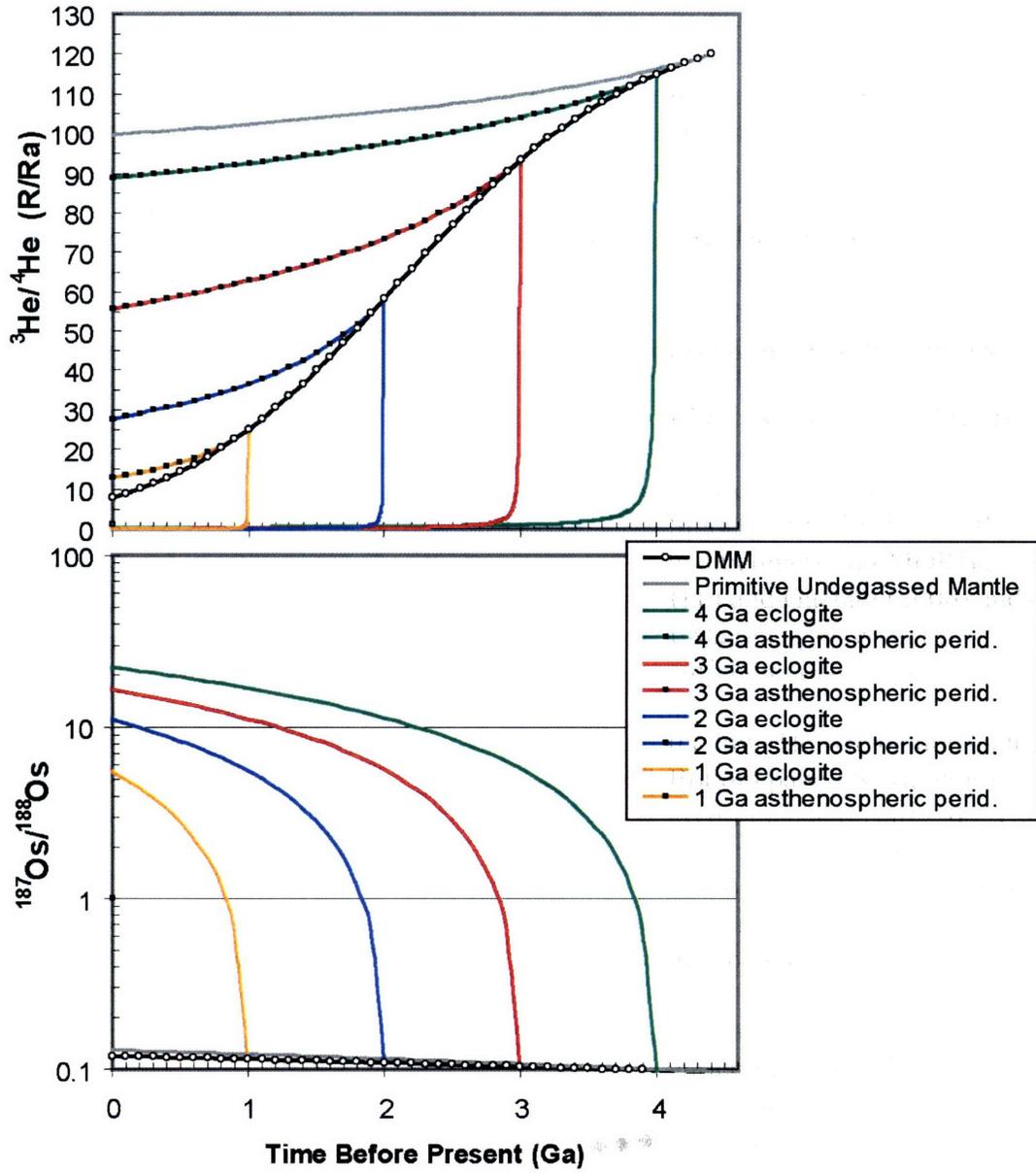


- ◆ Samoa
- Galapagoes
- Iceland
- ◆ Reunion
- ▲ Societies (Moorea)
- Heard and Kerguelen
- ⊠ Cameroon Line
- △ Cape Verde
- Comoros
- Pitcairn
- Hawaii (Koolau)
- ◆ Hawaii (HSDP2 drill core)
- Hawaii (Mauna Loa)
- Hawaii (Loihi)

