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LASER INSTABILITY IN A RELATIVISTIC ELECTRON  
BEAM WITH UNIFORM DENSITY

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A self-consistent description of the free electron laser instability is developed for a relativistic electron beam with uniform density propagating through a helical wiggler field  $\underline{B}^0 = -\hat{B} \cos k_0 z \hat{e}_x - \hat{B} \sin k_0 z \hat{e}_y$ . The analysis is carried out for the class of solutions to the Vlasov-Maxwell equations described by  $f_B(z, \underline{p}, t) = n_0 \delta(P_x) \delta(P_y) G(z, p_z, t)$  where  $P_x$  and  $P_y$  are the exact canonical momenta invariants perpendicular to the beam propagation direction. The linearized Vlasov-Maxwell equations lead to an exact matrix dispersion relation which is valid for perturbations about general beam equilibrium  $G_0(p_z)$  and includes coupling to arbitrary harmonic number ( $n$ ) of the fundamental wiggler wavenumber  $k_0$ . No a priori restriction is made to low beam density (as measured by  $\omega_p^2/c^2 k_0^2$ ) or small wiggler amplitude (as measured by  $\hat{\omega}_c/ck_0 = e\hat{B}/\bar{\gamma}mc^2 k_0$ ). Moreover, no assumption is made that any off-diagonal elements in the matrix dispersion relation are negligibly small. A detailed numerical analysis of the exact dispersion relation is presented for the case of a cold electron beam described by  $G_0(p_z) = \delta(p_z - p_0)$ . It is shown that the instability bandwidth increases rapidly with increasing wiggler amplitude  $\hat{\omega}_c/ck_0$ . Moreover, except for very modest values of wiggler amplitude, it is shown that the growth rate calculated from an approximate version of the dispersion relation can be in substantial error for large values of  $(k + nk_0)/k_0$ . Preliminary estimates of the influence of beam thermal effects are also presented.

## 1. INTRODUCTION

In recent years there have been several theoretical<sup>1-5</sup> and experimental<sup>6-8</sup> investigations of the free electron laser which generates coherent electromagnetic radiation using an intense relativistic electron beam as an energy source. With few exceptions, theoretical studies of the free electron laser instability are based on highly simplified models which often neglect beam kinetic effects and coupling to higher harmonics of the fundamental wiggler wavenumber  $k_0$ , or make use of very idealized approximations in analyzing the matrix dispersion relation. The purpose of the present paper is to develop a fully self-consistent description of the free electron laser instability based on the Vlasov-Maxwell equations. The final matrix dispersion relation [Eq. (45)] includes all beam kinetic effects and coupling to arbitrary harmonic number ( $n$ ) of the fundamental wiggler wavenumber  $k_0$ . Moreover, the final matrix dispersion relation [Eq. (45)] makes no a priori assumption that any off-diagonal elements are negligibly small.

The present analysis assumes a relativistic electron beam with uniform cross-section propagating in the  $z$ -direction through a helical wiggler field described by [Eq. (2)]

$$\vec{B}^0 = -\hat{B} \cos k_0 z \hat{e}_x - \hat{B} \sin k_0 z \hat{e}_y$$

where  $\hat{B} = \text{const.}$  is the field amplitude and  $\lambda_0 = 2\pi/k_0$  is the wiggler wavelength. Moreover, we consider the class of exact solutions to the Vlasov-Maxwell equations described by [Eq. (12)]<sup>5</sup>

$$f_b(z, \vec{p}, t) = n_0 \delta(p_x) \delta(p_y) G(z, p_z, t),$$

where  $n_0 = \text{const.}$ ,  $P_x$  and  $P_y$  are the exact canonical momenta invariants [Eqs. (6) and (7)], and spatial variations are assumed to be one-dimensional. A detailed analysis of the linearized Vlasov-Maxwell equations (Secs. 2.B and 3.A) leads to the matrix dispersion relation [Eq. (45)]

$$D_n^L(\omega) D_{n+1}^T(\omega) D_{n-1}^T(\omega) = \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} [D_{n+1}^T(\omega) + D_{n-1}^T(\omega)] \\ \times \{ [\chi_n^{(1)}(\omega)]^2 - D_n^L(\omega) [\alpha_3 \omega_p^2 + \chi_n^{(2)}(\omega)] \},$$

where  $\hat{\omega}_c = e\hat{B}/\bar{\gamma}mc$  is the relativistic cyclotron frequency,  $\omega_p^2 = 4\pi n_0 e^2/\bar{\gamma}m$  is the relativistic plasma frequency-squared,  $\bar{\gamma}mc^2$  is the characteristic electron energy,  $\alpha_3$  is a constant of order unity [Eq.(43)], the susceptibilities  $\chi_n^{(1)}(\omega)$  and  $\chi_n^{(2)}(\omega)$  are defined in Eqs. (37) and (38),  $D_n^L(\omega)$  is the longitudinal (electrostatic) dielectric function defined in Eq. (39), and  $D_{n+1}^T(\omega)$  are the transverse (electromagnetic) dielectric functions defined in Eqs. (40) and (41). The striking feature of Eq. (45) is that the dispersion relation is valid for arbitrary harmonic number  $n$ . (No a priori assumption has been made that  $n = \pm 1$ ). Moreover, the effective wavenumber variables that occur in the various factors in Eq. (45) are  $k + nk_0$ ,  $k + (n+1)k_0$  and  $k + (n-1)k_0$ . In addition, Eq. (45) describes stability behavior for perturbations about general beam equilibrium distribution  $G_0(p_z)$ , and no a priori restriction has been made to low beam density (as measured by  $\omega_p/c^2 k_0^2$ ) or small wiggler amplitude (as measured by  $\hat{\omega}_c/ck_0$ ). Apart from the linearization assumption, no approximations have been made in deriving Eq. (45). For example, we have not assumed a priori that  $D_n^L(\omega) \simeq 0$  and therefore neglected the corresponding term on the right-hand

side of Eq. (45). The latter point is very important. In Secs. 3.B and 3.C we present a detailed numerical analysis of the exact dispersion relation [Eq. (45)] for the case of a cold electron beam described by  $G_0(p_z) = \delta(p_z - p_0)$ . The exact stability results are then compared with the approximate results obtained from Eq. (45) by assuming at the outset that  $D_n^L(\omega) \simeq 0$  and  $D_{n-1}^T(\omega) \simeq 0$ , which gives the reference dispersion relation [Eq. (46)]

$$D_n^L(\omega) D_{n-1}^T(\omega) = \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} [\chi_n^{(1)}(\omega)]^2.$$

For sufficiently large wiggler amplitude (as measured by  $\hat{\omega}_c/ck_0$ ), it is shown in Sec. 3.C that the growth rate  $\omega_i = \text{Im}\omega$  obtained from the reference dispersion relation [Eq. (46)] can be in substantial error for large values of  $(k + nk_0)/k_0$ .

For completeness, and to orient the reader, it is useful to summarize here the interaction wavenumbers and frequencies pertinent to the free electron laser instability.<sup>9,10</sup> We consider a cold electron beam (Secs. 3.B and 3.C) and look for simultaneous solutions to  $D_n^L(\omega) = 0$  and  $D_{n-1}^T(\omega) = 0$ . Shown in Fig. 1 is a first-quadrant plot ( $\omega > 0$  and  $k + nk_0 > 0$ ) of

$$\omega = (k + nk_0)v_b \pm \omega_p/\bar{\gamma}$$

and

$$\omega = [c^2(k + nk_0 - k_0)^2 + \omega_p^2]^{1/2}$$

versus  $k + nk_0$ . Note that there are generally four intersection points. Solving the above equations simultaneously, the upshifted wavenumbers

(intersections with  $k + nk_0 > k_0$ ) are given by

$$(k+nk_0)_u^\pm = \bar{\gamma}^2 k_0 \left[ 1 + \beta_b \left( 1 \pm 2 \frac{\omega_p}{ck_0} \frac{1}{\bar{\gamma}\beta_b} \right)^{1/2} \right] \pm \bar{\gamma}\beta_b \frac{\omega_p}{ck_0} k_0,$$

and the downshifted wavenumbers (intersections with  $k+nk_0 < k_0$ ) are given by

$$(k+nk_0)_d^\pm = \bar{\gamma}^2 k_0 \left[ 1 - \beta_b \left( 1 \pm 2 \frac{\omega_p}{ck_0} \frac{1}{\bar{\gamma}\beta_b} \right)^{1/2} \right] \pm \bar{\gamma}\beta_b \frac{\omega_p}{ck_0} k_0,$$

where  $\bar{\gamma} = (1 - \beta_b^2)^{-1/2}$  and  $\beta_b = v_b/c$ . In the limit of low beam density with  $\omega_p^2/c^2 k_0^2 \ll 1$ ,  $\bar{\gamma}^2 \beta_b^2$ , the intersection wavenumbers can be approximated by

$$(k+nk_0)_u^\pm = (1+\beta_b)\bar{\gamma}^2 k_0 = k_0/(1-\beta_b),$$

and

$$(k+nk_0)_d^\pm = (1-\beta_b)\bar{\gamma}^2 k_0 = k_0/(1+\beta_b).$$

For sufficiently large  $\bar{\gamma}$ , the upshifted wavenumber  $(k+nk_0)_u^\pm = (1+\beta_b)\bar{\gamma}^2 k_0$  can correspond to very short wavelengths and is the intersection region of interest for the free electron laser instability.<sup>9,10</sup> Depending on the value of beam density ( $\omega_p^2/c^2 k_0^2$ ) and wiggler amplitude ( $\hat{\omega}_c/ck_0$ ), however, it is important to note that the dispersion relation (45) may also support instability in the long wavelength intersection region corresponding to  $(k + nk_0)_d^\pm$ . (See Sec. 3.C and Fig. 3).

## 2. THEORETICAL MODEL AND ASSUMPTIONS

### A. General Theoretical Model

The present analysis assumes a relativistic electron beam with uniform cross section propagating in the z-direction. The beam density is assumed to be sufficiently small that equilibrium space charge effects are negligibly small with

$$\vec{E}^0 = 0 . \quad (1)$$

Moreover, the electron beam propagates through a helical wiggler magnetic field described by <sup>1,2,5</sup>

$$\vec{B}^0 = -\hat{B} \cos k_0 z \hat{e}_{\nu_x} - \hat{B} \sin k_0 z \hat{e}_{\nu_y} , \quad (2)$$

where  $\hat{B} = \text{const.}$  is the field amplitude,  $\lambda_0 = 2\pi/k_0$  is the wavelength, and  $\hat{e}_{\nu_x}$  and  $\hat{e}_{\nu_y}$  are unit Cartesian vectors in the plane perpendicular to the propagation direction. The vector potential associated with Eq. (2) is given by

$$\vec{A}^0 = (\hat{B}/k_0) \cos k_0 z \hat{e}_{\nu_x} + (\hat{B}/k_0) \sin k_0 z \hat{e}_{\nu_y} . \quad (3)$$

It is also assumed that the beam density and current are sufficiently small that the equilibrium self magnetic field can be neglected in comparison with  $\vec{B}^0$ . <sup>11</sup>

We consider perturbations in which the spatial variations are one-dimensional in nature with  $\partial/\partial x = 0 = \partial/\partial y$ , and  $\partial/\partial z$  generally non-zero. Introducing the perturbed potentials,

$$\delta\phi(z, t)$$

and

$$\delta\vec{A}_{\nu}(x, t) = \delta A_x(z, t) \hat{e}_{\nu_x} + \delta A_y(z, t) \hat{e}_{\nu_y} ,$$

the electromagnetic field perturbations  $\delta\vec{E}$  and  $\delta\vec{B}$  can be expressed in the Coulomb gauge as

$$\delta\vec{E}(\vec{x}, t) = -\frac{\partial\delta\phi}{\partial z}(z, t)\hat{e}_z - \frac{1}{c}\frac{\partial}{\partial t}\delta A_x(z, t)\hat{e}_x - \frac{1}{c}\frac{\partial}{\partial t}\delta A_y(z, t)\hat{e}_y, \quad (4)$$

and

$$\delta\vec{B} = -\frac{\partial}{\partial z}\delta A_y(z, t)\hat{e}_x + \frac{\partial}{\partial z}\delta A_x(z, t)\hat{e}_y, \quad (5)$$

where  $\delta\vec{E} = -\nabla\delta\phi - (1/c)(\partial/\partial t)\delta\vec{A}$  and  $\delta\vec{B} = \nabla \times \delta\vec{A}$ . In the present geometry, there are two exact single-particle invariants in the combined equilibrium and perturbed field configuration. These are the canonical momenta,  $P_x$  and  $P_y$ , transverse to the beam propagation direction, i.e.,

$$P_x = p_x - \frac{e}{c}A_x^0(z) - \frac{e}{c}\delta A_x(z, t) = \text{const.}, \quad (6)$$

$$P_y = p_y - \frac{e}{c}A_y^0(z) - \frac{e}{c}\delta A_y(z, t) = \text{const.}, \quad (7)$$

where  $A_x^0(z) = (\hat{B}/k_0)\text{cos}k_0z$  and  $A_y^0(z) = (\hat{B}/k_0)\text{sin}k_0z$  [Eq. (3)], and  $p_x$  and  $p_y$  are the transverse mechanical momenta.