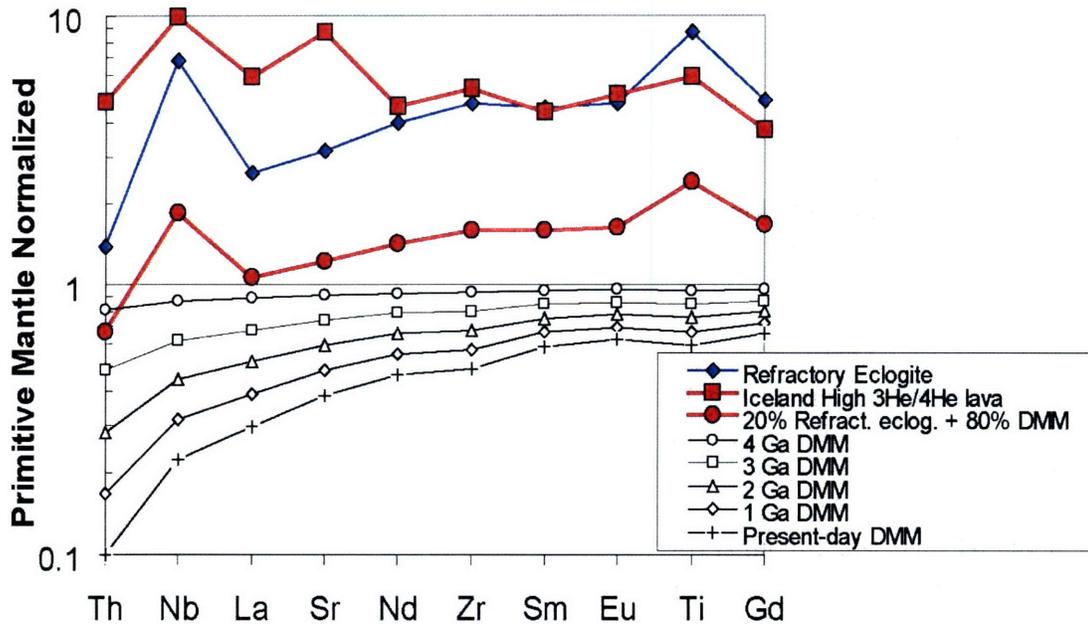


- ◆ Samoa
- Iceland (Brandon et al., 2007)
- Iceland
- Hawaii
- ◆ Hawaii HSDP2
- △ Cape Verde
- Pitcairn