The potential perturbations  $\delta\phi(z, t)$ ,  $\delta A_x(z, t)$  and  $\delta A_y(z, t)$  are determined self-consistently from the Maxwell equations

$$\left(\frac{1}{c^2}\frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2}\right)\delta A_x = -\frac{4\pi e}{c}\int d^3p v_x(f_b - f_b^0), \quad (8)$$

$$\left(\frac{1}{c^2}\frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2}\right)\delta A_y = -\frac{4\pi e}{c}\int d^3p v_y(f_b - f_b^0), \quad (9)$$

$$\frac{\partial^2}{\partial z^2}\delta\phi = 4\pi e\int d^3p(f_b - f_b^0), \quad (10)$$

where  $f_b^0(z, p)$  is the equilibrium ( $\partial/\partial t=0$ ) beam distribution, and  $f_b(z, p, t)$  in general solves the nonlinear Vlasov equation

$$\left\{\frac{\partial}{\partial t} + v_z\frac{\partial}{\partial z} - e\left[\delta\vec{E} + \frac{v \times (\vec{B}^0 + \delta\vec{B})}{c}\right] \cdot \frac{\partial}{\partial p}\right\} f_b(z, p, t) = 0. \quad (11)$$

In Eqs. (8)-(11), the particle velocity  $\vec{v}$  and momentum  $\vec{p}$  are related by



$m\gamma = p / (1 + p^2/m^2c^2)^{1/2}$ . Moreover,  $m$  is the electron rest mass,  $-e$  is the electron charge, and  $c$  is the speed of light in vacuo.

For present purposes, we examine the class of exact solutions to Eq. (11) of the form<sup>5</sup>

$$f_b(z, p, t) = n_0 \delta(P_x) \delta(P_y) G(z, p_z, t), \quad (12)$$

where  $n_0 = \text{const.}$ , and  $P_x$  and  $P_y$  are the exact invariants defined in Eqs. (6) and (7). Note from Eq. (12) that the effective transverse motion of the beam electrons is "cold". Substituting Eq. (12) into Eq. (11) and making use of Eqs. (6) and (7), we find that Eq. (12) solves Eq. (11) exactly provided  $G(z, p_z, t)$  evolves according to the one-dimensional Vlasov equation

$$\left\{ \frac{\partial}{\partial t} + v_z \frac{\partial}{\partial z} - \frac{\partial}{\partial z} \hat{H}(z, t) \frac{\partial}{\partial p_z} \right\} G(z, p_z, t) = 0, \quad (13)$$

where  $\hat{H}$  is defined by

$$\hat{H}(z, t) = \gamma_T mc^2 - e\delta\phi(z, t). \quad (14)$$

In Eq. (14),

$$\gamma_T mc^2 = [m^2 c^4 + c^2 p_z^2 + e^2 (A_x^0 + \delta A_x)^2 + e^2 (A_y^0 + \delta A_y)^2]^{1/2}, \quad (15)$$

is the particle energy for  $P_x = 0 = P_y$ . Moreover, substituting Eq. (12) into Eqs. (8)-(11), the equations describing the nonlinear evolution of the potential perturbations can be expressed as

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) \delta A_x = - \frac{4\pi n_0 e^2}{mc^2} \left( (A_x^0 + \delta A_x) \int \frac{dp_z}{\gamma_T} G - A_x^0 \int \frac{dp_z}{\gamma} G_0 \right), \quad (16)$$

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) \delta A_y = - \frac{4\pi n_0 e^2}{mc^2} \left( (A_y^0 + \delta A_y) \int \frac{dp_z}{\gamma_T} G - A_y^0 \int \frac{dp_z}{\gamma} G_0 \right), \quad (17)$$

$$\frac{\partial^2}{\partial z^2} \delta\phi = -4\pi e n_0 \int dp_z [G - G_0], \quad (18)$$

where  $\gamma_{Tmc}^2$  is defined in Eq. (15),  $G_0(p_z)$  is the equilibrium ( $\partial/\partial t = 0$ ) beam distribution,  $G(z, p_z, t)$  solves the nonlinear Vlasov equation (13), and

$$\gamma_{Tmc}^2 = (m^2 c^4 + c^2 p_z^2 + e^2 \hat{B}^2 / k_0^2)^{1/2}, \quad (19)$$

is the energy in the absence of perturbations ( $\delta A_x = 0 = \delta A_y$ ). In obtaining Eq. (19) from Eq. (15), use has been made of  $(A_x^0)^2 + (A_y^0)^2 = \hat{B}^2 / k_0^2 = \text{const.}$  [Eq. (3)].

Within the context of the present model, Eqs. (13) and (16)-(18) describe the exact nonlinear evolution of the system for perturbations about the general beam equilibrium  $G_0(p_z)$ . In this regard, we note from Eqs. (13) and (14) that the axial force  $F_z$  on an electron in the phase space  $(z, p_z)$  is given by

$$F_z = - \frac{\partial}{\partial z} \hat{H} = - \frac{e^2}{2\gamma_{Tmc}^2} \frac{\partial}{\partial z} [2A_x^0 \delta A_x + 2A_y^0 \delta A_y + (\delta A_x)^2 + (\delta A_y)^2] + e \frac{\partial \delta \phi}{\partial z}. \quad (20)$$

That is, the effective ponderomotive potential is proportional to

$$2A_x^0 \delta A_x + 2A_y^0 \delta A_y + (\delta A_x)^2 + (\delta A_y)^2.$$

### B. Linearization Approximation

For purposes of the stability analysis, we now consider the linearized version of Eqs. (13) and (16)-(18). In this regard, it is useful to introduce the dimensionless potentials defined by

$$\delta \hat{\phi} = \frac{e \delta \phi}{\gamma_{Tmc}^2}, \quad (21)$$

$$\delta \hat{A}_{\pm} = \frac{e}{\gamma_{Tmc}^2} (\delta A_x \pm i \delta A_y), \quad (22)$$

$$\hat{A}_{\pm}^0 = \frac{e}{\gamma_{Tmc}^2} (A_x^0 \pm i A_y^0) = \frac{\hat{\omega}_c}{ck_0} \exp(\pm i k_0 z). \quad (23)$$

In Eqs. (21)-(23),  $\gamma_{Tmc}^2 = \text{const.}$  denotes the characteristic mean energy

of the electron beam, and

$$\hat{\omega}_c = \frac{e\hat{B}}{\bar{\gamma}mc}, \quad (24)$$

is the relativistic cyclotron frequency associated with the wiggler field amplitude  $\hat{B}$ . Making use of Eqs. (21)-(24), the perturbed force  $\delta F_z$  [Eq. (20)] and inverse relativistic mass factor  $\gamma_T^{-1}$  [Eq. (15)] can be approximated by

$$\delta F_z = -\bar{\gamma}mc^2 \left\{ \frac{1}{2} \frac{\hat{\omega}_c}{ck_0} \frac{\bar{\gamma}}{\gamma} \frac{\partial}{\partial z} [\exp(ik_0 z) \delta \hat{A}_- + \exp(-ik_0 z) \delta \hat{A}_+] - \frac{\partial}{\partial z} \delta \hat{\phi} \right\}, \quad (25)$$

and

$$\frac{1}{\gamma_T} = \frac{1}{\gamma} - \frac{1}{2} \frac{\bar{\gamma}}{\gamma^3} \frac{\hat{\omega}_c}{ck_0} [\exp(ik_0 z) \delta \hat{A}_- + \exp(-ik_0 z) \delta \hat{A}_+], \quad (26)$$

where  $\gamma mc^2 = (m^2 c^4 + c^2 p_z^2 + e^2 \hat{B}^2 / k_0^2)^{1/2}$  [Eq. (19)]. Substituting Eqs. (25) and (26) into Eqs. (13) and (16)-(18), and combining Eqs. (16) and (17) to give equations for  $\delta A_x \pm i \delta A_y$ , the linearized equations for  $\delta G(z, p_z, t)$ ,  $\delta A_{\pm}(z, t)$  and  $\delta \phi(z, t)$  can be expressed as

$$\left( \frac{\partial}{\partial t} + v_z \frac{\partial}{\partial z} \right) \delta G = -\bar{\gamma}mc^2 \left\{ \frac{1}{2} \frac{\hat{\omega}_c}{ck_0} \frac{\bar{\gamma}}{\gamma} \frac{\partial}{\partial z} [\exp(ik_0 z) \delta \hat{A}_- + \exp(-ik_0 z) \delta \hat{A}_+] - \frac{\partial}{\partial z} \delta \hat{\phi} \right\} \frac{\partial}{\partial p_z} G_0(p_z), \quad (27)$$

$$\frac{\partial^2}{\partial z^2} \delta \phi = \frac{\omega_p^2}{c^2} \int dp_z \delta G, \quad (28)$$

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) \delta \hat{A}_+ = -\frac{\omega_p^2}{c^2} \left( \delta \hat{A}_+ \bar{\gamma} \int \frac{dp_z}{\gamma} G_0 + \frac{\hat{\omega}_c}{ck_0} \exp(ik_0 z) \left\{ \bar{\gamma} \int dp_z \frac{\delta G}{\gamma} - \frac{1}{2} \frac{\hat{\omega}_c}{ck_0} [\exp(ik_0 z) \delta \hat{A}_- + \exp(-ik_0 z) \delta \hat{A}_+] \times \bar{\gamma}^3 \int \frac{dp_z}{\gamma^3} G_0 \right\} \right), \quad (29)$$

$$\begin{aligned}
& \left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \frac{\partial^2}{\partial z^2} \right) \delta \hat{A}_- = - \frac{\omega_p^2}{c^2} \left( \delta \hat{A}_- \bar{\gamma} \int \frac{dp_z}{\gamma} G_0 \right. \\
& \left. + \frac{\hat{\omega}_c}{ck_0} \exp(-ik_0 z) \left\{ \bar{\gamma} \int dp_z \frac{\delta G}{\gamma} - \frac{1}{2} \frac{\hat{\omega}_c}{ck_0} [\exp(ik_0 z) \delta \hat{A}_- + \exp(-ik_0 z) \delta \hat{A}_+] \bar{\gamma}^{-3} \int \frac{dp_z}{\gamma^3} G_0 \right\} \right).
\end{aligned} \tag{30}$$

In Eqs. (28)-(30),

$$\omega_p^2 = \frac{4\pi n_0 e^2}{\bar{\gamma} m}, \tag{31}$$

is the relativistic plasma frequency-squared, and  $\delta \hat{\phi}$ ,  $\delta \hat{A}_\pm$  and  $\hat{\omega}_c$  are defined in Eqs. (21), (22), and (24). In the limit of zero wiggler amplitude ( $\hat{\omega}_c \rightarrow 0$ ), we note that Eqs. (27)-(30) give the usual uncoupled electromagnetic and electrostatic dispersion relations for perturbations about a one-dimensional relativistic plasma equilibrium.

### 3. STABILITY PROPERTIES

#### A. General Dispersion Relation

For the equilibrium configuration considered here, the electrons are constrained to move on surfaces with  $P_x^0 = p_x - (e/c)A_x^0 = 0 = p_y - (e/c)A_y^0 = P_y^0$  and  $\gamma m c^2 = (m^2 c^4 + c^2 p_z^2 + \hat{B}^2/k_0^2)^{1/2} = \text{const.}$  This implies that the axial momentum  $p_z$  is a constant in the equilibrium field configuration. Moreover, the corresponding electron trajectory that passes through  $(z, p_z)$  at time  $t' = t$  is given by

$$z' = z + \frac{p_z}{\gamma m} (t' - t), \quad (32)$$

$$p_z' = p_z,$$

where  $p_z/\gamma m = v_z$  is the axial velocity.

Without loss of generality we expand the field perturbations in Eqs. (27)-(30) according to

$$\begin{aligned} \delta \hat{\phi} &= \sum_n \phi_n \exp[i(k+nk_0)z - i\omega t], \\ \exp(ik_0 z) \delta \hat{A}_- &= \sum_n \hat{A}_{n-1}^- \exp[i(k+nk_0)z - i\omega t], \\ \exp(-ik_0 z) \delta \hat{A}_+ &= \sum_n \hat{A}_{n+1}^+ \exp[i(k+nk_0)z - i\omega t], \end{aligned} \quad (33)$$

where  $\lambda_0 = 2\pi/k_0$  is the periodicity length of the wiggler field. Here we assume that  $k$  is real and  $\text{Im}\omega > 0$ , corresponding to temporal growth. A completely parallel treatment can be developed for spatially growing perturbations ( $\text{Im}k < 0$ ) and real oscillation frequency  $\omega$ . Substituting Eq. (33) into Eq. (27) and integrating along characteristics, we obtain for the perturbed distribution function  $\delta G(z, p_z, t)$

$$\begin{aligned}
\delta G = & \bar{\gamma} m c^2 \frac{\partial G_0}{\partial p_z} \sum_n i(k+nk_0) \exp[i(k+nk_0)z] \\
& \times \left( \frac{1}{2} \frac{\hat{\omega}}{c k_0} \frac{\bar{\gamma}}{\gamma} (A_{n-1}^- + A_{n+1}^+) - \phi_n \right) \\
& \times \int_{-\infty}^t dt' \exp[i(k+nk_0)v_z(t'-t) - i\omega t'] .
\end{aligned} \tag{34}$$

Integrating with respect to  $t'$  in Eq. (34) with  $\text{Im}\omega > 0$  gives

$$\begin{aligned}
\delta G = & \bar{\gamma} m c^2 \frac{\partial G_0}{\partial p_z} \sum_n \frac{(k+nk_0)}{\omega - (k+nk_0)v_z} \exp[i(k+nk_0)z - i\omega t] \\
& \times \left( \phi_n - \frac{1}{2} \frac{\hat{\omega}}{c k_0} \frac{\bar{\gamma}}{\gamma} (A_{n-1}^- + A_{n+1}^+) \right) ,
\end{aligned} \tag{35}$$

where  $v_z = p_z / \gamma m$  and  $\gamma m c^2 = (m^2 c^4 + c^2 p_z^2 + \hat{B}^2 / k_0^2)^{1/2}$ .

From Eqs. (28)-(30) for the field perturbations, it is evident that integrals of the form  $\omega_p^2 \int dp_z \delta G$  and  $\omega_p^2 \int dp_z \delta G / \gamma$  are required. Therefore, comparing with Eq. (35), it is convenient to introduce the effective susceptibilities  $\chi_n^{(j)}(\omega)$  defined by

$$\chi_n^{(0)}(\omega) = \bar{\gamma} m c^2 \omega_p^2 \int dp_z \frac{(k+nk_0) \partial G_0 / \partial p_z}{\omega - (k+nk_0)v_z} , \tag{36}$$

$$\chi_n^{(1)}(\omega) = \bar{\gamma} m c^2 \omega_p^2 \int \frac{dp_z}{\gamma} \frac{(k+nk_0) \partial G_0 / \partial p_z}{\omega - (k+nk_0)v_z} . \tag{37}$$

$$\chi_n^{(2)}(\omega) = \bar{\gamma} m c^2 \omega_p^2 \int \frac{dp_z}{\gamma^2} \frac{(k+nk_0) \partial G_0 / \partial p_z}{\omega - (k+nk_0)v_z} . \tag{38}$$

For future notational convenience we also introduce the longitudinal and transverse dielectric functions defined by

$$D_n^L(\omega) = c^2 (k+nk_0)^2 + \chi_n^{(0)}(\omega) , \tag{39}$$

$$D_{n+1}^T(\omega) = \omega^2 - c^2 [k+(n+1)k_0]^2 - \alpha_1 \omega_p^2 , \tag{40}$$

$$D_{n-1}^T(\omega) = \omega^2 - c^2 [k+(n-1)k_0]^2 - \alpha_1 \omega_p^2 , \tag{41}$$

where

$$\alpha_1 = \bar{\gamma} \int \frac{dp_z}{\gamma} G_0(p_z) \quad (42)$$

and

$$\alpha_3 = \bar{\gamma}^3 \int \frac{dp_z}{\gamma^3} G_0(p_z) \quad (43)$$

for future reference. Note that the constants  $\alpha_1$  and  $\alpha_3$  are of order unity whenever  $G_0$  is strongly peaked around  $\gamma = \bar{\gamma}$ .

We substitute Eqs. (33) and (35) into the field equations (28)-(30) and make use of the definitions in Eqs. (36)-(43). After some straightforward algebra, we obtain the matrix equation relating  $A_{n+1}^+$ ,  $A_{n-1}^-$ , and  $\phi_n$

$$\begin{pmatrix} D_{n+1}^T + \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} (\alpha_3 \omega_p^2 + \chi_n^{(2)}) & \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} (\alpha_3 \omega_p^2 + \chi_n^{(2)}) & -\frac{\hat{\omega}_c}{ck_0} \chi_n^{(1)} \\ \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} (\alpha_3 \omega_p^2 + \chi_n^{(2)}) & D_{n-1}^T + \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} (\alpha_3 \omega_p^2 + \chi_n^{(2)}) & -\frac{\hat{\omega}_c}{ck_0} \chi_n^{(1)} \\ -\frac{\hat{\omega}_c}{ck_0} \chi_n^{(1)} & -\frac{\hat{\omega}_c}{ck_0} \chi_n^{(1)} & 2D_n^L \end{pmatrix} \begin{pmatrix} A_{n+1}^+ \\ A_{n-1}^- \\ \phi_n \end{pmatrix} = 0. \quad (44)$$

The condition for a nontrivial solution to Eq. (44) is that the determinant of the matrix vanish. This gives the dispersion relation

$$\begin{aligned} D_n^L(\omega) D_{n+1}^T(\omega) D_{n-1}^T(\omega) &= \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} [D_{n+1}^T(\omega) + D_{n-1}^T(\omega)] \\ &\times \{ [\chi_n^{(1)}(\omega)]^2 - D_n^L(\omega) [\alpha_3 \omega_p^2 + \chi_n^{(2)}(\omega)] \}, \end{aligned} \quad (45)$$

which determines the complex eigenfrequency  $\omega$  in terms of  $k+nk_0$ ,  $k_0$ ,  $\omega_p^2$ , and  $\hat{\omega}_c^2/c^2 k_0^2$ .

The striking feature of Eq. (45) is that it is valid for arbitrary harmonic number  $n$ . (No a priori assumption has been made that  $n=\pm 1$ .) Moreover, the effective wavenumber variable  $k'=k+nk_0$  occurs in every factor in Eq. (45). In addition, Eq. (45) describes stability behavior

for general equilibrium distribution  $G_0(p_z)$ , and no a priori restriction has been made to low beam density (as measured by  $\omega_p^2/c^2k_0^2$ ) or small wiggler amplitude (as measured by  $\hat{\omega}_c/ck_0$ ). Since Eq. (45) contains no approximations apart from the linearization approximation, we refer to Eq. (45) as the full dispersion relation (FDR).

In circumstances where the beam density is low and the wiggler amplitude is very small, Eq. (45) supports solutions near  $D_n^L(\omega)=0$ ,  $D_{n+1}^T(\omega)=0$  and  $D_{n-1}^T(\omega)=0$ . In this case, it is instructive to simplify Eq. (45) near the simultaneous zeroes of  $D_n^L(\omega)=0$  and  $D_{n-1}^T(\omega)=0$ . (Here we assume  $D_{n+1}^T \neq 0$ .) Equation (45) then reduces to the simplified approximate form

$$D_n^L(\omega)D_{n-1}^T(\omega) = \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2k_0^2} [\chi_n^{(1)}(\omega)]^2. \quad (46)$$

In the subsequent analysis, we refer to Eq. (46) as the reference dispersion relation (RDR).

### B. Dispersion Relation for a Cold Electron Beam

As a first application, we consider the dispersion relation (45) for the case of a cold beam equilibrium described by

$$G_0(p_z) = \delta(p_z - p_0), \quad (47)$$

where  $p_0$  is related to the mean energy  $\bar{\gamma}mc^2$  by

$$\bar{\gamma}mc^2 = (m^2c^4 + c^2p_0^2 + e^2B^2/k_0^2)^{1/2}. \quad (48)$$

It is straightforward to show from Eqs. (47) and (48) that  $\int dp_z \gamma^n G_0(p_z) = \bar{\gamma}^n$  so that

$$\alpha_1 = \alpha_3 = 1, \quad (49)$$

follows from Eqs. (42) and (43). Substituting Eq. (47) into Eqs. (36)-(41), the effective susceptibilities and dielectric functions for a



cold electron beam can be expressed as

$$\chi_n^{(0)}(\omega) = -c^2 (k+nk_0)^2 \frac{\omega_p^2 / \gamma^{-2}}{[\omega - (k+nk_0)v_b]^2}, \quad (50)$$

$$\chi_n^{(1)}(\omega) = c(k+nk_0) \omega_p^2 \frac{[\omega \beta_b - c(k+nk_0)]}{[\omega - (k+nk_0)v_b]^2}, \quad (51)$$

$$\chi_n^{(2)}(\omega) = c(k+nk_0) \omega_p^2 \frac{[2\omega \beta_b - c(k+nk_0)(1+\beta_b^2)]}{[\omega - (k+nk_0)v_b]^2}, \quad (52)$$

$$D_n^L(\omega) = c^2 (k+nk_0)^2 \left\{ 1 - \frac{\omega_p^2 / \gamma^{-2}}{[\omega - (k+nk_0)v_b]^2} \right\}, \quad (53)$$

$$D_{n+1}^T(\omega) = \omega^2 - c^2 [k+(n+1)k_0]^2 - \omega_p^2, \quad (54)$$

$$D_{n-1}^T(\omega) = \omega^2 - c^2 [k+(n-1)k_0]^2 - \omega_p^2. \quad (55)$$

In Eqs. (50)-(55),  $v_b = p_0 / \gamma m$  is the beam velocity, and  $\beta_b$  is defined by  $\beta_b = v_b / c$ . Substituting Eqs. (50)-(55) into Eq. (45) and carefully combining and simplifying terms on the right-hand side, the full dispersion relation (FDR) for the free electron laser instability can be expressed as

$$\begin{aligned} & \{[\omega - (k+nk_0)v_b]^2 - \omega_p^2 / \gamma^{-2}\} \{\omega^2 - c^2 [(k+nk_0) - k_0]^2 - \omega_p^2\} \\ & \times \{\omega^2 - c^2 [(k+nk_0) + k_0]^2 - \omega_p^2\} \\ & = - \frac{\hat{\omega}_c^2}{c^2 k_0^2} \omega_p^2 [\omega^2 - c^2 (k+nk_0)^2 - c^2 k_0^2 - \omega_p^2] \\ & \times \{[\omega - (k+nk_0)v_b]^2 + 2v_b (k+nk_0) [\omega - (k+nk_0)v_b] \\ & \quad - c^2 (k+nk_0)^2 / \gamma^{-2} - \omega_p^2\}. \end{aligned} \quad (56)$$

Note the exact cancellations that have occurred on the right-hand side of

Eq. (56). Making use of  $\bar{\gamma}^{-2} = 1 - \beta_b^2$ , the full dispersion relation (FDR) in Eq. (56) can be expressed in the equivalent form

$$\begin{aligned} & \{[\omega - (k + nk_0)V_b]^2 - \omega_p^2/\bar{\gamma}^2\} \{\omega^2 - c^2[(k + nk_0) - k_0]^2 - \omega_p^2\} \\ & \quad \times \{\omega^2 - c^2[(k + nk_0) + k_0]^2 - \omega_p^2\} \\ & = -\frac{\hat{\omega}_c^2}{c^2 k_0^2} \omega_p^2 [\omega^2 - c^2(k + nk_0)^2 - c^2 k_0^2 - \omega_p^2] \\ & \quad \times [\omega^2 - c^2(k + nk_0)^2 - \omega_p^2]. \end{aligned} \quad (57)$$

Moreover, substituting Eqs. (50), (51), (53), and (54) into Eq. (46), the reference dispersion relation (RDR), which is approximately valid for low beam density and very low wiggler amplitude, can be expressed as

$$\begin{aligned} & \{[\omega - (k + nk_0)V_b]^2 - \omega_p^2/\bar{\gamma}^2\} \{\omega^2 - c^2[(k + nk_0) - k_0]^2 - \omega_p^2\} \\ & = \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} \omega_p^4 \frac{[\omega \beta_b - c(k + nk_0)]^2}{[\omega - (k + nk_0)V_b]^2}. \end{aligned} \quad (58)$$

Approximating  $[\omega - (k + nk_0)V_b]^2 = \omega_p^2/\bar{\gamma}^2$  in the denominator on the right-hand side of Eq. (57), the reference dispersion relation (RDR) becomes

$$\begin{aligned} & \{[\omega - (k + nk_0)V_b]^2 - \omega_p^2/\bar{\gamma}^2\} \{\omega^2 - c^2[(k + nk_0) - k_0]^2 - \omega_p^2\} \\ & = \frac{1}{2} \frac{\hat{\omega}_c^2}{c^2 k_0^2} \omega_p^2 \bar{\gamma}^{-2} [\omega \beta_b - c(k + nk_0)]^2. \end{aligned} \quad (59)$$

### C. Numerical Analysis of Dispersion Relation

In this section, we summarize the results of numerical studies of the reference dispersion relation (59) and the full dispersion relation (57) for a broad range of dimensionless system parameters  $\hat{\omega}_c/ck_0$ ,  $\omega_p^2/c^2 k_0^2$ , and  $\bar{\gamma}$ .