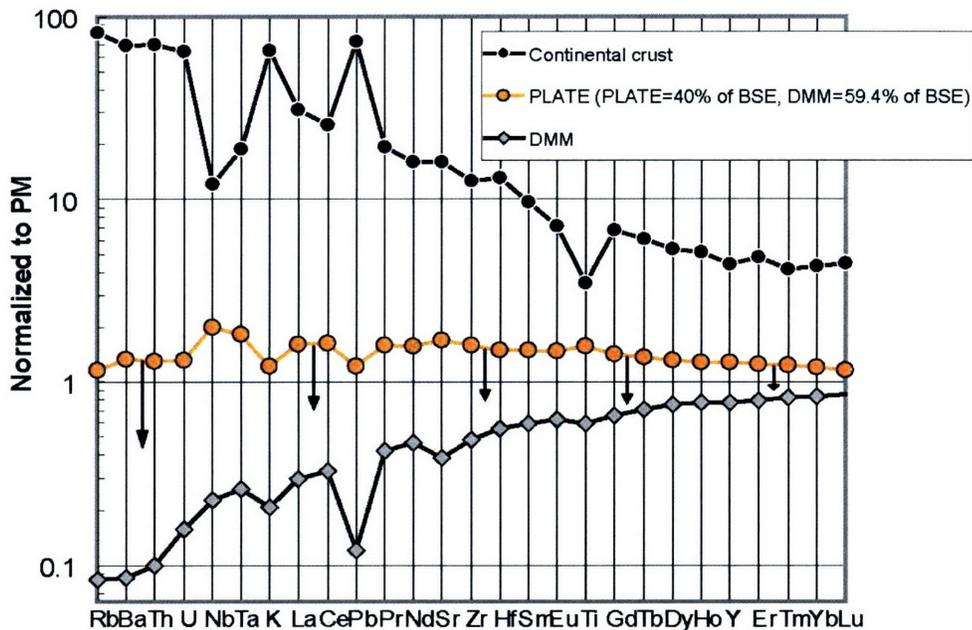
**Figure 2.** Relationships between TITAN anomalies,  $^{187}\text{Os}/^{188}\text{Os}$  and  $^3\text{He}/^4\text{He}$  in representative hotspot lavas. High  $^3\text{He}/^4\text{He}$  lavas all have elevated  $\text{Ti}/\text{Ti}^*$  (top panel),  $\text{Nb}/\text{Nb}^*$  (middle panel) and  $^{187}\text{Os}/^{188}\text{Os}$  (bottom panel). With increasing positive TITAN anomalies and  $^{187}\text{Os}/^{188}\text{Os}$  ratios in the OIBs, the maximum observed  $^3\text{He}/^4\text{He}$  increases. All samples with high  $^3\text{He}/^4\text{He}$  ( $>20$  Ra) host radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  ( $>0.130$ ). The model curves describe mixing between the high  $^3\text{He}/^4\text{He}$  (low TITAN anomaly) peridotitic and low  $^3\text{He}/^4\text{He}$  (high TITAN anomaly) refractory eclogite portions of ancient oceanic plates. More recently isolated plates, and/or more contribution from eclogite, both tend to generate lower  $^3\text{He}/^4\text{He}$  ratios in the peridotite-eclogite mixture over time. The model curves are not meant to describe the global OIB array. Instead, the model curves are only intended to constrain the mixing proportions of eclogite and peridotite in the high  $^3\text{He}/^4\text{He}$  mantle sampled by the highest  $^3\text{He}/^4\text{He}$  lavas ( $>30$  Ra). In the model, more recently isolated plates (the 1 Ga plate example is shown) can generate TITAN anomalies, but cannot generate high  $^3\text{He}/^4\text{He}$ . However, the mixing model does match the  $^3\text{He}/^4\text{He}$  ratios, moderately radiogenic  $^{187}\text{Os}/^{188}\text{Os}$  and large positive TITAN anomalies observed in the highest  $^3\text{He}/^4\text{He}$  lava (sample SEL 97 from Iceland, see Table 1) when the eclogitic (20-25%) and peridotitic (75-80%) portions of a 3 Ga plate are mixed. Mixing is marked at 10% intervals, with increasing contribution from eclogite. The most extreme lava in the figure (sample SEL 97) is the highest  $^3\text{He}/^4\text{He}$  lava available that also has a complete suite of trace elements measured by ICP. Unfortunately,  $^{187}\text{Os}/^{188}\text{Os}$  is not available for this sample, and the  $^{187}\text{Os}/^{188}\text{Os}$  of the highest  $^3\text{He}/^4\text{He}$  mantle is instead approximated using lavas (with  $^3\text{He}/^4\text{He} > 30$  Ra) from Hawaii and Samoa. Samoan post-erosional (from Savai'i) data are excluded; all  $^{187}\text{Os}/^{188}\text{Os}$  data in the figure are  $>50$  ppt Os, except Cape Verde ( $>10$  ppt) and Pitcairn ( $>20$  ppt). Data sources for the highest  $^3\text{He}/^4\text{He}$  mantle endmembers can be found in Table 1 and Appendix A.  $\text{Nb}/\text{Nb}^* = \text{Nb}_N / (\text{Th}_N \times \text{La}_N)^{0.5}$  and  $\text{Ti}/\text{Ti}^* = \text{Ti}_N / (\text{Sm} \times \text{Tb})^{0.5}$ , where N means normalized to primitive mantle.