Typical results are summarized in Fig. 2 for  $\bar{\gamma}=2$  and moderate values of wiggler amplitude ( $\hat{\omega}_c/ck_0=0.5$ ) and beam density ( $\omega_p^2/c^2k_0^2=0.4$ ). For such large beam density and wiggler amplitude, we note from Fig. 2(a) that the growth rate  $\omega_i=Im\omega$  obtained from the reference dispersion relation (59) can be in substantial error for large values of  $(k+nk_0)/k_0$ . This follows since the dashed curve in Fig. 2(a) [Eq. (59)] deviates significantly from the solid curve [Eq. (57)] for large values of  $(k+nk_0)/k_0$ . On the other hand, from Fig. 2(b), both the reference dispersion relation (59) and full dispersion relation (57) give very similar values of the real oscillation frequency  $\omega_r=Re\omega$ . Moreover, the oscillation frequency  $\omega_r$  is linearly proportional to  $(k+nk_0)$  over the entire range of unstable wavenumbers. We also note from Fig. 2(a) that the maximum growth assumes a relatively large value,  $[\omega_i]_{\max} \approx 0.3 ck_0$ , for  $(\hat{\omega}_c/ck_0, \omega_p^2/c^2k_0^2)=(0.5, 0.4)$ .

To contrast with Fig. 2, we show in Fig. 3 a plot of normalized growth rate  $\omega_i/ck_0$  versus  $(k+nk_0)/k_0$  for  $\bar{\gamma}=2$ , moderate beam density ( $\omega_p^2/c^2k_0^2=0.4$ ), and small wiggler amplitude ( $\hat{\omega}_c/ck_0=0.1$ ). In this case the reference dispersion relation (59) and full dispersion relation (57) give nearly identical growth rate curves. Note also from Fig. 3 that there are two unstable wavenumber bands [corresponding to two intersection regions of the  $\omega$  versus  $k+nk_0$  curves from  $D_n^L(\omega) = 0$  and  $D_{n-1}^T(\omega)=0$ ] for  $(\hat{\omega}_c/ck_0, \omega_p^2/c^2k_0^2)=(0.1, 0.4)$ . Moreover, the maximum growth rate assumes a relatively small value,  $[\omega_i]_{\max} \approx 0.06 ck_0$ , for the small value of wiggler amplitude chosen in Fig. 3.

Shown in Fig. 4, for  $\bar{\gamma}=2$  and small wiggler amplitude ( $\hat{\omega}_c/ck_0=0.1$ ), is a plot of normalized growth rate  $\omega_i/ck_0$  versus  $(k+nk_0)/k_0$  [Eq. (57)] for several values of normalized beam density  $\omega_p^2/c^2k_0^2$ . An important feature of Fig. 4 is the fact that the instability bandwidth remains

reasonably narrow for  $\omega_p^2/c^2k_0^2$  in the range  $0.1 \leq \omega_p^2/c^2k_0^2 \leq 1$ . This is in contrast to Fig. 5 where  $\omega_i/ck_0$  is plotted versus  $(k+nk_0)/k_0$  [Eq. (57)] for  $\bar{\gamma}=2$ , moderate wiggler amplitude ( $\hat{\omega}_c/ck_0=0.5$ ), and values of  $\omega_p^2/c^2k_0^2$  identical to Fig. 4. The rapid increase of instability bandwidth with normalized wiggler amplitude  $\hat{\omega}_c/ck_0$  is also evident from Fig. 6 where the growth rate is plotted for  $\bar{\gamma}=2$ ,  $\omega_p^2/c^2k_0^2=0.4$  and several values of  $\hat{\omega}_c/ck_0$  [Eq. (57)].

We now examine the case where  $\bar{\gamma}$  is large. Figures 7-11 show plots of normalized growth rate  $\omega_i/ck_0$  versus  $(k+nk_0)/k_0$  for  $\bar{\gamma}=50$  and a wide range of values of normalized wiggler amplitude  $\hat{\omega}_c/ck_0$  and beam density  $\omega_p^2/c^2k_0^2$ . In all cases, we note from Figs. 7-11 that the instability bandwidth is large. This is in contrast with the low- $\bar{\gamma}$  case where the bandwidth is relatively narrow for small values of  $\hat{\omega}_c/ck_0$  (Fig. 3). We also find (Figs. 7 and 8) that the growth rates determined from the reference dispersion relation (59) and the full dispersion relation (57) are different for both small wiggler amplitude (Fig. 8) and moderate wiggler amplitude (Fig. 7). This is in contrast with the low- $\bar{\gamma}$  case where the reference dispersion relation (59) gives a good estimate of the growth rate for small values of  $\hat{\omega}_c/ck_0$  (Fig. 3). Comparing Figs. 4 and 9, for small wiggler amplitude ( $\hat{\omega}_c/ck_0=0.1$ ), it is evident that the growth rate is considerably larger for large values of  $\bar{\gamma}$ . On the other hand, comparing Figs. 5 and 10, for moderate wiggler amplitude ( $\hat{\omega}_c/ck_0=0.5$ ), the growth rate is only slightly larger for large values of  $\bar{\gamma}$ .

Finally, for  $(k+nk_0)^2 v_b^2 \gg \omega_p^2/\bar{\gamma}^2$ , we expect the strongest interaction between the longitudinal and transverse modes,  $D_n^L(\omega)=0$  and  $D_{n-1}^T(\omega)=0$ , to occur for the critical value of wavenumber given in Sec. 1, i.e., for

$$(k+nk_0)_u \approx (1+\beta_b)^{-2} \bar{\gamma}^2 k_0. \quad (60)$$

For  $\beta_b \approx 1$  and  $\bar{\gamma} = 50$ , Eq. (60) gives  $(k+nk_0)_s \approx 5000 k_0$ . For  $\bar{\gamma} = 50$ , we note from Figs. 7-11 that the maximum growth rate indeed occurs for  $(k+nk_0)/k_0 \approx 5000$ .

#### D. Influence of Beam Thermal Effects on Stability Properties

The main purpose of this paper is to develop a stability formalism for general beam equilibrium distribution  $G_0(p_z)$  (Secs. 2.A-3.A) and to study detailed stability properties for the case of a cold electron beam with  $G_0(p_z) = \delta(p_z - p_0)$  (Secs. 3.B and 3.C). For completeness, however, in this section we make some preliminary estimates of the influence of beam thermal effects on stability behavior. As a simple example, we assume a beam equilibrium distribution of the form

$$G_0(p_z) = \frac{\Delta_0}{\pi} \frac{1}{(p_z - p_0)^2 + \Delta_0^2}, \quad (61)$$

where  $\Delta_0$  is the characteristic momentum spread about the mean momentum  $p_0$ . We further assume that the momentum spread  $\Delta_0$  is small in comparison with the directed momentum  $p_0$ , i.e.,

$$\Delta_0 \ll p_0. \quad (62)$$

Making use of Eqs. (36), (39), (61), and (62), the longitudinal dielectric function can be approximated by

$$D_n^L(\omega) = c^2 (k+nk_0)^2 \left( 1 - \frac{\omega_p^2 / \bar{\gamma}^2}{[\omega - (k+nk_0)V_b + i|k+nk_0|\Delta_0/\bar{\gamma}m]^2} \right), \quad (63)$$

for  $\Delta_0 \ll p_0$ . In Eq. (63),  $\bar{\gamma}mc^2 = (m^2c^4 + c^2p_0^2 + e^2B^2/k_0^2)^{1/2}$  is the mean energy [Eq. (48)], and  $V_b = p_0/\bar{\gamma}m$  is the directed beam velocity. We note from Eq. (63) that the fundamental longitudinal mode obtained from  $D_n^L(\omega) = 0$  is heavily Landau damped by thermal effects whenever

$$(k+nk_0)^2 \frac{\Delta_0^2}{\bar{\gamma}^2 m^2} > \frac{\omega_p^2}{\bar{\gamma}^2}. \quad (64)$$

It is useful to express the effective momentum spread  $\Delta_0$  in terms of an equivalent energy spread  $\Delta\gamma$ . From  $\gamma mc^2 = (m^2 c^4 + c^2 p_z^2 + e^2 B^2 / k_0^2)^{1/2}$ , we estimate  $\Delta\gamma = p_0 \Delta_0 / \bar{\gamma} m^2 c^2$ , or equivalently  $\Delta\gamma / \bar{\gamma} = \beta_b \Delta_0 / \bar{\gamma} mc$ . Equation (64) can then be expressed as

$$c^2 (k + nk_0)^2 (\Delta\gamma / \bar{\gamma})^2 > \beta_b^2 \frac{\omega_p^2}{\bar{\gamma}^2}, \quad (65)$$

for  $\Delta_0 \ll p_0$ , or equivalently for  $\Delta\gamma / \bar{\gamma} \ll \beta_b^2$ . If we estimate the wavenumber of interest for the free electron laser instability by  $(k + nk_0) = (1 + \beta_b) \bar{\gamma}^{-2} k_0$  [Eq. (60)], then Eq. (65) gives

$$(\Delta\gamma / \bar{\gamma})^2 > \frac{\beta_b^2}{(1 + \beta_b)^2} \frac{1}{\bar{\gamma}^6} \frac{\omega_p^2}{c^2 k_0^2}. \quad (66)$$

Equation (66) gives an estimate of the energy spread required for heavy Landau damping of the fundamental longitudinal mode obtained from  $D_n^L(\omega) = 0$ . Whenever Eq. (66) is satisfied, the waves are heavily damped (linearly) and the free electron laser instability does not occur. Equation (66) constitutes a very stringent requirement on energy spread. For example, for  $\bar{\gamma} = 5$ ,  $\beta_b = 1$  and  $\omega_p^2 / c^2 k_0^2 = 0.5$ , Eq. (66) predicts stabilization for  $(\Delta\gamma / \bar{\gamma})^2 > 8 \times 10^{-6}$ , i.e., for a fractional energy spread  $\Delta\gamma / \bar{\gamma}$  of about 0.3%.

From a practical point of view, for given values of  $\bar{\gamma}$  and energy spread  $\Delta\gamma$ , the reverse of the inequality in Eq. (66) can be used to determine the range of normalized beam density  $\omega_p^2 / c^2 k_0^2$  required for the free electron laser instability to occur, i.e.,

$$\frac{\omega_p^2}{c^2 k_0^2} > \bar{\gamma}^{-6} (\Delta\gamma / \bar{\gamma})^2 \frac{(1 + \beta_b)^2}{\beta_b^2}. \quad (67)$$

Whenever Eq. (67) is satisfied, it is important to note that the inequality in Eq. (65) will provide stabilization of the instability growth rate for sufficiently large values of wavenumber  $(k + nk_0)$ . This implies a natural tendency of beam thermal effects to limit the instability bandwidth, which should be contrasted with the cold-beam stability results in Sec. 3.C.