**Figure 3.** Time evolution of  $^3\text{He}/^4\text{He}$  and  $^{187}\text{Os}/^{188}\text{Os}$  of DMM, ancient subducted (asthenospheric) DMM and eclogite. Top panel: At 4.4 Ga, DMM starts forming by melt extraction from primitive mantle to form continental and oceanic crust, and the  $^3\text{He}/^4\text{He}$  trajectories of DMM and the (hypothetical) primitive undegassed mantle separate immediately. Due to continuous depletion by melt extraction, DMM evolves low  $^3\text{He}/^4\text{He}$ . Portions of the ancient (asthenospheric) DMM, coupled to the bottomside of downgoing slabs, are sent into the lower mantle throughout geologic time (model curves are shown at 1 Ga intervals, and this is not intended to imply that the isolation of oceanic plates is episodic). These isolated peridotite portions of the downgoing slabs are modeled as having exactly the same U, Th/U,  $^{238}\text{U}/^3\text{He}$  and  $^3\text{He}/^4\text{He}$  as ambient DMM at the time of isolation, and they preserve higher  $^3\text{He}/^4\text{He}$  than DMM due to their isolation from further melt depletion. Also shown is the concomitant subduction of the eclogitic portions of the same slabs, which begin with the same  $^3\text{He}/^4\text{He}$  as DMM at the time of subduction; the U and Th of the subducted refractory eclogite are from McDonough (1991) and shown in Table 2. To simulate degassing, the  $^{238}\text{U}/^3\text{He}$  of the eclogite is increased by a factor of 1000 (Parman et al., 2005) relative to the contemporary DMM composition, and as a result the  $^3\text{He}/^4\text{He}$  ratios of the subducted eclogites rapidly decrease. Bottom Panel: The  $^{187}\text{Os}/^{188}\text{Os}$  of Primitive Mantle (0.130; Meisel et al., 2001) and DMM (0.125; Standish et al., 2002) are very similar, thus ancient subducted asthenospheric DMM is well-approximated by the trajectory of DMM in the figure (a continuous depletion model is not used to describe the evolution of  $^{187}\text{Os}/^{188}\text{Os}$  of DMM). The present-day  $^{187}\text{Re}/^{188}\text{Os}$  of the refractory eclogite is from Becker (2000), and this composition generates increasingly radiogenic present-day  $^{187}\text{Os}/^{188}\text{Os}$  for earlier isolation times. Refer to Appendix B and Table 2 for all parameters used in the model. When 20-25% of a refractory eclogite from an oceanic plate subducted at 3 Ga is mixed together with 75-80% of the asthenospheric portion of the same plate, the models generate the  $^3\text{He}/^4\text{He}$  and  $^{187}\text{Os}/^{188}\text{Os}$  of the mantle sampled by the highest  $^3\text{He}/^4\text{He}$  (>30 Ra) OIBs (see Fig. 2 for mixing results). Note that the very radiogenic Os of the eclogite is vastly reduced in the mixture due to the very low Os contents of the eclogite (6 ppt) and the high Os contents of the isolated DMM peridotite (3,000 ppt).



**Figure 4.** Spidergram of McDonough's (1991) refractory eclogite used in the modeling, including spidergrams demonstrating the time-dependent trace element composition of DMM. In the mixing models shown in Fig. 2, the asthenospheric DMM and refractory eclogitic portions of a 3 Ga subducted plate are mixed together such that the final mixture has 20-25% eclogite. This mixing calculation is shown in the spidergram of this figure, where 20% refractory eclogite has been added to the composition of DMM at 3 Ga (See Table 2 for compositions of the refractory eclogite and the 3 Ga DMM peridotite). Excluding Sr (see text for discussion of shallow lithospheric contamination, and the resulting Sr-anomalies in some Icelandic lavas), the shape of the hybrid peridotite-eclogite spidergram, including the positive Nb and Ti anomalies, is very similar in shape to the spidergram of the high  $^3\text{He}/^4\text{He}$  Icelandic lava. This similarity is consistent with the hybrid peridotite being the source of the high  $^3\text{He}/^4\text{He}$  mantle sample by OIBs.



**Figure 5.** Spidergram of a reservoir composed of recycled oceanic plates (called PLATE) hosting the elements missing in the CC and DMM reservoirs. If BSE (McDonough and Sun, 1995) is chondritic, then CC (continental crust; Rudnick and Gao, 2003) and DMM (depleted MORB mantle; Workman and Hart, 2005) are not wholly complementary reservoirs in the earth. Regardless of the relative proportions of CC and DMM, there is a shortage of the elements Ti, Ta and Nb in the silicate earth. The trace element composition of a PLATE reservoir is calculated to balance the budget of elements missing in CC and DMM (the PLATE spidergram is calculated such that the three reservoirs, PLATE, CC and DMM, sum to the total silicate mantle). The PLATE reservoir is thus composed of everything in the silicate earth that is not in the DMM and CC reservoirs, and is likely composed of mostly subducted plates and (a small portion of) sediment. The plotted spidergram of the PLATE reservoir assumes that it constitutes 40% of the mass of the silicate earth, the minimum mass of oceanic crust and lithosphere subducted over the past 2.5 Ga. If the PLATE reservoir constitutes more than 40% of BSE, its spidergram becomes more depleted (note direction of arrows in figure), but it will still exhibit positive TITAN anomalies. The PLATE spidergram shares the positive TITAN anomalies with the high  $^3\text{He}/^4\text{He}$  OIB lavas, consistent with the hypothesis that high  $^3\text{He}/^4\text{He}$  OIBs sample the domain of the PLATE reservoir composed of subducted, refractory, rutile-bearing eclogites.