#### 4. CONCLUSIONS

In this paper, we have developed a fully self-consistent description of the free electron laser instability based on the Vlasov-Maxwell equations. As summarized in Sec. 2, the present analysis assumes a relativistic electron beam with uniform cross section propagating through a helical wiggler field  $\mathbf{B}_0 = -\hat{B} \cos k_0 z \hat{e}_x - \hat{B} \sin k_0 z \hat{e}_y$  [Eq.(2)]. Moreover, we consider the class of exact solutions to the Vlasov-Maxwell equations described by  $f_b(z, \mathbf{p}, t) = n_0 \delta(P_x) \delta(P_y) G(z, p_z, t)$  [Eq.(12)]. A detailed analysis of the linearized Vlasov-Maxwell equations (Secs. 2.B and 3.A) leads to the exact matrix dispersion relation in Eq. (45). The striking feature of Eq. (45) is that the dispersion relation is valid for arbitrary harmonic number  $n$ . Moreover, Eq. (45) describes stability behavior for perturbations about general beam equilibrium distribution  $G_0(p_z)$ , and no a priori restriction has been made to low beam density (as measured by  $\omega_p^2/c^2 k_0^2$ ) or small wiggler amplitude (as measured by  $\hat{\omega}_c/ck_0$ ). In Secs. 3.B and 3.C we present a detailed numerical analysis of the full dispersion relation [Eq.(57)] and the reference dispersion relation [Eq.(59)] for the case of a cold electron beam described by  $G_0(p_z) = \delta(p_z - p_0)$ . Except for very modest values of wiggler amplitude, it is shown in Sec. 3.C that the growth rate  $\omega_i = \text{Im}\omega$  obtained from the reference dispersion relation can be in substantial error for large values of  $(k+nk_0)/k_0$ . An important feature of the stability analysis in Sec. 3.C is the fact that the instability bandwidth increases rapidly with increasing  $\hat{\omega}_c/ck_0$ . Moreover, the instability bandwidth is considerably larger for large values of  $\bar{\gamma}$  (Figs. 7 - 11) than for moderate values of  $\bar{\gamma}$  (Figs. 2 - 6).

Finally, in Sec. 3.D, we have made preliminary estimates of the influence of beam thermal effects on stability behavior. In particular, assuming that  $G_0(p_z)$  is given by the Lorentzian distribution in Eq.(61), it is found that waves with wavenumber  $(k+nk_0) \approx (1+\beta_b)\bar{\gamma}^2 k_0$  are heavily Landau damped whenever the fractional energy spread  $\Delta\gamma/\bar{\gamma}$  exceeds the rather modest value in Eq. (66).

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FIGURE CAPTIONS

- Fig. 1. Intersection of dispersion curves  $D_n^L(\omega) = 0$  and  $D_{n-1}^T(\omega) = 0$ .
- Fig. 2. Plot of normalized (a) growth rate  $\omega_i/k_0c$  and (b) real oscillation frequency  $\omega_r/k_0c$  versus normalized wavenumber  $(k+nk_0)/k_0$  for  $\gamma = 2$ ,  $\hat{\omega}_c/ck_0 = 0.5$  and  $\omega_p^2/c^2k_0^2 = 0.4$ . Solid curves are obtained from full dispersion relation [Eq. (57)]. Dashed curves are obtained from reference dispersion relation [Eq. (59)].
- Fig. 3. Plot of normalized growth rate  $\omega_i/k_0c$  versus  $(k+nk_0)/k_0$  for  $\gamma = 2$ ,  $\hat{\omega}_c/ck_0 = 0.1$  and  $\omega_p^2/c^2k_0^2 = 0.4$  obtained from Eq. (57) (solid curves) and Eq. (59) (dashed curves).
- Fig. 4. Plot of normalized growth rate  $\omega_i/k_0c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 2$ ,  $\hat{\omega}_c/ck_0 = 0.1$  and several values of  $\omega_p^2/c^2k_0^2$ .
- Fig. 5. Plot of normalized growth rate  $\omega_i/k_0c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 2$ ,  $\hat{\omega}_c/ck_0 = 0.5$  and several values of  $\omega_p^2/c^2k_0^2$ .
- Fig. 6. Plot of normalized growth rate  $\omega_i/k_0c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 2$ ,  $\omega_p^2/c^2k_0^2 = 0.4$  and several values of  $\hat{\omega}_c/ck_0$ .
- Fig. 7. Plot of normalized growth rate  $\omega_i/k_0c$  versus  $(k+nk_0)/k_0$  for  $\gamma = 50$ ,  $\hat{\omega}_c/ck_0 = 0.5$  and  $\omega_p^2/c^2k_0^2 = 0.4$  obtained from Eq. (57) (solid curve) and Eq. (59) (dashed curve).

- Fig. 8. Plot of normalized growth rate  $\omega_i/k_0 c$  versus  $(k+nk_0)/k_0$  for  $\gamma = 50$ ,  $\hat{\omega}_c/ck_0 = 0.1$  and  $\omega_p^2/c^2k_0^2 = 0.4$  obtained from Eq. (57) (solid curve) and Eq. (59) (dashed curve).
- Fig. 9. Plot of normalized growth rate  $\omega_i/k_0 c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 50$ ,  $\hat{\omega}_c/ck_0 = 0.1$  and several values of  $\omega_p^2/c^2k_0^2$ .
- Fig. 10. Plot of normalized growth rate  $\omega_i/k_0 c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 50$ ,  $\hat{\omega}_c/ck_0 = 0.5$  and several values of  $\omega_p^2/c^2k_0^2$ .
- Fig. 11. Plot of normalized growth rate  $\omega_i/k_0 c$  versus  $(k+nk_0)/k_0$  obtained from Eq. (57) for  $\gamma = 50$ ,  $\omega_p^2/c^2k_0^2 = 0.4$  and several values of  $\hat{\omega}_c/ck_0$ .

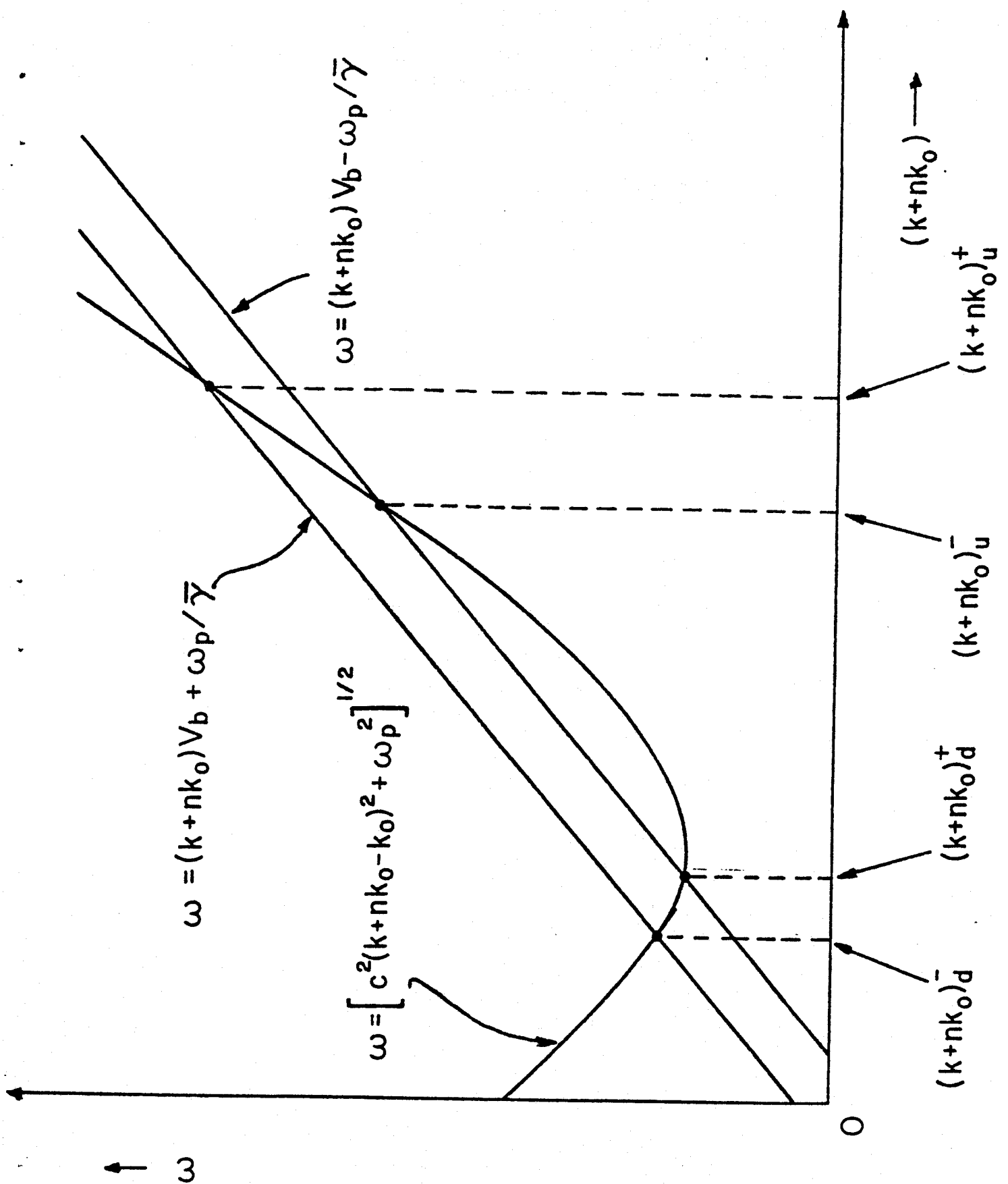
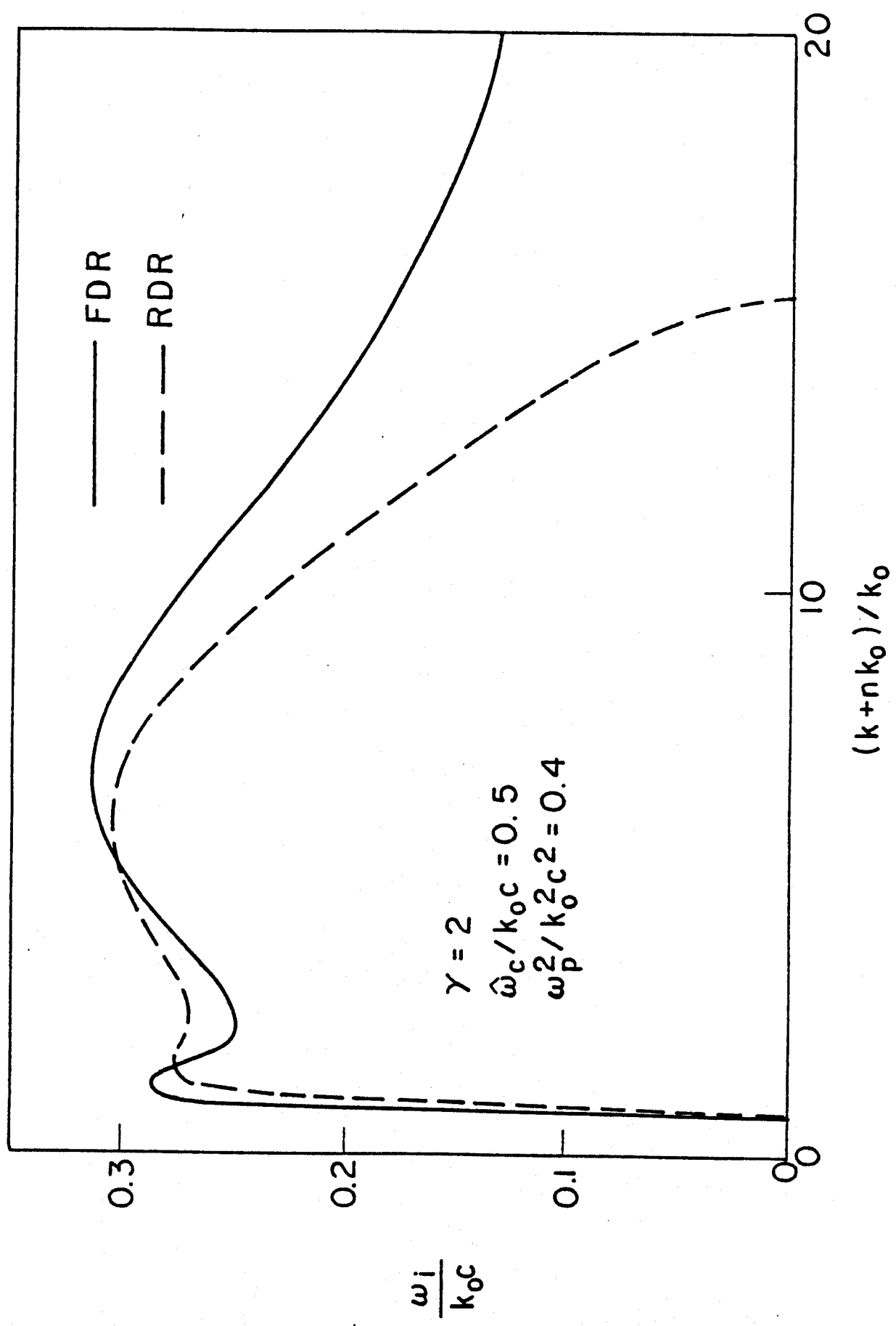


Fig. 1



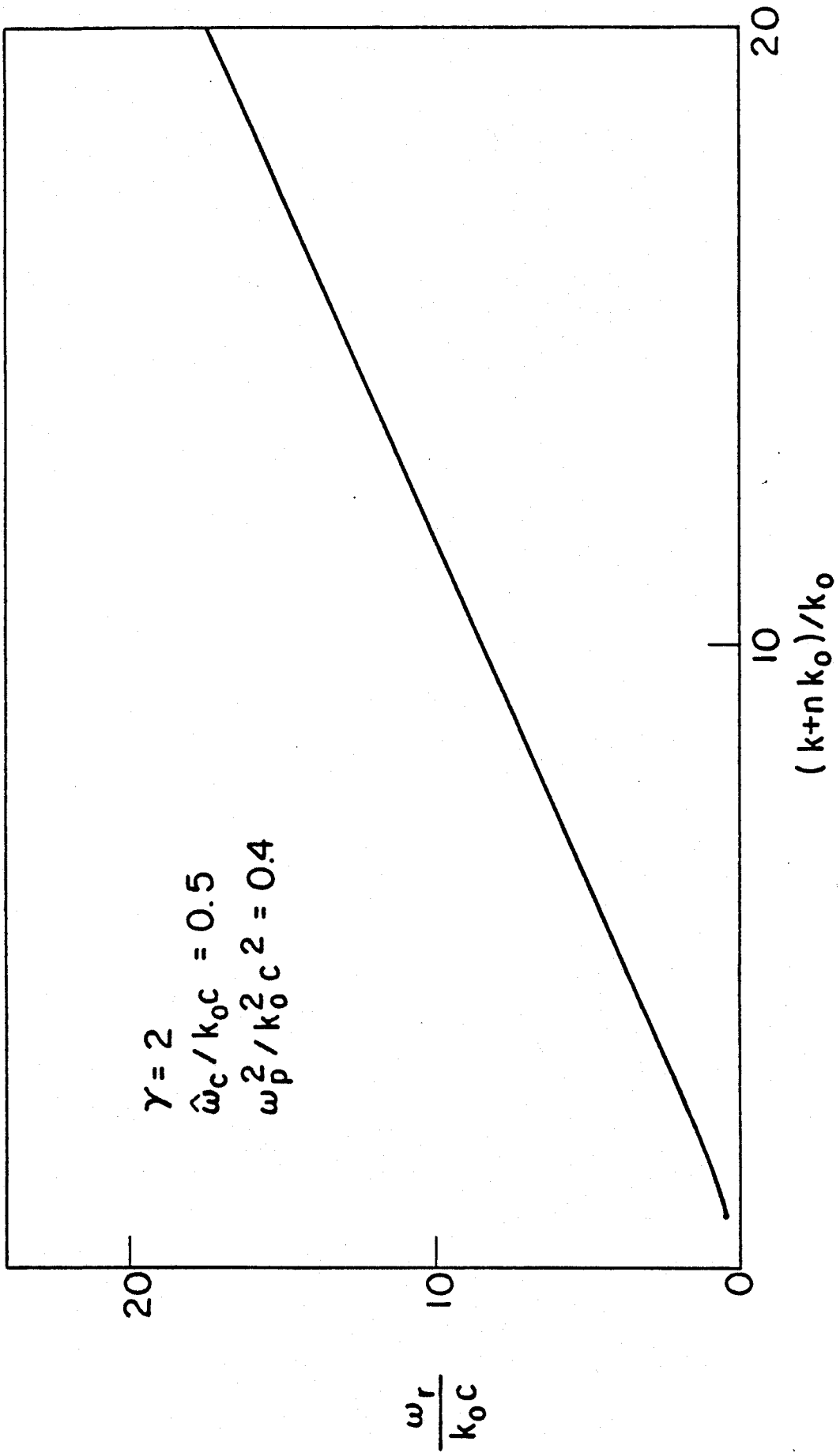
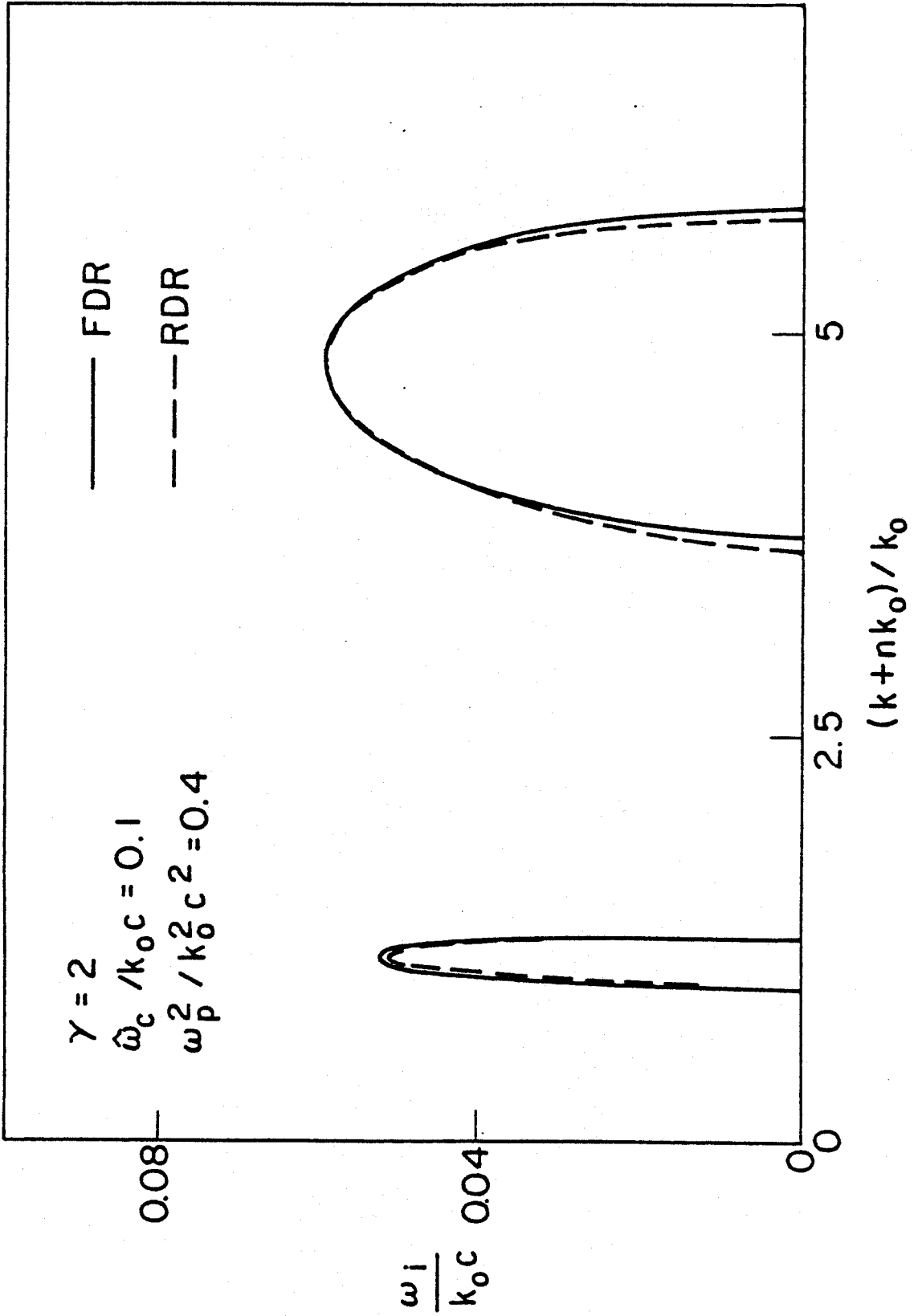


Fig. 2(b)