**Supplementary Table 1. Major and trace element and isotopic composition of the highest  $^3\text{He}/^4\text{He}$  lavas from 4 hotspots.**

Hotspot Volcano Sample No. rock type	Iceland Selardalur SEL 97 Tholeiite	Samoa Ofu OFU-04-06 Alkali basalt	Hawaii Loihi KK18-8 Tholeiite	Galapagos Fernandina NSK 97-214 Tholeiite
SiO <sub>2</sub>	45.94	44.84	48.43	49.44
Al <sub>2</sub> O <sub>3</sub>	9.72	11.04	12.62	14.77
TiO <sub>2</sub>	1.19	4.95	2.88	3.12
FeOT	10.37	12.78	11.90	11.3
MnO	0.15	0.18	0.18	0.19
CaO	10.15	12.42	12.82	11.6
MgO	18.74	9.81	8.08	6.41
K <sub>2</sub> O	0.05	1.14	0.50	0.39
Na <sub>2</sub> O	1.18	2.24	2.34	2.46
P <sub>2</sub> O <sub>5</sub>	0.06	0.59	0.50	0.31
<sup>87</sup> Sr/ <sup>86</sup> Sr	0.703465	0.704584	0.703680	0.703290
<sup>143</sup> Nd/ <sup>144</sup> Nd	0.512969	0.512827	0.512945	0.512937
<sup>206</sup> Pb/ <sup>204</sup> Pb	18.653	18.189	18.448	19.080
<sup>207</sup> Pb/ <sup>204</sup> Pb	15.473	15.571	15.447	15.537
<sup>208</sup> Pb/ <sup>204</sup> Pb	38.453	39.202	18.189	38.710
$^3\text{He}/^4\text{He}$	37.7	33.8	32.3	30.3
Ni	759	201	80	49
Cr	1728	533	384	155
V	203	364	350	366
Ga		21		26
Cu		80		83
Zn		132		110
Cs	0.02	0.33	0.09	0.07
Rb	0.915	25.64	9.12	7.60
Ba	19.03	248.9	123.1	89.0
Th	0.382	3.956	0.994	1.310
U	0.072	0.991	0.285	0.380
Nb	6.50	50.56	16.35	20.00
Ta	0.434	3.84		1.50
La	3.85	37.93	13.82	14.1
Ce	8.52	80.5	32.9	32.3
Pb	0.38	2.61	1.25	0.93
Pr	1.23	9.99	4.62	4.23
Nd	5.80	44.08	20.65	19.5
Sr	173	599	349	351
Zr	57	293	125	155
Hf	1.65	7.56	3.48	4.15
Sm	1.78	10.75	5.33	5.77
Eu	0.79	3.39	1.83	2.04
Gd	2.05	9.93	5.50	6.01
Tb	0.35	1.40	0.87	1.03
Dy	2.16	7.34	4.89	6.08
Ho	0.42	1.23	0.92	1.21
Y	10.03	30.29	21.87	30.1
Er	1.06	2.70	2.24	3.04
Tm	0.15	0.32	0.29	0.41
Yb	0.88	1.66	1.69	2.46
Lu	0.13	0.23	0.24	0.36
Sc	33.9	31.2	35.4	44.3

Trace element data (excluding Ni, Cr, V, Ga, Cu, Zn and Sc) are by ICP-MS and are here reported for the first time on the high  $^3\text{He}/^4\text{He}$  samples from Hawaii, Iceland and Samoa; the trace element data on the Galapagos sample are reported elsewhere (Saal et al., 2007). The isotope data for all four lavas are also reported elsewhere (Hilton et al., 1999; Jackson et al., 2007a; Saal et al., 2007; Kurz et al., 1983; Staudigel et al., 1984). Major element data for the Samoan lava are reported here for the first time; major element data for the Hawaii, Iceland and Galapagos samples are available elsewhere (Hilton et al., 1999; Saal et al., 2007; Frey and Clague, 1983). Samoan major element (and Ni, Cr, V, Ga, Cu, Zn and Sc) data reported here were measured by XRF. Trace element (for Samoa, Hawaii and Iceland) data reported here were analyzed by ICP-MS. Major and trace element data reported here were analyzed at the Geoanalytical Lab at Washington State University. Internal and (estimated) external precision for major and trace element analyses at the Geoanalytical Lab are given in Jackson et al. (2007b). The Loihi sample was powdered in Tungsten-carbide, and Ta is therefore not reported for this sample.

Table 2. Components used to model He, Os and TITAN anomalies in the high  $^3\text{He}/^4\text{He}$  mantle sampled by OIBs.

	$^{238}\text{U}/^3\text{He}^1$	$^3\text{He}$ atoms/g <sup>1</sup>	$^3\text{He}/^4\text{He}$ (Ra) <sup>5</sup>	$^4\text{He}$ atoms/g	$^{232}\text{Th}/^{238}\text{U}$	$^{187}\text{Re}/^{186}\text{Os}$	Os (ppt)	$^{187}\text{Os}/^{186}\text{Os}$	Th (ppm)	U (ppm)	Nb (ppm)	La (ppm)	Sr (ppm)	Nd (ppm)	Zr (ppm)	Sm (ppm)	Eu (ppm)	Ti (ppm)	Gd (ppm)	Tb (ppm)	
Undegassed Mantle <sup>2</sup>	7.00E+01	7.3E+11	99.6	5.3E+15	4.05	0.396	3400	0.130	0.0795	0.0203	0.658	0.648	19.9	1.25	10.5	0.406	0.154	1205	0.54	0.099	
Present-day DMM <sup>3</sup>	5.40E+04	1.5E+08	8.0	1.3E+13	2.55	0.314	3000	0.120	0.0079	0.0032	0.149	0.192	7.664	0.581	5.08	0.239	0.096	716	0.36	0.070	
<b>Pure peridotite component<sup>4</sup></b>																					
DMM peridotite (isolated at 1 Ga)	1.19E+04	1.0E+09	12.8	5.8E+13	2.81	0.314	3000	0.125	0.0134	0.0049	0.207	0.253	9.52	0.692	5.99	0.270	0.107	806	0.39	0.076	
DMM peridotite (isolated at 2 Ga)	2.63E+03	7.1E+09	27.7	1.8E+14	3.12	0.314	3000	0.125	0.0228	0.0074	0.291	0.334	11.83	0.823	7.07	0.304	0.119	907	0.43	0.082	
DMM peridotite (isolated at 3 Ga)	5.80E+02	4.9E+10	55.6	6.3E+14	3.47	0.314	3000	0.125	0.0381	0.0113	0.409	0.440	14.69	0.980	8.34	0.343	0.132	1021	0.48	0.089	
DMM peridotite (isolated at 4 Ga)	1.28E+02	3.4E+11	88.6	2.7E+15	3.85	0.314	3000	0.125	0.0644	0.0172	0.574	0.580	18.25	1.166	9.83	0.387	0.148	1149	0.52	0.086	
<b>Pure eclogite component<sup>6,7</sup></b>																					
Refrac. eclogite (isolated at 1 Ga)	1.19E+07	9.4E+06	0.026	2.6E+14	2.81	325	6	5.5	0.11	0.044	4.5	1.7	63	5.0	50	1.9	0.7	10500	2.7	0.5	
Refrac. eclogite (isolated at 2 Ga)	2.63E+06	4.2E+07	0.053	5.8E+14	3.12	325	6	11.0	0.11	0.044	4.5	1.7	63	5.0	50	1.9	0.7	10500	2.7	0.5	
Refrac. eclogite (isolated at 3 Ga)	5.80E+05	1.9E+08	0.137	1.0E+15	3.47	325	6	16.6	0.11	0.044	4.5	1.7	63	5.0	50	1.9	0.7	10500	2.7	0.5	
Refrac. eclogite (isolated at 4 Ga)	1.28E+05	8.7E+08	0.385	1.6E+15	3.85	325	6	22.3	0.11	0.044	4.5	1.7	63	5.0	50	1.9	0.7	10500	2.7	0.5	

All values are calculated for the present-day. The  $^{238}\text{U}/^3\text{He}$ , U,  $^{232}\text{Th}/^{238}\text{U}$ ,  $^3\text{He}/^4\text{He}$  and trace element budgets of DMM at all times (4 Ga to present) are calculated using the continuous transport equations in Appendix B and initial values provided in this table.