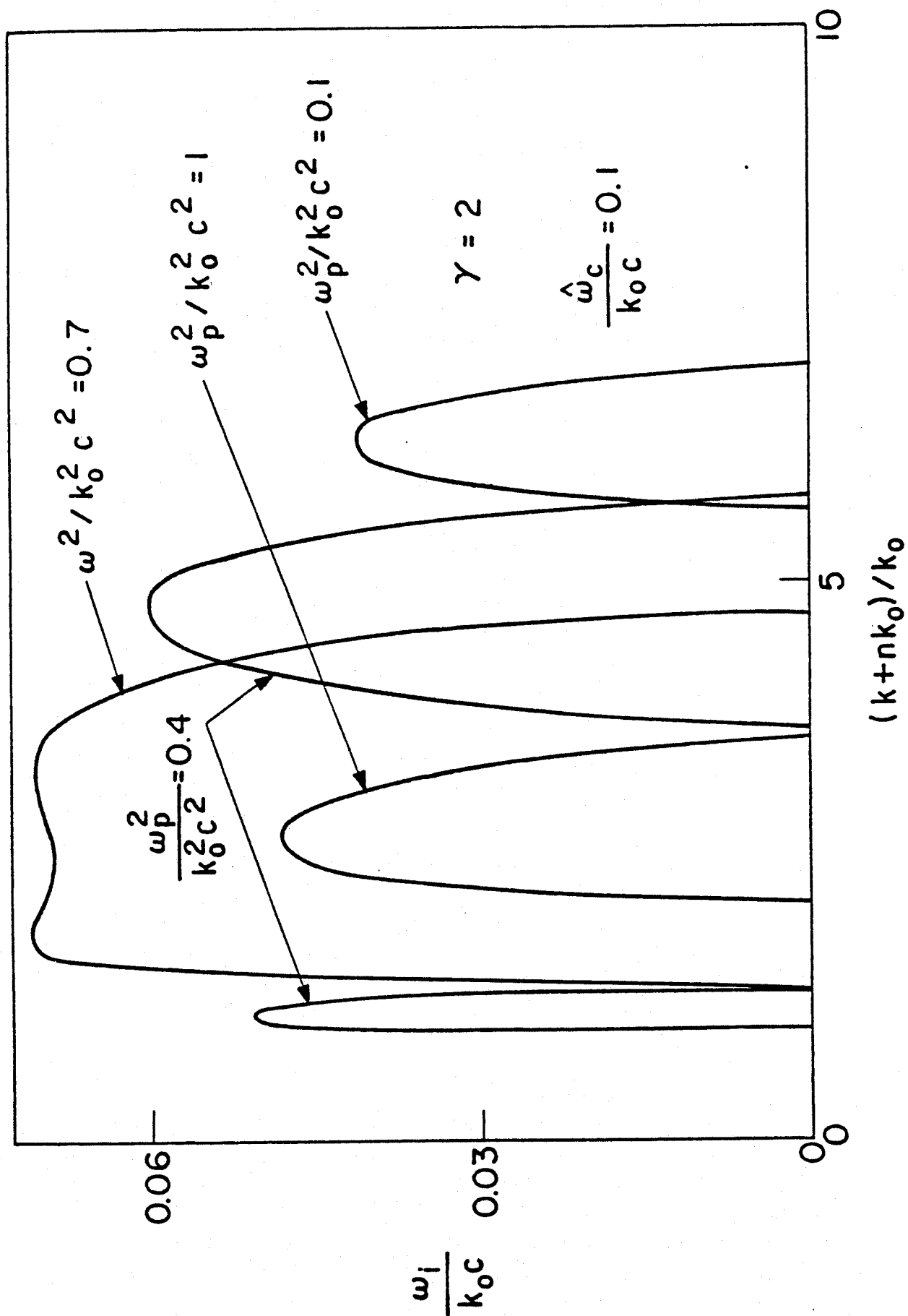
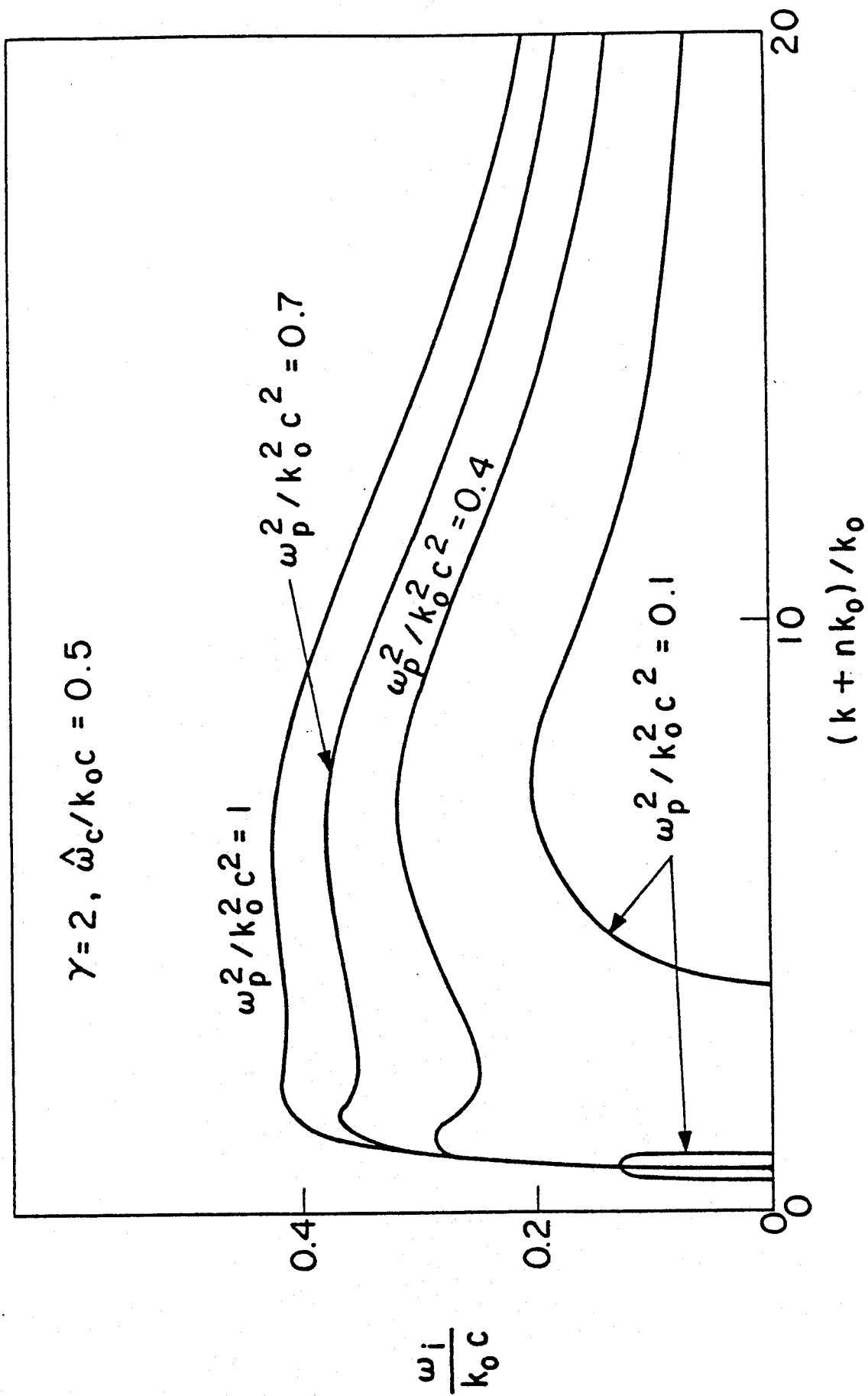
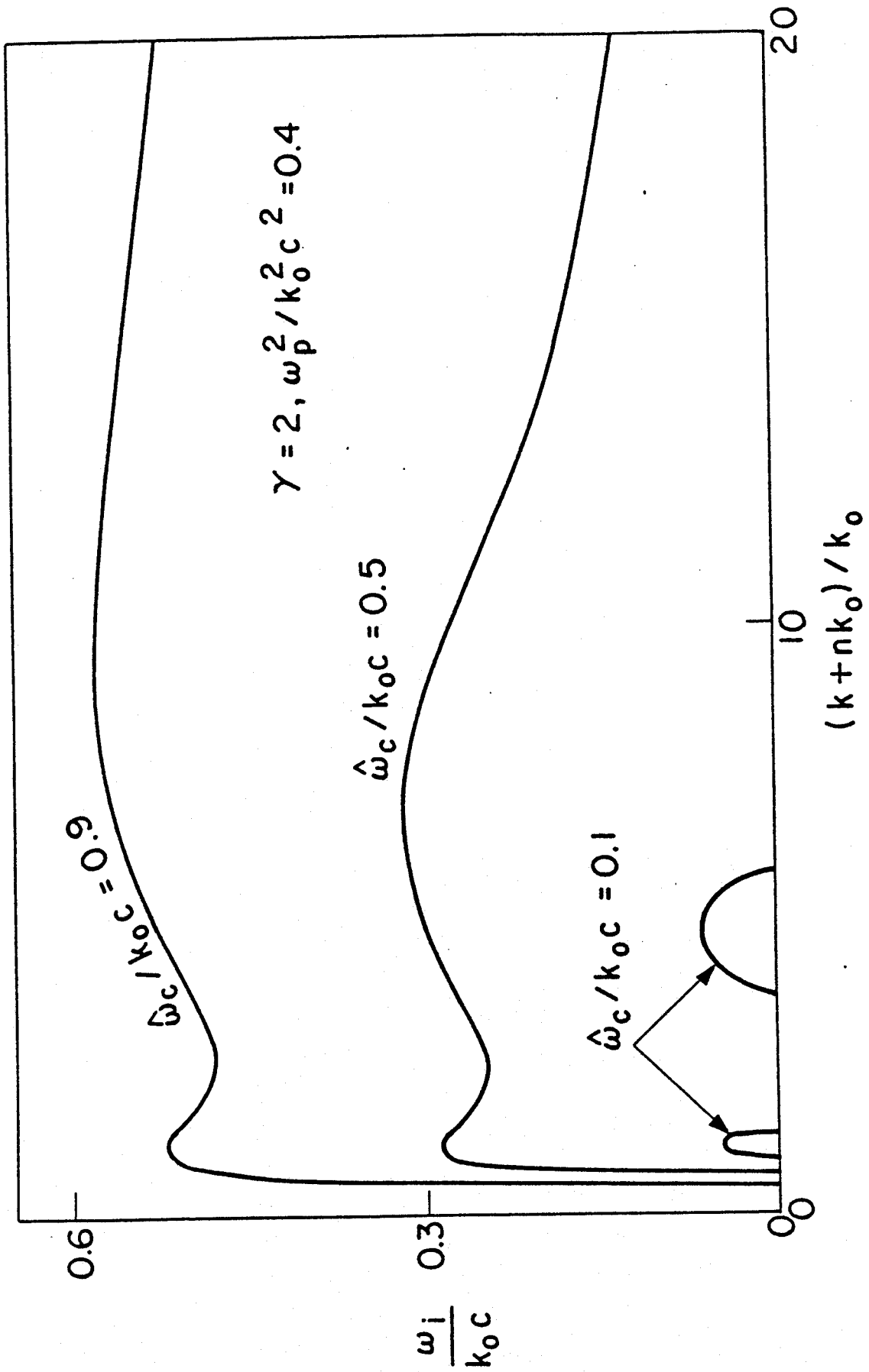


Fig. 4







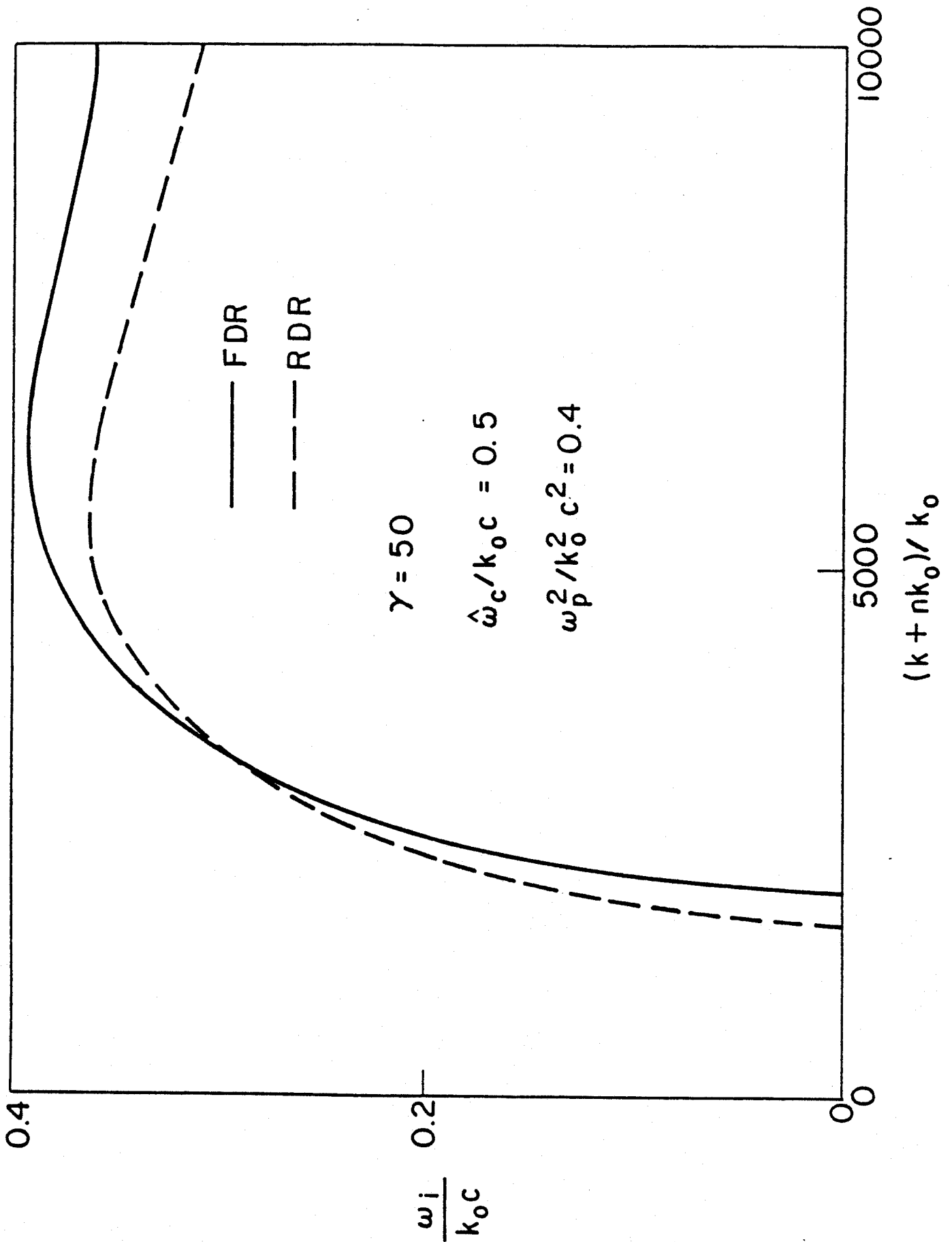


Fig. 7

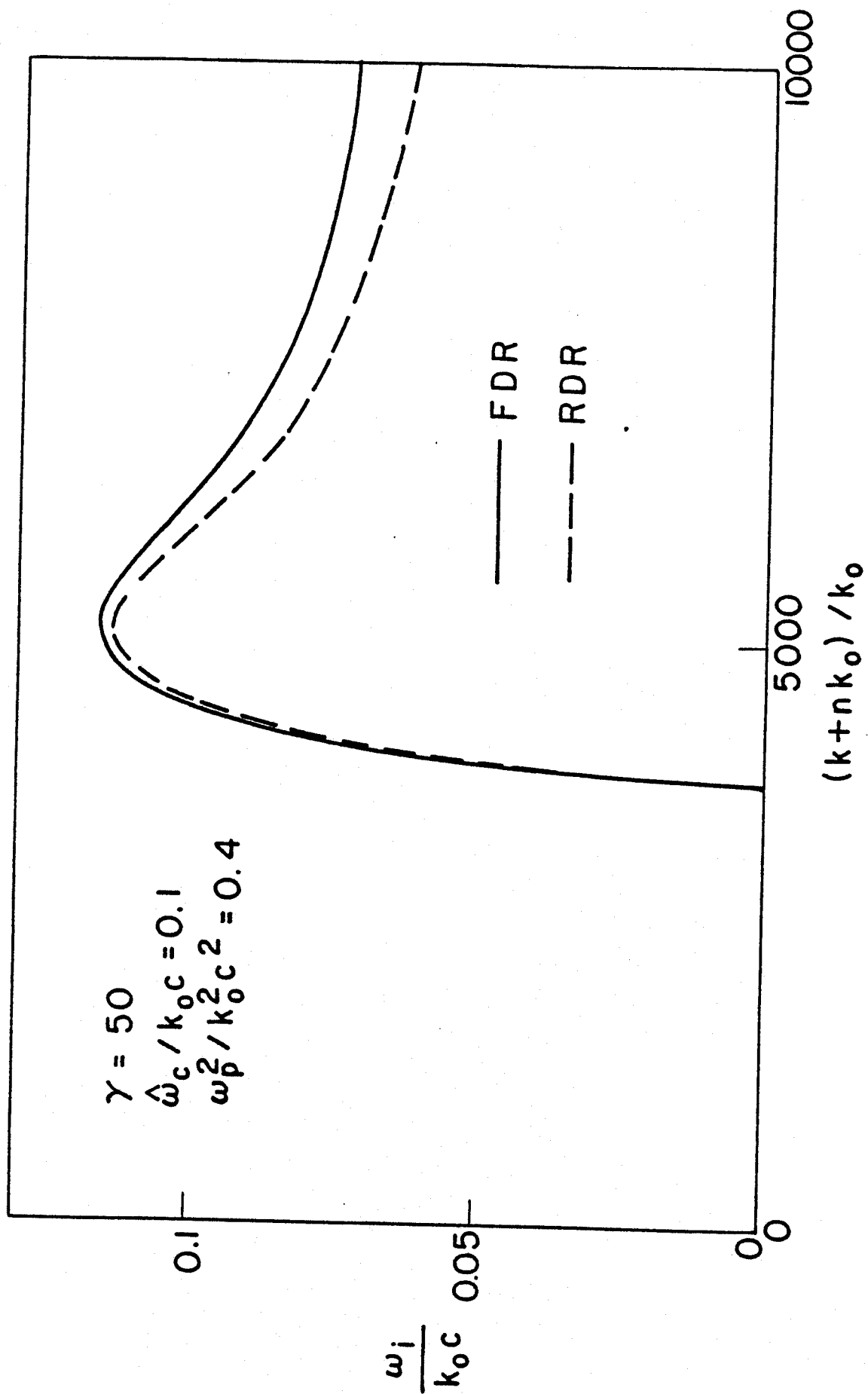
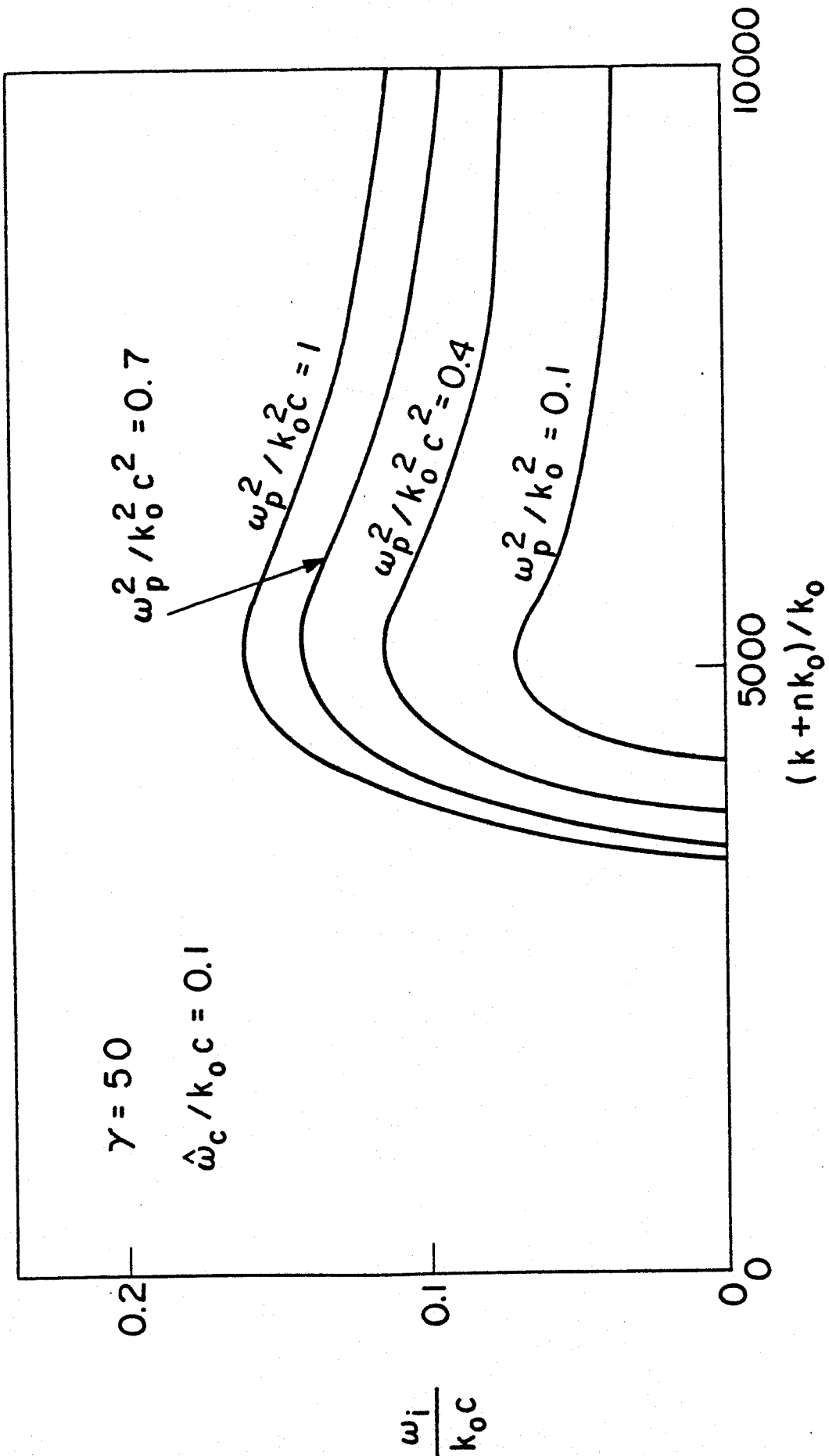
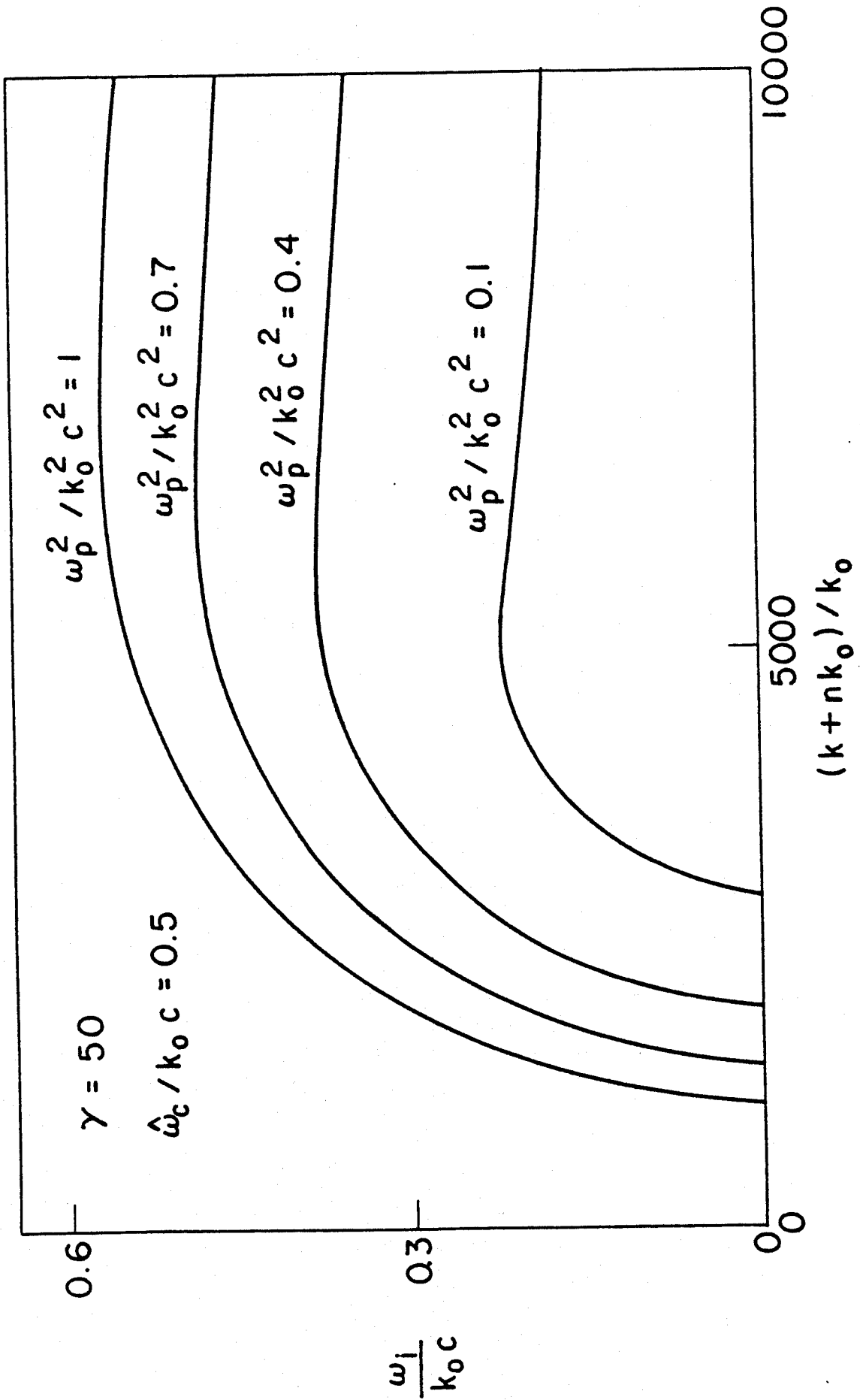


Fig. 8





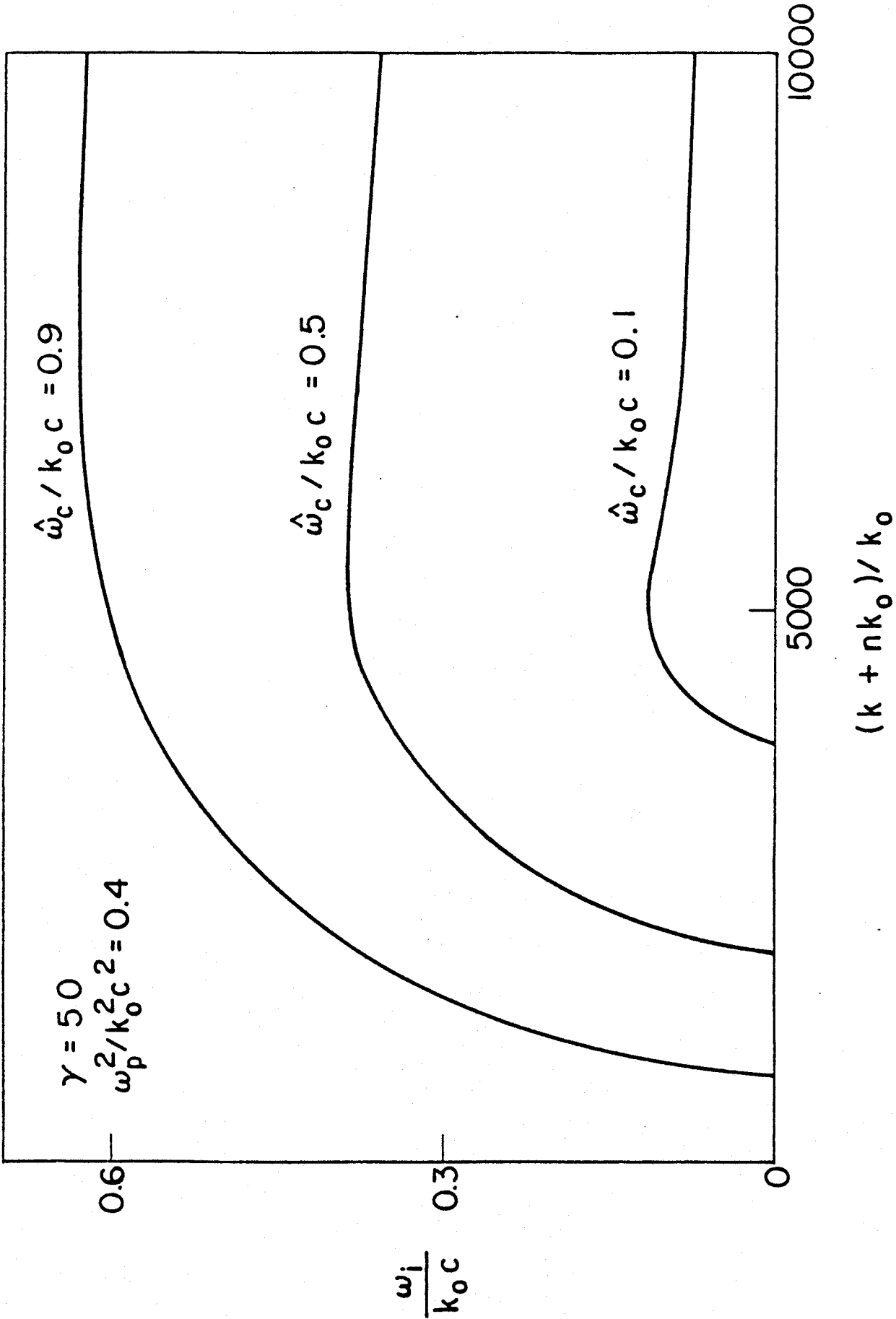


Fig. 11