<sup>1</sup> The undegassed mantle has been a closed system to all elements, volatile and non-volatile, since 4.4 Ga. The  $^{238}\text{U}/^3\text{He}$  of undegassed mantle is based on a primitive mantle U concentration (McDonough and Sun, 1995) and a  $^3\text{He}$  concentration of  $7.3 \times 10^{11}$  atoms/g, a  $^3\text{He}$  value that is well within the range suggested for the undegassed mantle in the literature ( $1.1 \times 10^{11}$  atoms/g in Class and Goldstein [2005] and  $\sim 1 \times 10^{12}$  atoms/g in the D'' layer in Tostikhin and Hofmann [2006]). To simulate >99% degassing of the slab in the subduction zone, the  $^{238}\text{U}/^3\text{He}$  value of the refrac. (refractory) eclogite is taken to be a factor of 1,000 times greater than the  $^{238}\text{U}/^3\text{He}$  of the contemporary DMM peridotite (Parman et al., 2005).

<sup>2</sup> The trace element concentrations for undegassed mantle (excluding Os and He) are from McDonough and Sun (1995). The  $^{187}\text{Os}/^{186}\text{Os}$  is from Meisel et al. (2001), and the Os and  $^{187}\text{Re}/^{186}\text{Os}$  are from McDonough and Sun (1995).

<sup>3</sup> The trace element concentrations (excluding He and Os) for present-day DMM are from Workman and Hart (2005). The  $^{187}\text{Os}/^{186}\text{Os}$  of DMM is from Standish et al. (2002).

<sup>4</sup> Using the continuous transport equations in Appendix B, the U and Th contents of the DMM peridotite are calculated to be the same as ambient upper mantle during the time of subduction and isolation; all other trace elements concentrations (except Os) are calculated the same way. Helium concentrations are derived from the calculated  $^{238}\text{U}/^3\text{He}$  and U concentrations.

<sup>5</sup> At 4.4 Ga, the mantle is assumed to have had a  $^3\text{He}/^4\text{He}$  ratio of 120 Ra, a ratio that is similar to the present-day atmosphere of Jupiter (Niemann et al., 1996); given the assumed  $^{238}\text{U}/^3\text{He}$  and  $^{232}\text{Th}/^{238}\text{U}$  primitive mantle ratios, closed system evolution yields a present-day  $^3\text{He}/^4\text{He}$  ratio of 99.6 Ra (see Appendix B). Using the continuous transport equations in Appendix B, the  $^3\text{He}/^4\text{He}$  ratios of isolated DMM peridotite (asthenosphere) throughout time are the same as ambient upper mantle DMM at the time of subduction and isolation.

<sup>6</sup> The  $^{187}\text{Re}/^{186}\text{Os}$  and Os concentrations for refractory eclogite are the median values reported for metamorphosed metabasalts in Becker (2000). The various  $^{187}\text{Os}/^{186}\text{Os}$  values shown for the recycled eclogites are then calculated assuming different isolation times of the same eclogite.

<sup>7</sup> Four of the trace elements for the refractory eclogite (Nb, La, Zr and Ti) were provided explicitly in McDonough (1991). Th, Sr, Nd, Sm, Eu and Gd were estimated from Fig. 5 in the same source. U concentrations for the refractory eclogite are not provided by McDonough (1991). U concentrations are estimated assuming a Th/U ratio of 2.5 and using the Th concentration provided; Tb (used to calculate Ti/Ti\*) is not provided by McDonough (1991), and its concentration in the refractory eclogite is set by a linear (on a Primitive Mantle normalized basis) extrapolation from Gd and Sm, which are provided. The trace element composition of McDonough's (1991) hypothetical eclogite is quite similar to the eclogites with the highest Nb/Th and Ti/Sm (proxies for Nb and Ti anomalies) in the eclogite dataset of Becker et al. (2000). Considering only the eclogites with Ti/Sm > 3000 and Nb/Th > 24, and excluding all samples that lack data for all the elements of interest (Nb, Th, Ti, Sm, U, etc.), the average trace element composition of the TITAN-enriched eclogites in Becker et al. (2000) is: Th = 0.159 ppm; U = 0.165 ppm; Nb = 10 ppm; Nd = 7.58 ppm; Sr = 88.7 ppm; Zr = 152 ppm; Sm = 2.16 ppm; Ti = 10,080 ppm.